

Properties of nuclear masses for heavy and superheavy nuclei

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I. Introduction

II. Bulk properties of nuclear masses

III. Nuclear mass formulae

IV. Application:

i) r-process nucleosynthesis (heavy nuclei)

ii) Superheavy nuclei

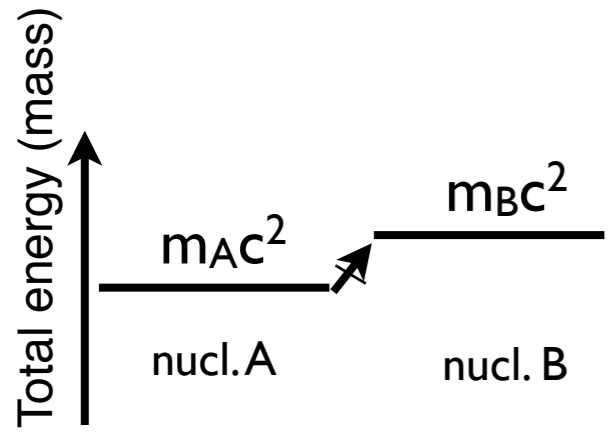
V. Summary

I. Introduction

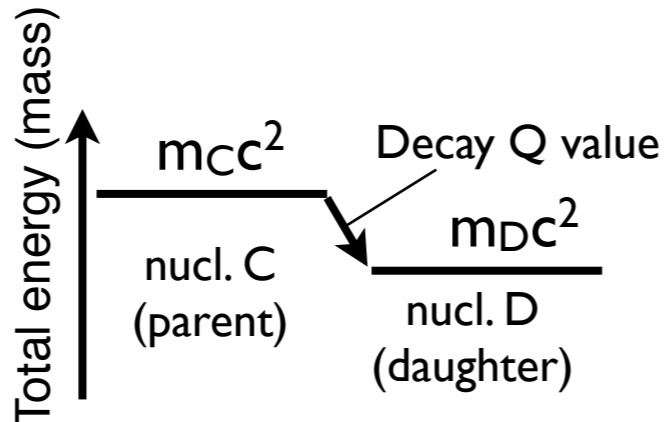
Why nuclear mass?

- Equivalence to **total energy** of nucleus: $E = mc^2$
- Governing nuclear reaction and decay modes

$$E = mc^2$$

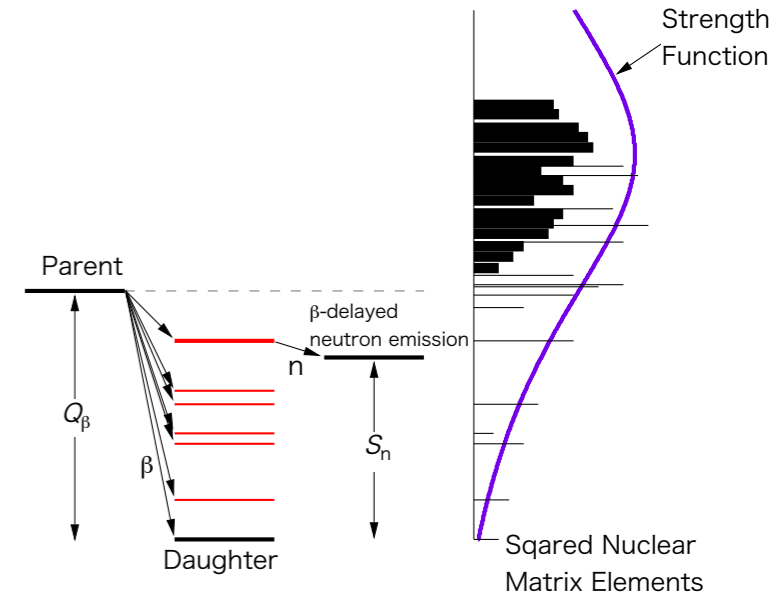


Nucleus A can not decay.



Nucleus C can decay.

Diff. of mass(total energy) determine the direction of nuclear decay.

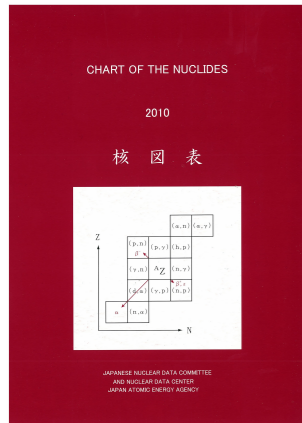


$$\lambda = \frac{1}{2\pi^3} \int_{-Q}^0 \sum_{\Omega} |g_{\Omega}|^2 \cdot |M_{\Omega}(E_g)|^2 f(-E_g + 1) dE_g$$

Decay rate of beta-decay

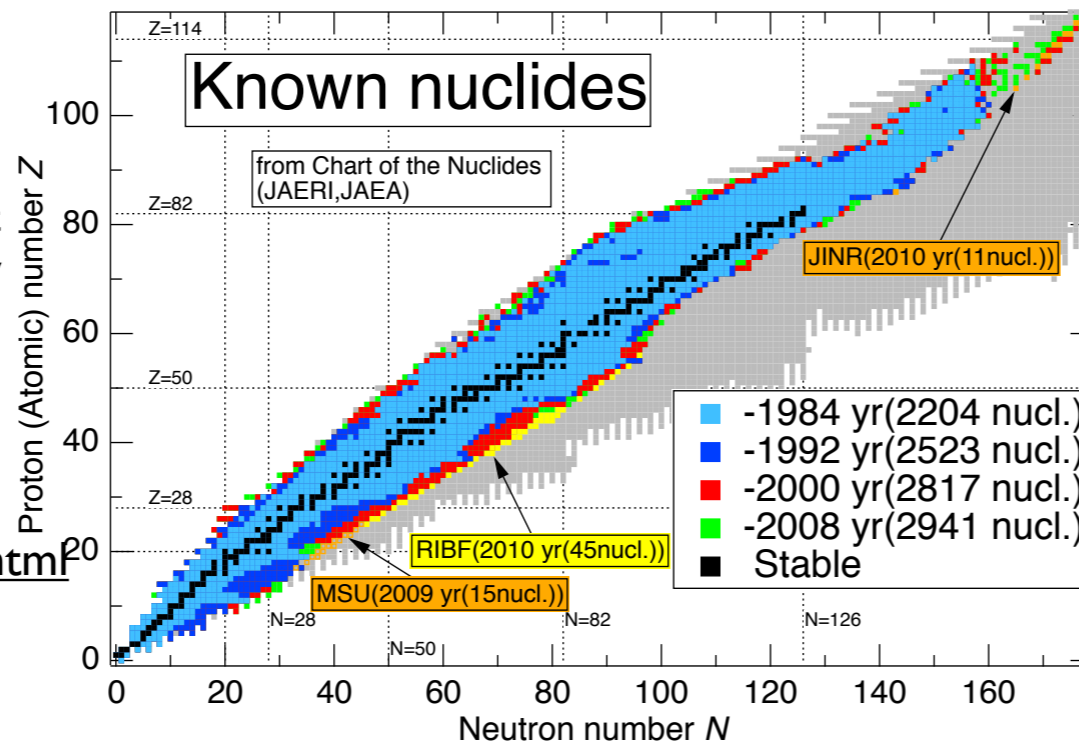
Mass-measured nuclei: current understandings

Identified



taken from Chart of the nuclides by JAERI and JAEA (Tachibana, HK, Katakura, 2010)

www.ndc.jaea.go.jp/CN10/index.html



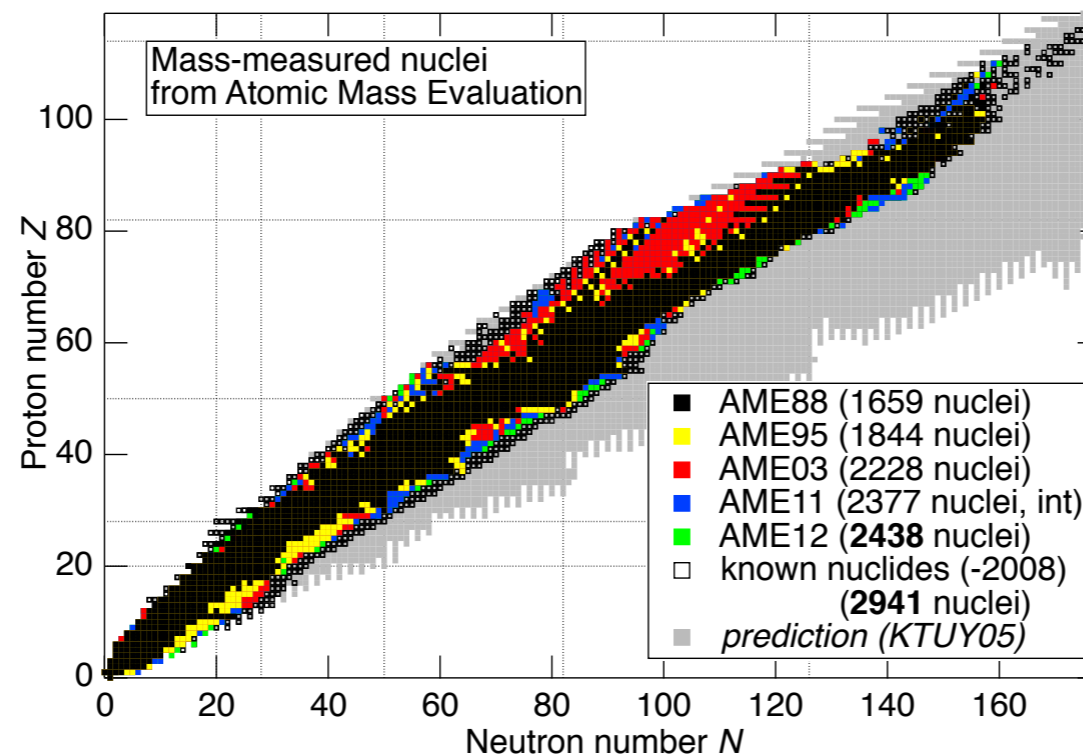
~3000 nuclei

Mass-measured



AME2012 is updated last year!

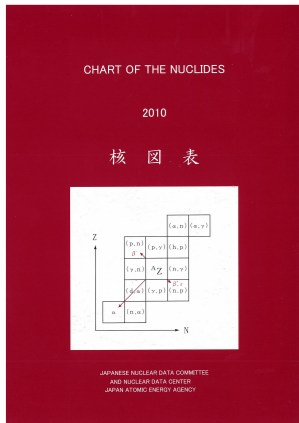
amdc.in2p3.fr/mastables/file1.html



~2400 nuclei

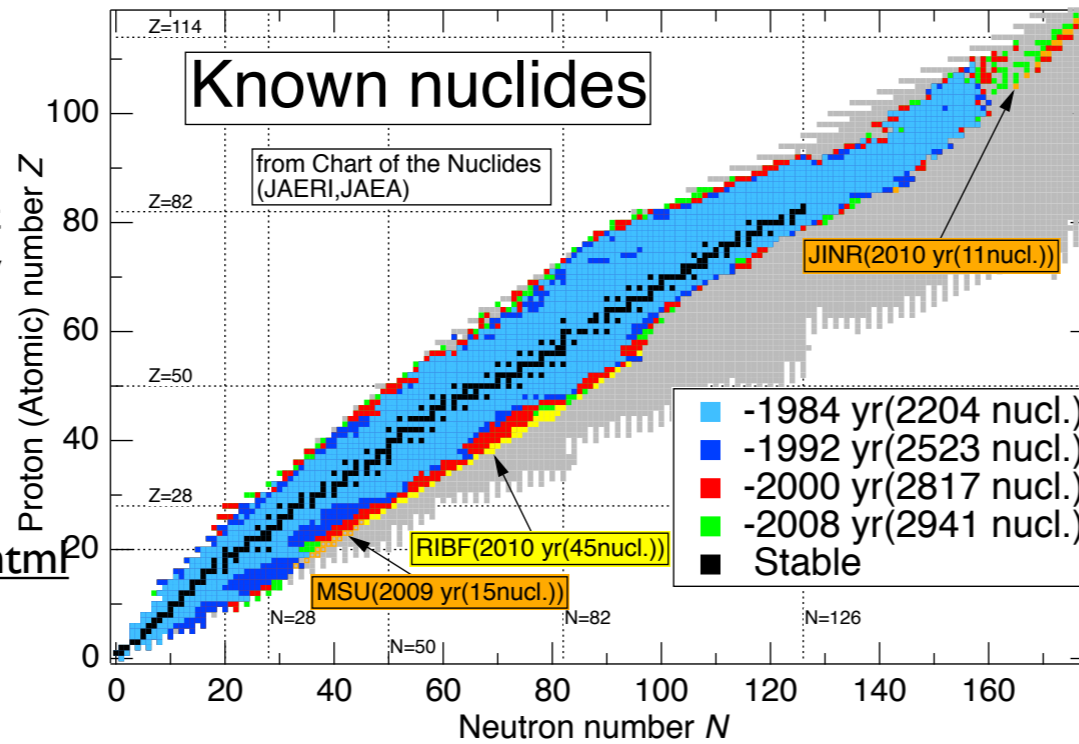
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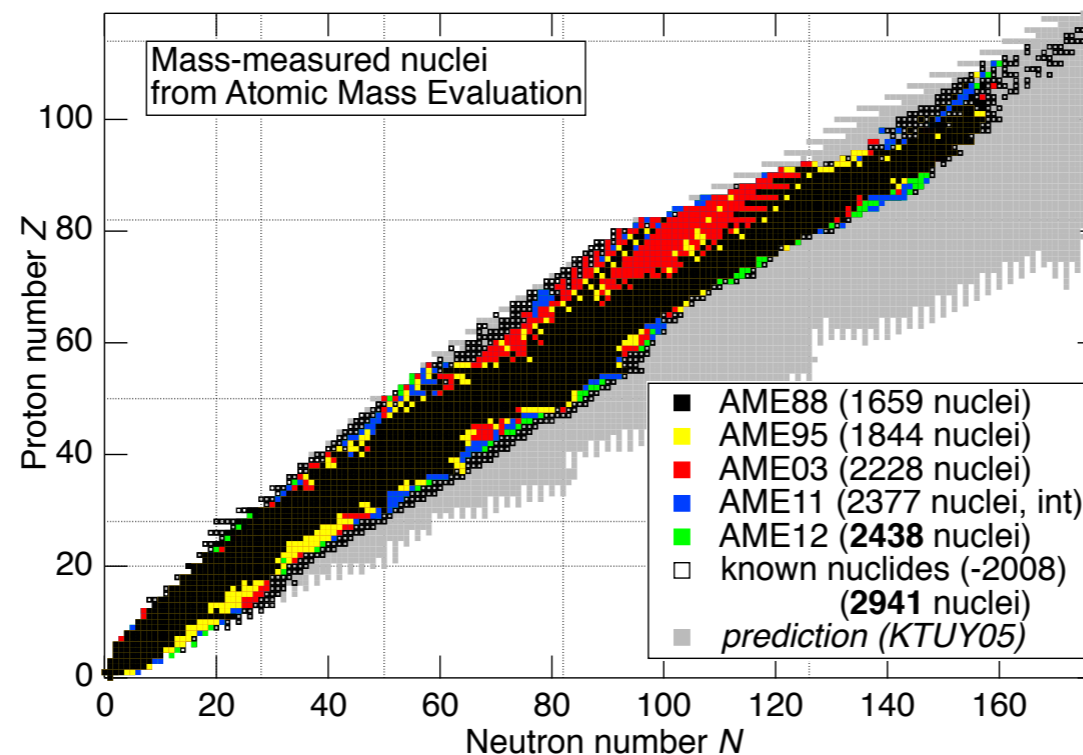
We know now
~3000 nuclei
and

Mass-measured



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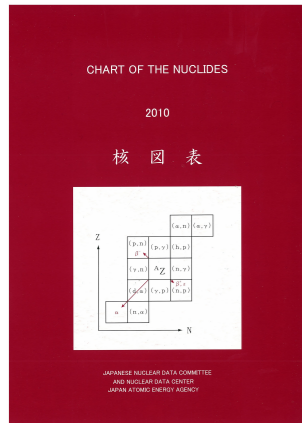
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We know now masses of
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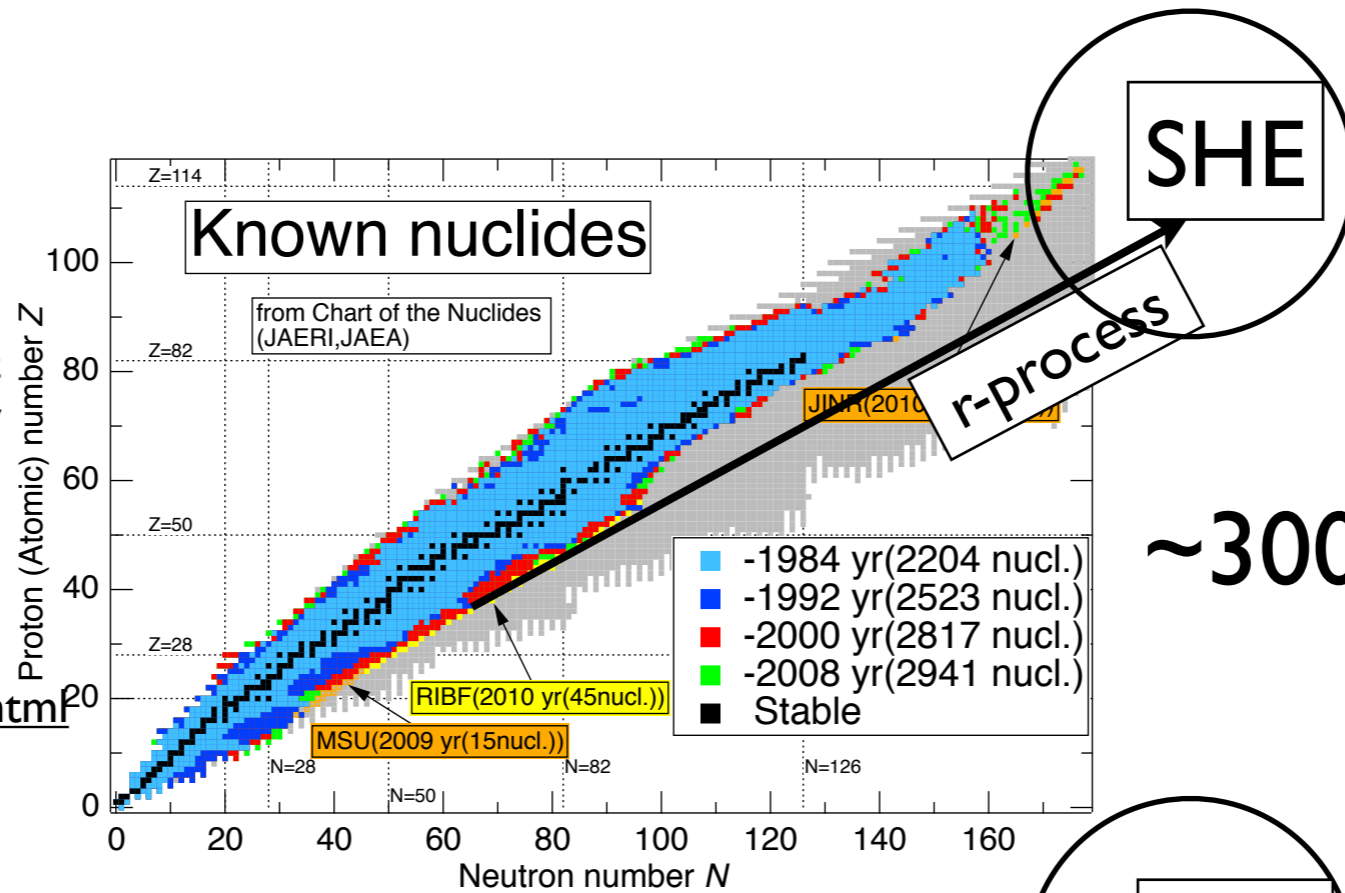
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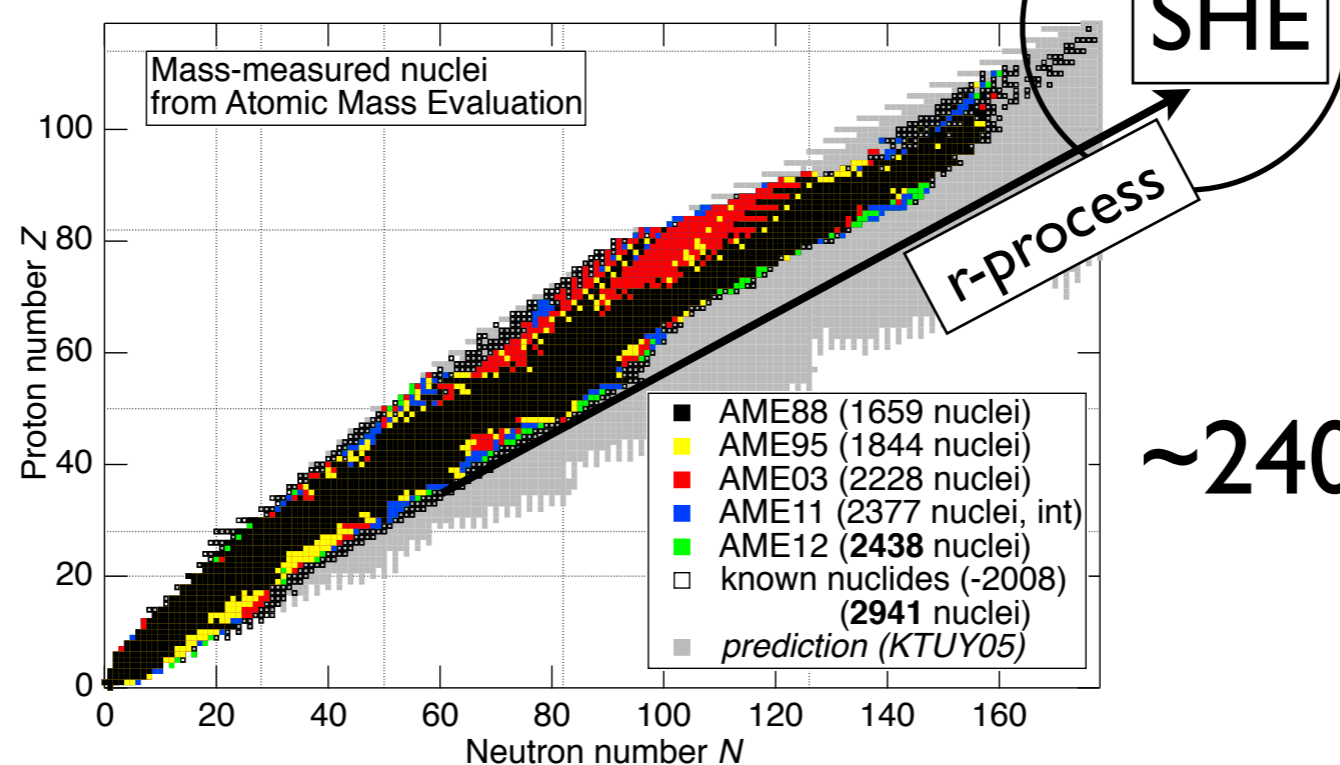
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We know now masses of

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First measurement of ground-state di-proton decay (from ^{45}Fe)

VOLUME 89, NUMBER 10 PHYSICAL REV

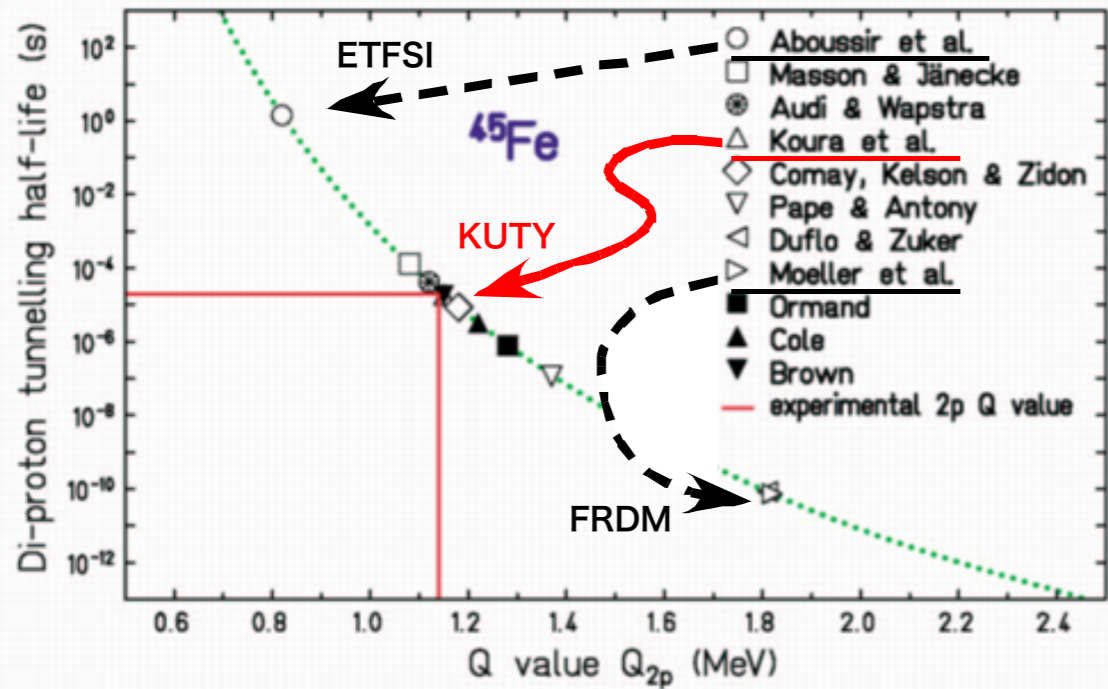
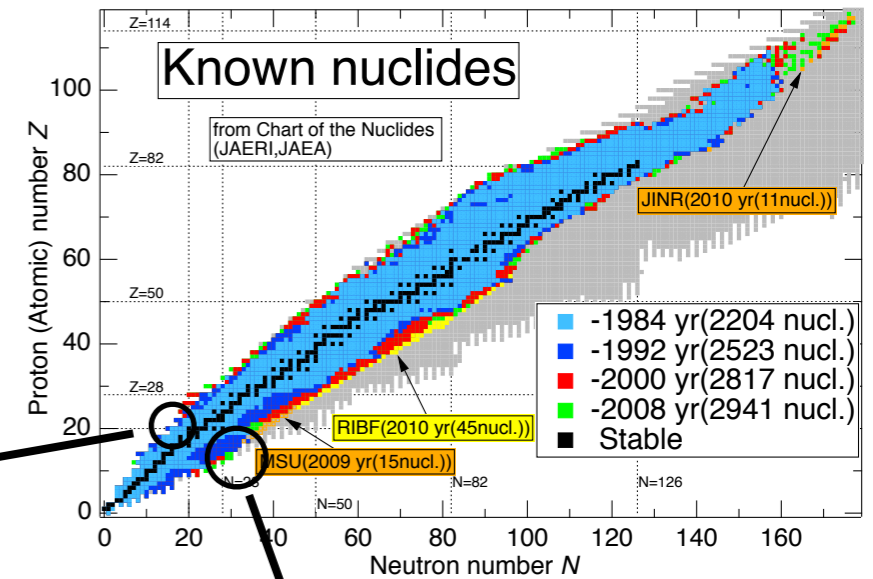


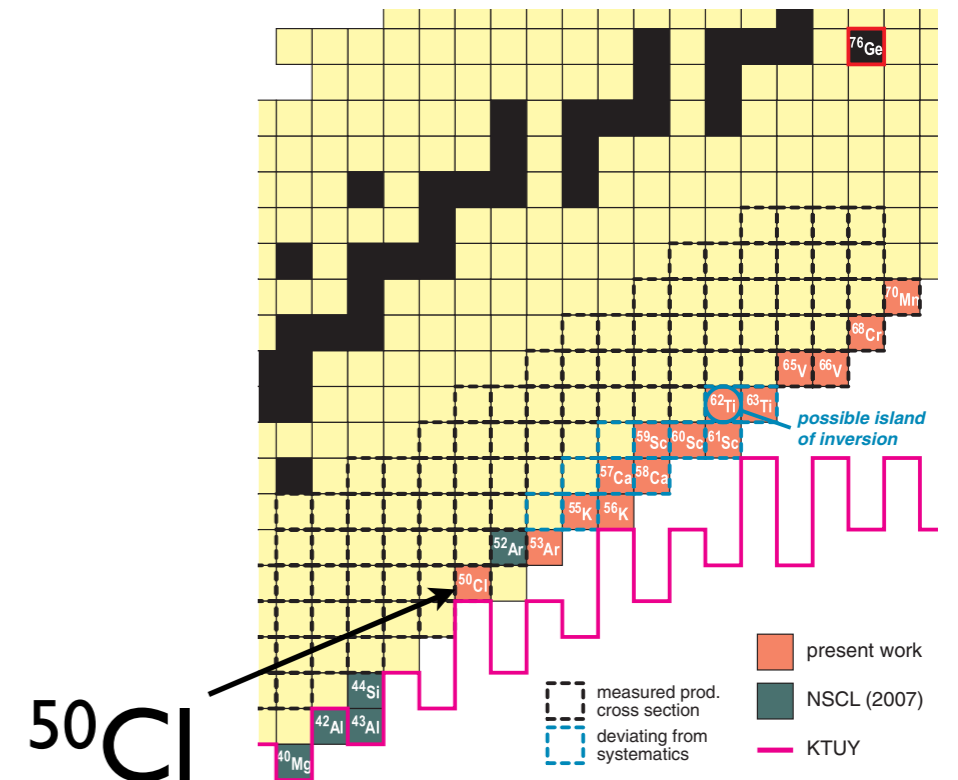
FIG. 5 (color online). Barrier-penetration half-life as a function of the two-proton Q value, Q_{2p} , for ^{45}Fe . The barrier-penetration was calculated by assuming a spectroscopic factor of unity. Different model predictions [5–7,16–21] were used for Q_{2p} . The experimentally observed Q value of ^{45}Fe implies a diproton barrier-penetration half-life of 0.024 ms.

*Gioviazzo et al., PRC 89, 2002 (GANIL)
(and also done in GSI by Pfützner(2002))*



New isotope ^{50}Cl etc. near the n-drip line

PRL 102, 142501 (2009) PHYSICAL R



O.V. Tarsov, et al., PRL 102, 2009 (MSU)

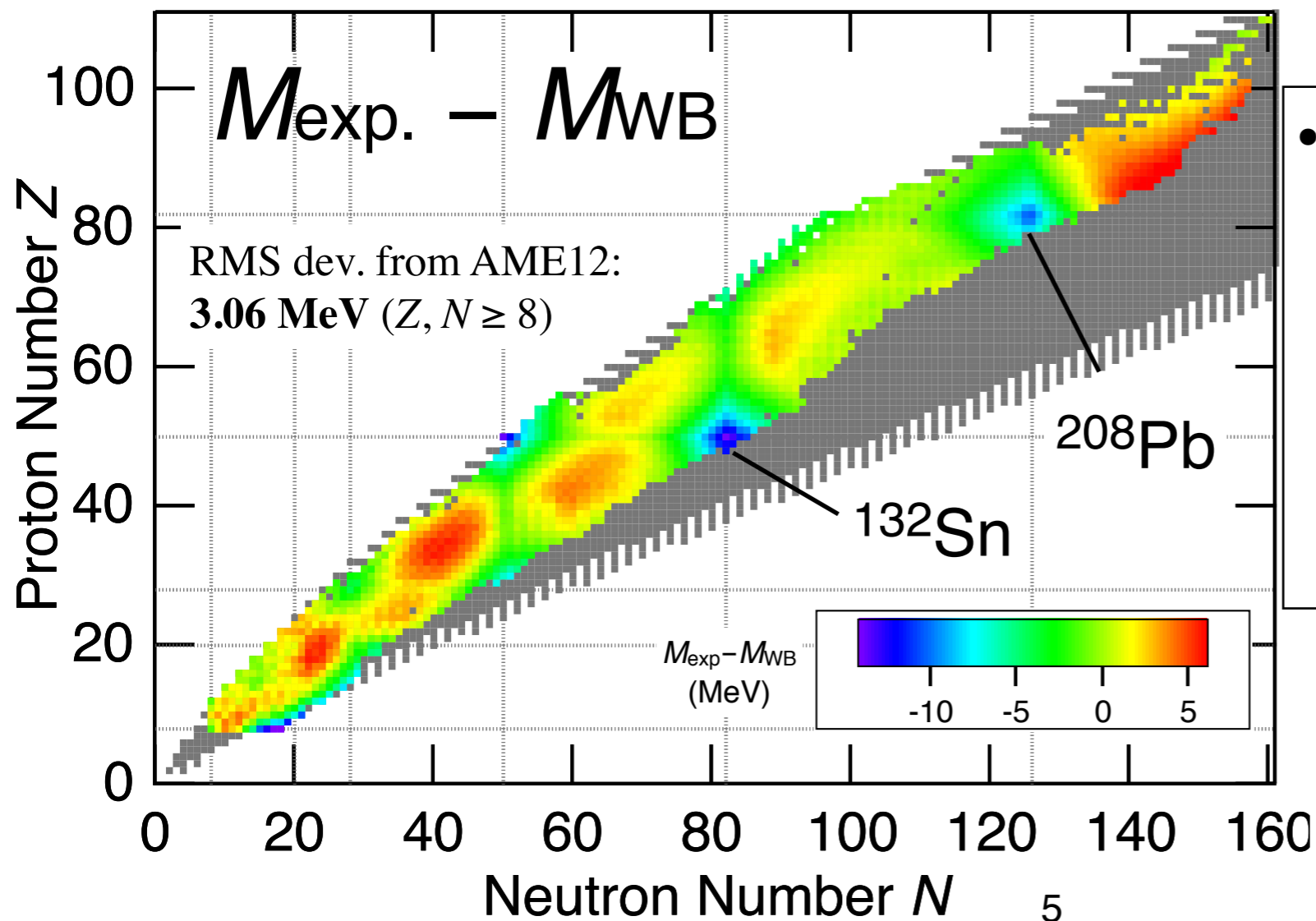
Weizsäcker-Bethe semi-empirical atomic mass formula

$$M_{WB}(Z, N) = Z m_H + N m_n - B(Z, N)$$

$$= Z m_H + N m_n - a_V A + a_S A^{2/3} + a_I (N-Z)^2 / A + a_C Z(Z-1) / A^{1/3} + \delta_{eo}$$

a_V	a_S	a_I	a_C	a_{eo}	(MeV)
15.604	17.472	22.99	0.7073	12.338	

$$\delta_{eo} = \begin{cases} -a_{eo}/A^{1/2} & \text{for even-}Z \text{ and even-}N \\ 0 & \text{for odd-}A \\ +a_{eo}/A^{1/2} & \text{for odd-}Z \text{ and odd-}N \end{cases}$$



From the 'exp.' shell energy:

- Existence of magic number
N=28,50,82,126
Z=28,50,82

Mass data : 2012 Atomic mass evaluation
(M. Wang, G. Audi, A.H. Wapstra *et al.*)

Trend in MeV-order

Weizsäcker-Bethe semi-empirical atomic mass formula

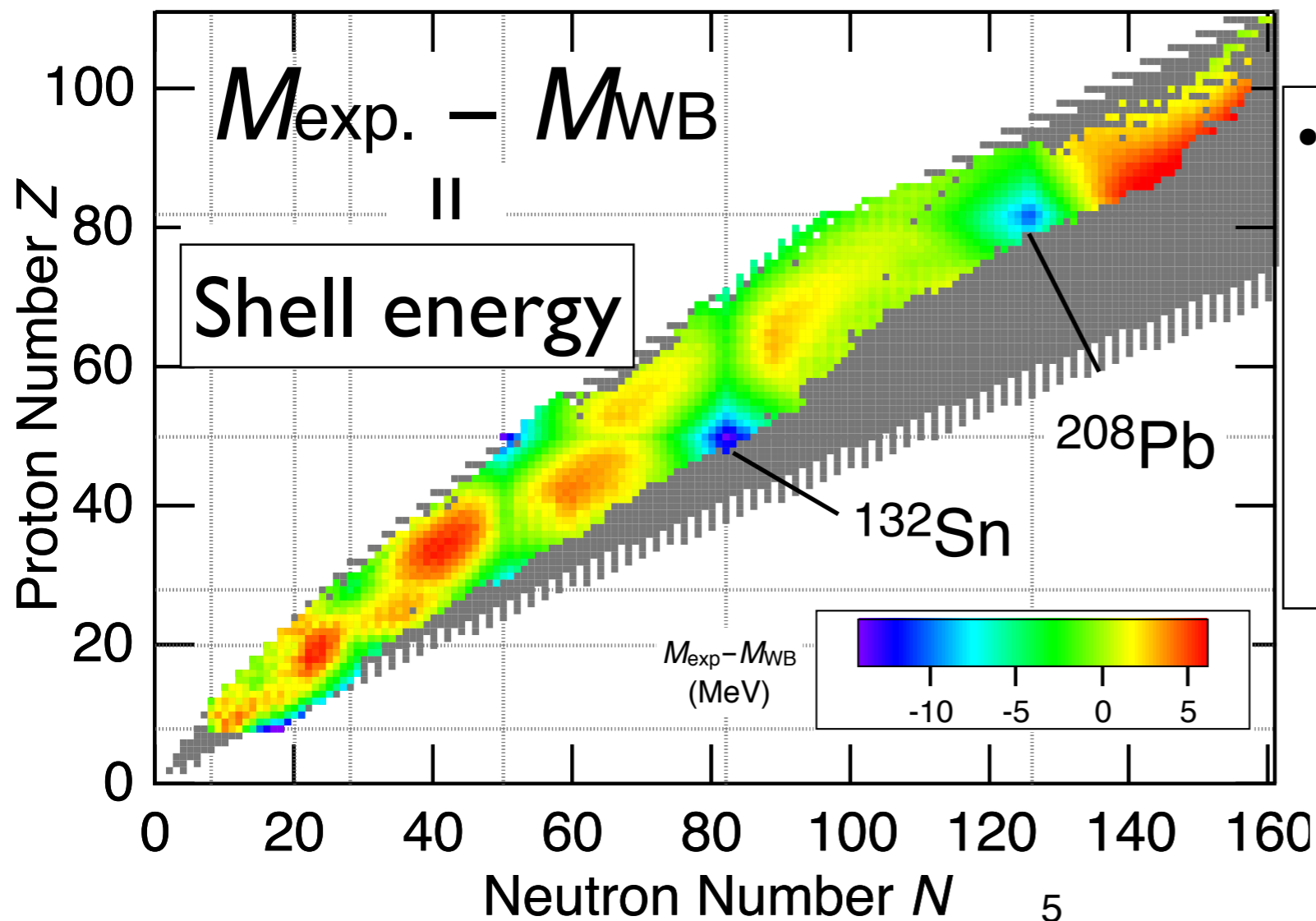
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From the 'exp.' shell energy:

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 $N=2, 8, 20, 28, 50, 82, 126$
 $Z=2, 8, 20$

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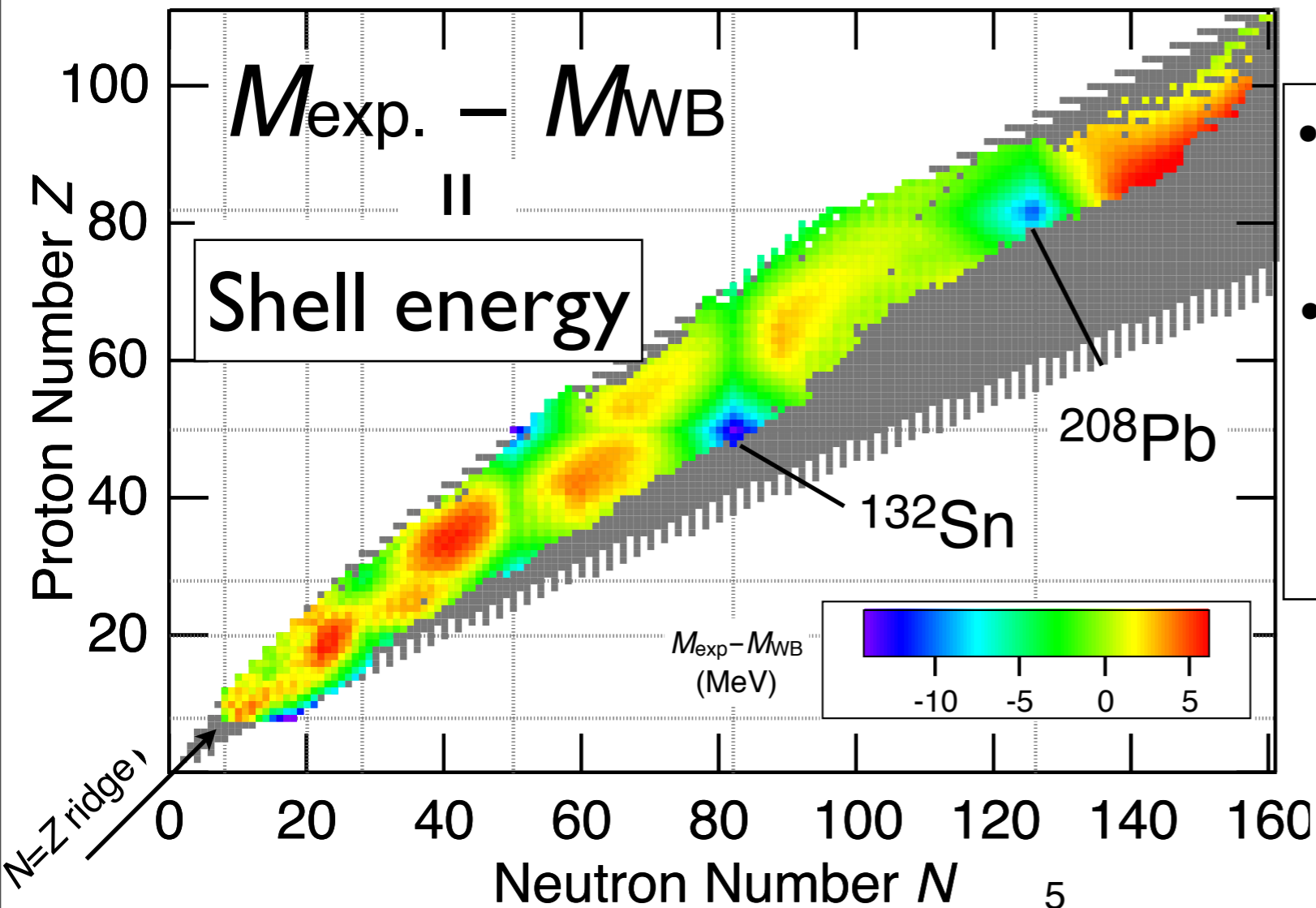
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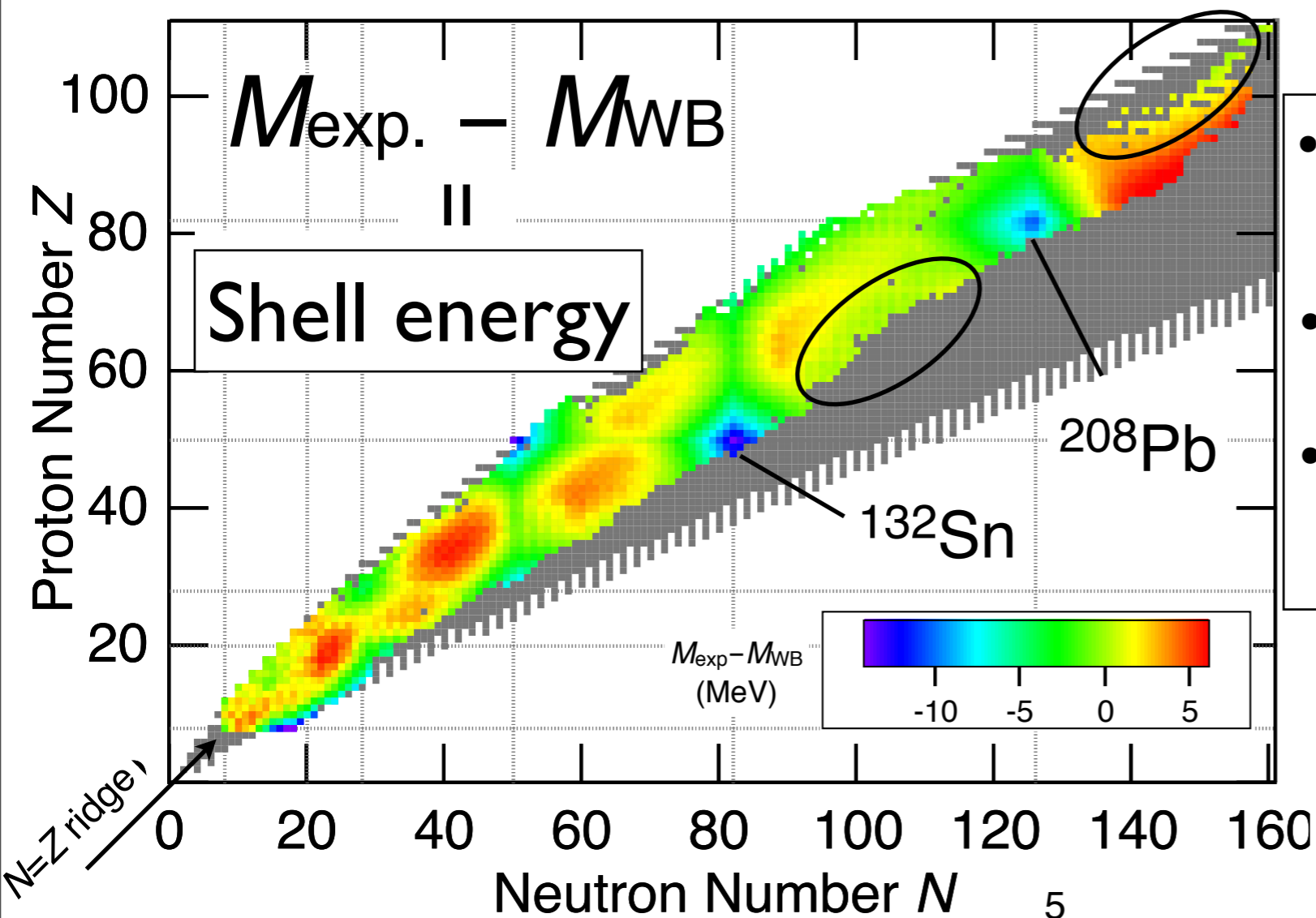
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From the 'exp.' shell energy:

- Existence of magic number
 $N=28, 50, 82, 126$
 $Z=28, 50, 82$
- Wigner energy
 $N=Z$ ridge
- Depression due to the deform.
 rare-earth, actinide

Mass data : 2012 Atomic mass evaluation
 (M. Wang, G. Audi, A.H. Wapstra *et al.*)

Trend in MeV-order

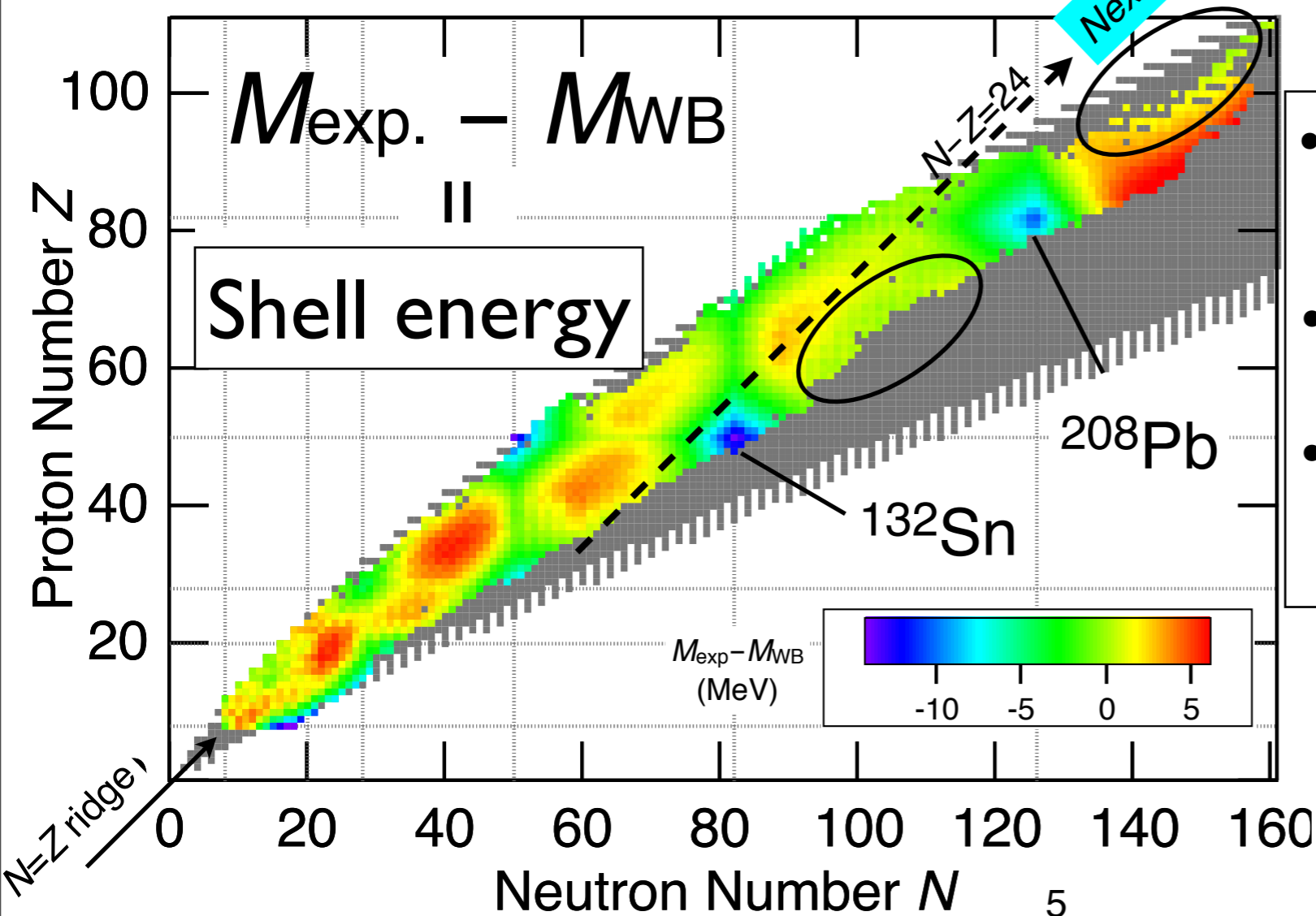
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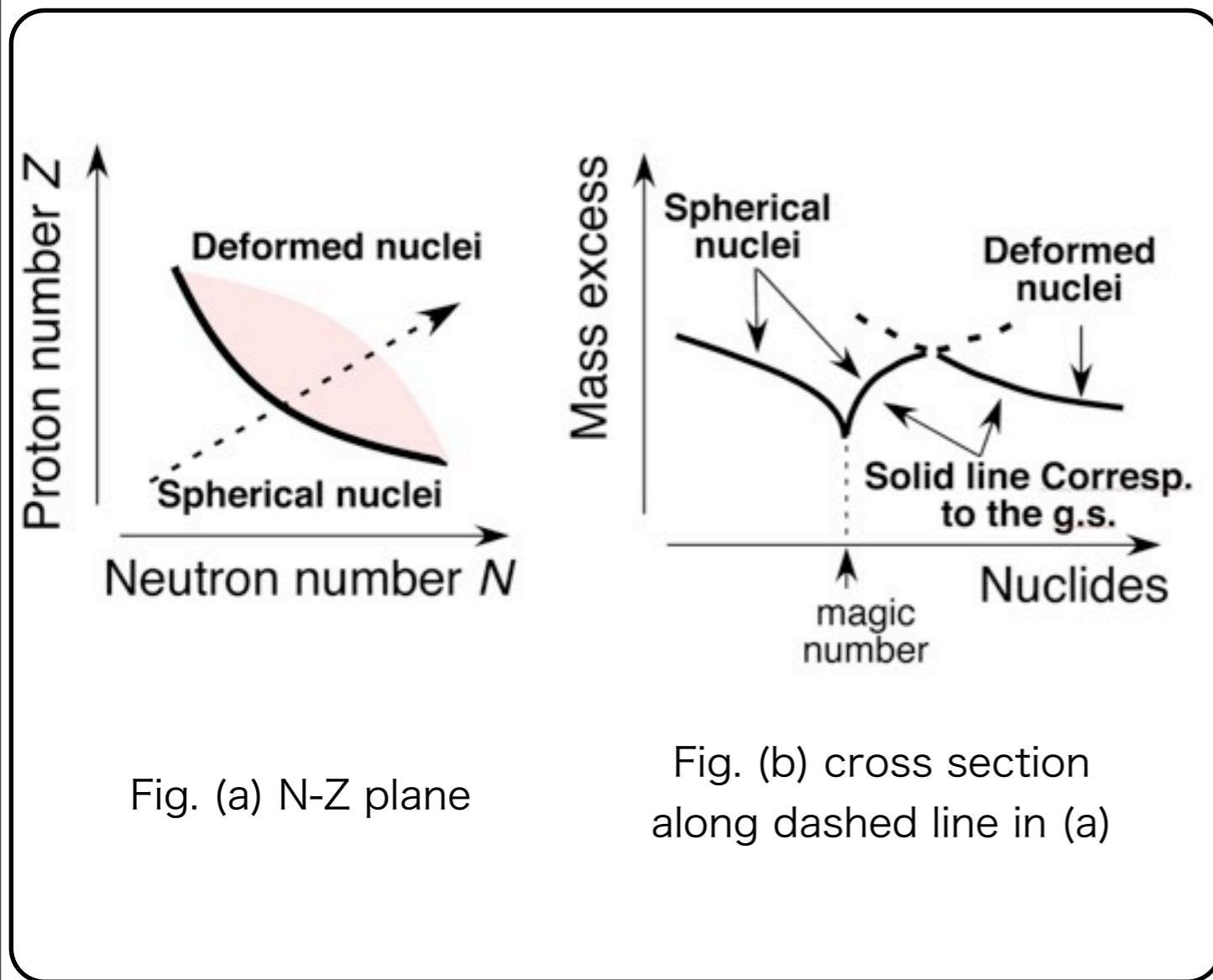
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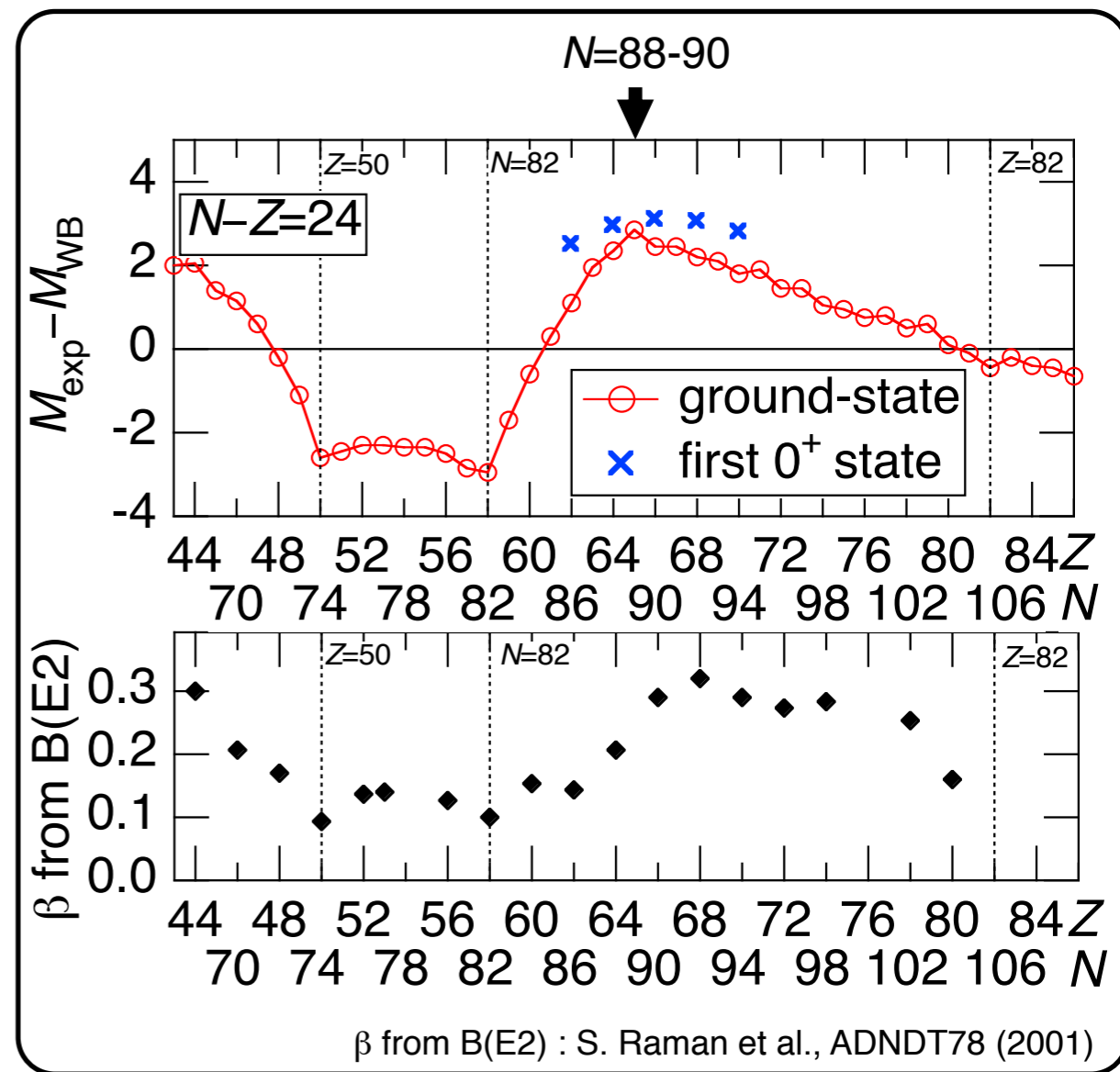
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Trend in MeV-order

- Schematic -



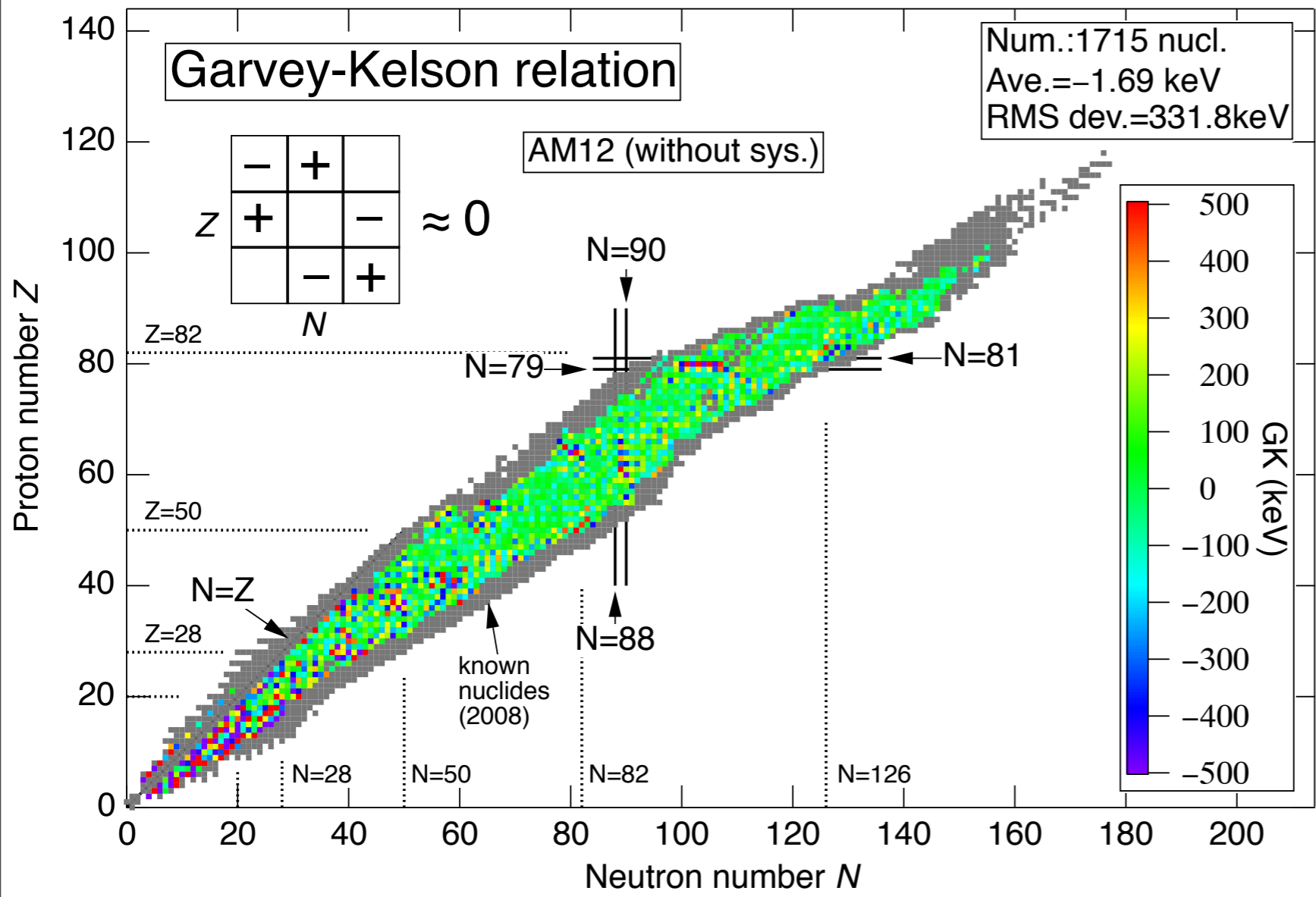
- Experiment -



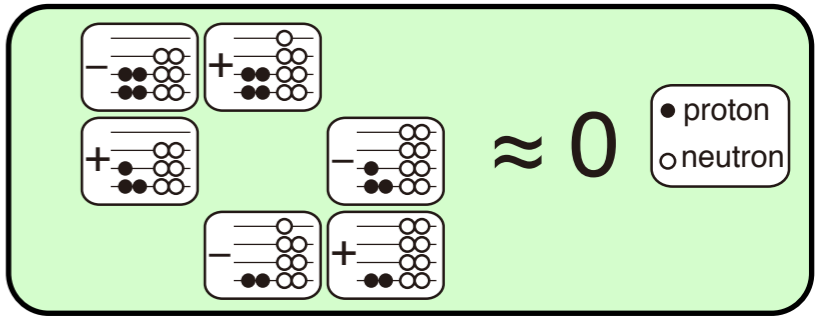
Notable feature on discontinuity of derivative of mass values

- $Z=50$, $N=82$ and $Z=82$ discontinuity of derivative: Spherical single-particle shell closure
- $N=88-90$ discontinuity: Shape transition

Mass relation: Garvey-Kelson systematics



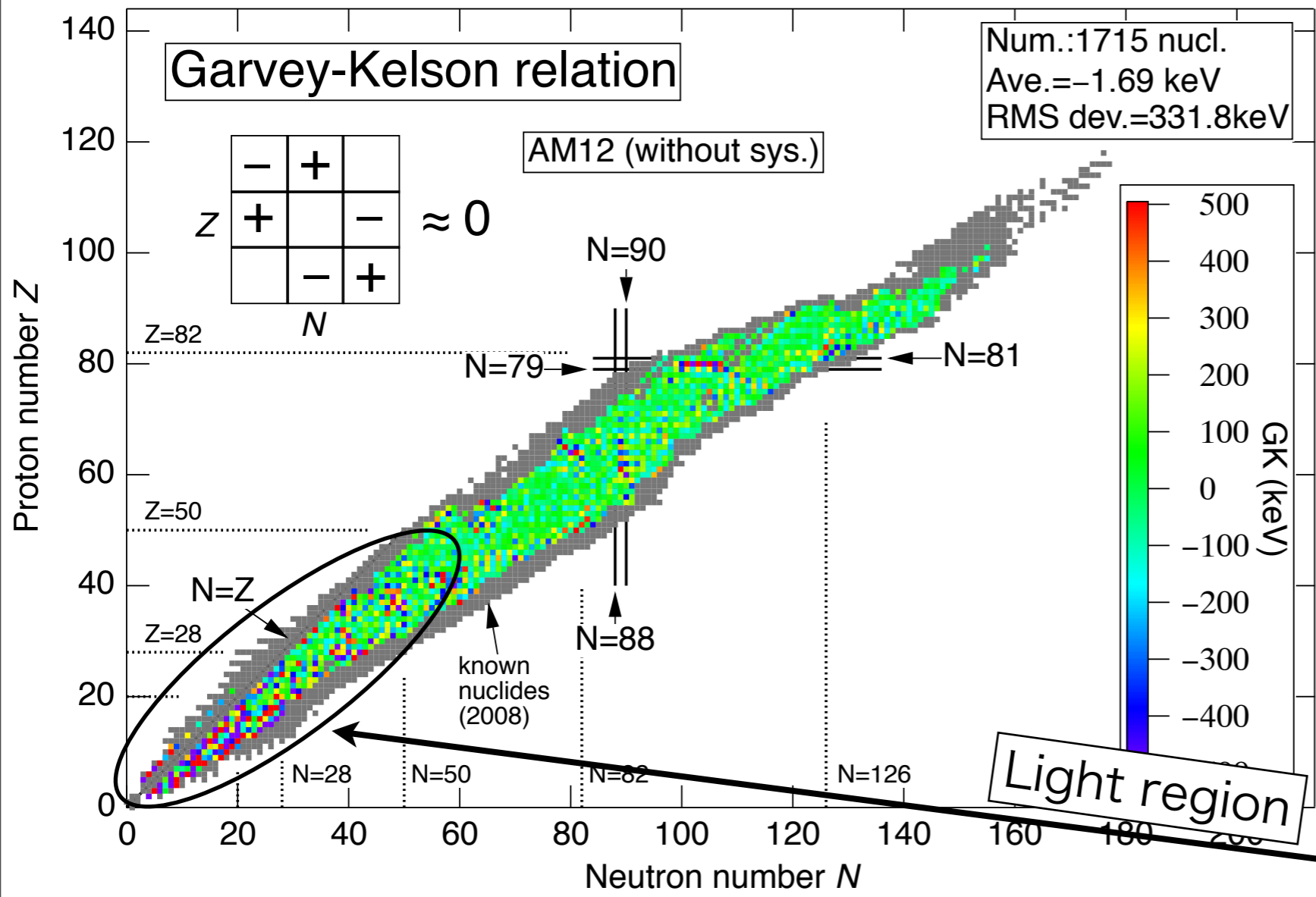
A consideration of cancellation of core + valence nucleons (based on the shell model)



• Assumption: Cores among related six nuclei are the same.

Region	num.	Average	RMS dev.
All	1715	-1.7 (keV)	331.8(keV)
$A > 100$	1202	-0.03	161.2
$A \leq 100$	513	-29.3	554.1

Mass relation: Garvey-Kelson systematics



A consideration of cancellation of core + valence nucleons (based on the shell model)

≈ 0

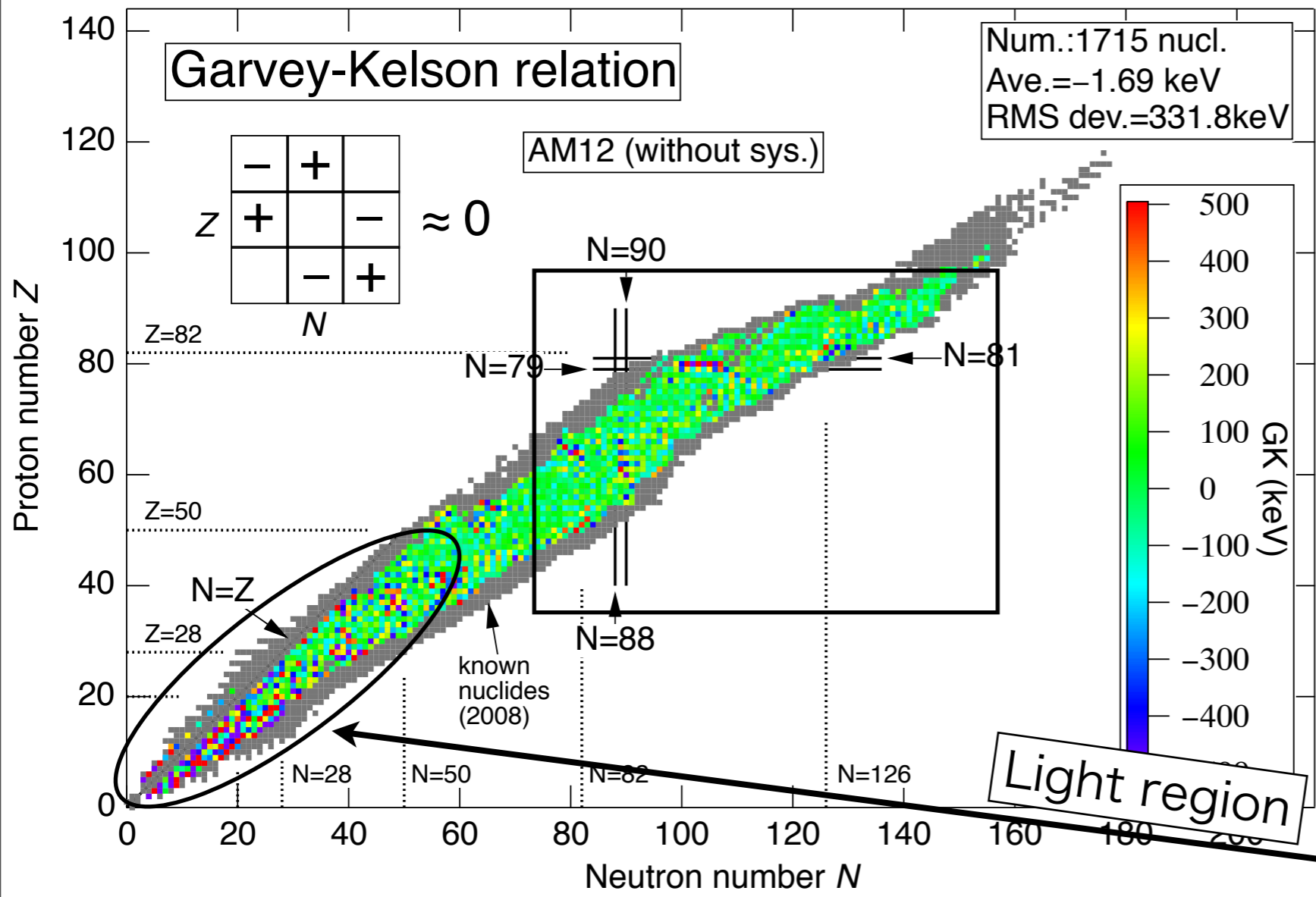
● proton
○ neutron

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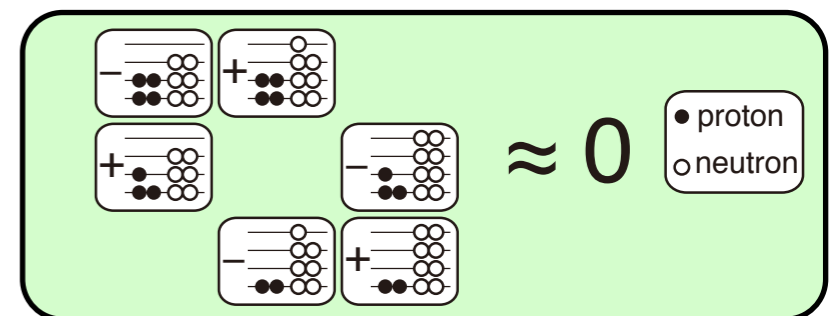
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larger

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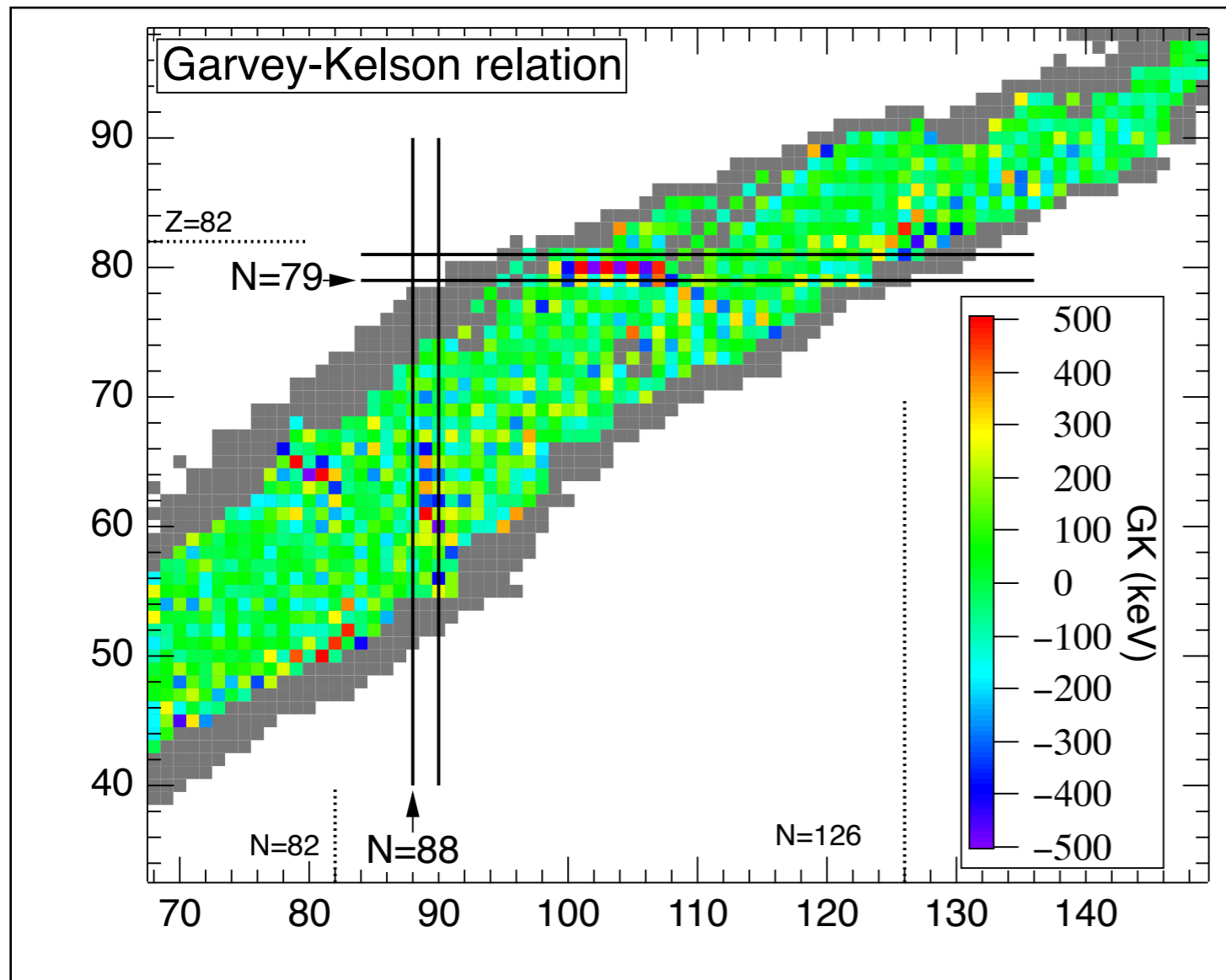
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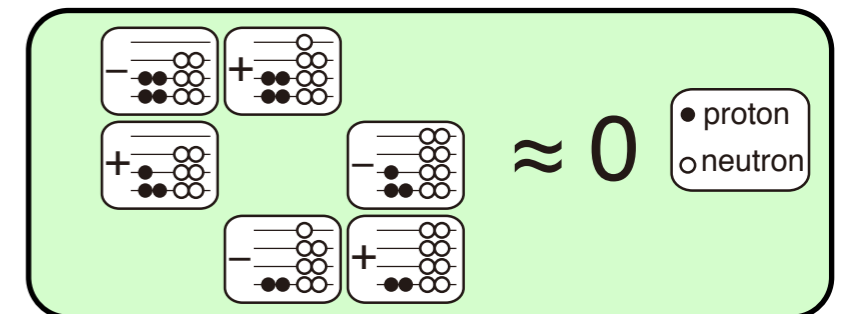
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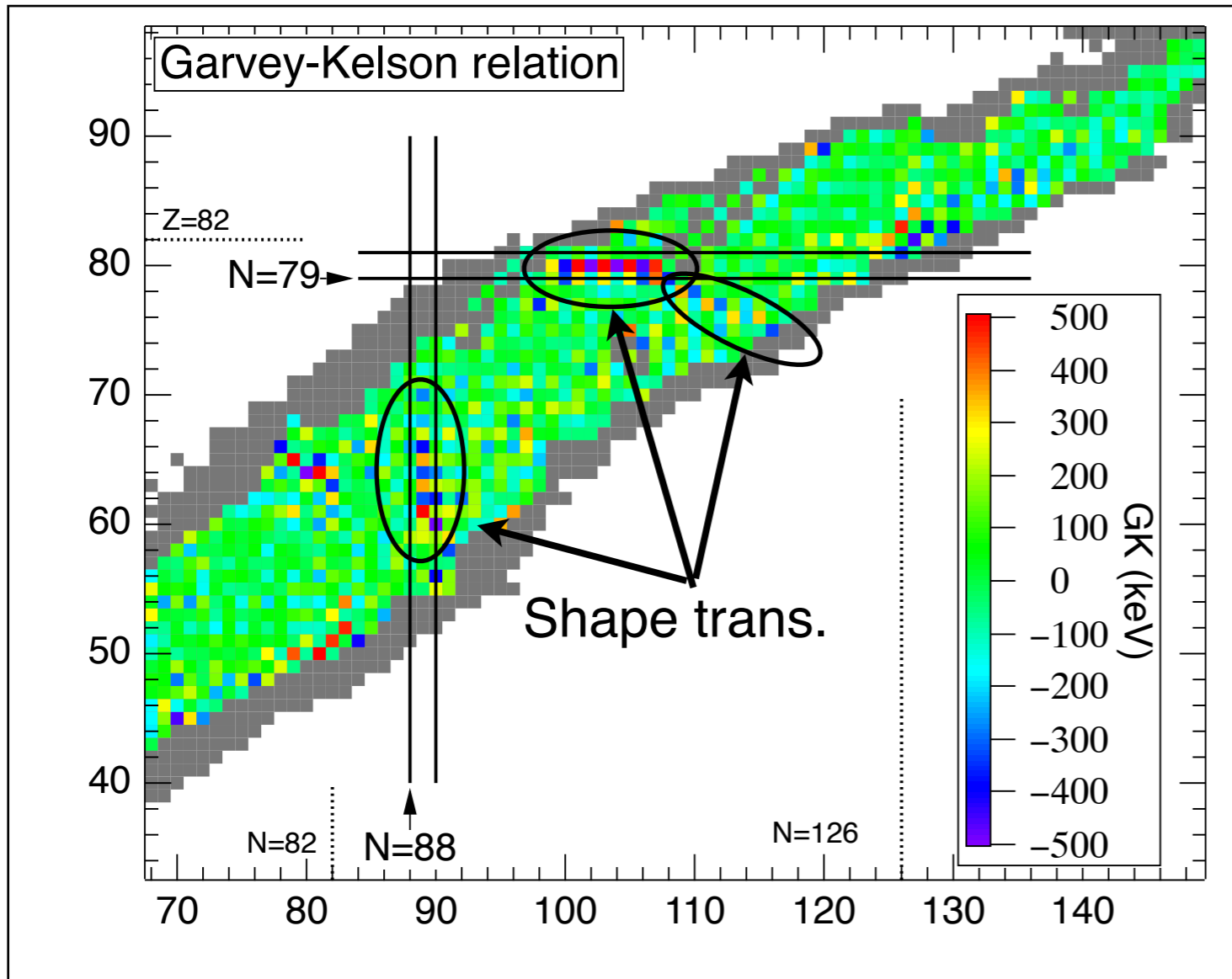


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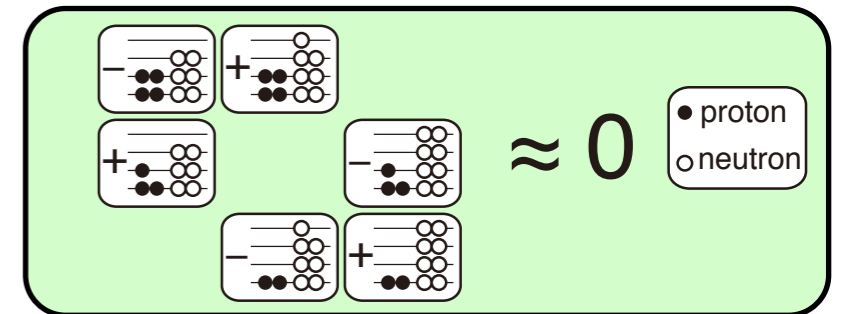
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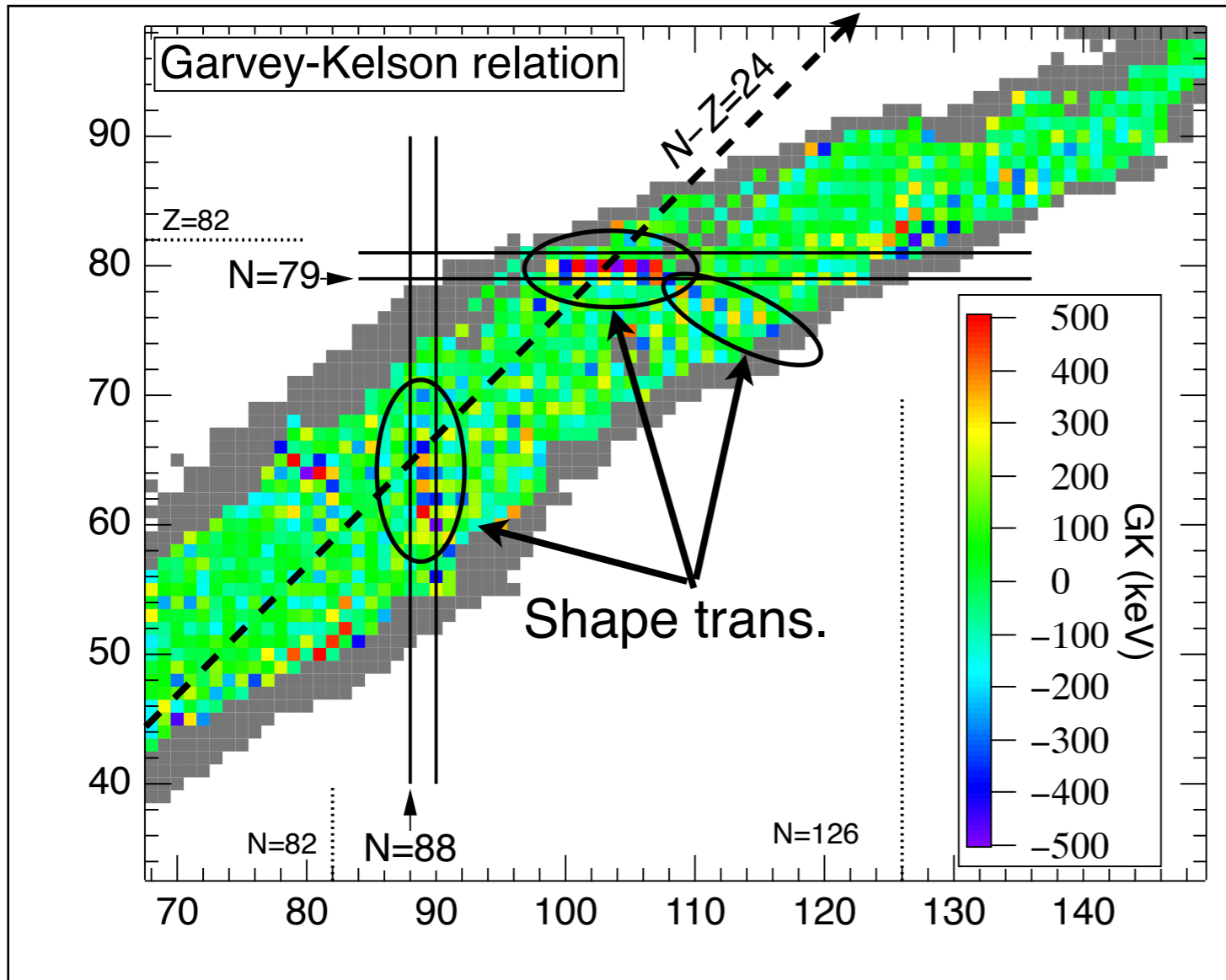


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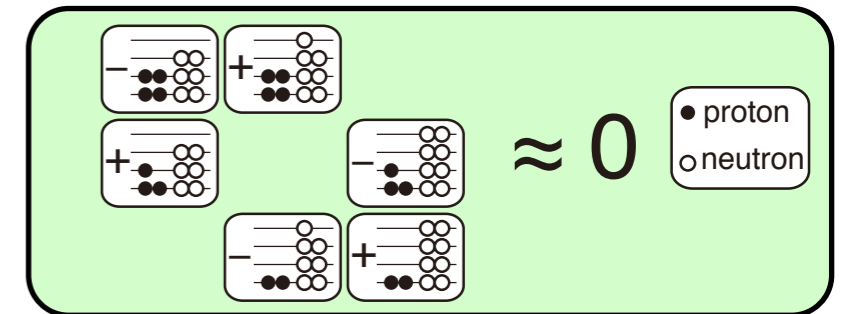
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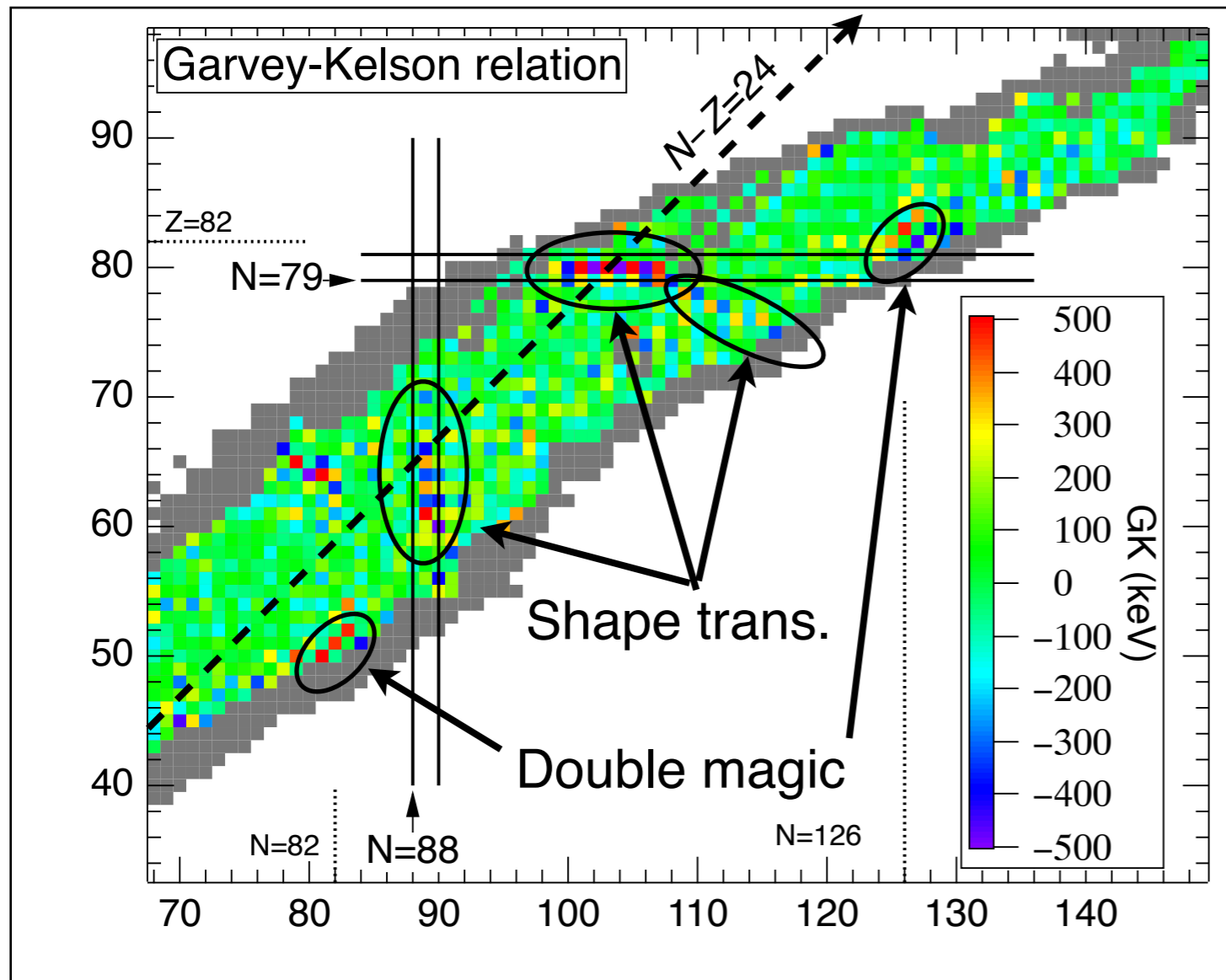
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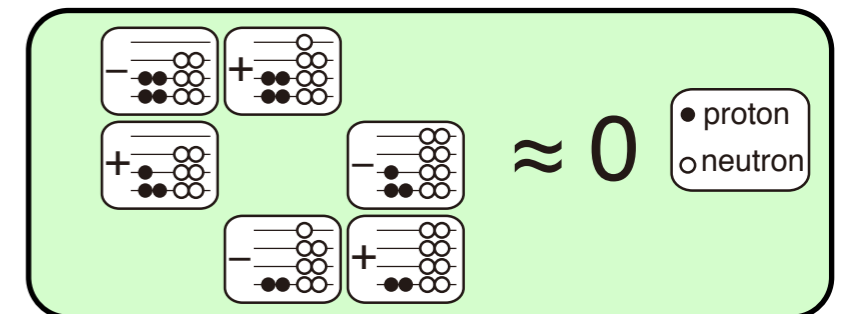
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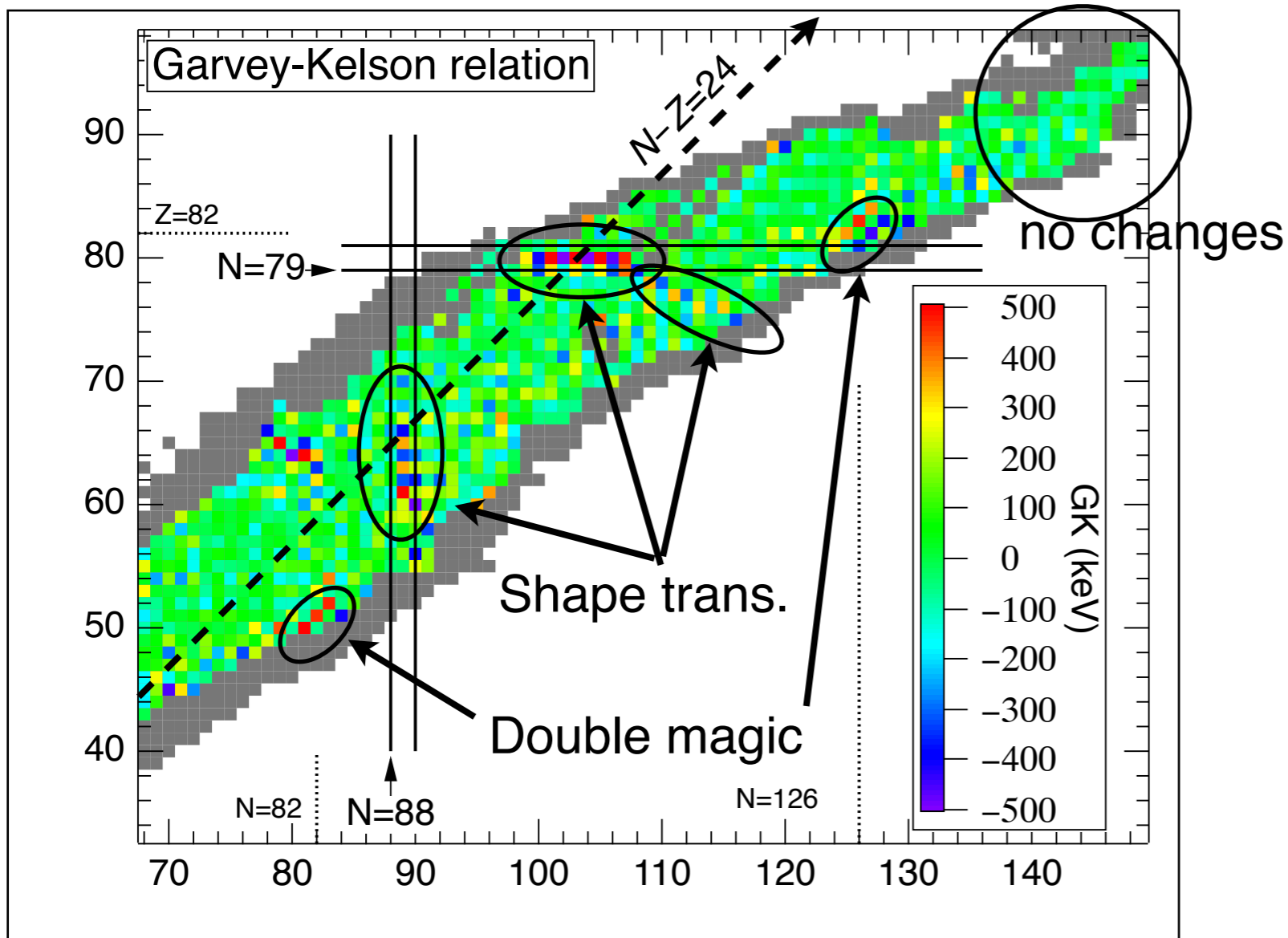


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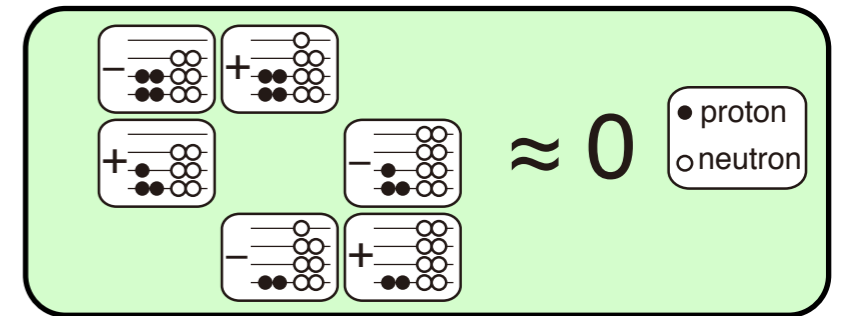
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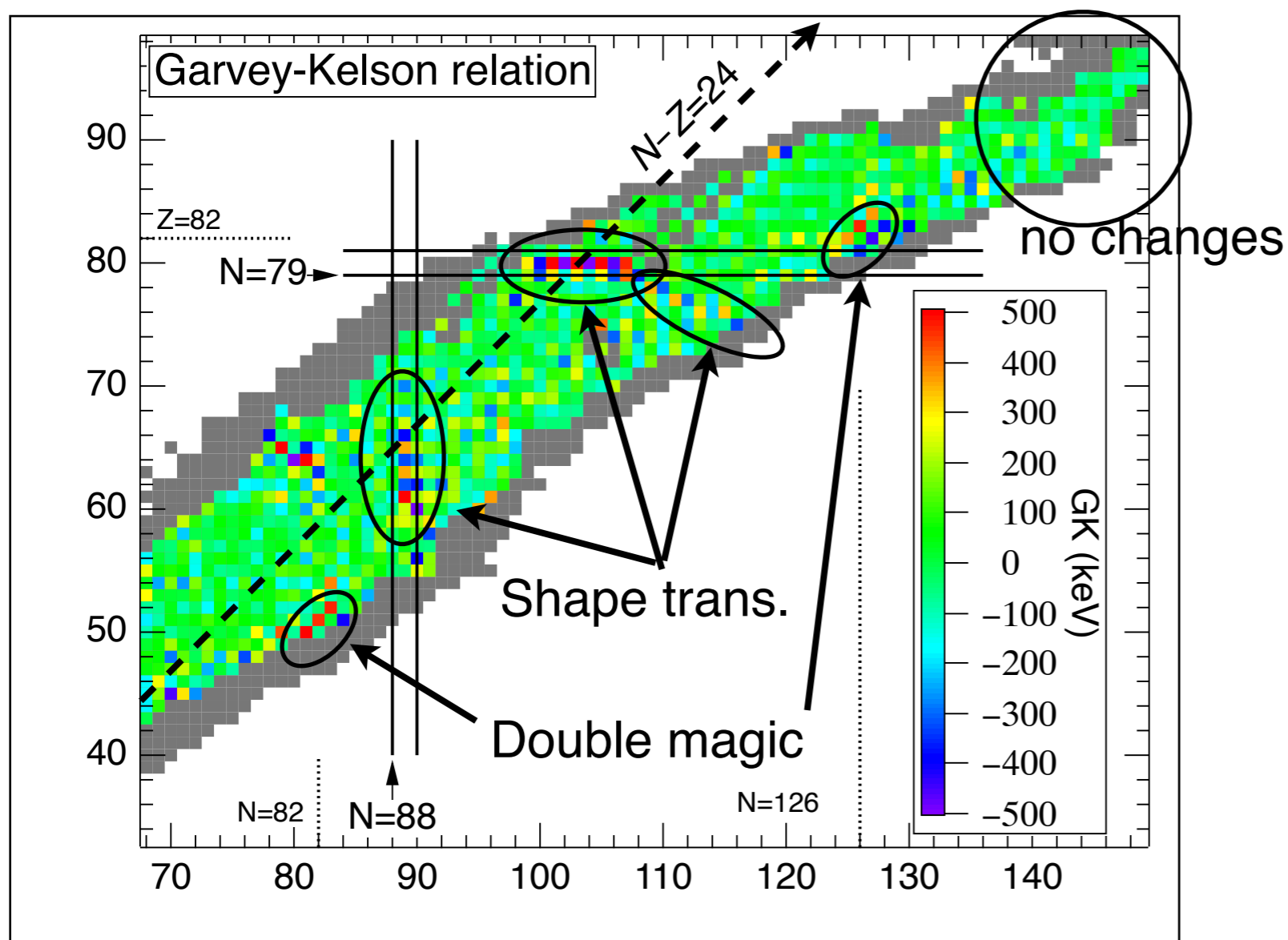
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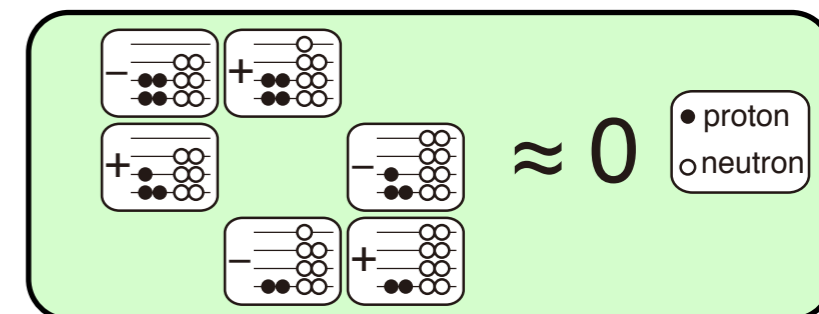
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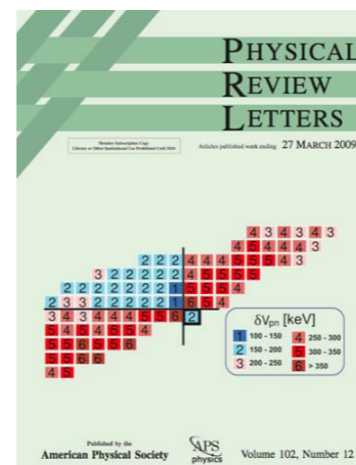
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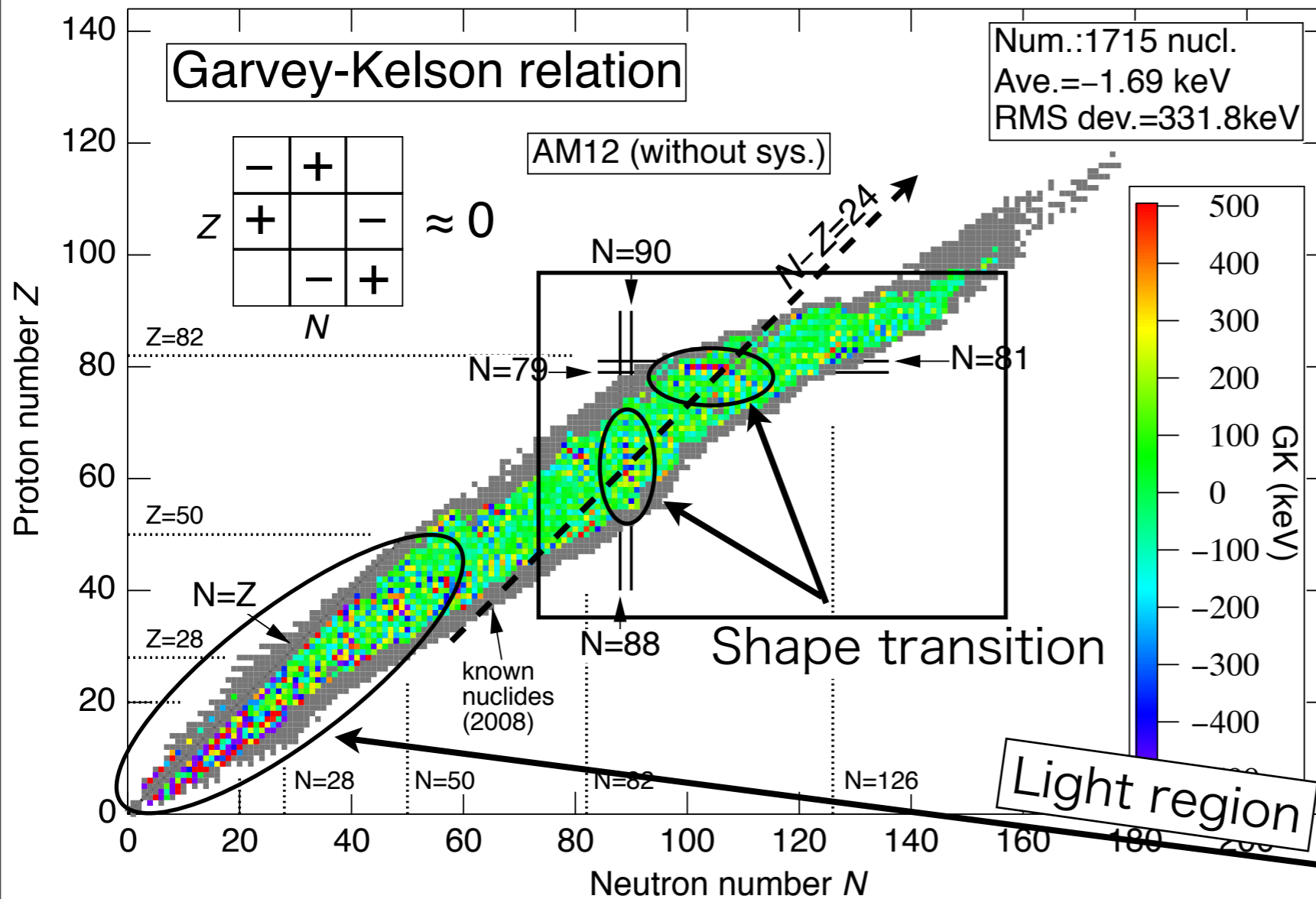
larger

cf. Study for pairing and proton-neutron interaction: neighboring doubly-magic nuclei.

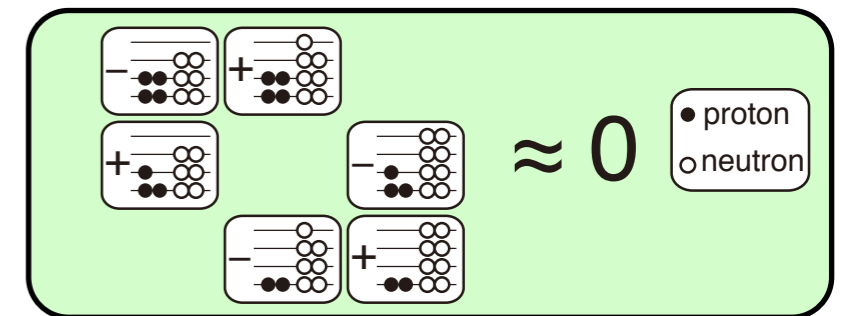
Systematical trend of average p-n interaction crossing N=126.



PRL 102, 122503(2009)



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larger

GK type mass formula:
 Comay-Kelson-Zidon (CKZ₁₉₈₈)
 Jänecke-Masson (JM₁₉₈₈)
 Masson-Jänecke (MJ₁₉₈₈)
 (ADNDT39, 1988)

In heavy region, the GK sys. gives some transition as changes of structures, especially nuclear shape.

- **Bulk properties of mass surface:**

- Volume energy, surface energy, symmetry energy, ...

- **Shell gaps:**

- $N, Z=20, 28, 50, 82, 126$ (only N) and a change of magicities (ex. $N=14$ to 16)

- **Transition of sphere to deformation:**

- Discontinuity of derivatives at $N=88$ to 90 near the β -stable region.

- **Wigner term:**

- Discontinuity at $N=Z$.

- **Averaged even-odd effect:**

- Staggering change of masses alternates even and odd- N/Z .

There are many and various mass models.

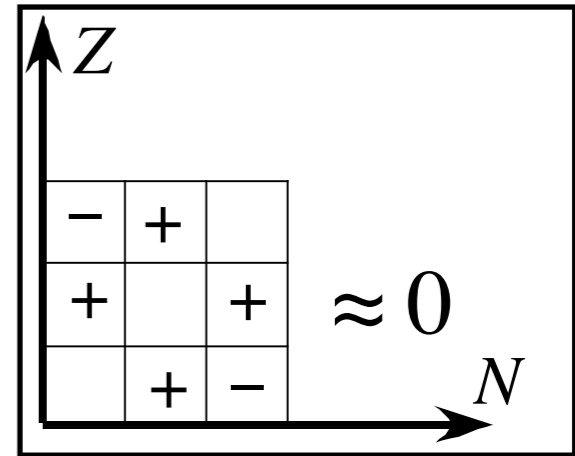
- **Systematics:**
 - Construction by focusing mass relation
- **Mass Model, Approximation:**
 - Macro-micro, Hybrid, or micro-like framework

Only mass data available to obtain are adopted.
(RMF, EDF mass formula are not included.)

- **Garvey-Kelson-type mass systematics**

focusing on relation between mass values and Z, N

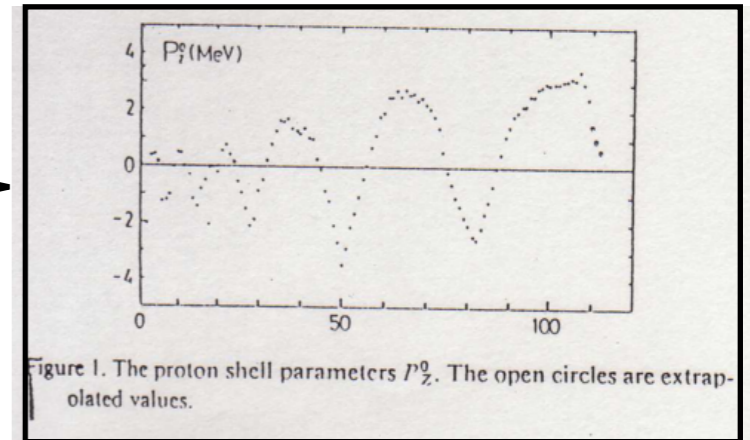
Comay-Kelson-Zidon, Jänecke-Masson (1988)



- **Empirical shell term**

focusing on Bulk part (WB-like)+deviation (Shell term) →

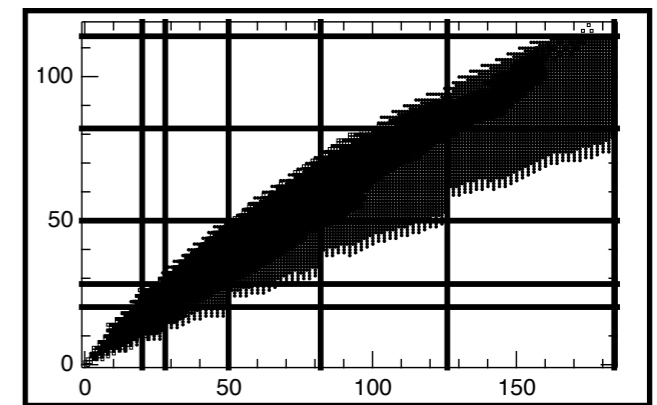
Tachibana-Uno-Yamada-Yamada (1988)



- **Phenomenological shell model calculation**

Polynomials of particle and hole numbers, obliged to assume magic numbers in advance.

Liran-Zeldes (1976), Duflo-Zuker (1995)



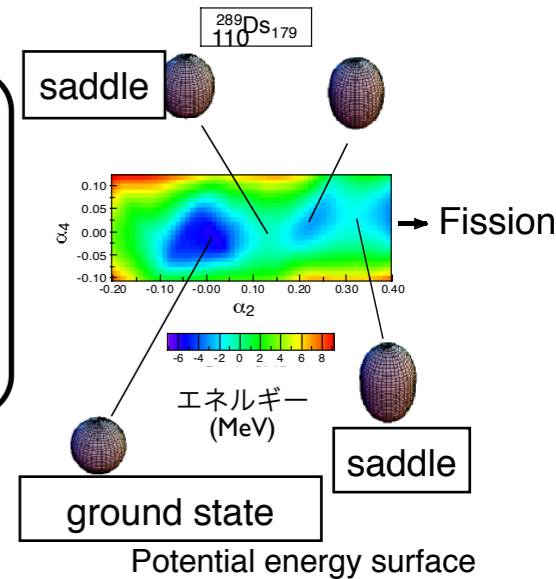
• ...

Properties

- Good reproduction of masses for known nuclei + good prediction for unknown nuclei (quite near mass-measured nuclides. (300-600 keV))
- No predictable power for superheavy nuclei (next magic number, etc.)
- No deformation is obtained.

Recent mass formulas:

- are designed for nuclei with $Z, N=8$ to $^{310}[126]_{184}$ or more
- have the RMS dev. from exp. masses. of 600-800 keV
- give deformation parameters $\beta_2, \beta_4...$ and fission barriers



• **Density functional theory** <- recent project

• **Hartree-Fock method with Skyrme force**

by Dobaczewski et al.

Strong short-range force $\Rightarrow \delta$ -function \Rightarrow HF calc.
[ETFSI](#) (1995), [HFBCS](#) (2001), [HFB](#) (2002-)

• **Liquid-drop model**

by S.Goriely et al.

Deformed liquid-drop part+Micro. (folded Yukawa)
[FRDM](#) (1995), [FRLDM](#) (2002),

• **Mass formula with spherical-basis shell term**

by P.Möller et al.

Phenom. gross (WB-like)+spherical-basis shell part
[KUTY](#) (2000), [KTUY](#) (2005) Koura, Uno, Tachibana, Yamada

micro (-like)
 macro+micro
 phenom.

Skyrme-Hartree-Fock-Bogoliubov mass formula (2002-2010)

by S. Goriely et al.

$$E_{\text{tot}} = E_{\text{HFB}} + E_{\text{wigner}}$$

Bsk21 force parameter set:

$t_0 = -3961.39 \text{ MeV fm}^3$, $t_1 = 396.131 \text{ MeV fm}^5$
 $t_2 = 0 \text{ MeV fm}^5$, $t_3 = 22588.2 \text{ MeV fm}^{3+3\alpha}$
 $t_4 = -100.000 \text{ MeV fm}^{5+3\beta}$, $t_5 = -150.000 \text{ MeV fm}^{5+3\gamma}$
 $x_0 = 0.885231$, $x_1 = -0.0648452$, $t_2 x_2 = 1390.38 \text{ MeV fm}^5$
 $x_3 = 1.03928$, $x_4 = 2.00000$, $x_5 = -11.0000$
 $W_0 = 109.622 \text{ MeV fm}^5$, $\alpha = 1/12$, $\beta = 1/2$, $\gamma = 1/12$
 $f_n^+ = 1.00$, $f_p^+ = 1.07$, $f_n^- = 1.05$, $f_p^- = 1.13$
 $V_W = -1.80 \text{ MeV}$, $\lambda = 280$, $V'_W = 0.96$, $A_0 = 24$

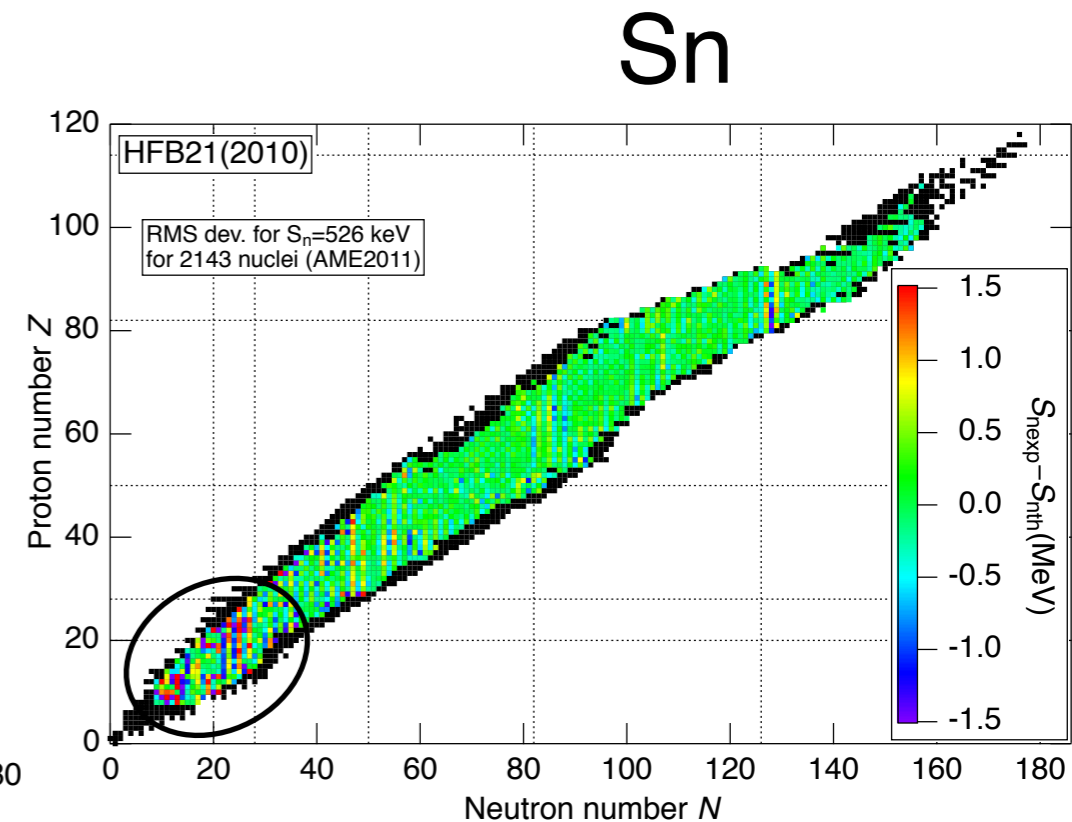
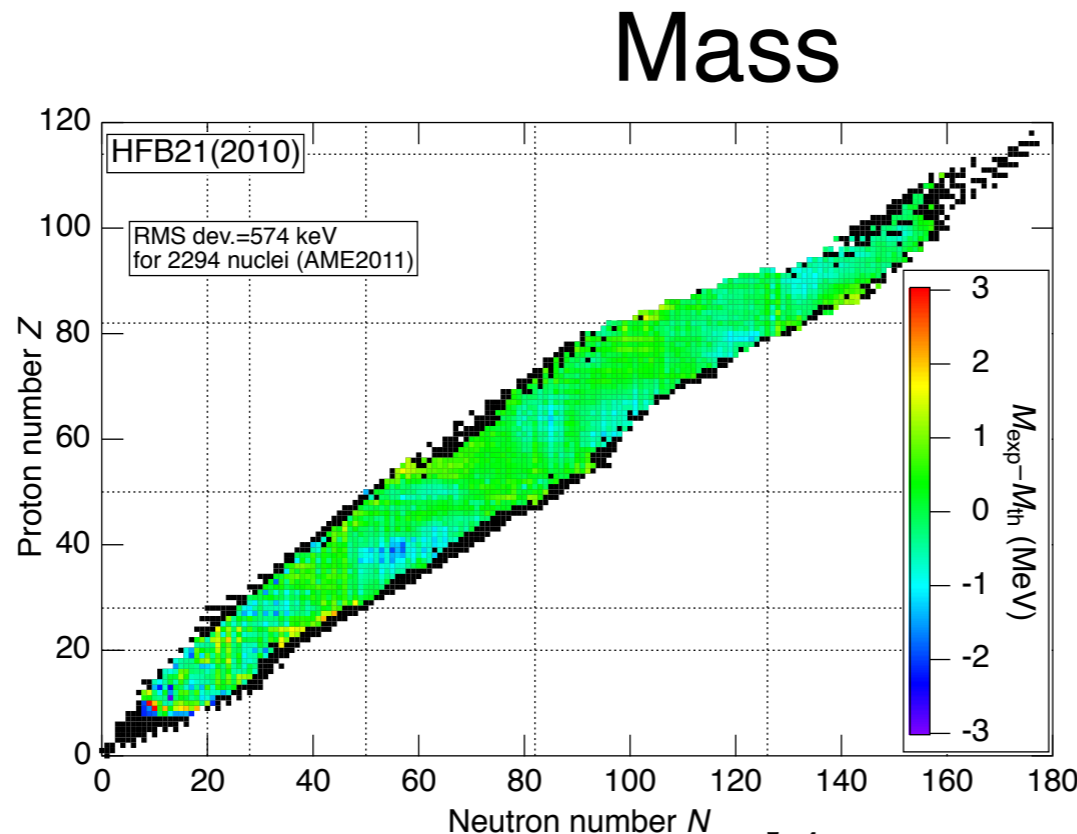
The long road in the HFB mass model development

		Accuracy σ_{rms} (2149 nuc)
HFB-2:	Possible to fit all 2149 exp masses $Z \geq 8$	659 keV
HFB-3:	Volume versus surface pairing	635 keV
HFB-4-5:	Nuclear matter EoS: $M_s^* = 0.92$	660 keV
HFB-6-7:	Nuclear matter EoS: $M_s^* = 0.80$	657 keV
HFB-8:	Particle-number projection	635 keV
HFB-9:	Neutron matter EoS	733 keV
HFB-10-13:	Low pairing & NLD	717 keV
HFB-14:	Collective correction and Fission B_f	729 keV
HFB-15:	Coulomb correlations / CSB	678 keV
HFB-16:	Pairing constrained to NM	632 keV

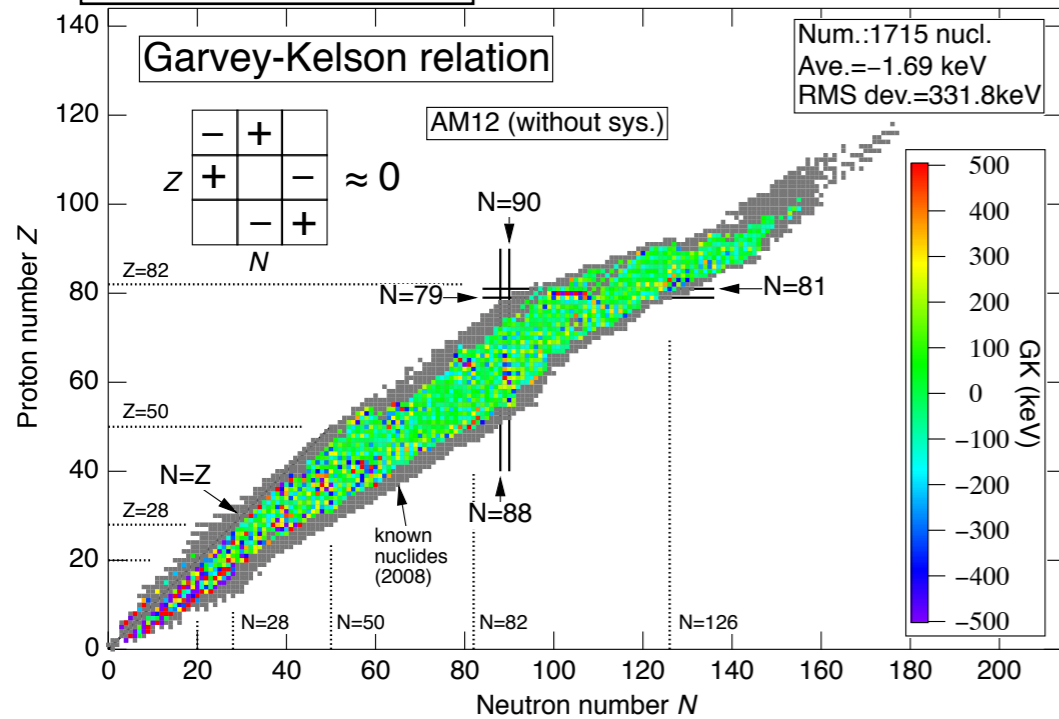
Current version: HFB-21 (2010)

HFB21 gives a less than 600 keV of the RMS dev. In the light region there is some discrepancy in derivatives as S_n .

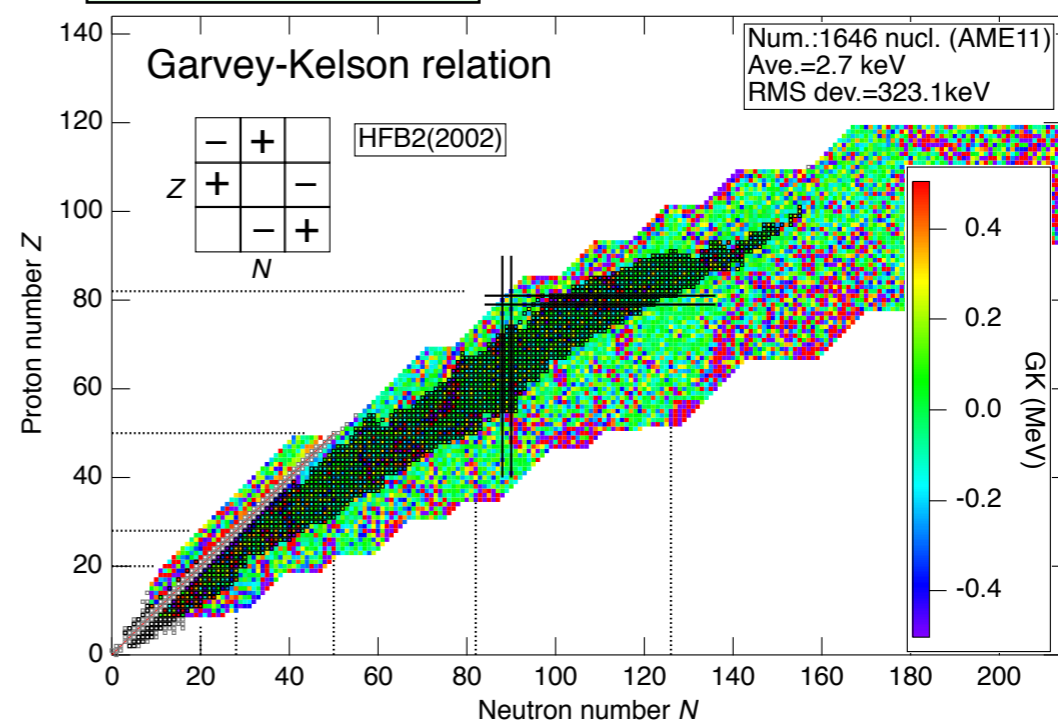
Referred mass data: AME03



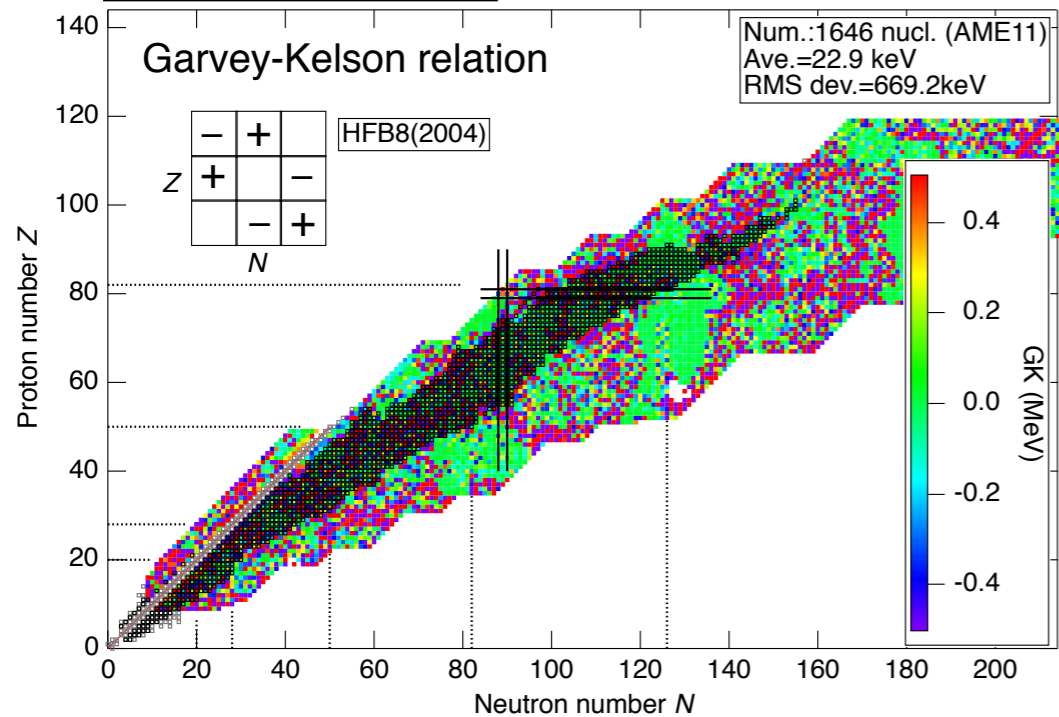
AME11(exp)



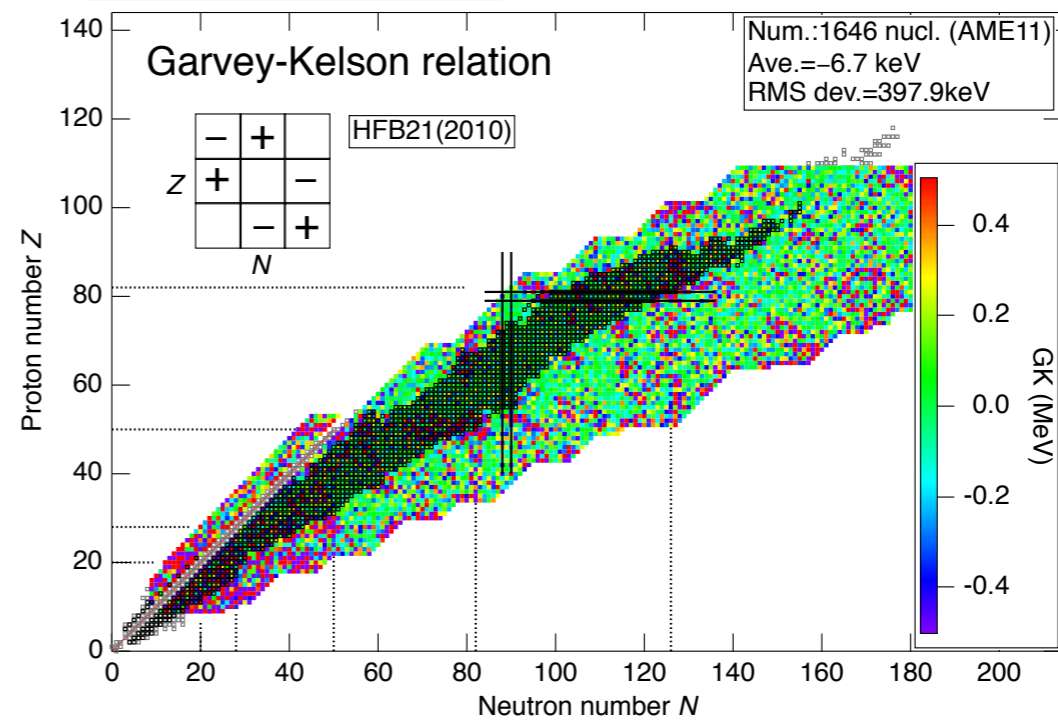
HFB2(2002)



HFB8(2004)



HFB21(2010)



- All the HFB calc. violate the GK summation everywhere!

Finite-Range-Droplet Model (FRDM) mass formula (1995)

by P. Möller et al.

Current version is FRLDM (2003-)

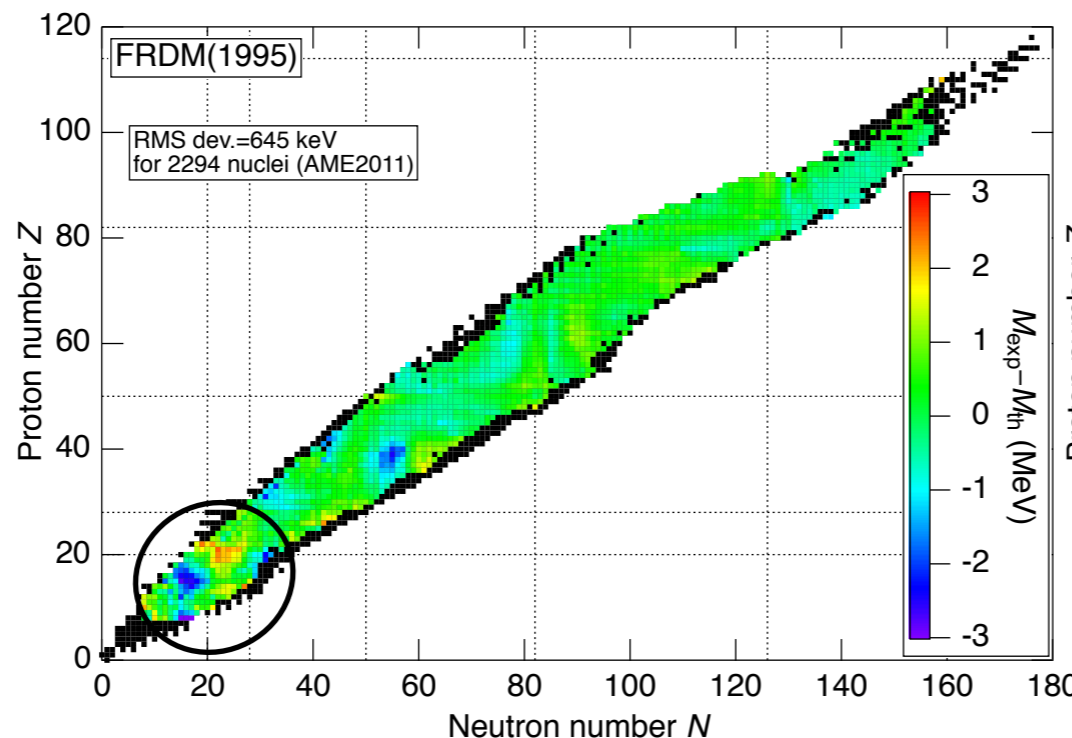
$$E(Z, N, \text{shape}) = E_{\text{macro}}(Z, N, \text{shape}) + E_{\text{micro}}(Z, N, \text{shape})$$

E_{macro} : Droplet part as a function of Z and N

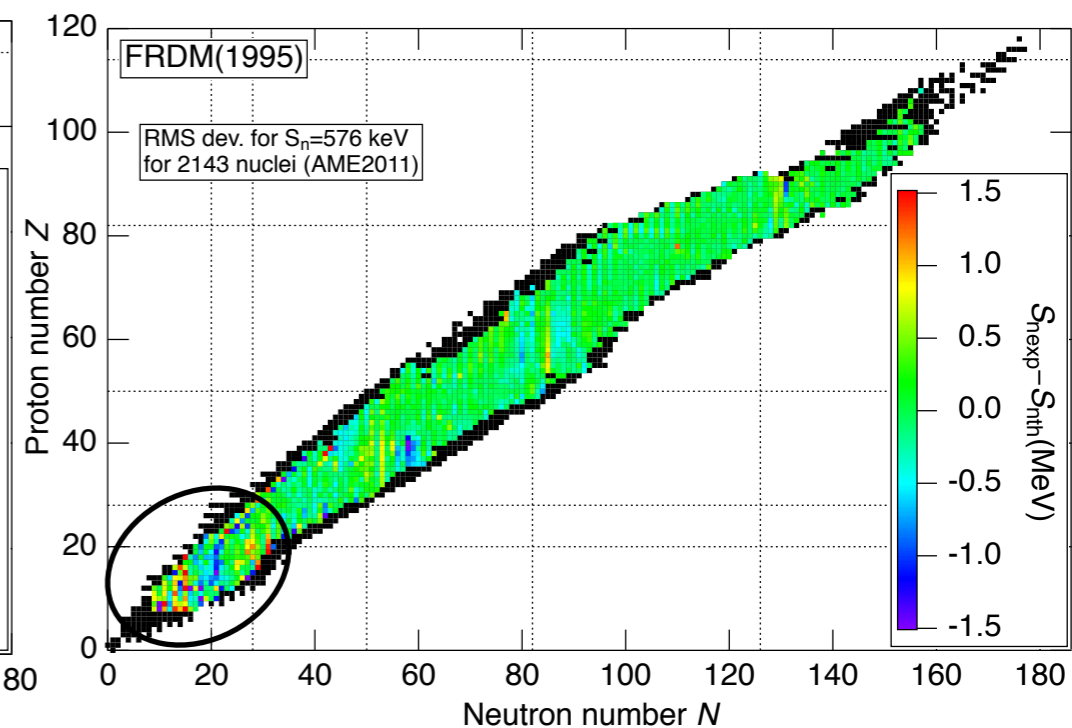
E_{micro} : Folded Yukawa-type potential + Nilsson-Strutinsky method

- Deformation, fission barrier is obtained
- Good prediction on fission properties.

Mass



Sn



Good for the heavier mass region. Some large discrepancies appear in the light region.

Referred mass data: AME93

Spherical-Basis (KTUY) mass formula (2005)

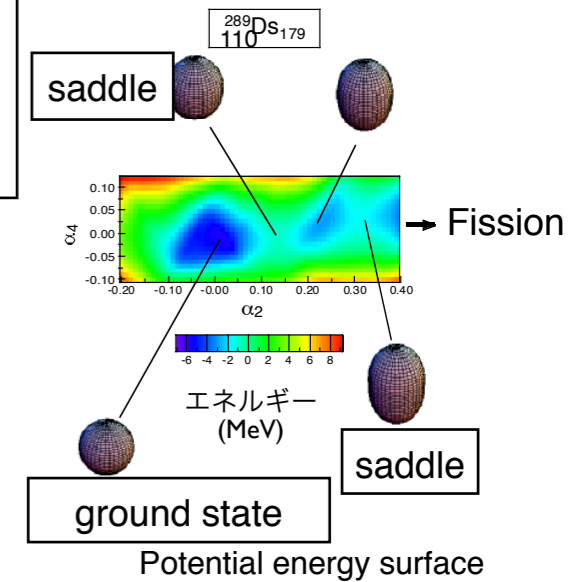
$$M(Z, N) = M_{\text{gross}}(Z, N) + M_{\text{eo}}(Z, N) + M_{\text{shell}}(Z, N)$$

by H. Koura et al.,
PTP113 (2005)

M_{gross} smooth function of N and Z. (same as the TUIY formula)

M_{shell} : modified Woods-Saxon pot.+BCS+deform. config.

- Deformation, fission barrier is obtained
- Change of magicities in the n-rich nuclei is predicted. (N=20 → 16, etc.)
- Topic: decay modes for superheavy nuclei are applied for.



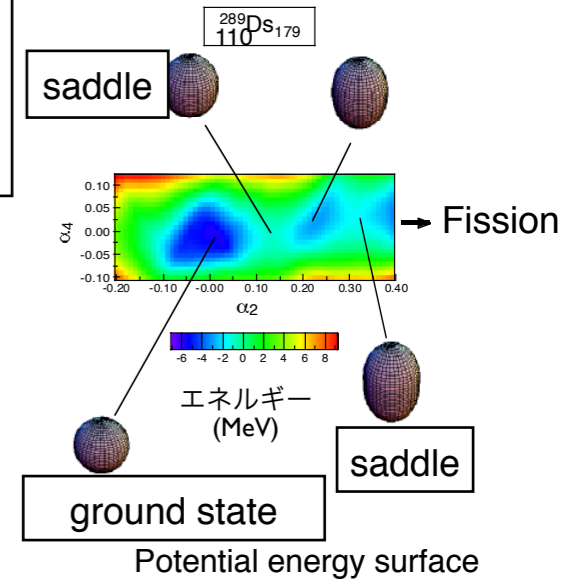
$$M(Z, N) = M_{\text{gross}}(Z, N) + M_{\text{eo}}(Z, N) + M_{\text{shell}}(Z, N)$$

by H. Koura et al.,
PTP113 (2005)

M_{gross} smooth function of N and Z. (same as the TUYU formula)

M_{shell} : modified Woods-Saxon pot.+BCS+deform. config.

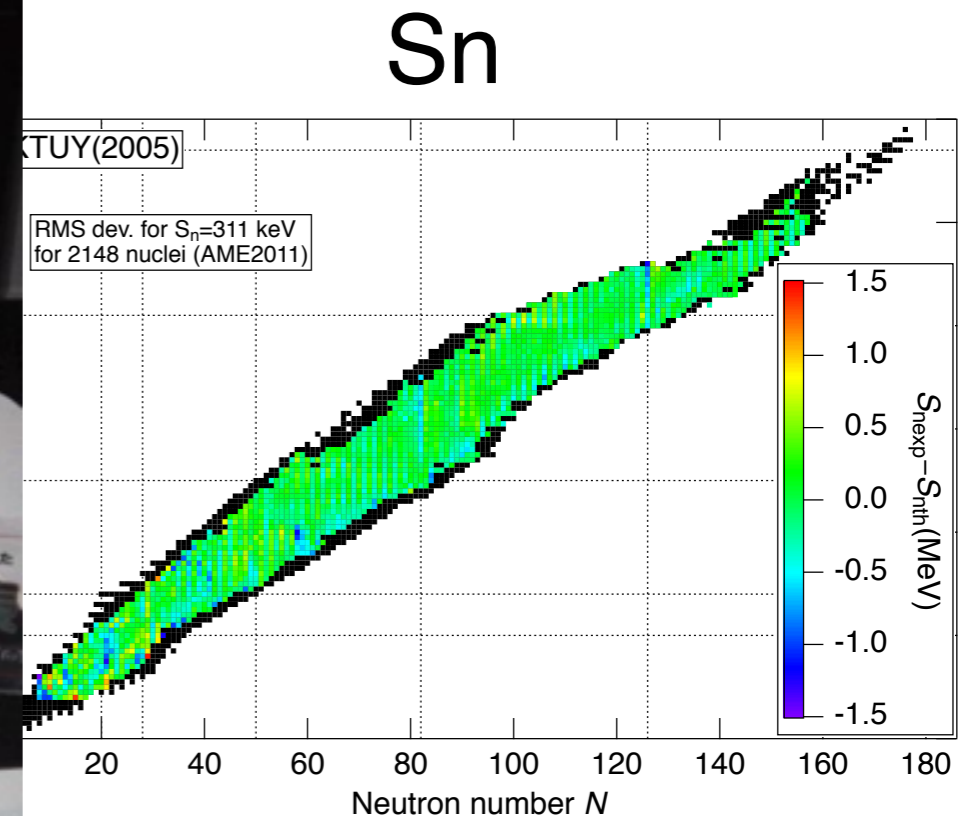
- Deformation, fission barrier is obtained
- Change of magicities in the n-rich nuclei is predicted. (N=20 → 16, etc.)
- Topic: decay modes for superheavy nuclei are applied for.



Derivatives of mass like S_n , Q_α , Q_β , gives a good reproduction.



Referred Mass data: AME03



Spherical-Basis (KTUY) mass formula (2005)

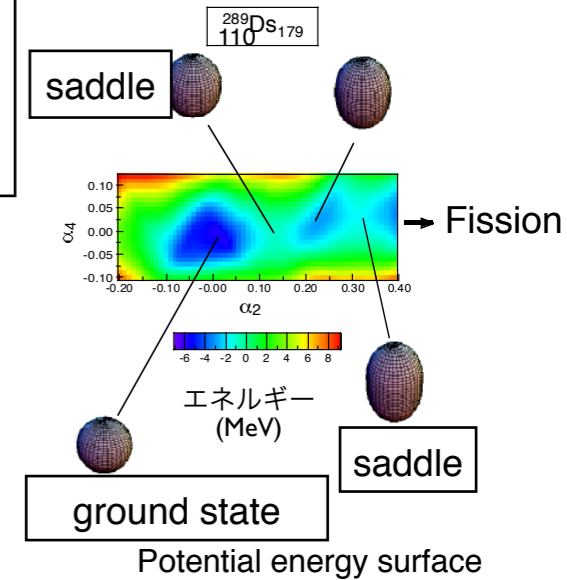
$$M(Z, N) = M_{\text{gross}}(Z, N) + M_{\text{eo}}(Z, N) + M_{\text{shell}}(Z, N)$$

by H. Koura et al.,
PTP113 (2005)

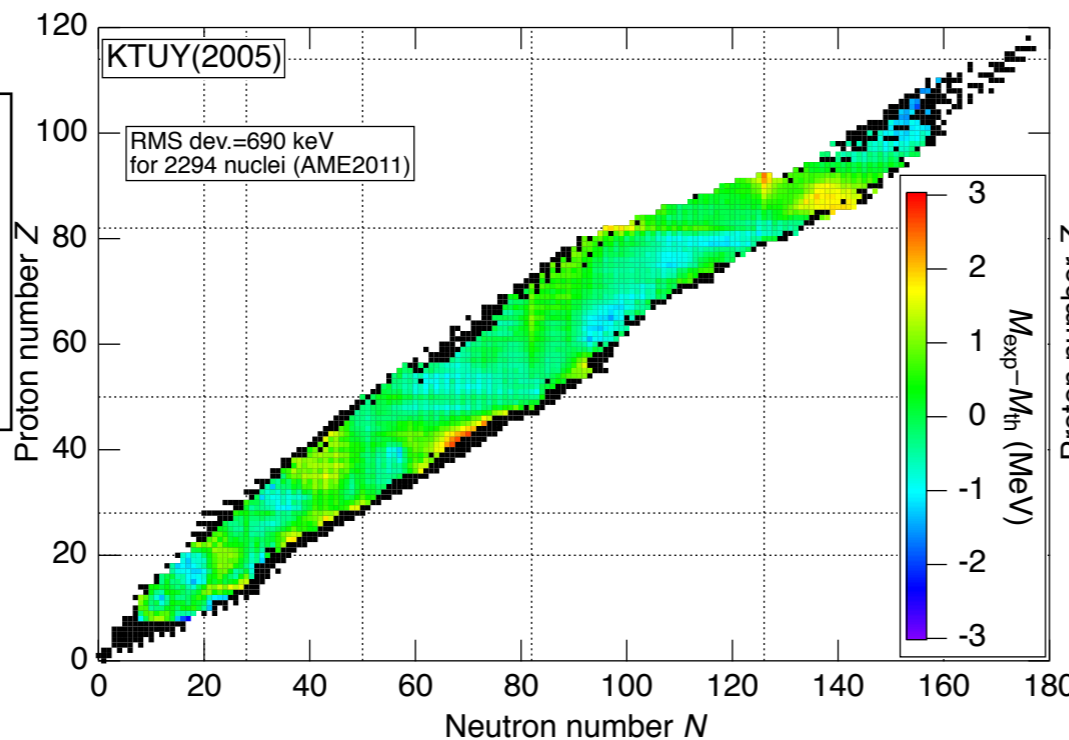
M_{gross} smooth function of N and Z. (same as the TUYU formula)

M_{shell} : modified Woods-Saxon pot.+BCS+deform. config.

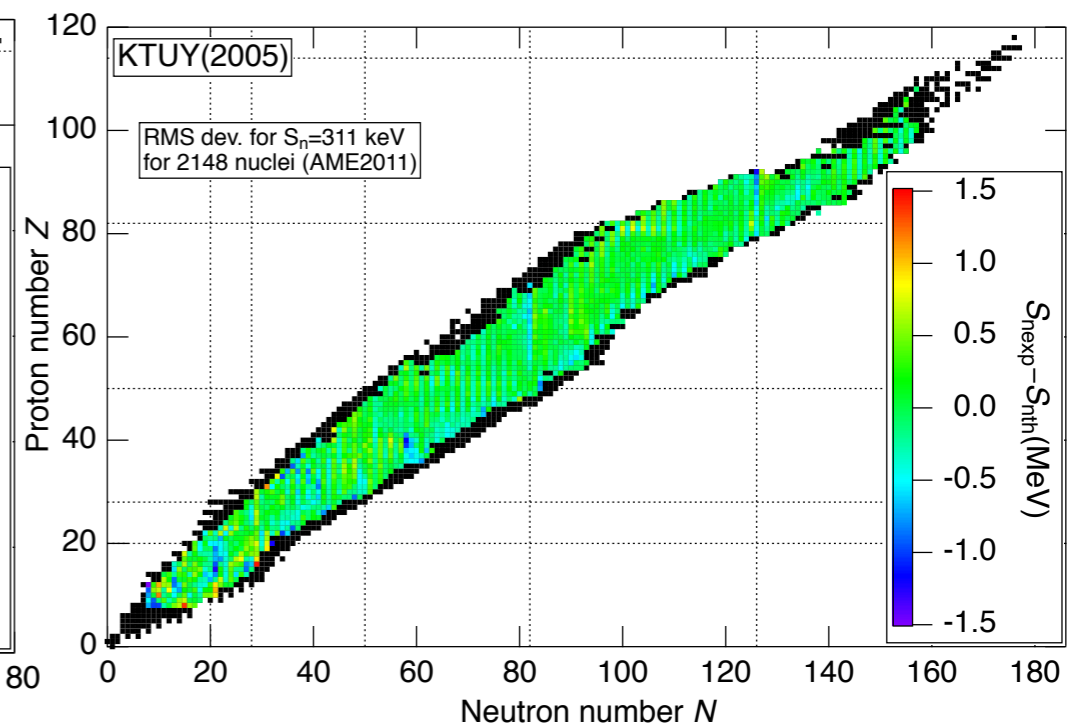
- Deformation, fission barrier is obtained
- Change of magicities in the n-rich nuclei is predicted. (N=20 → 16, etc.)
- Topic: decay modes for superheavy nuclei are applied for.



Mass



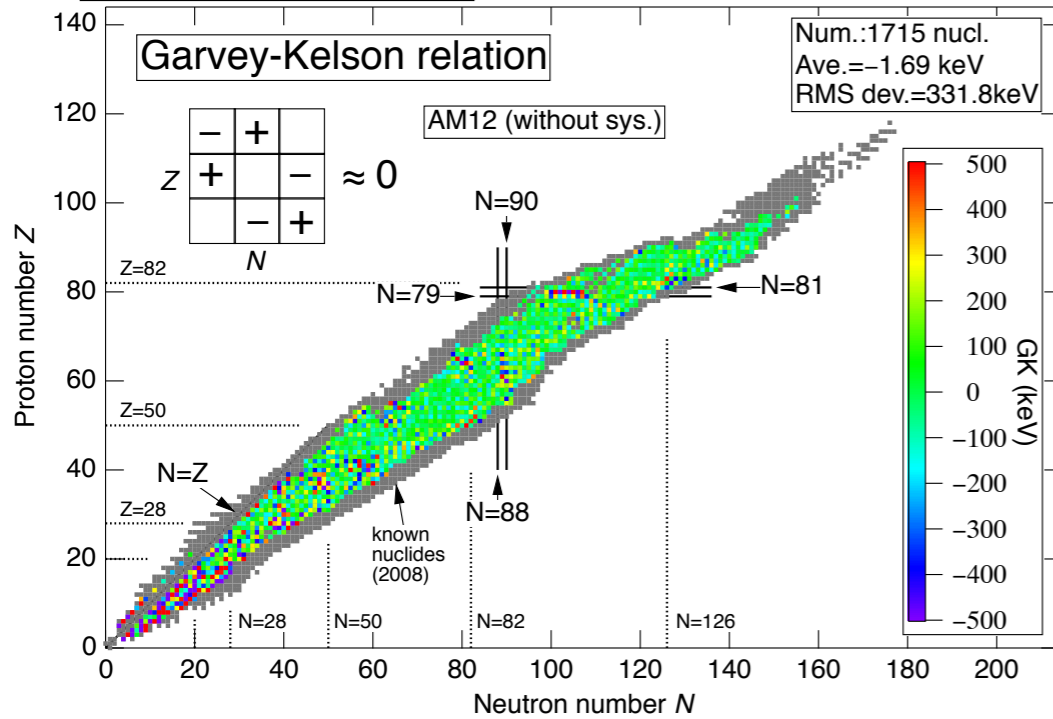
Sn



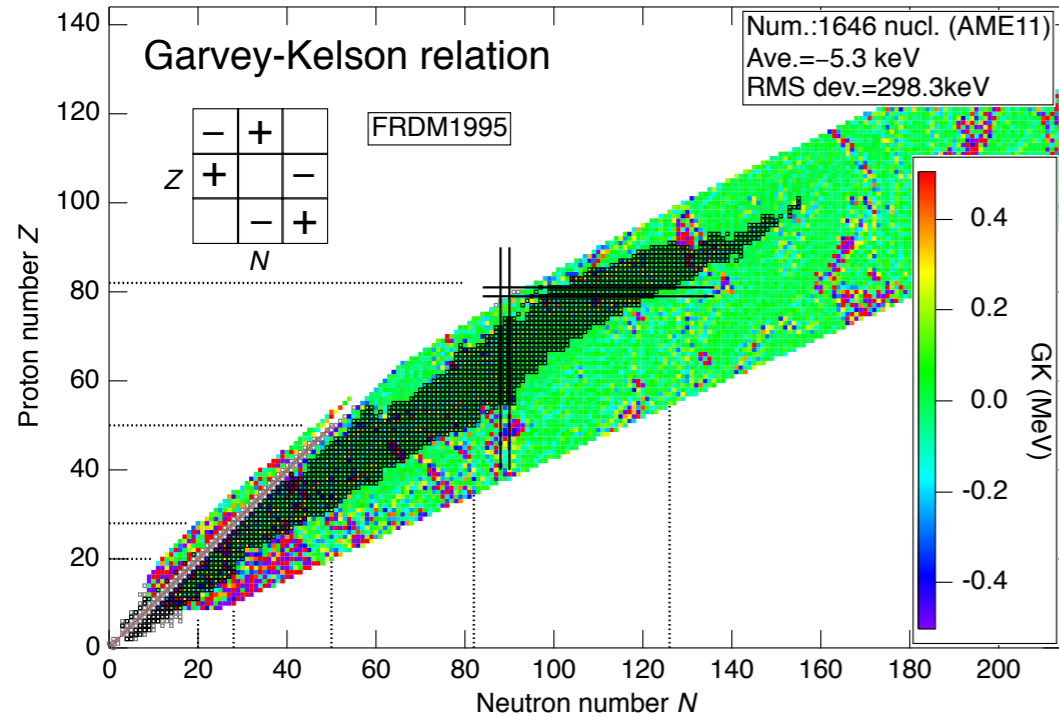
Derivatives of mass
like S_n , Q_α , Q_β ,
gives a good
reproduction.

Referred Mass
data: AME03

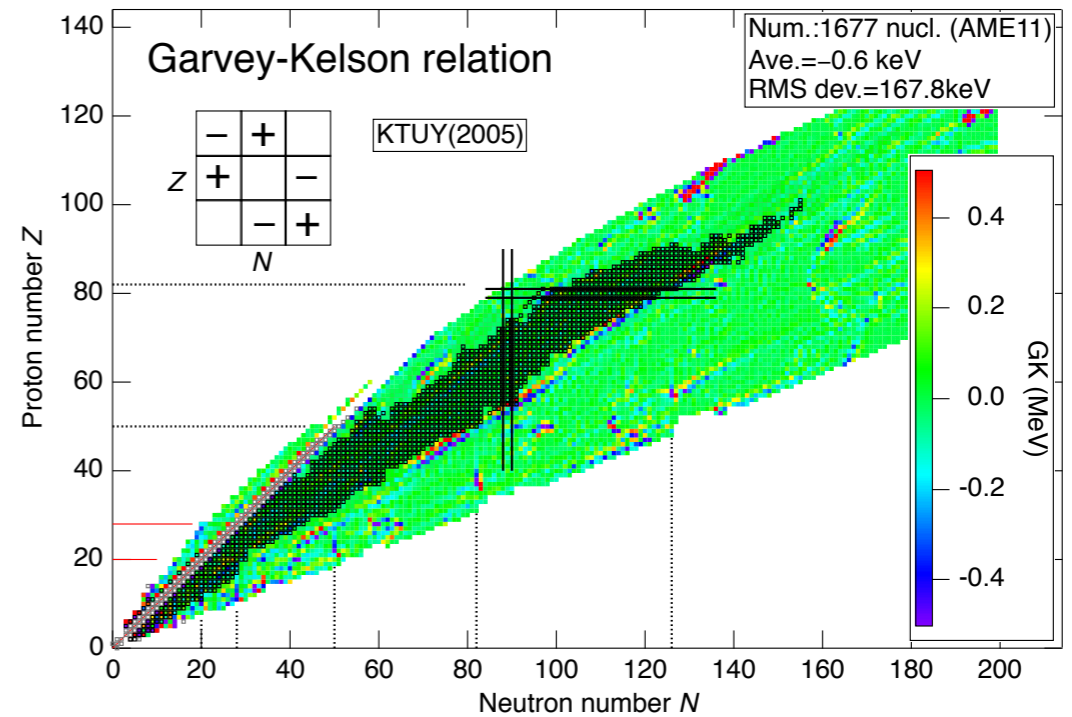
AME11(exp)



FRDM1995 : gives a smooth trends



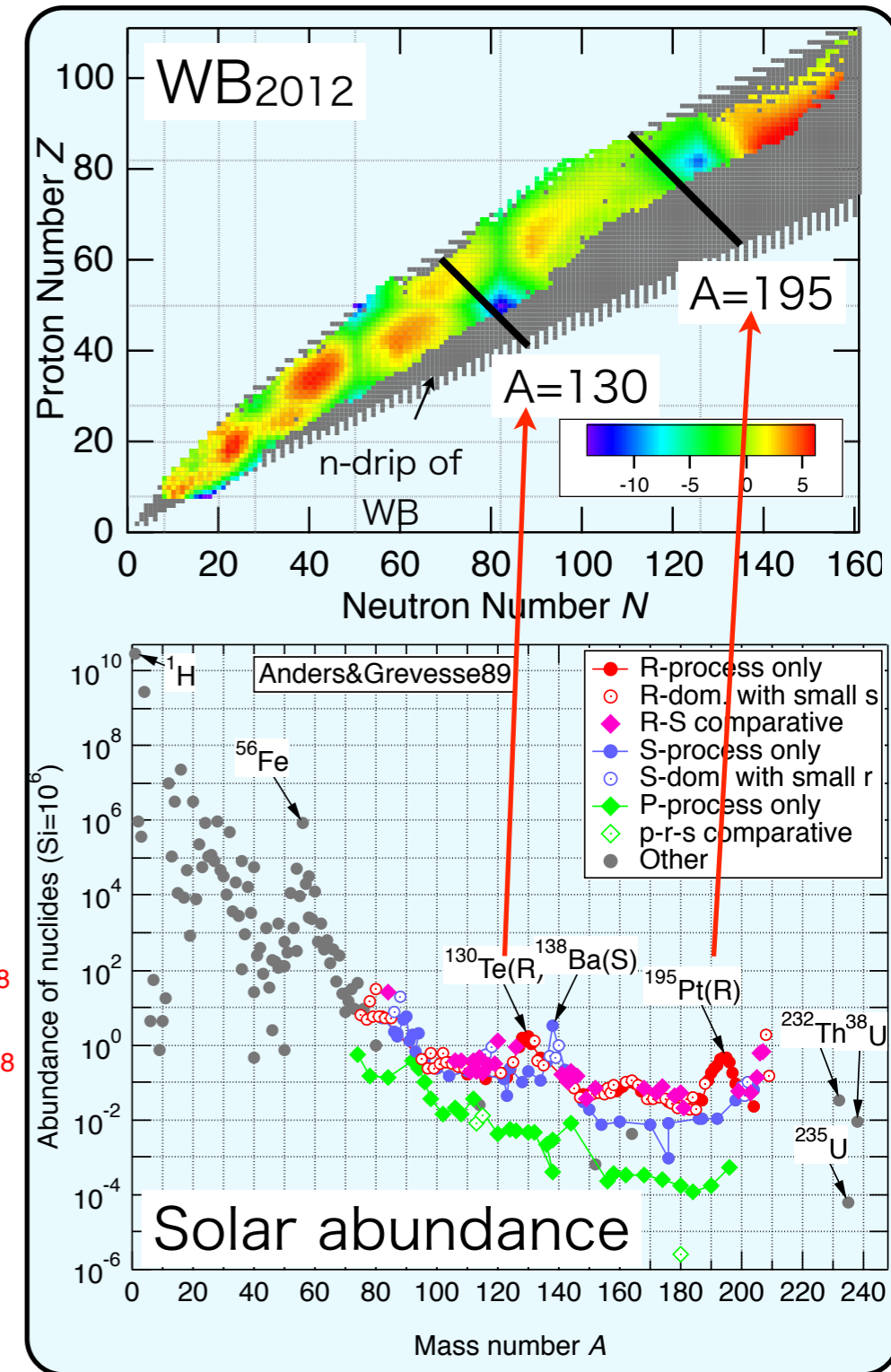
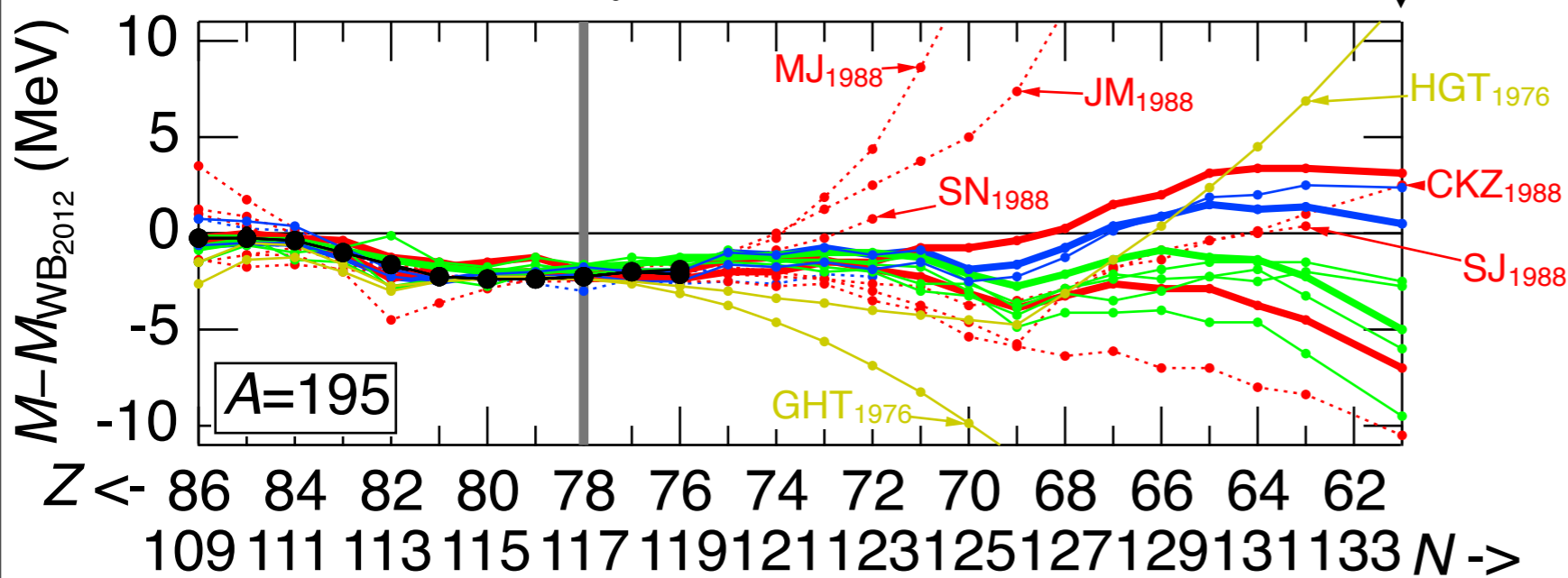
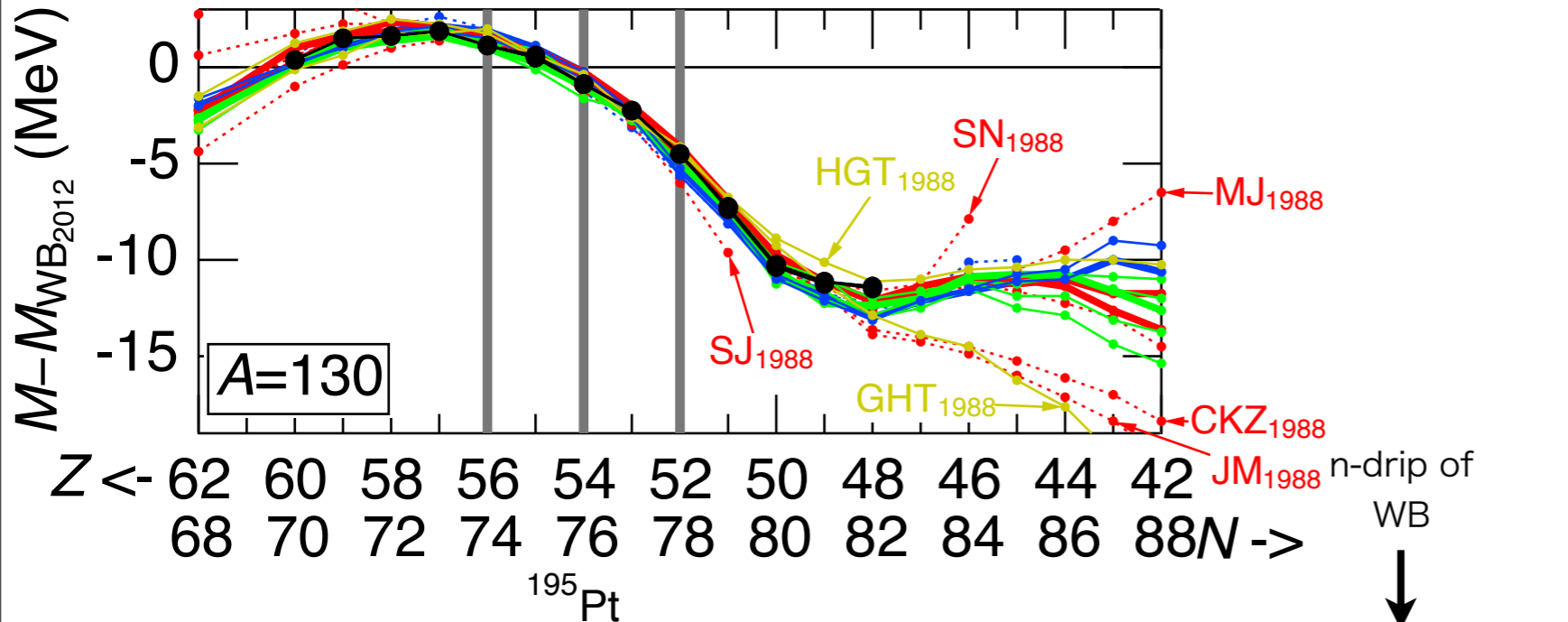
KTUY2005 : gives a smooth trends



- FRDM and KTUY give a smooth trends.

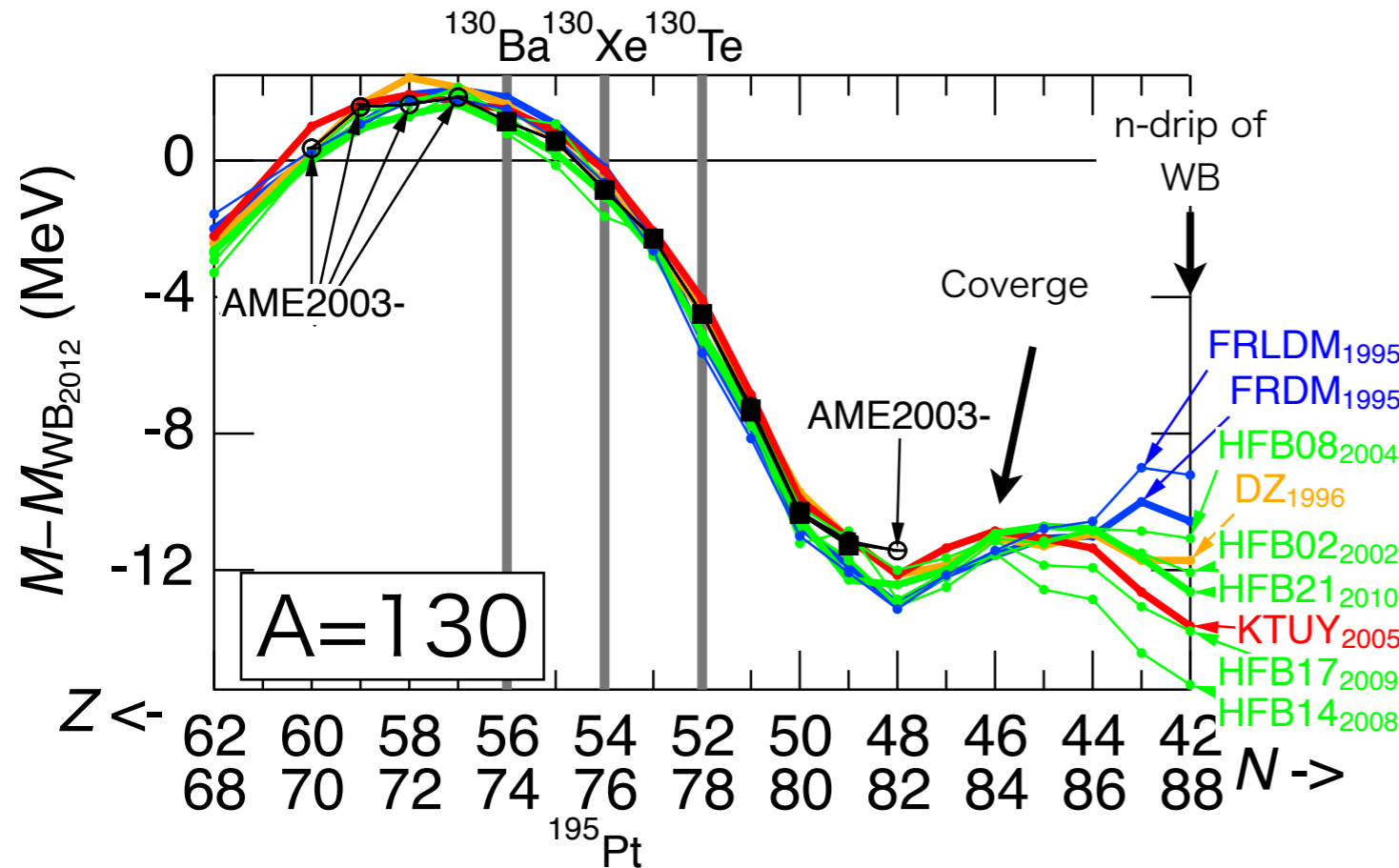
Mass value is subtracted by WB₂₀₁₂.

n-drip of WB



- In old-type mass formulae (-1988), mass values extremely diverge in the very neutron-rich region

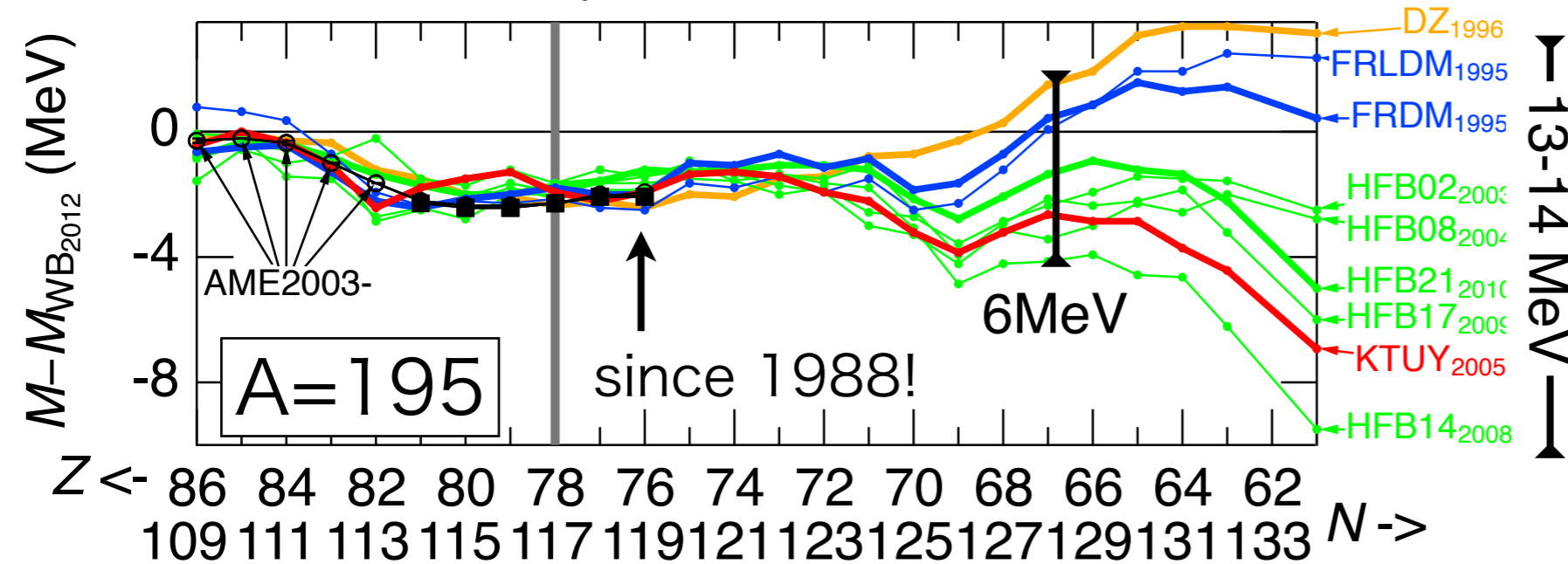
(Only since 1990-)



7-8 MeV

Predicted mass values still diverge.
Even among HFB's, mass values diverge in the n-rich region.
(several MeV)

Poor experimental mass data.



13-14 MeV

-Check the mass formulae as astrophysical data-

● Canonical model

Steady flow + Waiting point Approximation

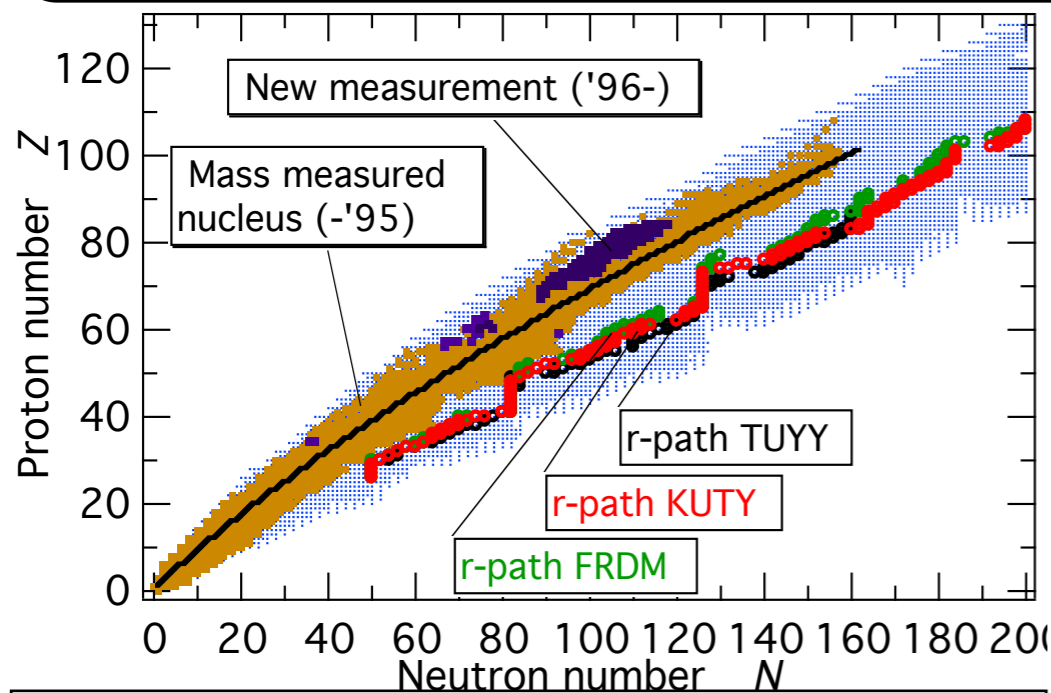
Neutron-number density (N_n) and temperature (T_9) are constants

(n, γ)-(γ, n) equilibrium is established over an irradiation time τ

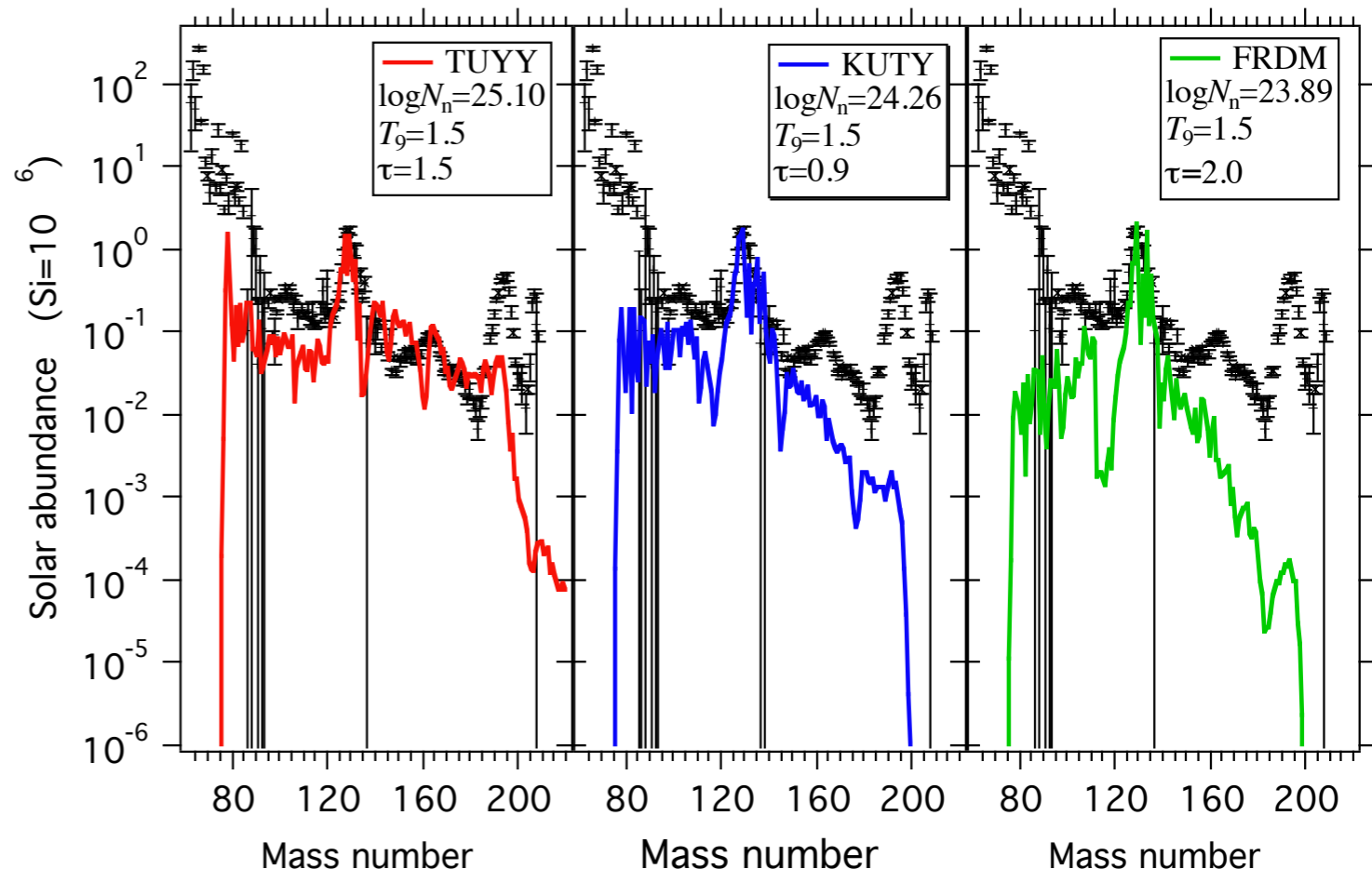
N_n, T_9, τ : chosen to reproduce the abundance peak at $A=130$ (obs.)

S_{2n} for equilibrium eq. (determine the path) and Q_β for λ_β :
 estimated from mass formulae (TUYU, KUTY, FRDM)

+ $N_r = N(\text{Solar abund.}) - N_s$
 × N_r r-only nuclei



- **TUYU**: gross term (WV-like with higher expansion) + empirical shell term.
- **KUTY**: TUYU gross term + deformed shell with a modified Woods-Saxon pot.
- **FRDM**: Macroscopic Droplet + microscopic deformed shell with a folded Yukawa pot.



21 **TUYU+GT2**

KUTY+GT2

FRDM+GT2

-Check the mass formulae as astrophysical data-

● Canonical model

Steady flow + Waiting point Approximation

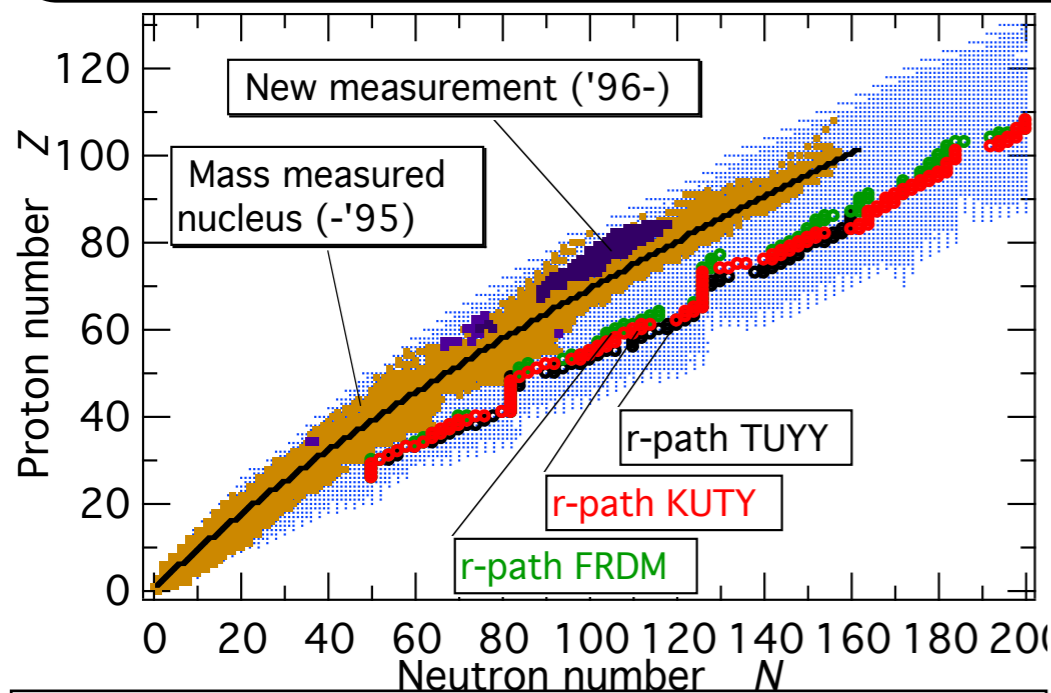
Neutron-number density (N_n) and temperature (T_9) are constants

(n, γ)-(γ ,n) equilibrium is established over an irradiation time τ

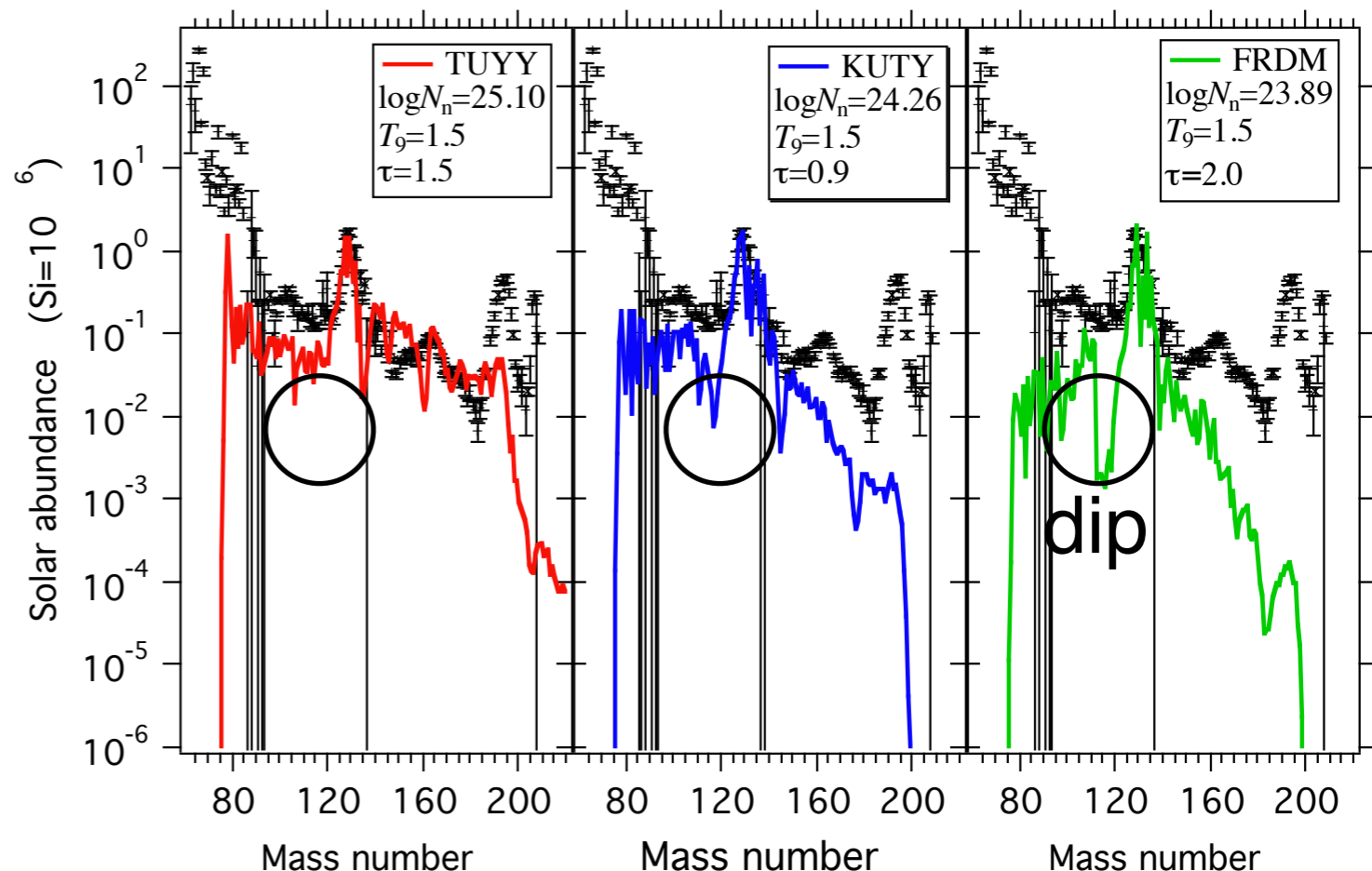
N_n, T_9, τ : chosen to reproduce the abundance peak at $A=130$ (obs.)

S_{2n} for equilibrium eq. (determine the path) and Q_β for λ_β :
 estimated from mass formulae (TUYU, KUTY, FRDM)

+ $N_r = N(\text{Solar abund.}) - N_s$
 × N_r r-only nuclei



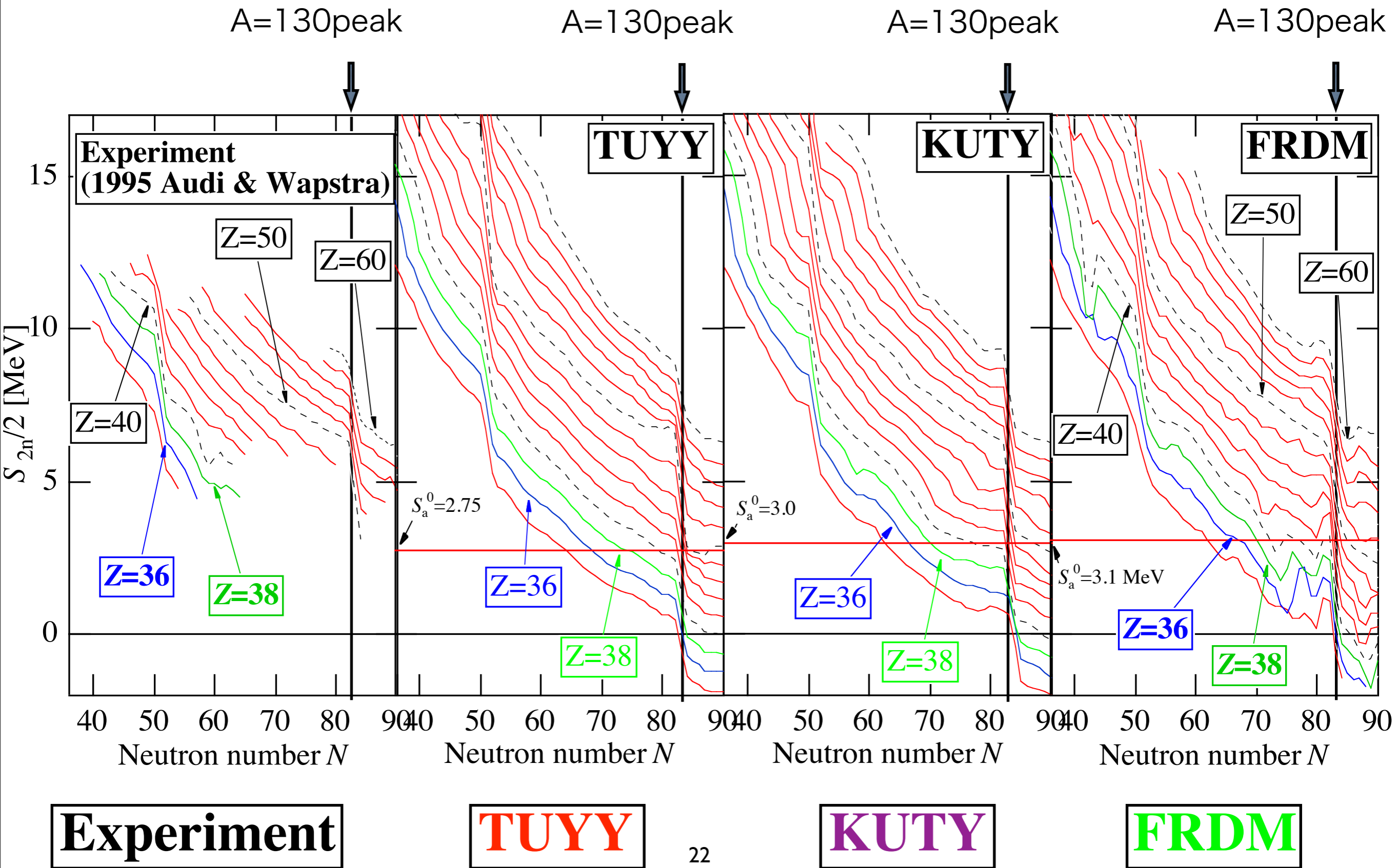
- **TUYU**: gross term (VWB-like with higher expansion) + empirical shell term.
- **KTUY**: TUYU gross term + deformed shell with a modified Woods-Saxon pot.
- **FRDM**: Macroscopic Droplet + microscopic deformed shell with a folded Yukawa pot.

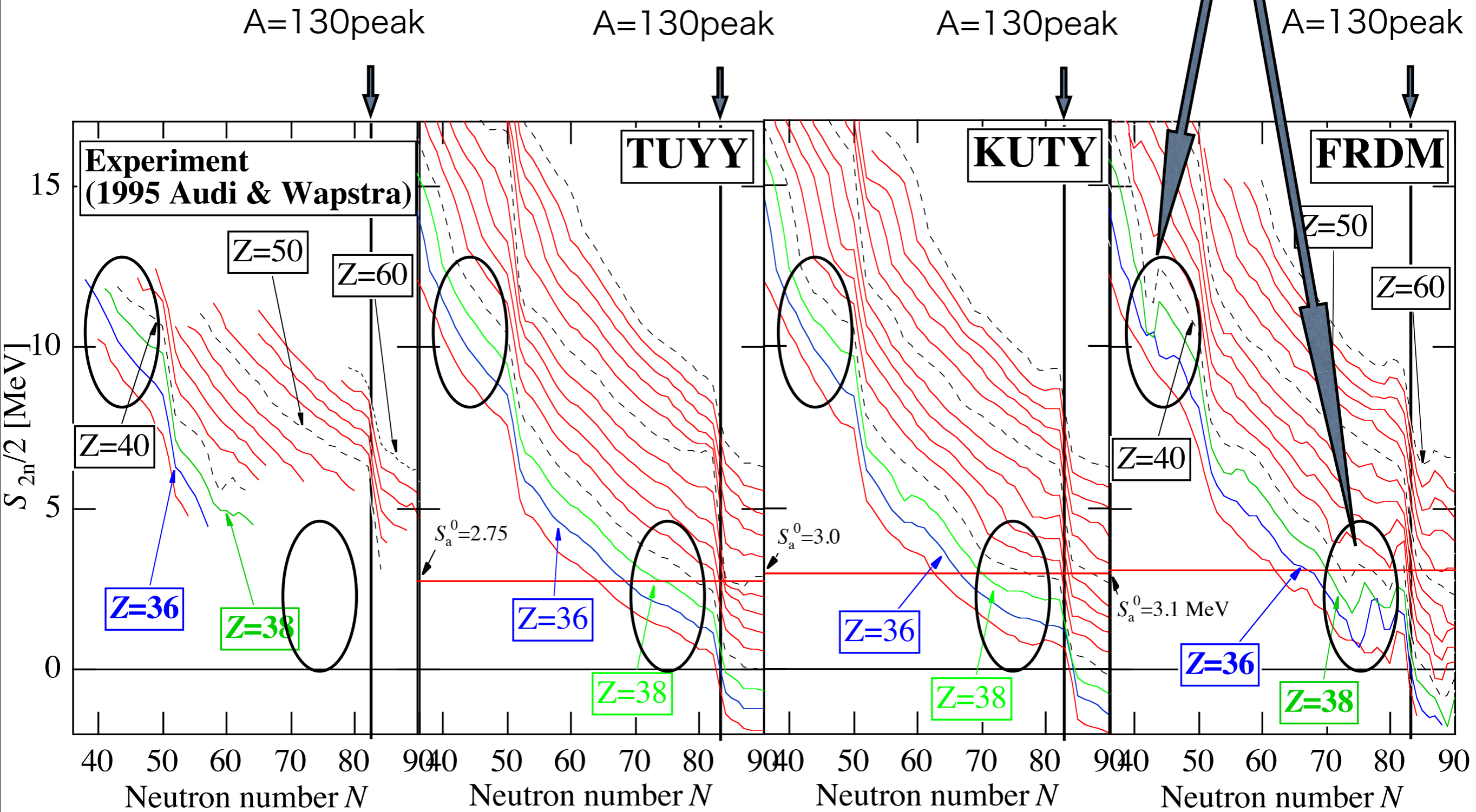


21 **TUYU+GT2**

KUTY+GT2

FRDM+GT2



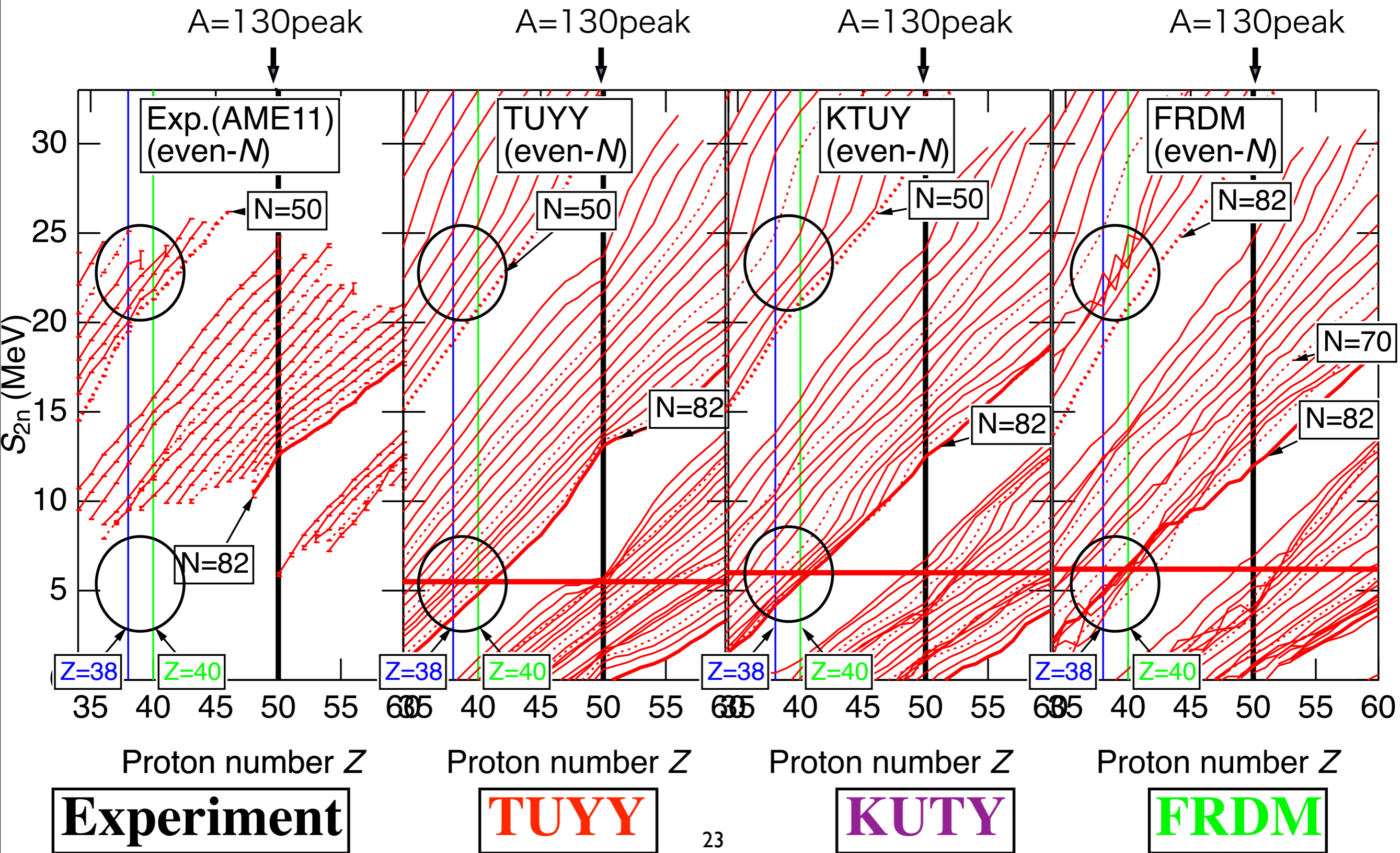


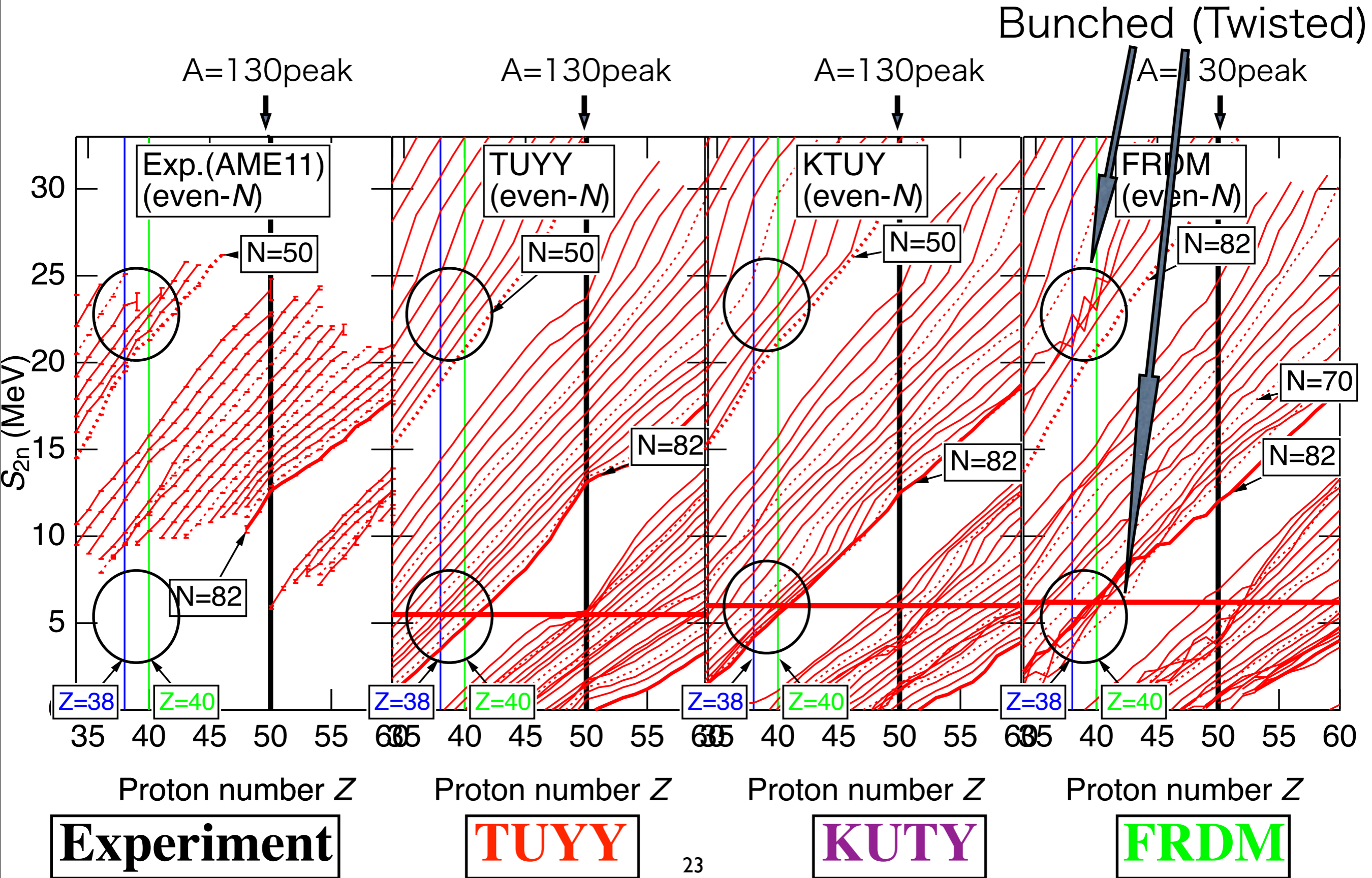
Experiment

TUYU

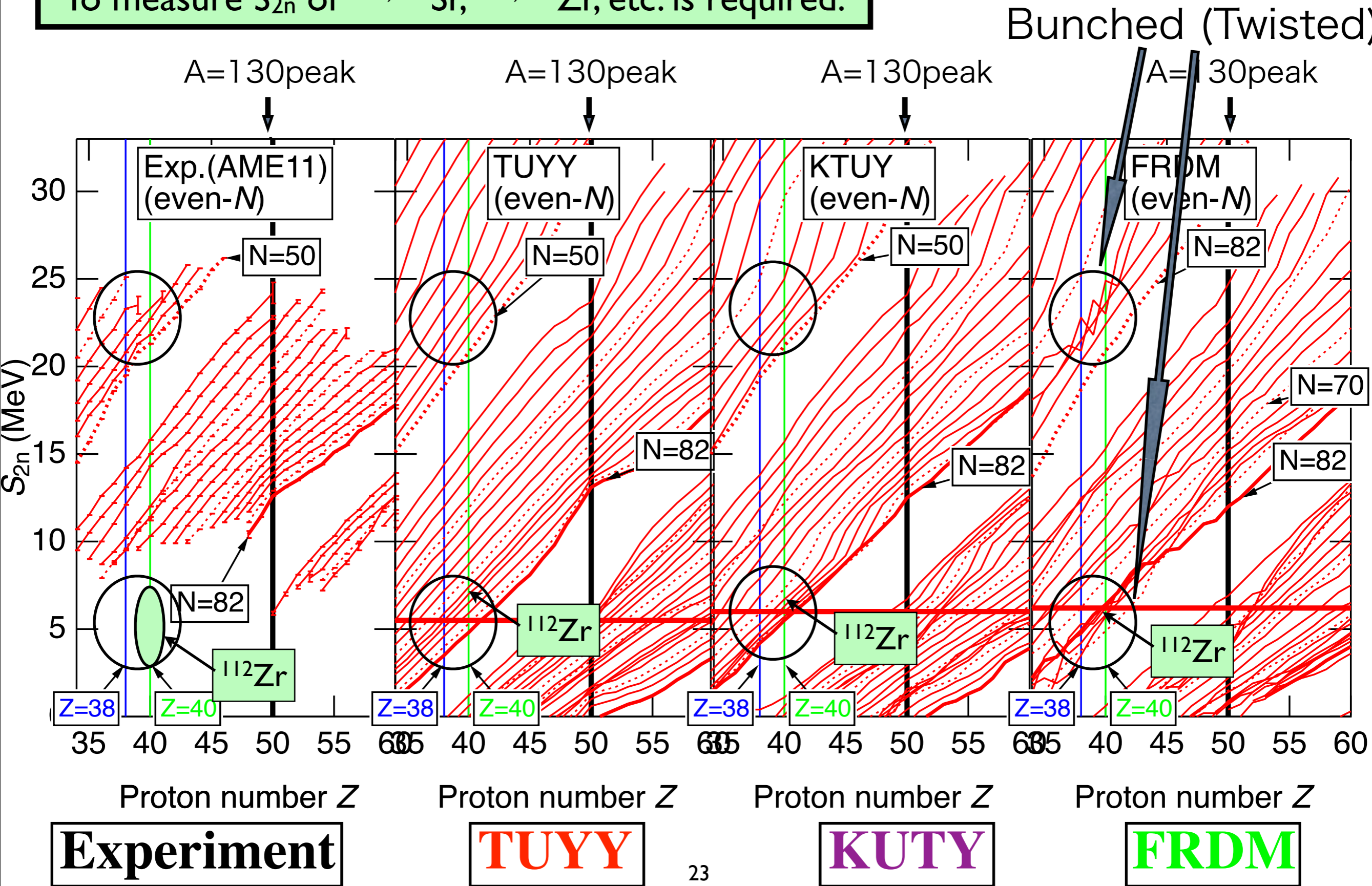
KUTY

FRDM

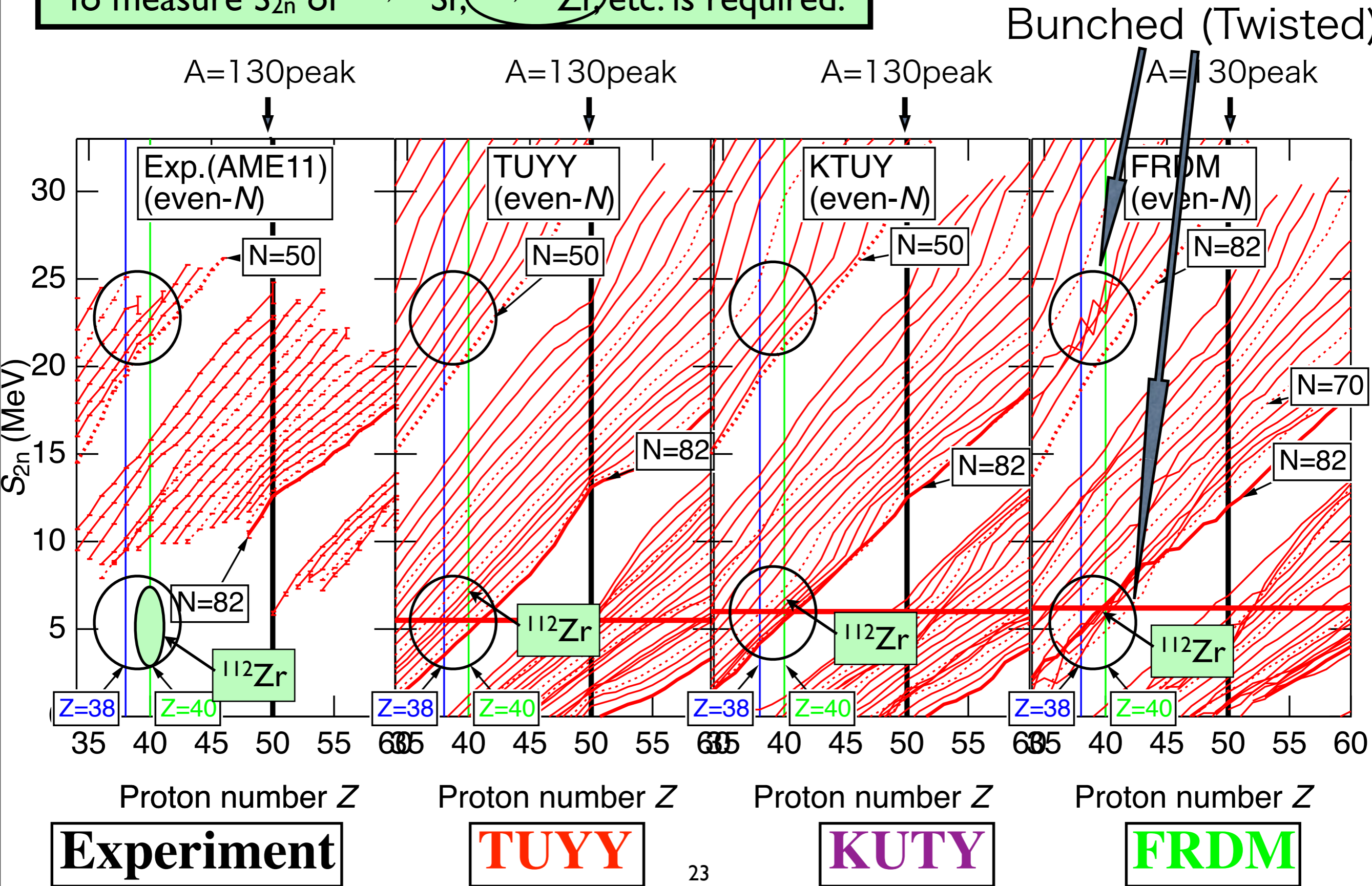


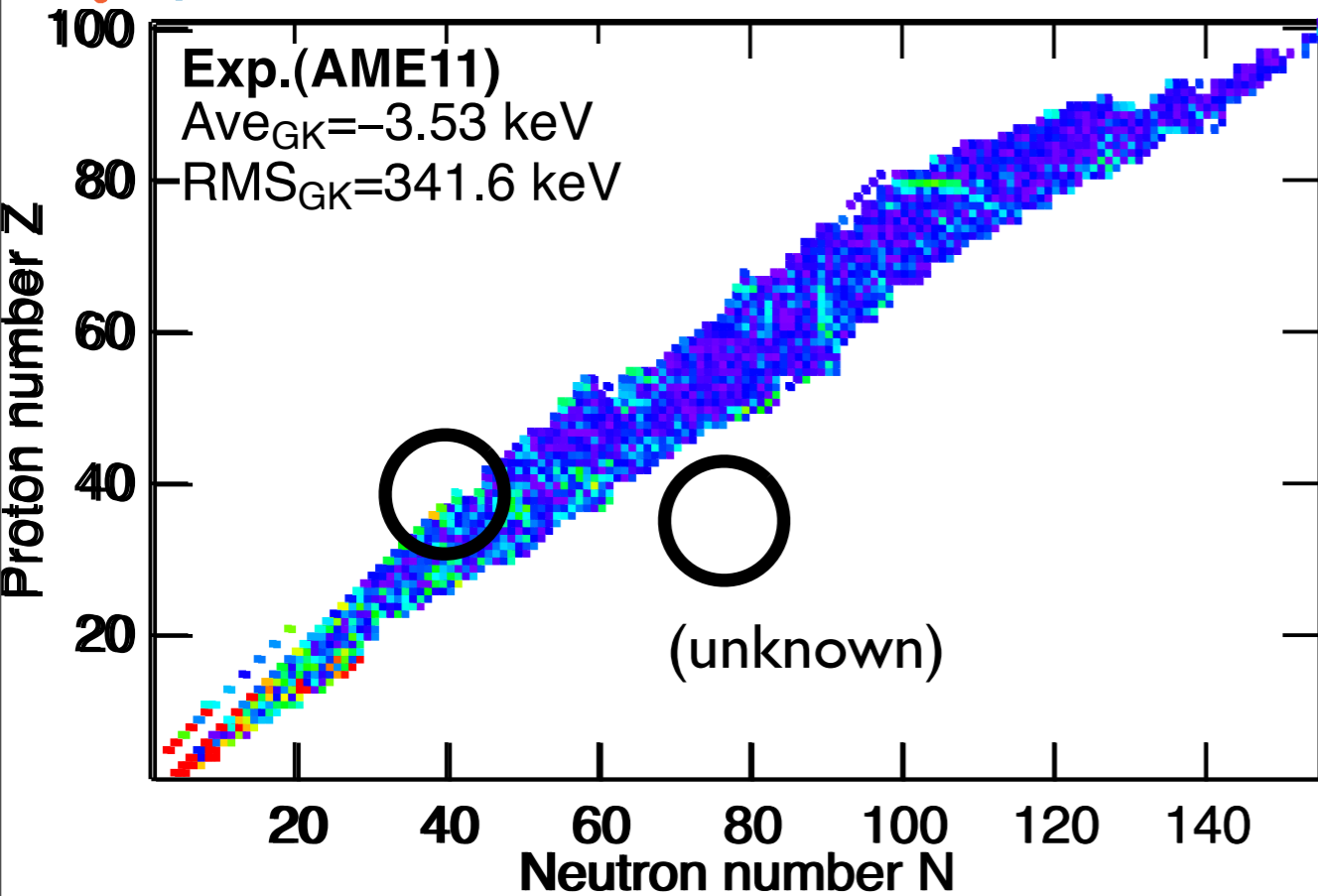


To measure S_{2n} of $^{108,110}\text{Sr}$, $^{110,112}\text{Zr}$, etc. is required.



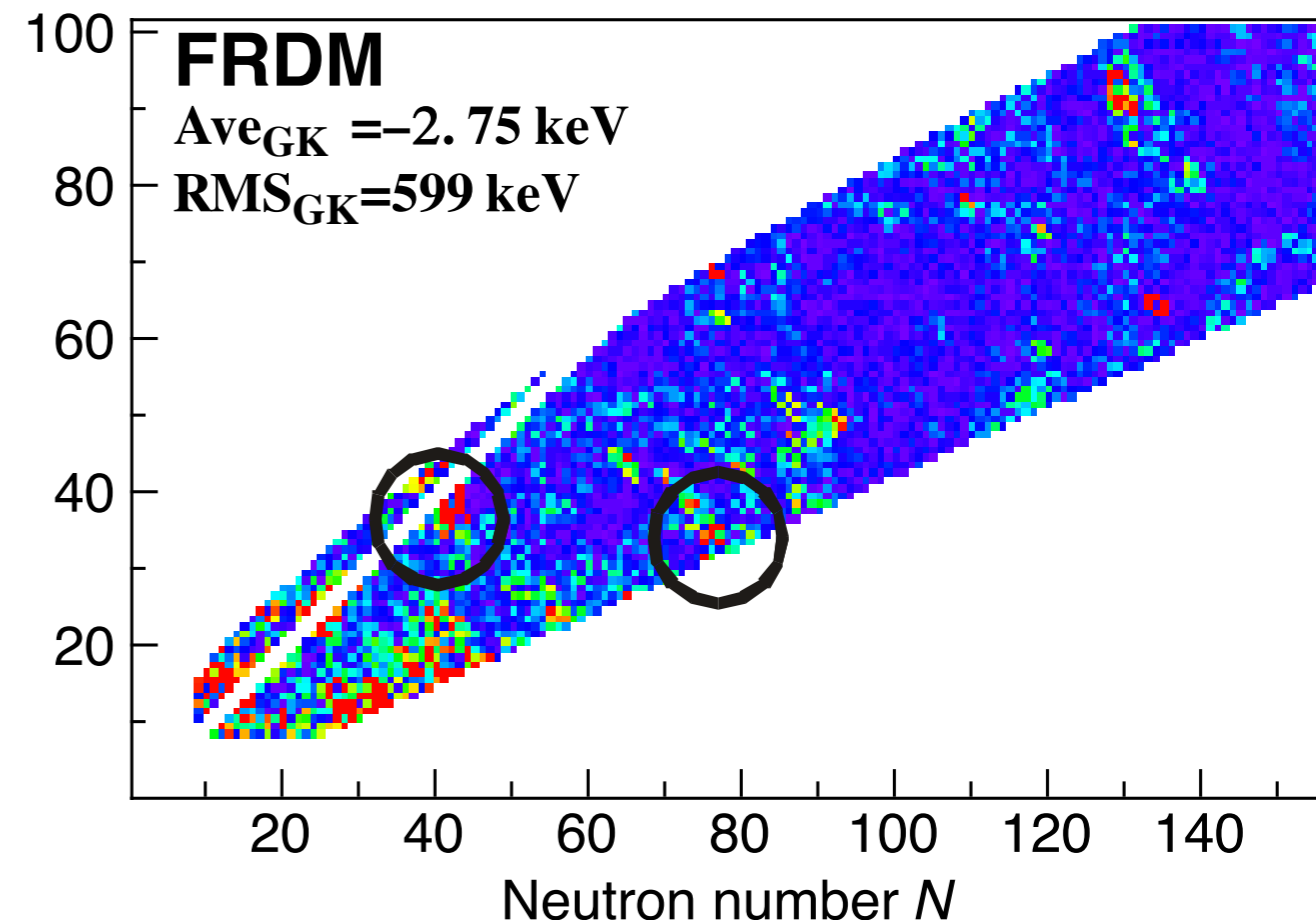
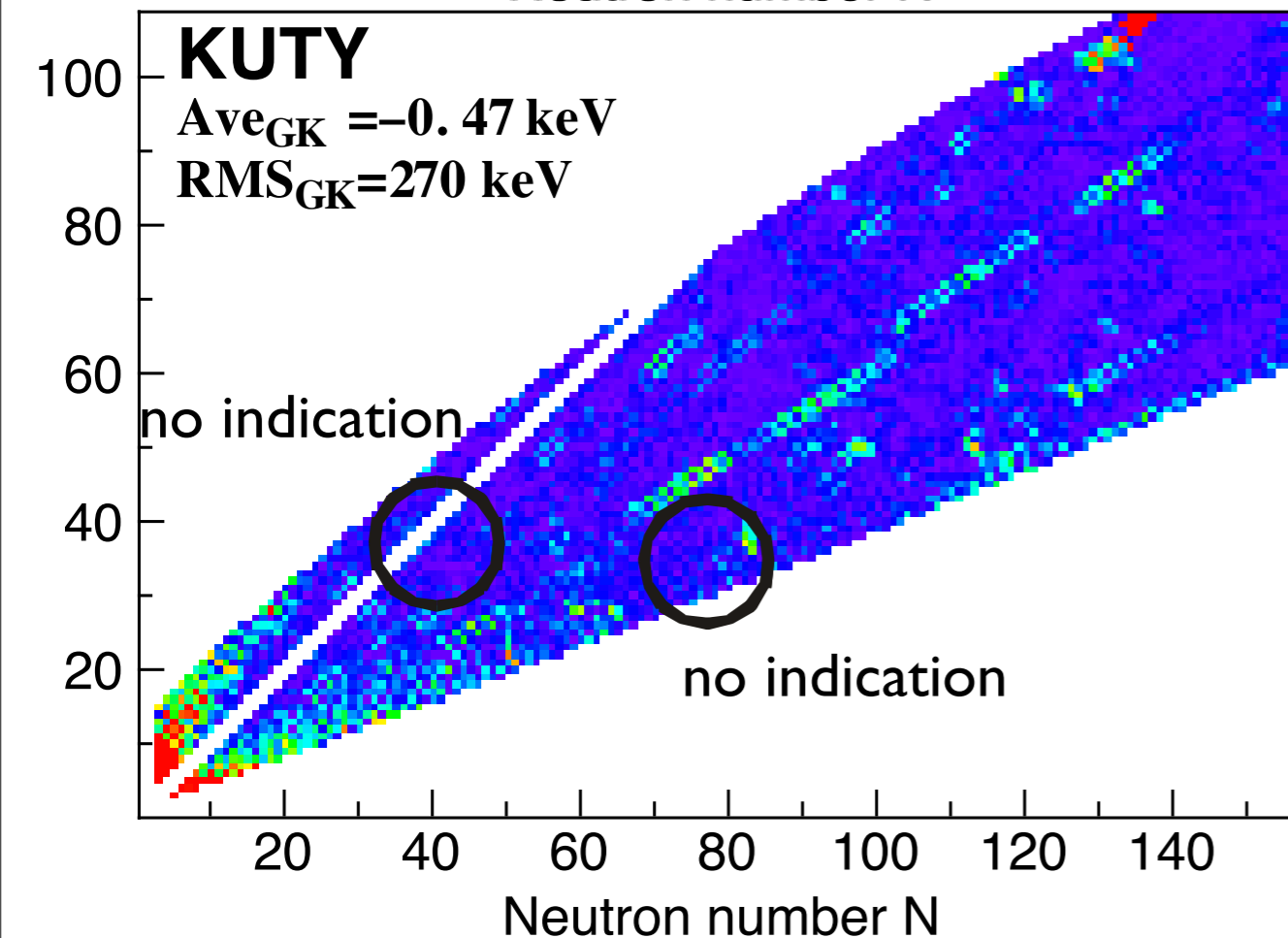
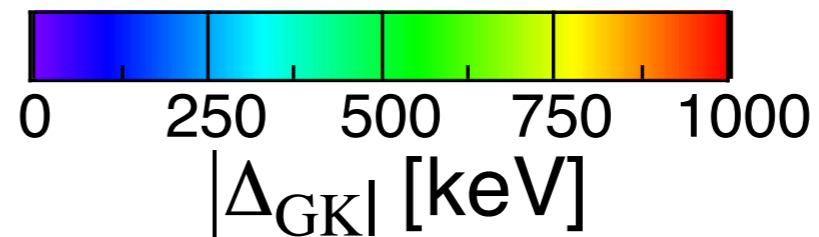
To measure S_{2n} of $^{108,110}\text{Sr}$, $^{110,112}\text{Zr}$, etc. is required.

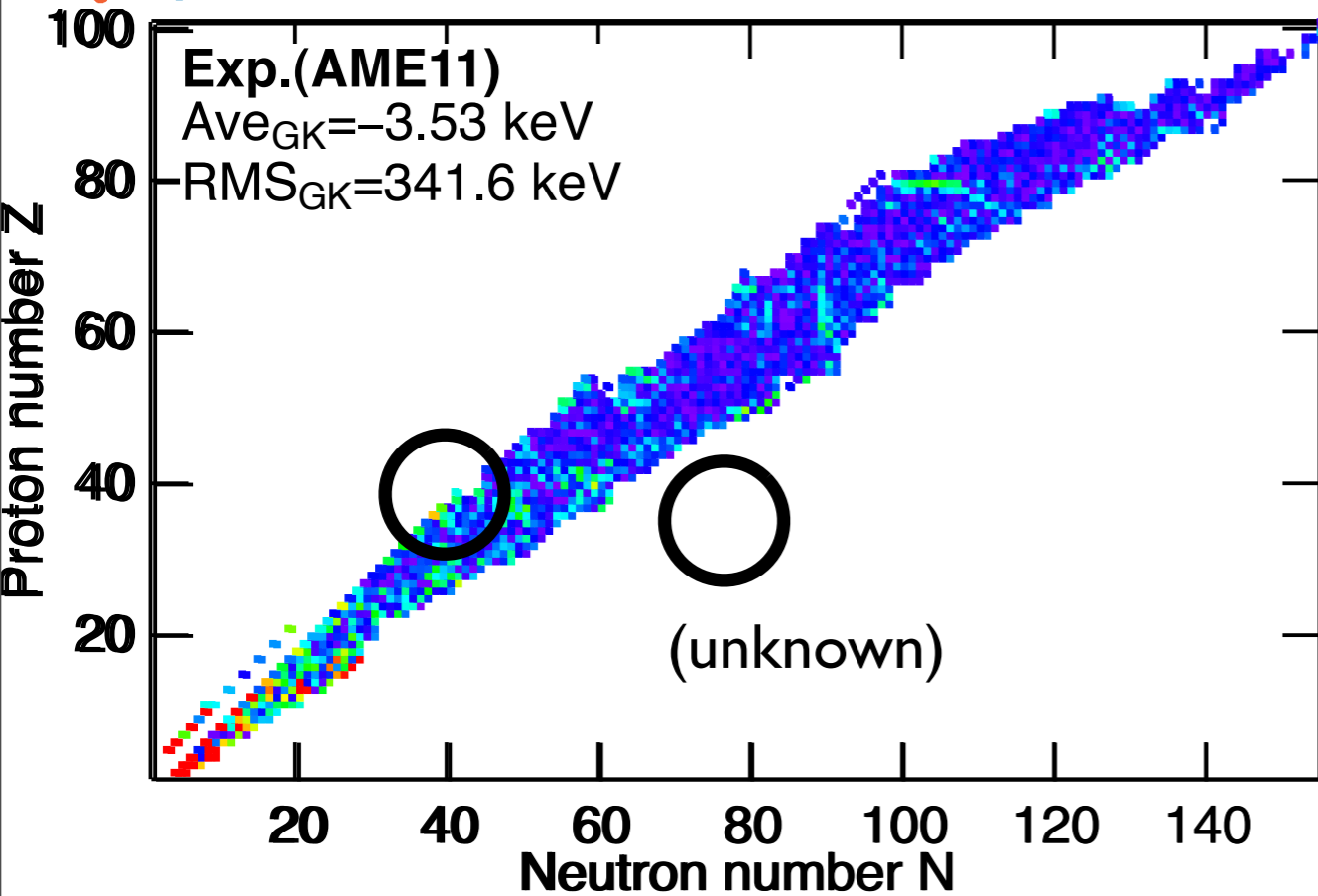




Garvey-Kelson eq.

$$\begin{matrix} Z \\ \begin{matrix} - & + & \\ + & & - \\ & - & + \end{matrix} \\ N \end{matrix} \equiv \Delta_{GK} \approx 0$$

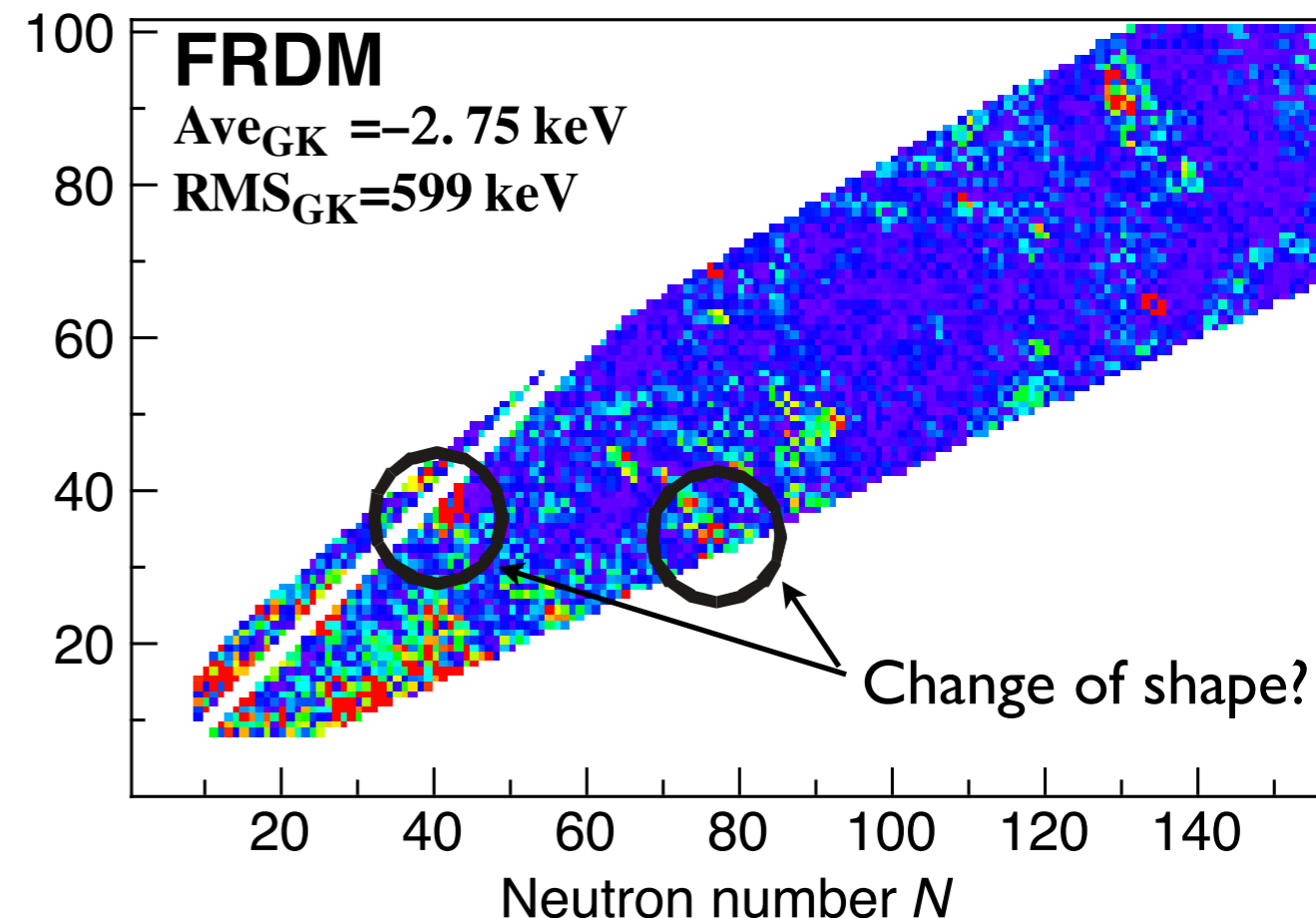
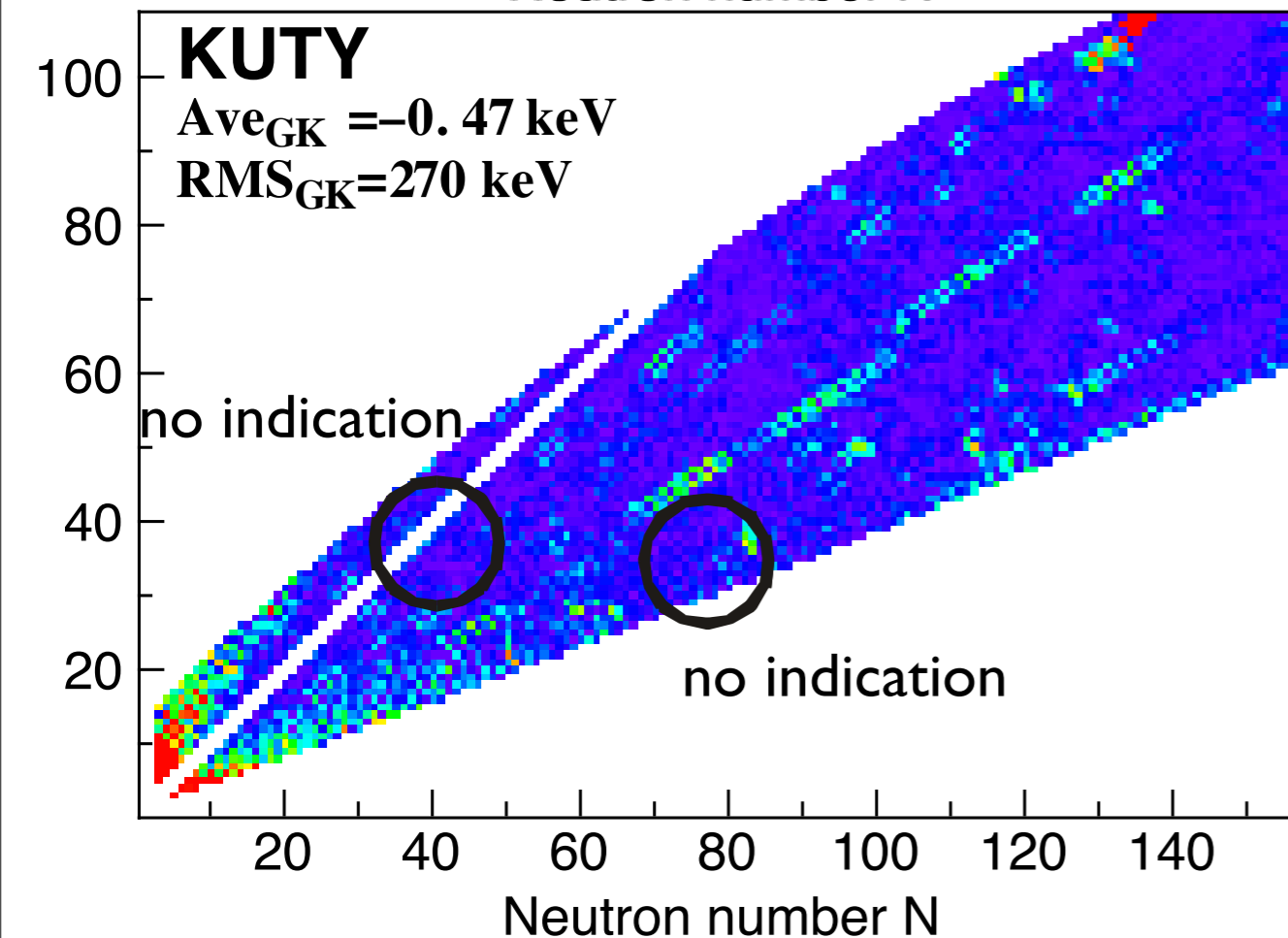
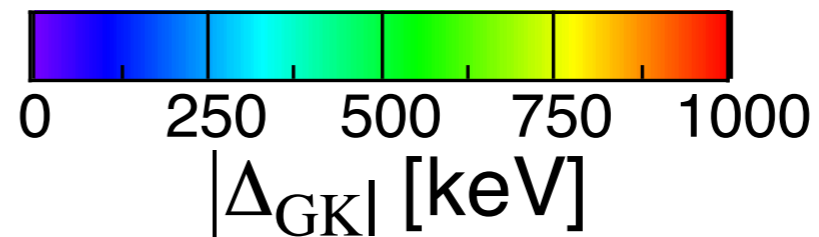




Garvey-Kelson eq.

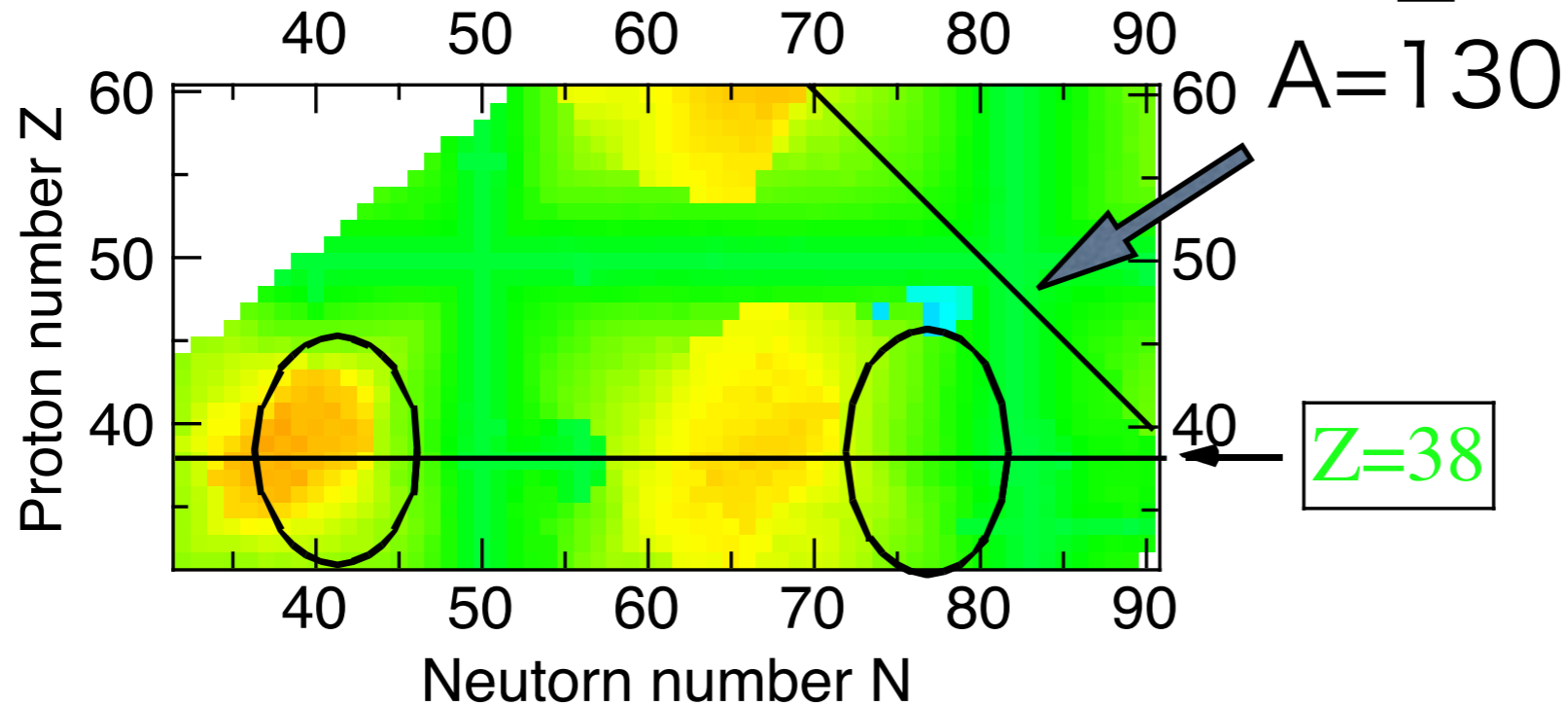
$$\begin{array}{|c|c|c|} \hline - & + & \\ \hline + & & - \\ \hline & - & + \\ \hline \end{array} \equiv \Delta_{GK} \approx 0$$

N

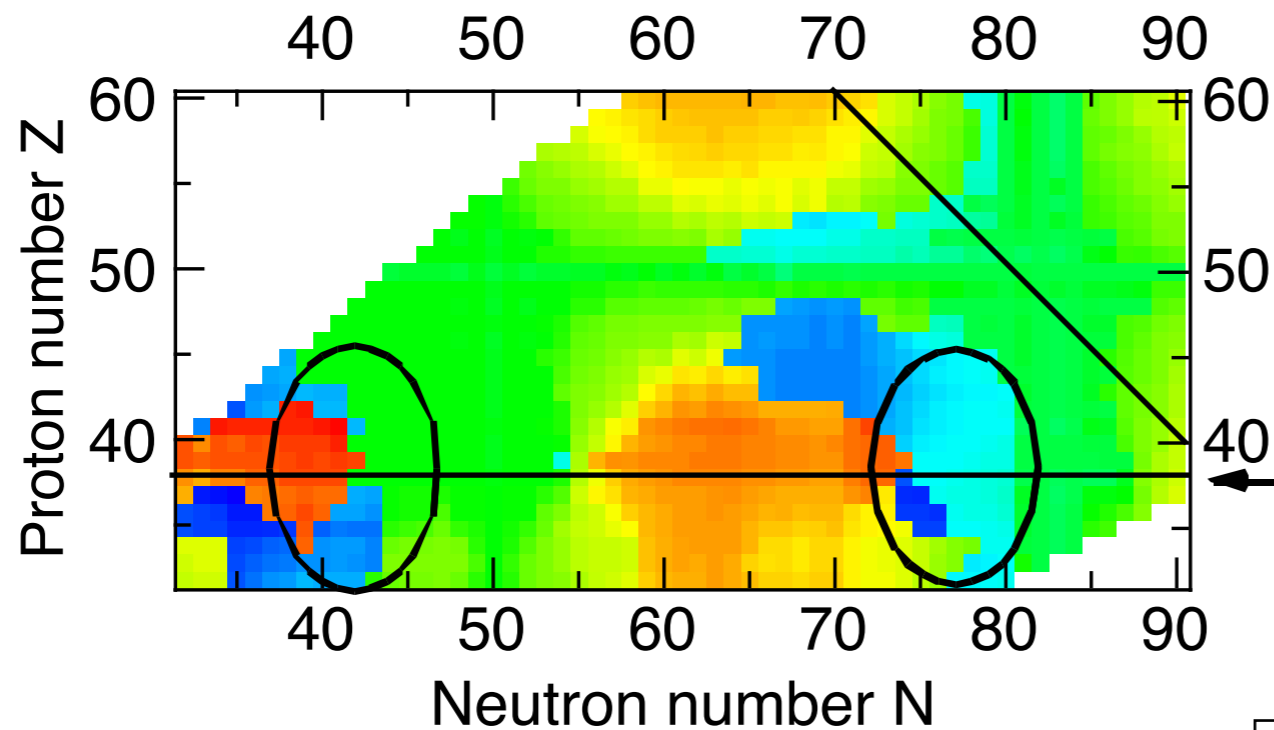


Deformation parameter α_2

KUTY



FRDM



dip of $S_{2n} \iff$ change of α_2
(FRDM)

Discontinuous change of shape would give kink of S_n . (FRDM case)
Theoretical (numerical) problem?

Chen, Dabaczewski, Kratz, Langanke, Thieleman, Vogt, Phys. Lett. **B355** (1995)37-44

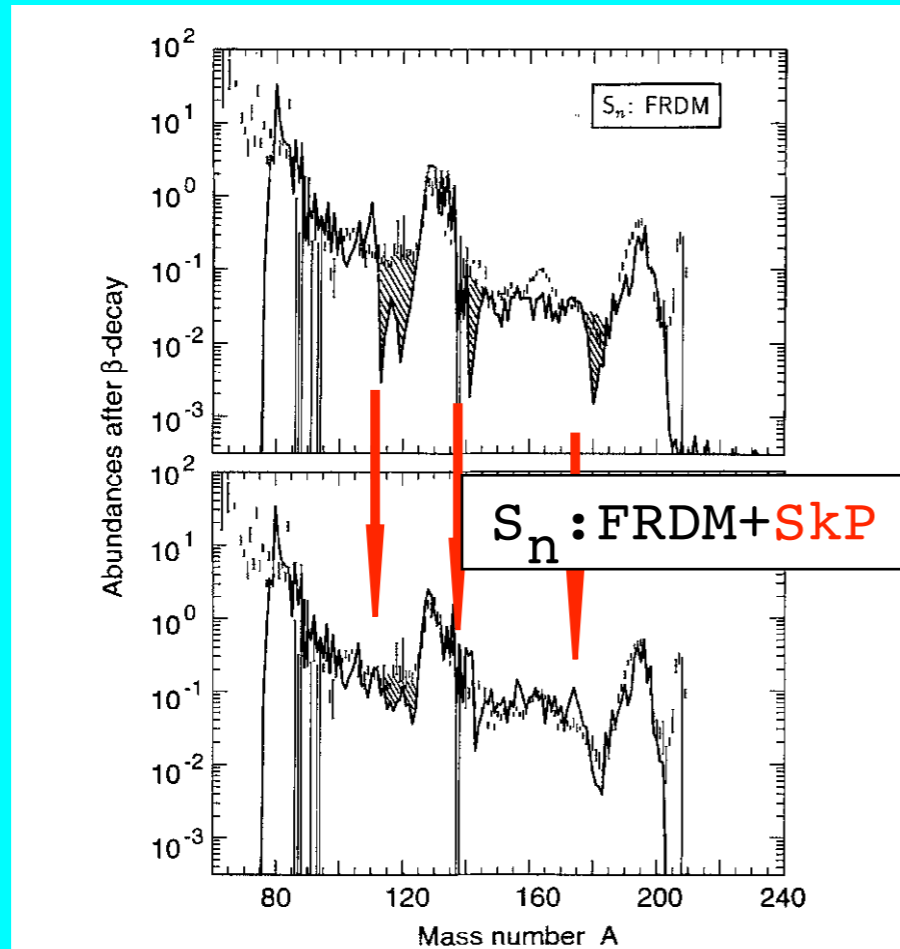


Fig. 2. r-process abundance fits obtained with ten equidistant neutron-density components from 10^{20} cm^{-3} to $3 \times 10^{24} \text{ cm}^{-3}$ according to Fig. 1. In the upper part, the result is presented for FRDM [10] masses with the $T_{1/2}$ and P_n values from the QRPA calculations according to Ref. [11]. In the lower part, masses of spherical nuclei around $N = 82$ have been replaced by masses from HFB calculations with the Skyrme force SkP. The quenching of the $N = 82$ shell gap (see Fig. 4) leads to a filling of the abundance troughs around $A \simeq 120$ and 140 , and to a better overall reproduction of the heavy-mass region.

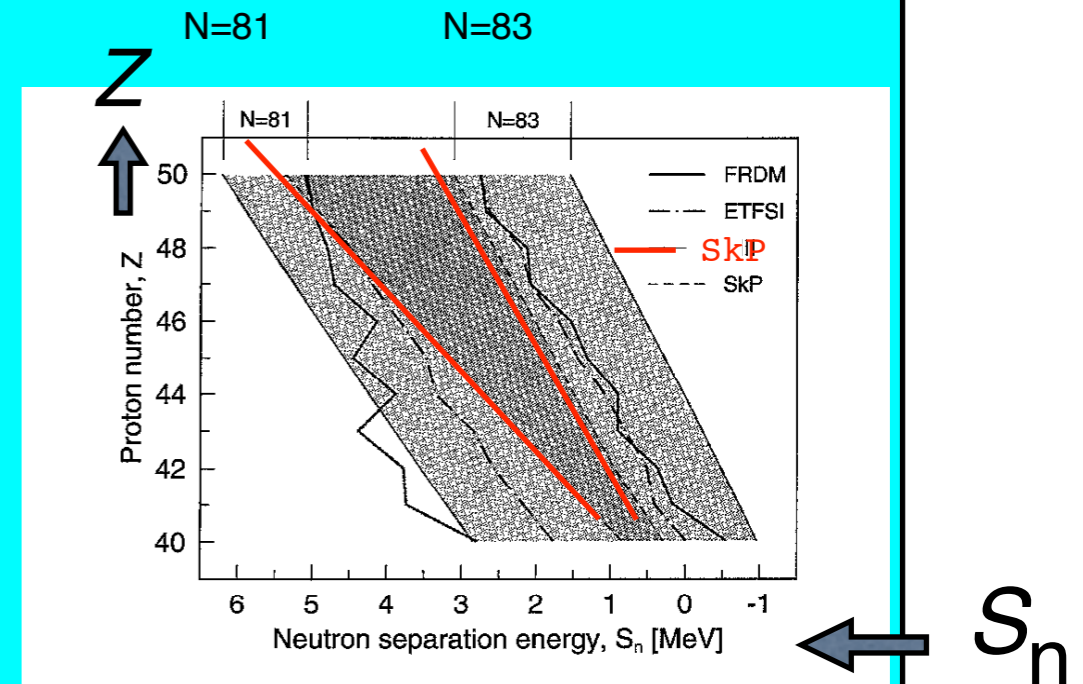
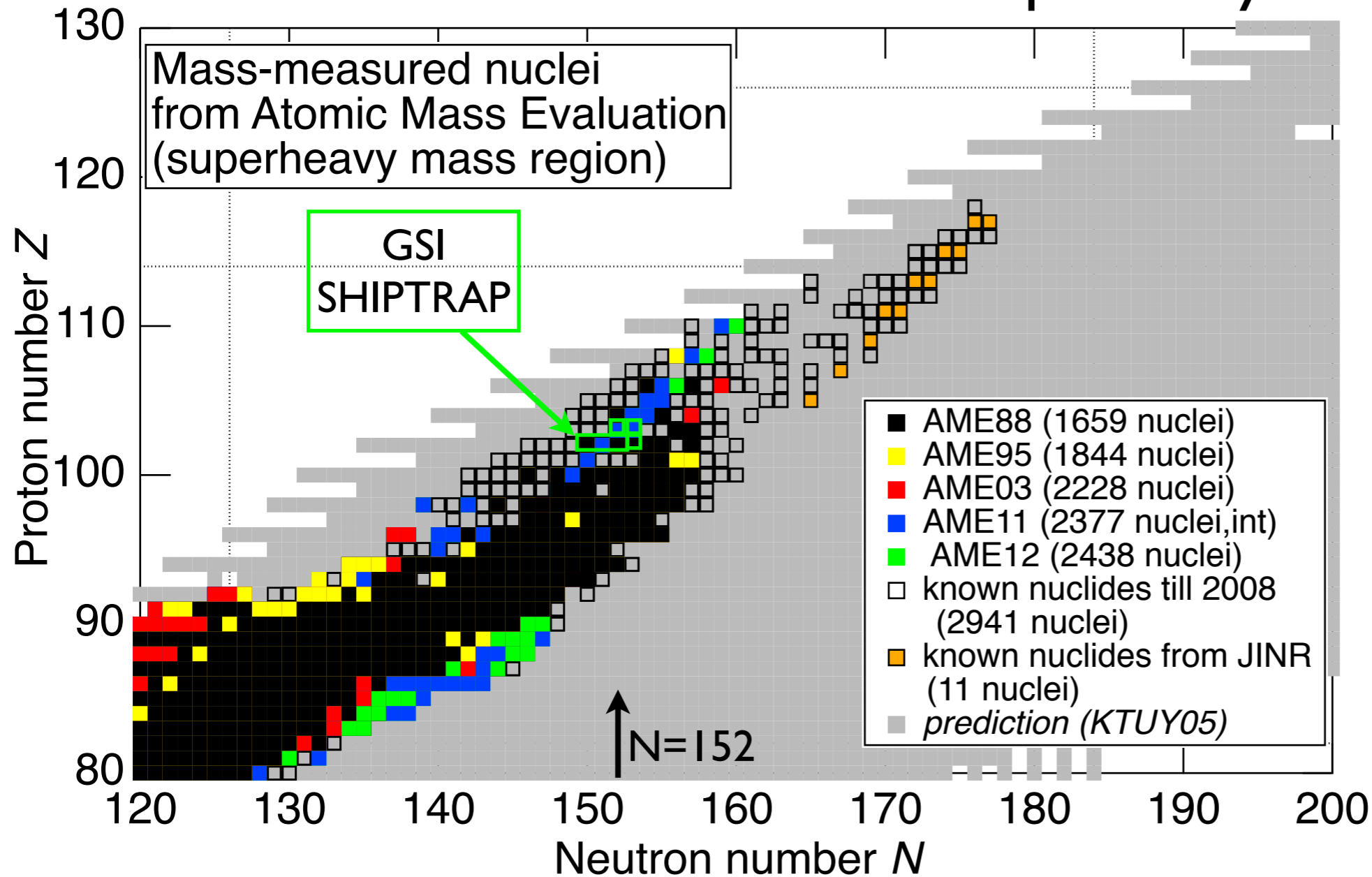


Fig. 4. Comparison of S_n values for the isotones $N = 81$ and 83 as predicted by different mass models. The difference $\delta S_n = [S_n(N = 81) - S_n(N = 83)]$ is a measure of the $N = 82$ shell strength and is shaded for SIII (light) and SkP (dark). The shell quenching with distance from stability for SkP, in contrast to SIII, can be recognized. Masses of odd-odd nuclei have not been calculated in our SIII study.

shell-quenching
 => decreasing of dips
 around the peaks
 (Chen et al. 1995)

Kink of S_{2n} , or shell quenching?

Mass-measured nuclei in the superheavy mass region



Poor mass-measurement in the superheavy mass region

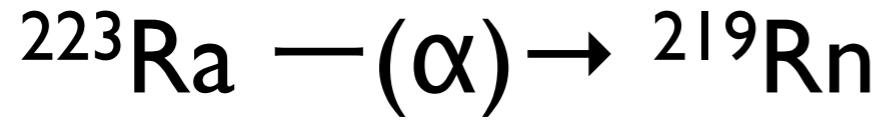
Why is direct mass measurement?

Why is direct mass measurement?

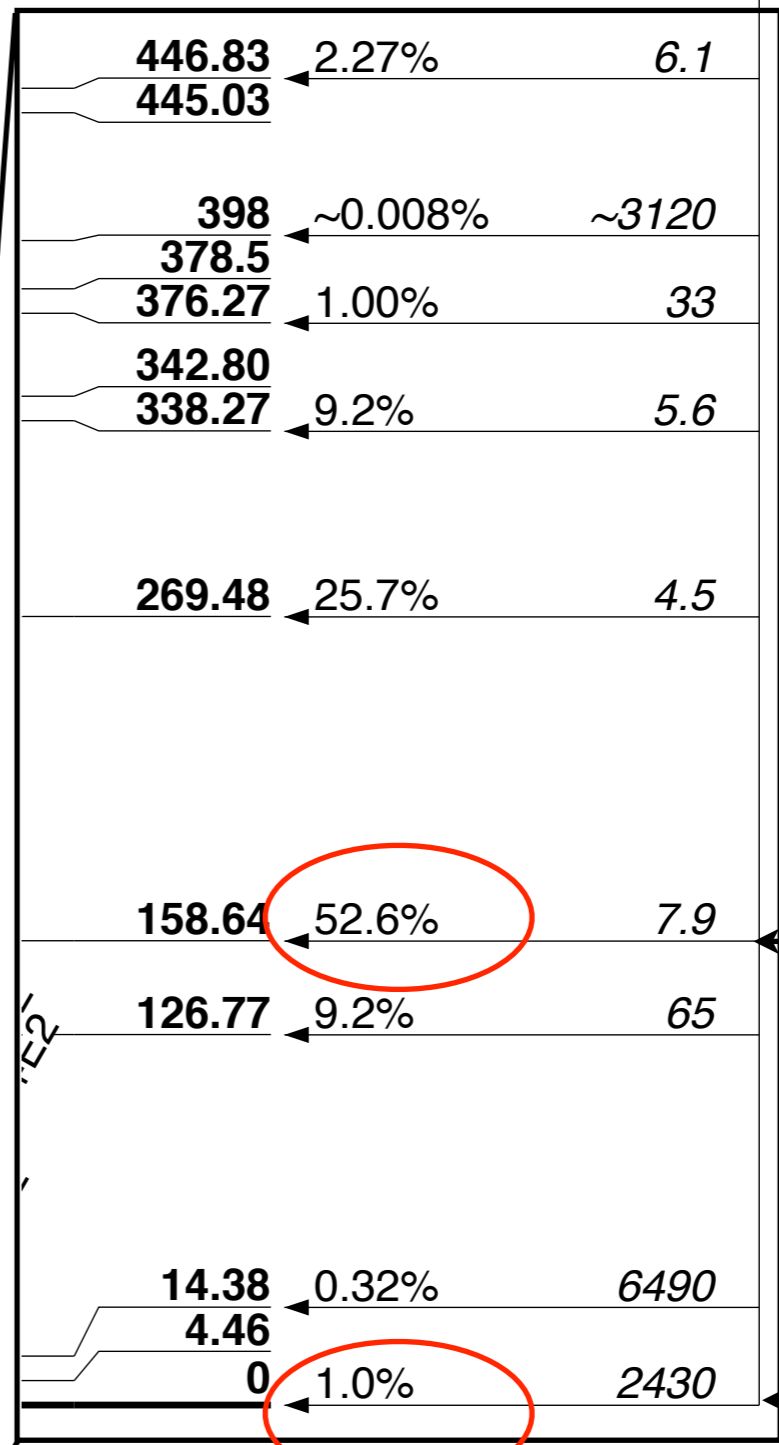
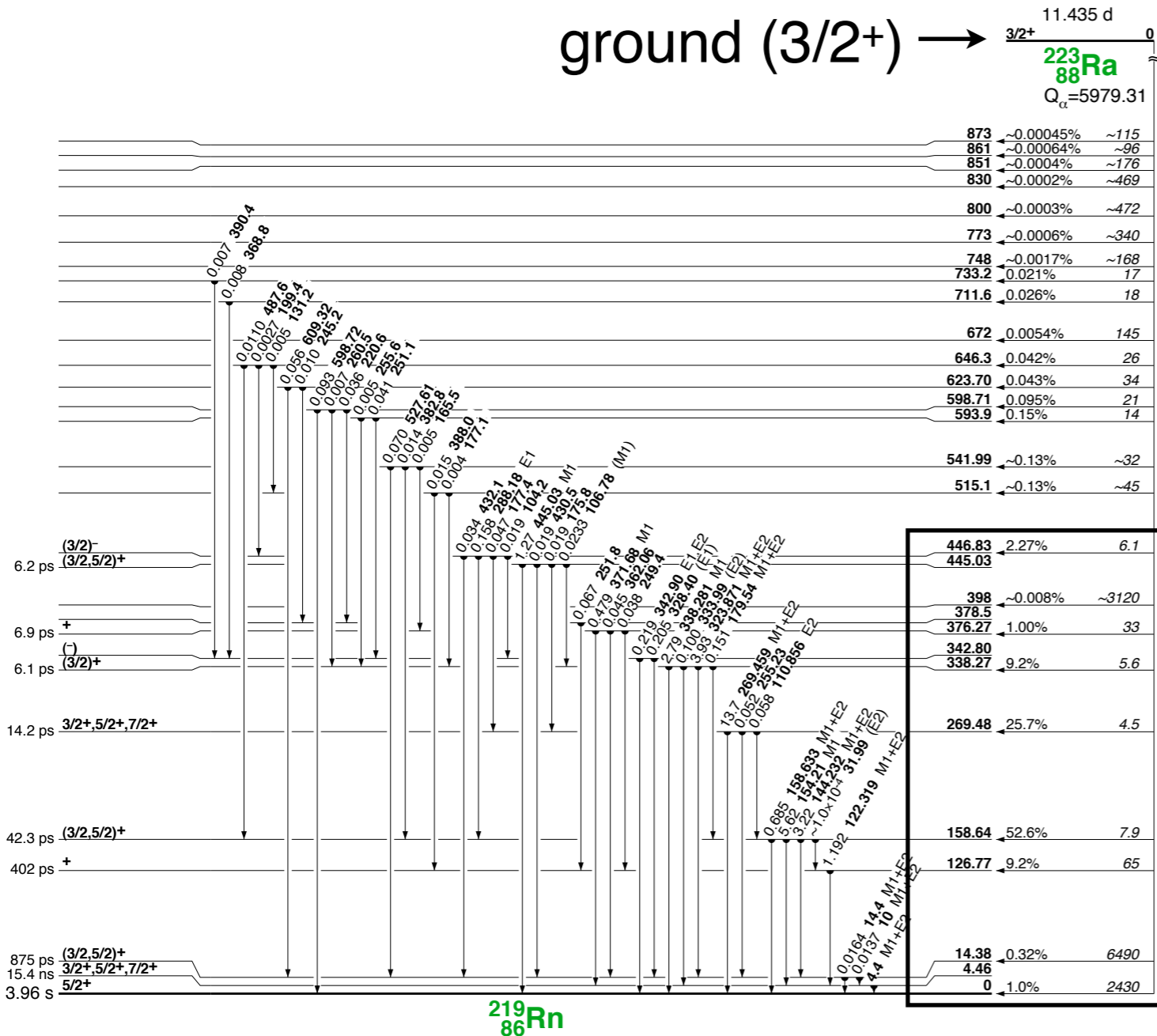


To obtain ground-state energies (essentially important!)◎

Example: ground to excited states for odd-A



ground ($3/2^+$) \rightarrow $^{223}_{88}\text{Ra}$ $3/2^+$ $Q_\alpha = 5979.31$



excited ($3/2^+, 5/2^+$)

ground ($5/2^+$)

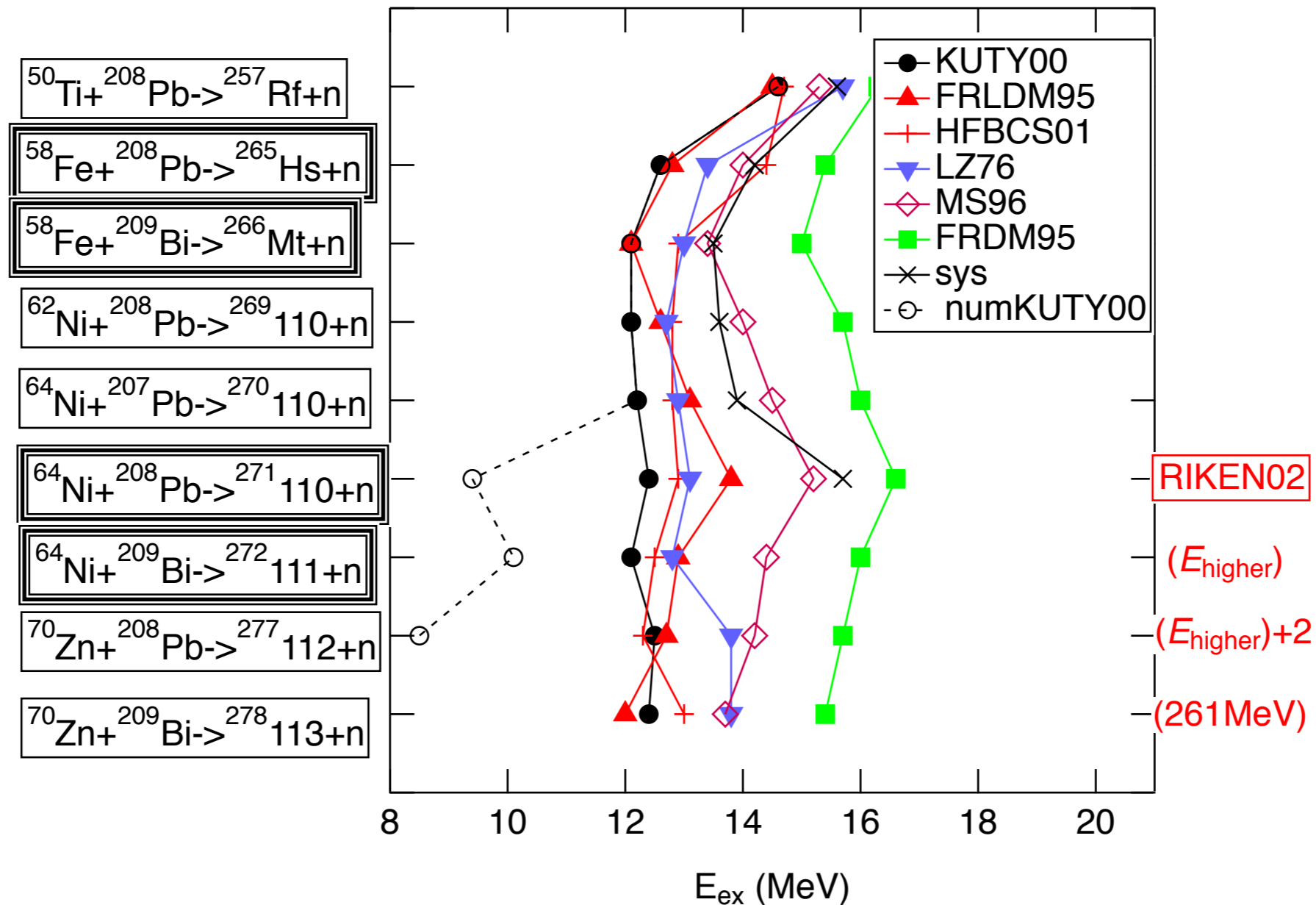
a-few-hundred-keV-order

Alpha decay chains in SHE

												120		295	296	297	298																	
													292	293	294	295	296	297																
												119																						
													289	290	291	292	293	294	295	296														
												118																						
													286	287	288	289	290	291	292	293	294	295												
												117																						
												278	279	280	281	282	283	284	285	286	287	288	289	290	291	292	293	294						
												116																						
												277	278	279	280	281	282	283	284	285	286	287	288	289	290	291	292	293	294	295				
												115																						
												273	274	275	276	277	278	279	280	281	282	283	284	285	286	287	288	289	290	291	292			
												114																						
												272	273	274	275	276	277	278	279	280	281	282	283	284	285	286	287	288	289	290	291	292		
												113																						
												268 Cn	269 Cn	270 Cn	271 Cn	272 Cn	273 Cn	274 Cn	275 Cn	276 Cn	277 Cn	278 Cn	279 Cn	280 Cn	281 Cn	282 Cn	283 Cn	284 Cn	285 Cn	286 Cn	287 Cn	288 Cn	289 Cn	290 Cn
												112																						
												267 Rg	268 Rg	269 Rg	270 Rg	271 Rg	272 Rg	273 Rg	274 Rg	275 Rg	276 Rg	277 Rg	278 Rg	279 Rg	280 Rg	281 Rg	282 Rg	283 Rg	284 Rg	285 Rg	286 Rg	287 Rg	288 Rg	289 Rg
												111																						
263 Ds	264 Ds	265 Ds	266 Ds	267 Ds	268 Ds	269 Ds	270 Ds	271 Ds	272 Ds	273 Ds	274 Ds	275 Ds	276 Ds	277 Ds	278 Ds	279 Ds	280 Ds	281 Ds	282 Ds	283 Ds	284 Ds	285 Ds	286 Ds	287 Ds	288 Ds									
262 Mt	263 Mt	264 Mt	265 Mt	266 Mt	267 Mt	268 Mt	269 Mt	270 Mt	271 Mt	272 Mt	273 Mt	274 Mt	275 Mt	276 Mt	277 Mt	278 Mt	279 Mt	280 Mt	281 Mt	282 Mt	283 Mt	284 Mt	285 Mt	286 Mt	287 Mt									
261 Hs	262 Hs	263 Hs	264 Hs	265 Hs	266 Hs	267 Hs	268 Hs	269 Hs	270 Hs	271 Hs	272 Hs	273 Hs	274 Hs	275 Hs	276 Hs	277 Hs	278 Hs	279 Hs	280 Hs	281 Hs	282 Hs	283 Hs	284 Hs	285 Hs	286 Hs									
260 Bh	261 Bh	262 Bh	263 Bh	264 Bh	265 Bh	266 Bh	267 Bh	268 Bh	269 Bh	270 Bh	271 Bh	272 Bh	273 Bh	274 Bh	275 Bh	276 Bh	277 Bh	278 Bh	279 Bh	280 Bh	281 Bh	282 Bh	283 Bh	284 Bh	178									
259 Sg	260 Sg	261 Sg	262 Sg	263 Sg	264 Sg	265 Sg	266 Sg	267 Sg	268 Sg	269 Sg	270 Sg	271 Sg	272 Sg	273 Sg	274 Sg	275 Sg	276 Sg	277 Sg	278 Sg	279 Sg	280 Sg	281 Sg	176	177										
258 Db	259 Db	260 Db	261 Db	262 Db	263 Db	264 Db	265 Db	266 Db	267 Db	268 Db	269 Db	270 Db	271 Db	272 Db	273 Db	274 Db	275 Db	276 Db	277 Db	278 Db	174	175												
257 Rf	258 Rf	259 Rf	260 Rf	261 Rf	262 Rf	263 Rf	264 Rf	265 Rf	266 Rf	267 Rf	268 Rf	269 Rf	270 Rf	271 Rf	272 Rf	273 Rf	274 Rf	275 Rf	276 Rf	173														
256 Lr	257 Lr	258 Lr	259 Lr	260 Lr	261 Lr	262 Lr	263 Lr	264 Lr	265 Lr	266 Lr	267 Lr	268 Lr	269 Lr	270 Lr	271 Lr	272 Lr	273 Lr	274 Lr	172															
255 No	256 No	257 No	258 No	259 No	260 No	261 No	262 No	263 No	264 No	265 No	266 No	267 No	268 No	269 No	270 No	271 No	170	171																

- Almost energies are obtained from alpha-decay chains.
- No relations of other directions are obtained.
- Direct mass-measurements are required.

Beam energy from (predicted) excitation energy



Energy at maximum cross section
 (derived from injected beam energy with Q of mass formulas)

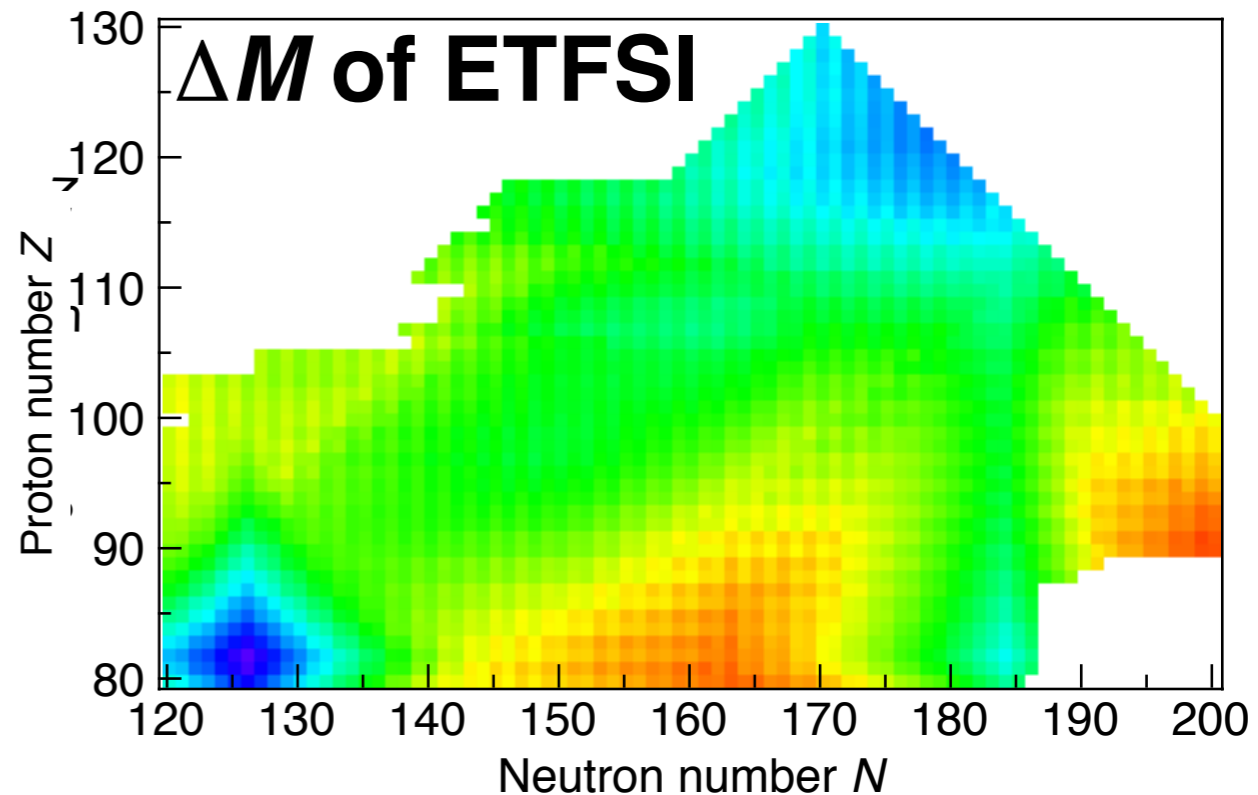
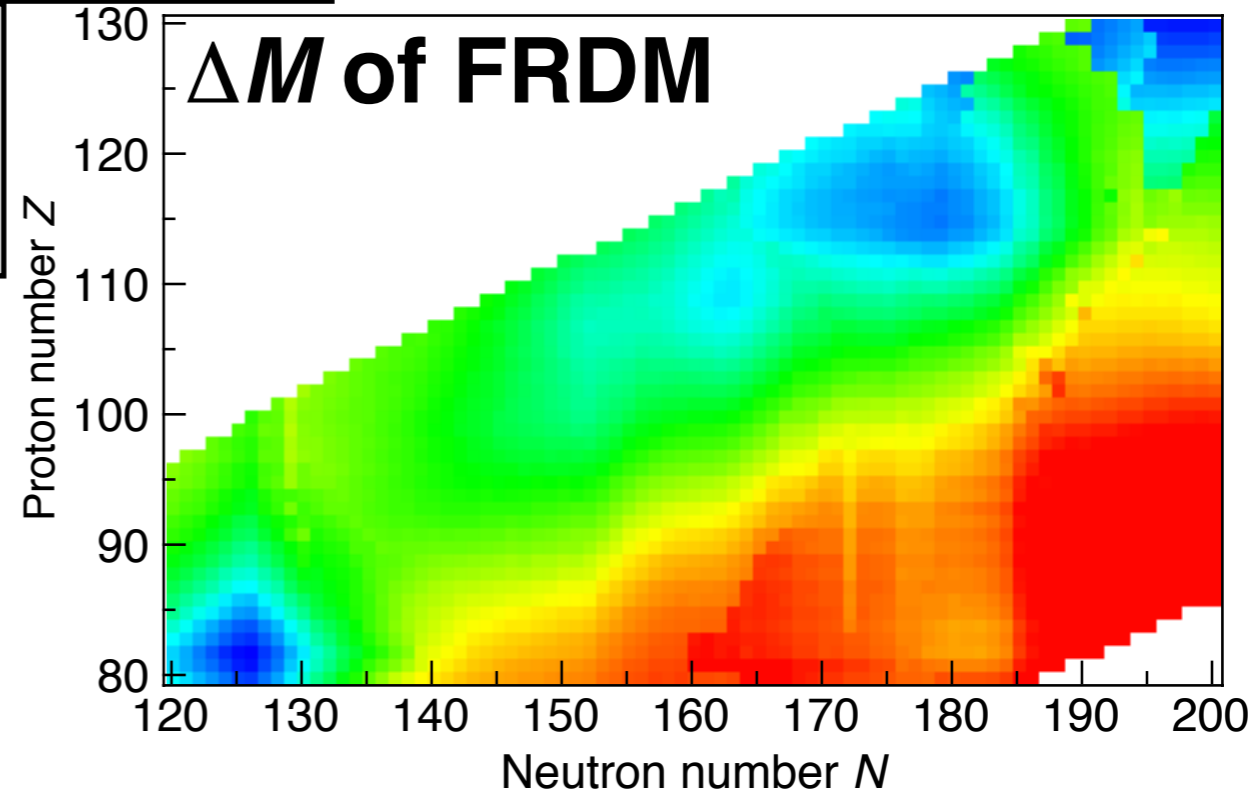
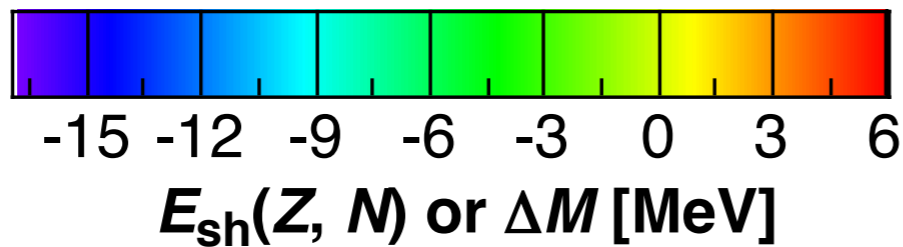
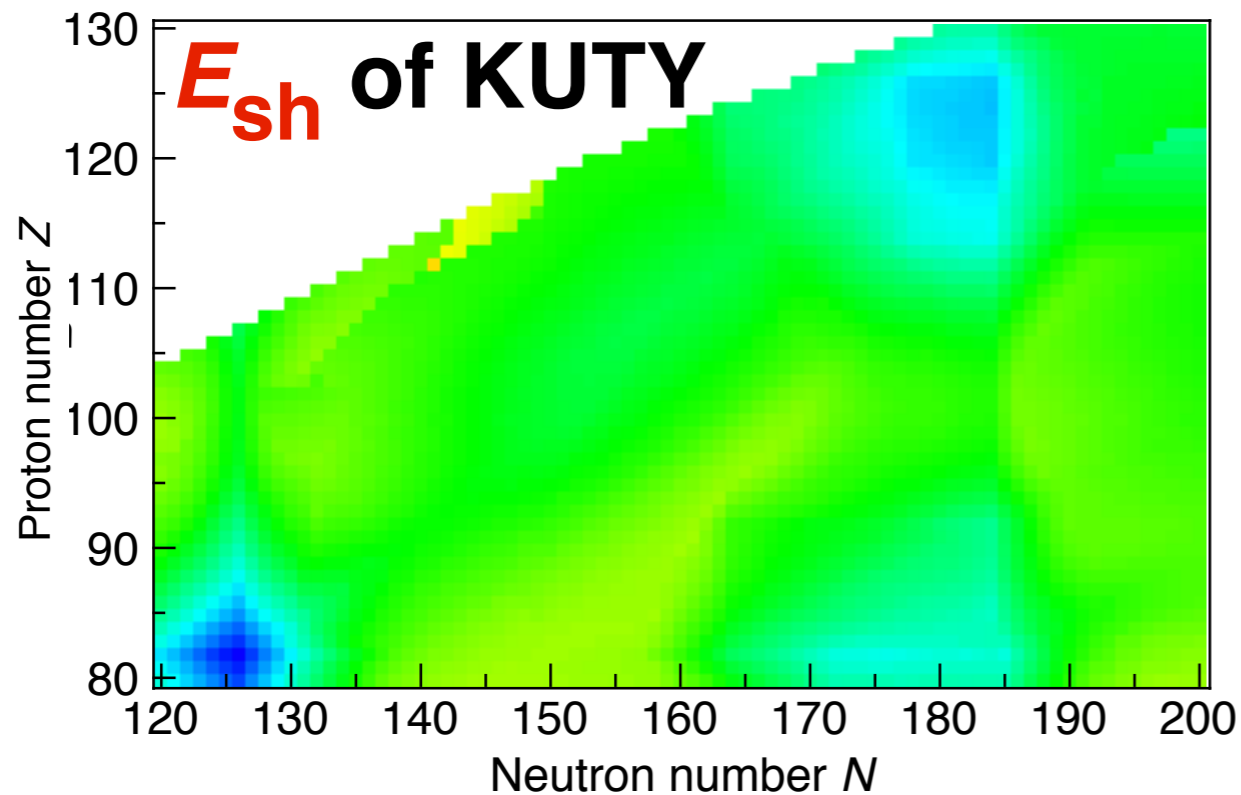
Estimation of absolute values of beam energy depends on
 (unknown) masses of compound nuclei

$$\Delta M(Z, N) = M_{\text{FRDM}}(Z, N) - (M_{\text{gross}}(Z, N) + M_{\text{eo}}(Z, N))$$

ETFSI

$M_{\text{gross}}(Z, N)$: KUTY gross term

$M_{\text{eo}}(Z, N)$: KUTY average even-odd term

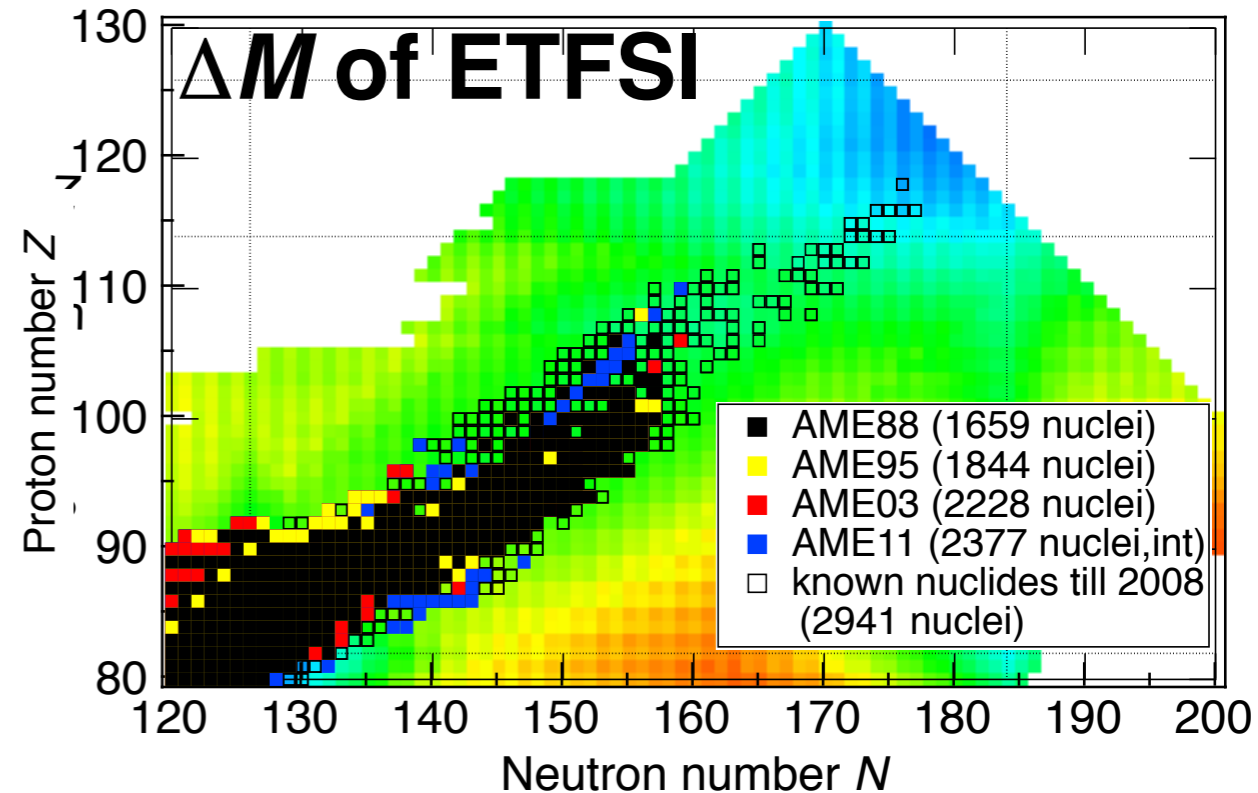
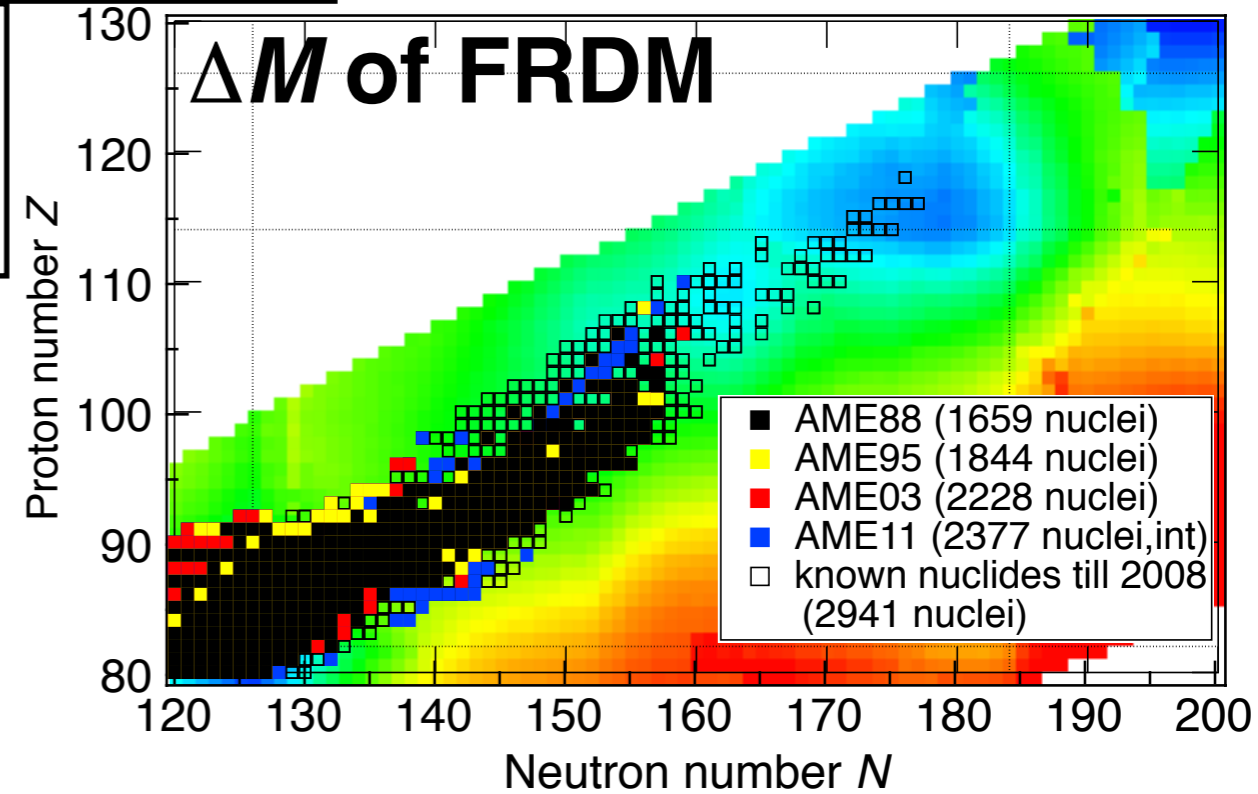
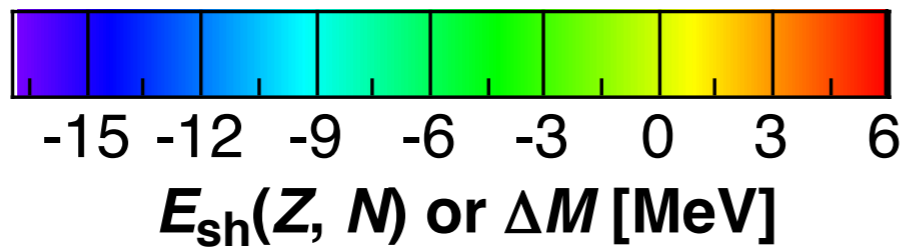
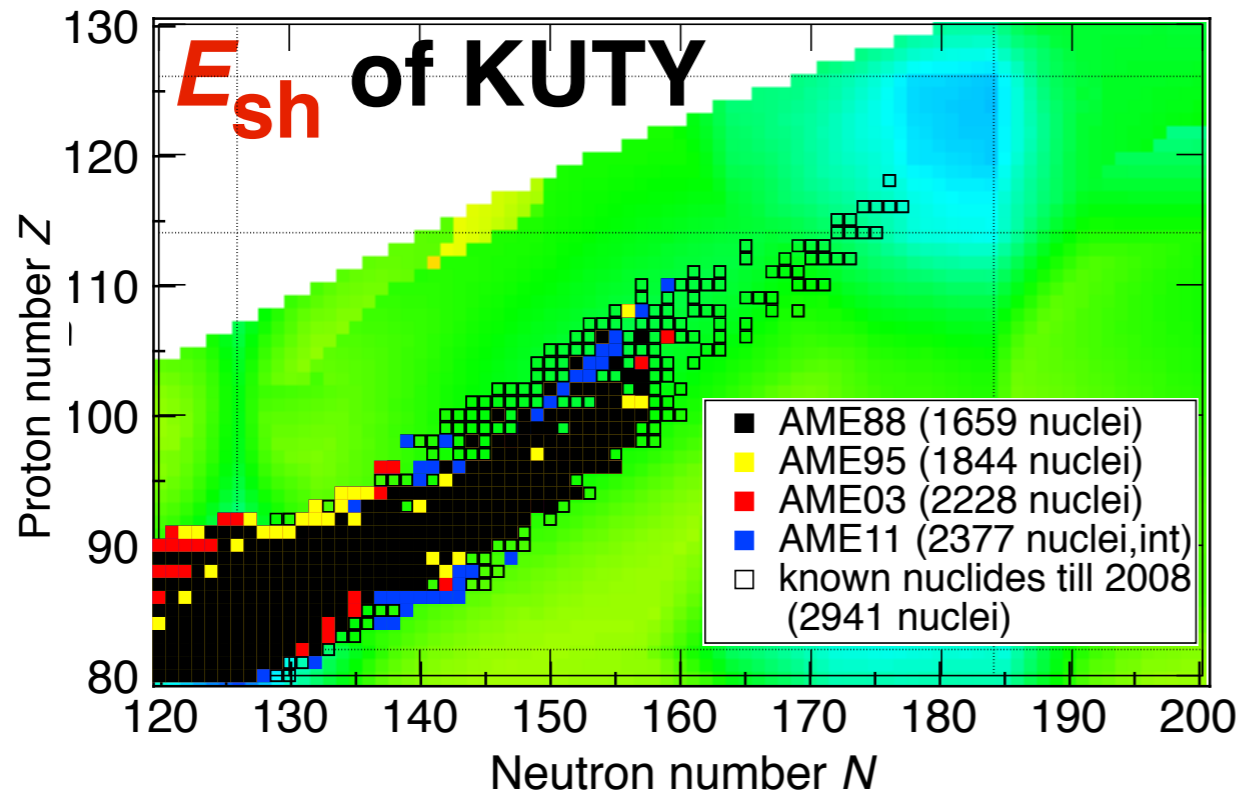


$$\Delta M(Z, N) = M_{\text{FRDM}}(Z, N) - (M_{\text{gross}}(Z, N) + M_{\text{eo}}(Z, N))$$

ETFSI

$M_{\text{gross}}(Z, N)$: KUTY gross term

$M_{\text{eo}}(Z, N)$: KUTY average even-odd term

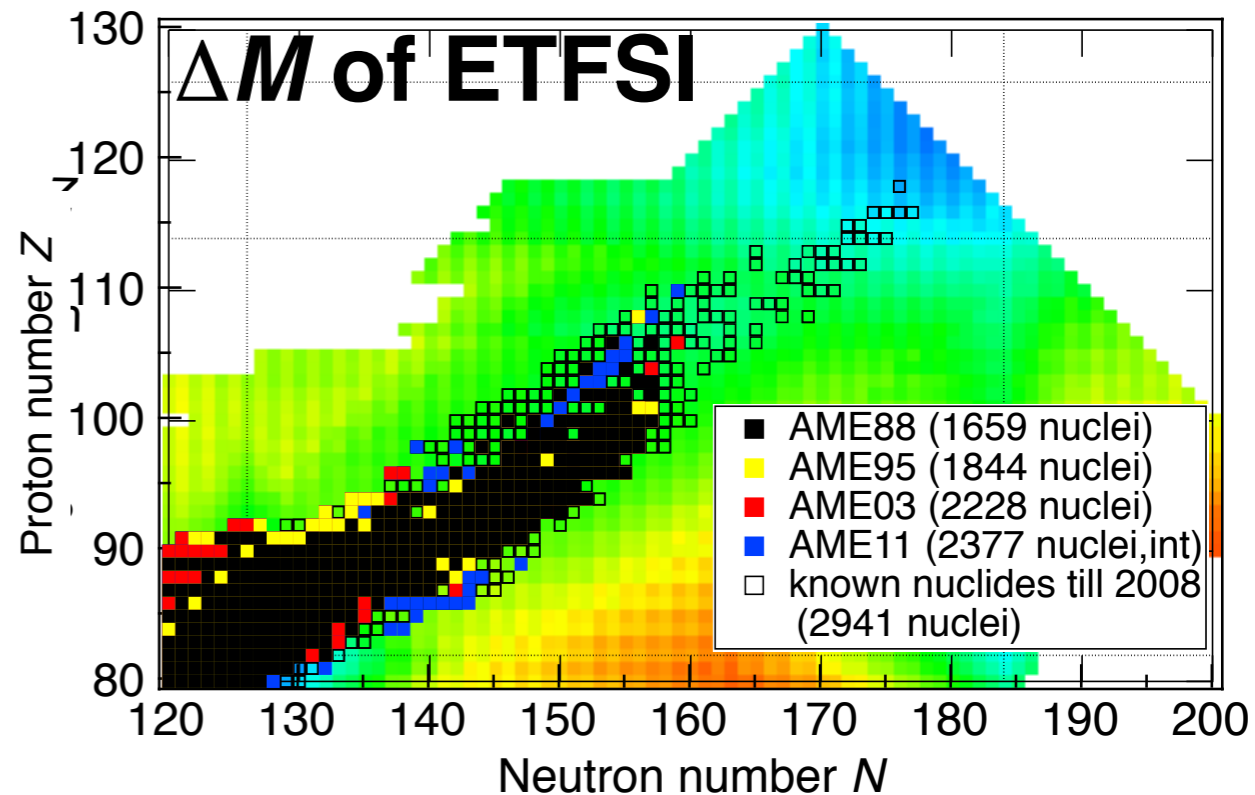
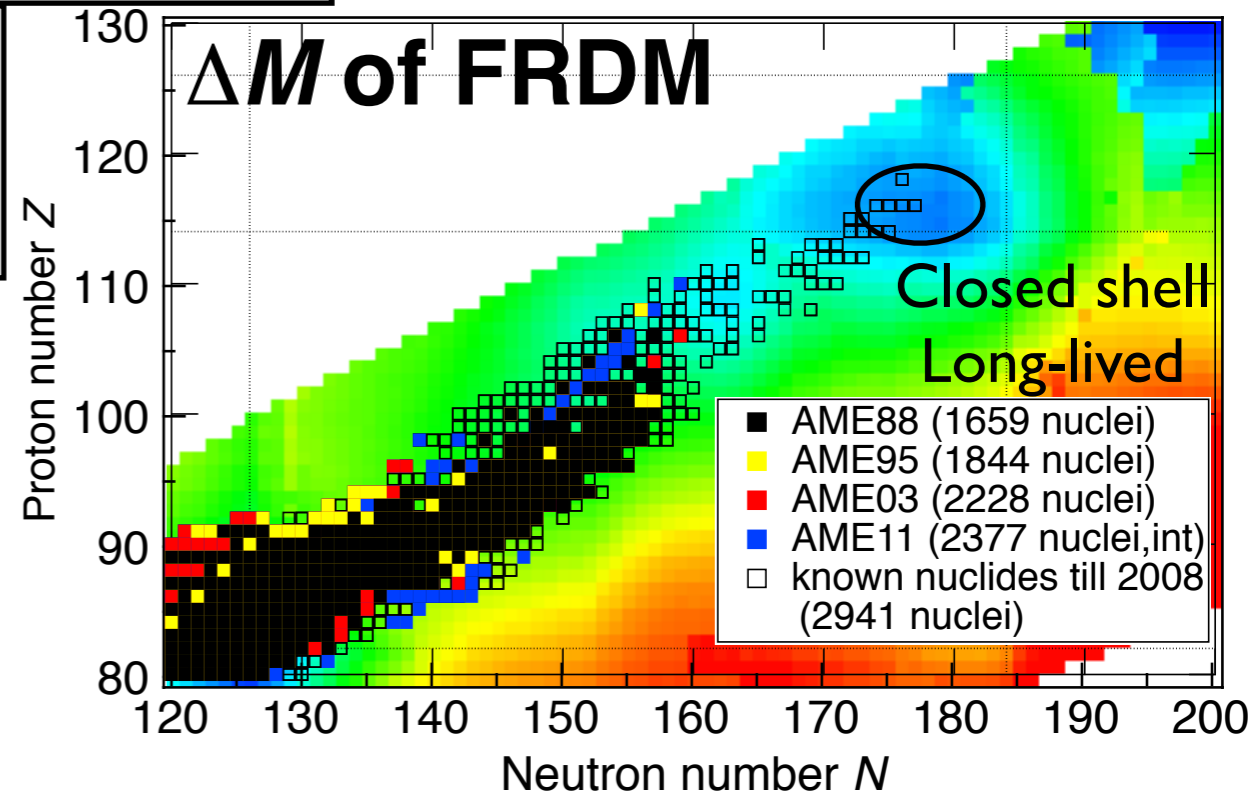
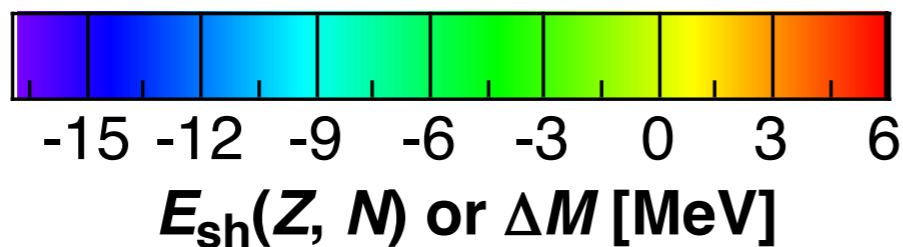
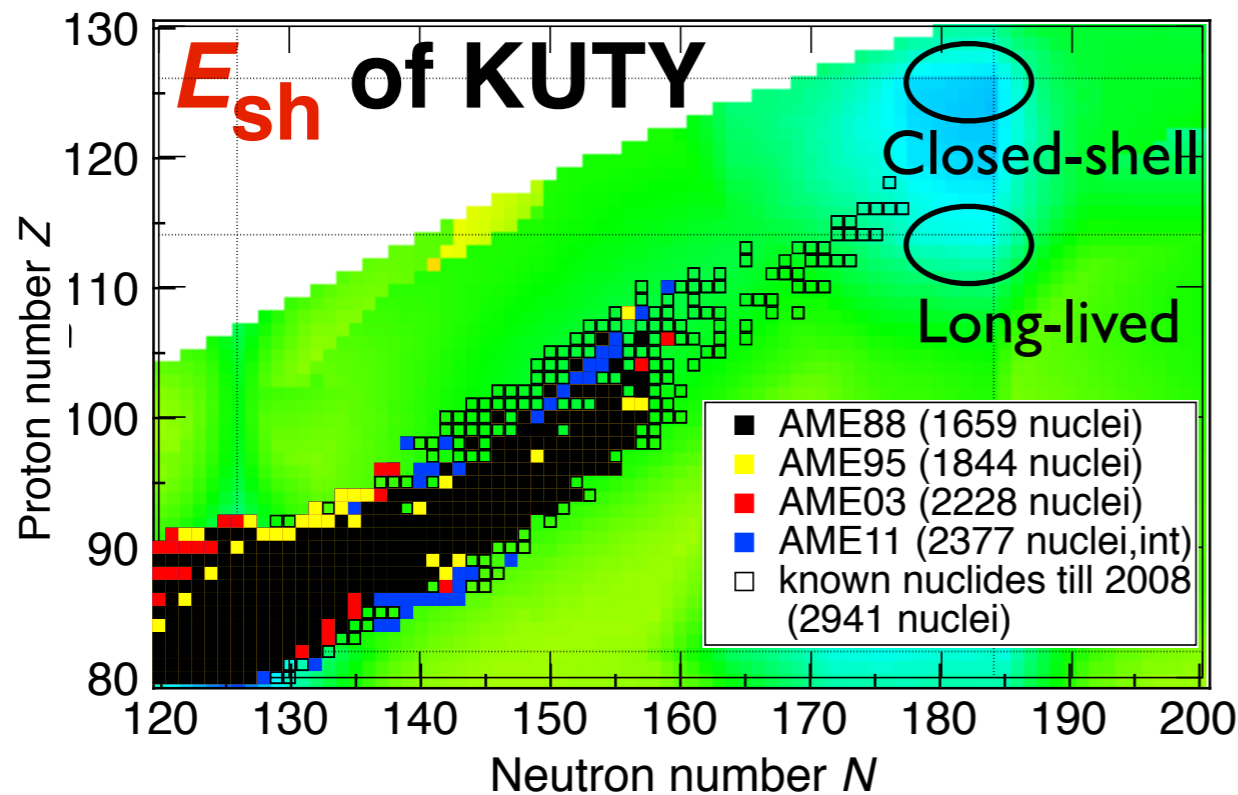


$$\Delta M(Z, N) = M_{\text{FRDM}}(Z, N) - (M_{\text{gross}}(Z, N) + M_{\text{eo}}(Z, N))$$

ETFSI

$M_{\text{gross}}(Z, N)$: KUTY gross term

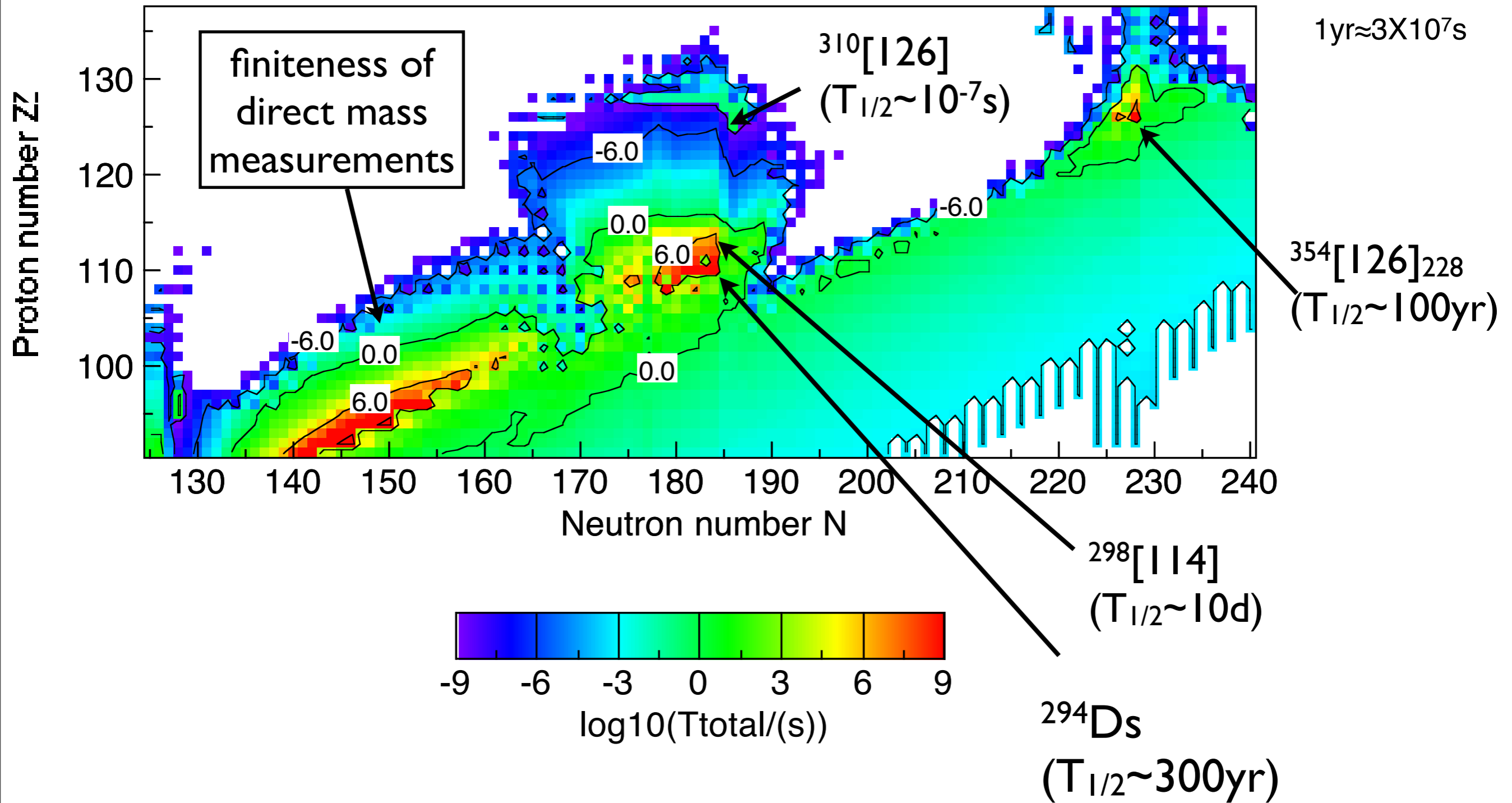
$M_{\text{eo}}(Z, N)$: KUTY average even-odd term



Theoretical total half-lives (α, β, p, sf)

KTUY+decay models

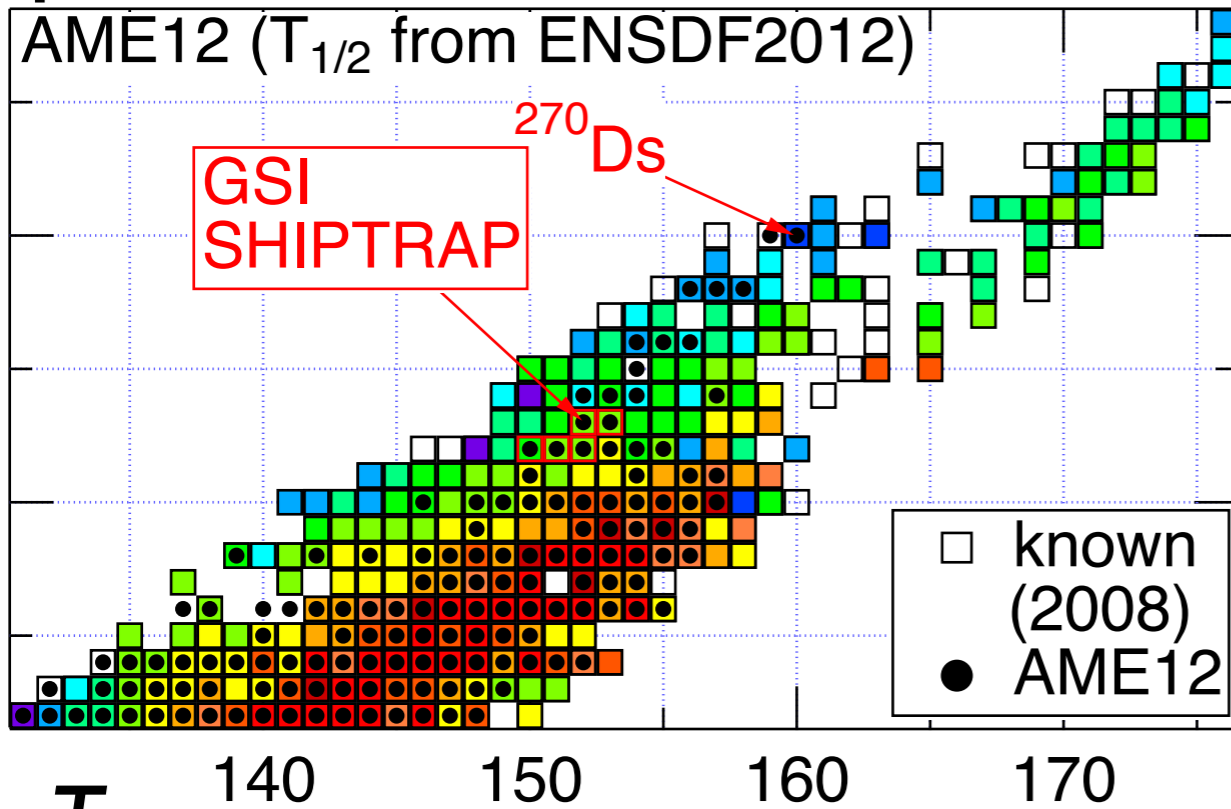
1yr $\approx 3 \times 10^7$ s



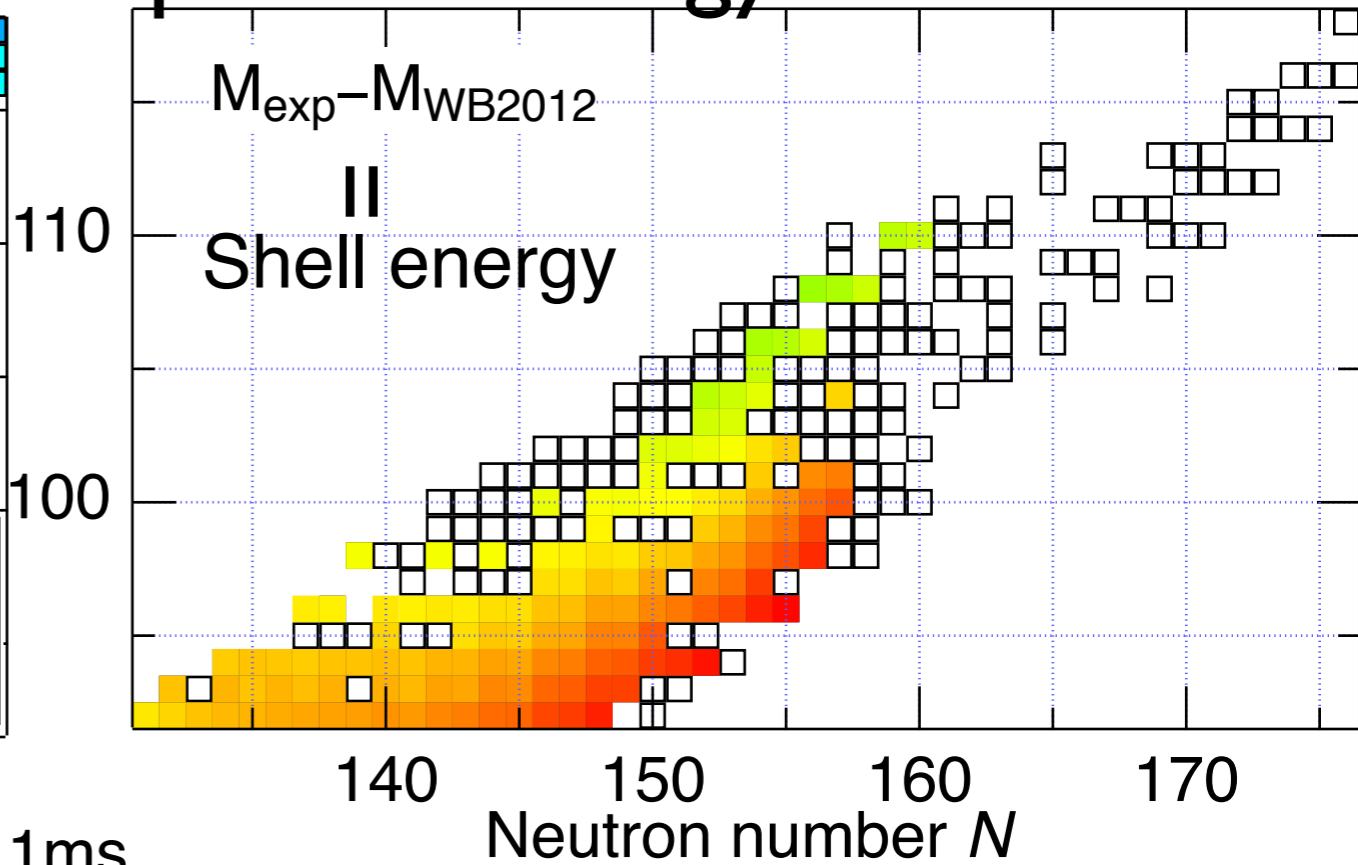
(Long-lived superheavy nuclei are located near the β -stability line)

Total half-lives (α, β, p, sf)

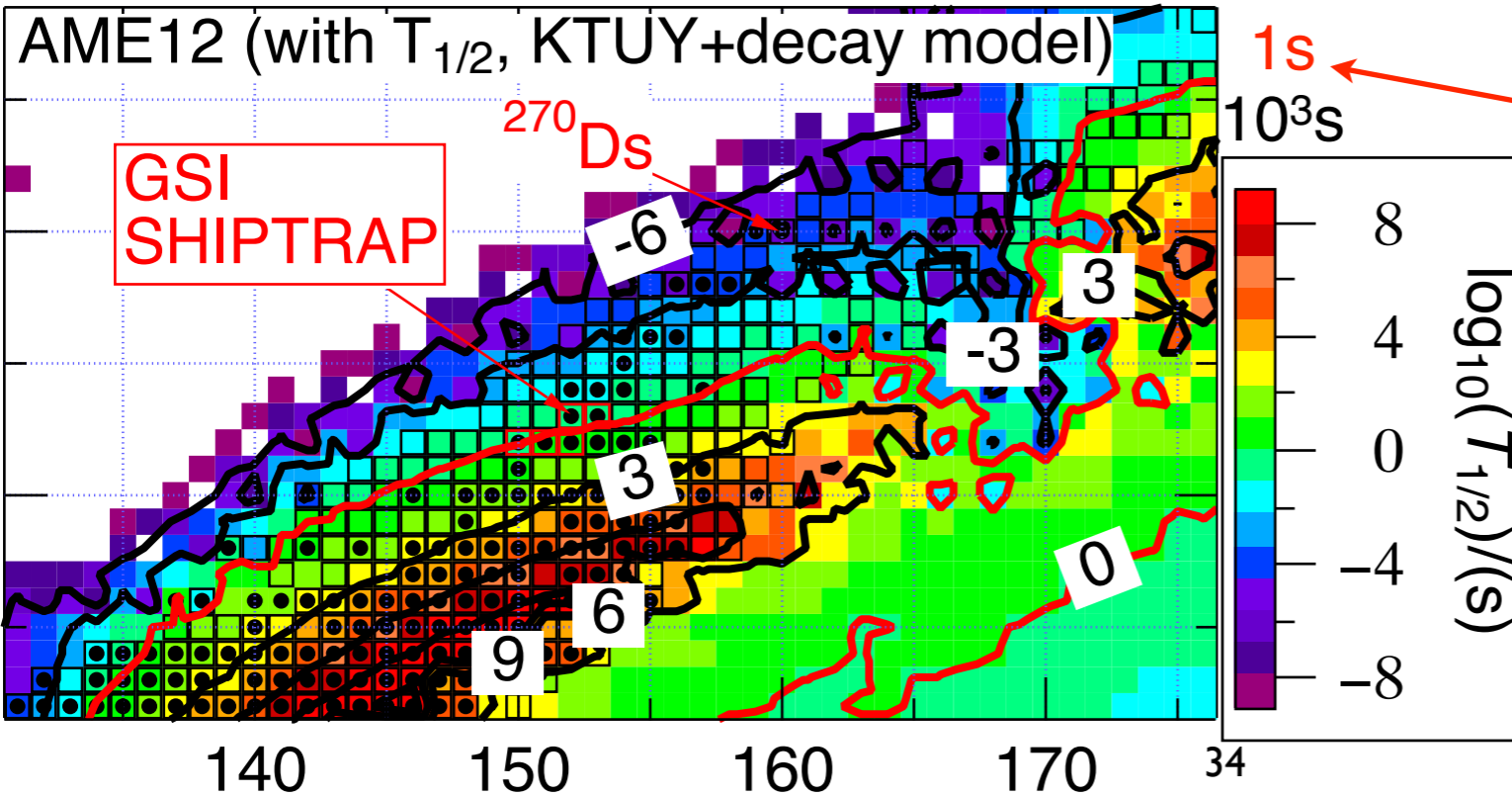
Exp. $T_{1/2}$



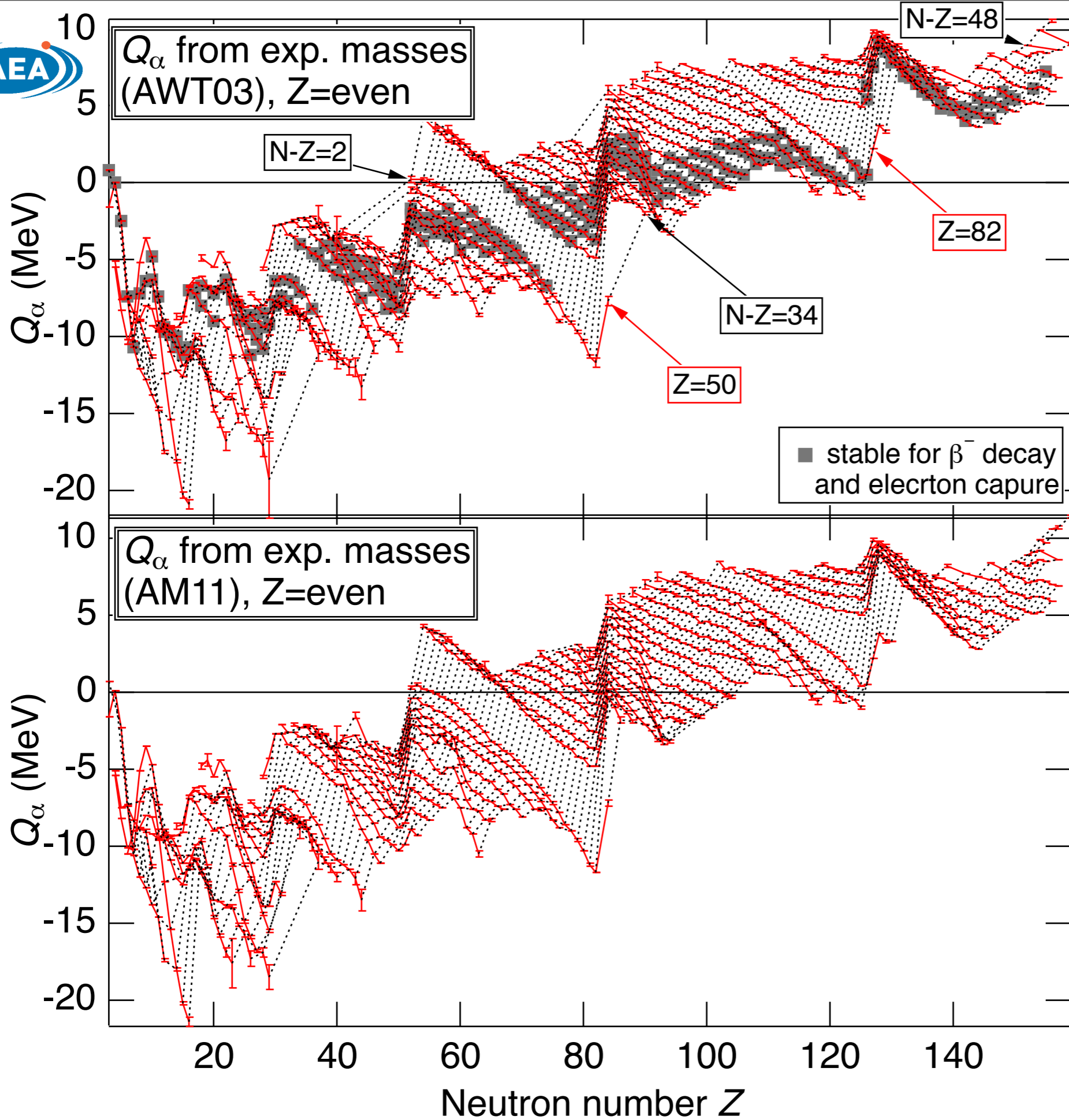
'Exp.' shell Energy



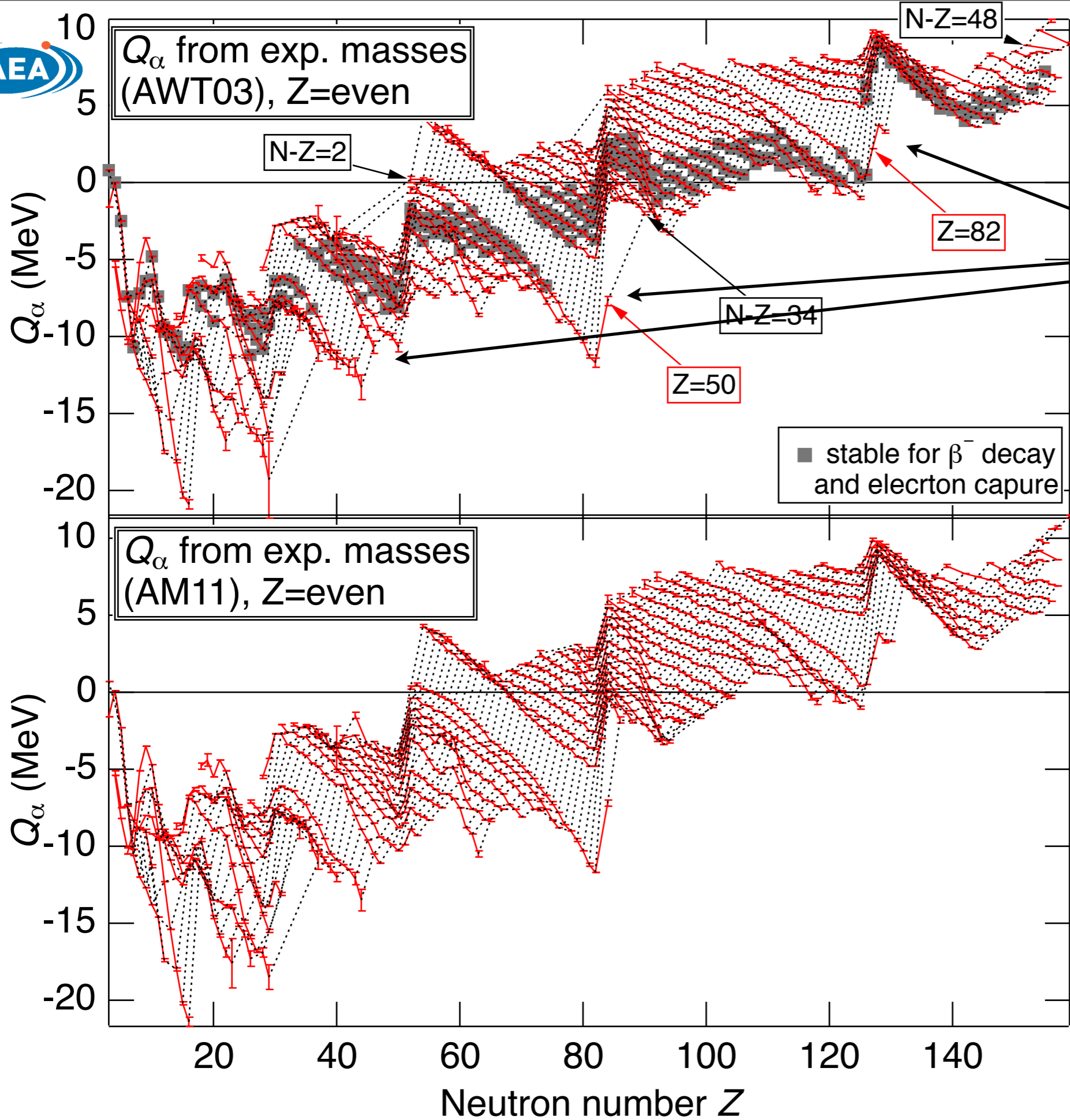
Th. $T_{1/2}$



Nuclei along the red line (1 sec) are experimentally-provable candidates for RIKENTRAP or SHIPTRAP.



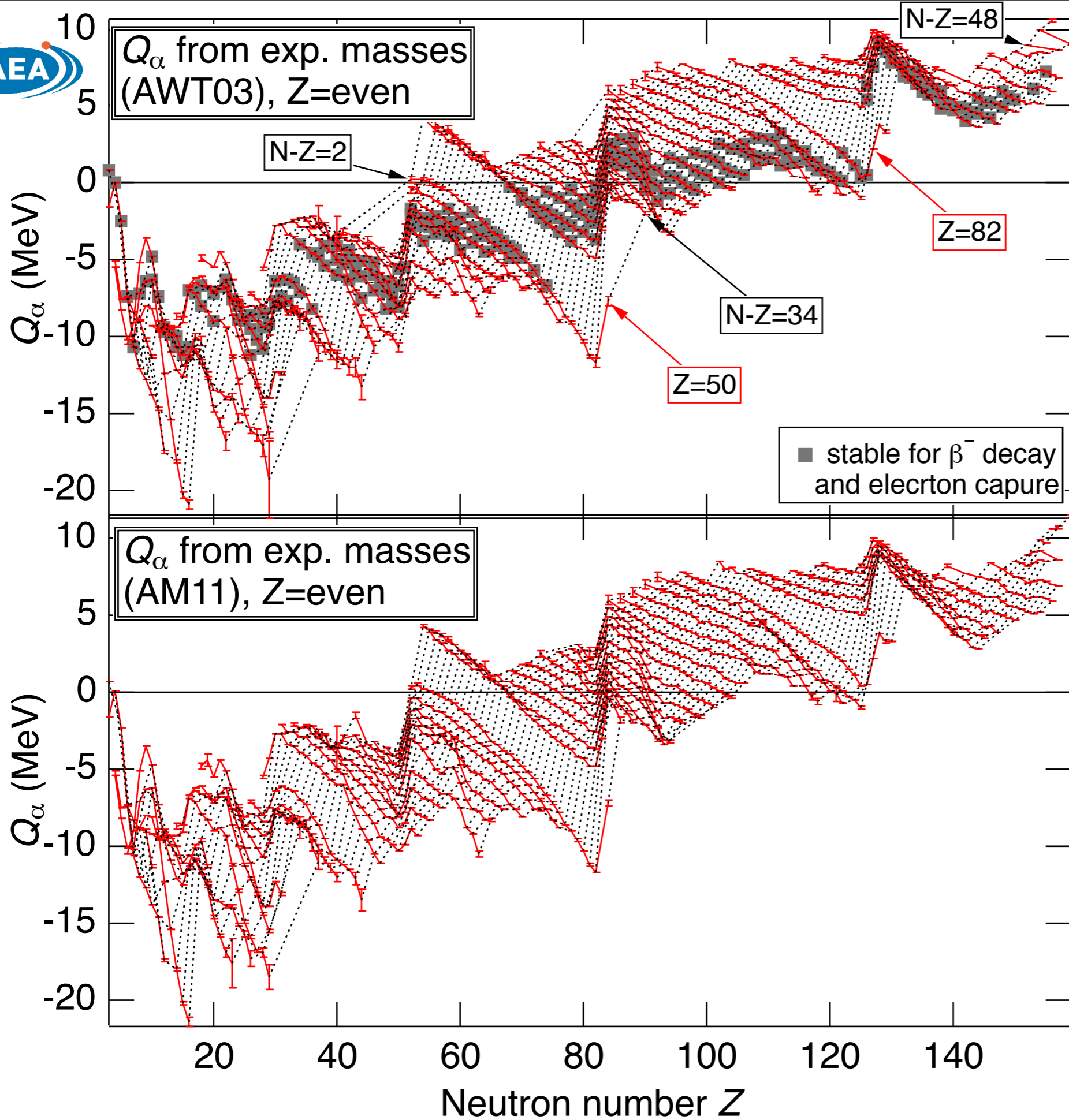
shell gaps are seen



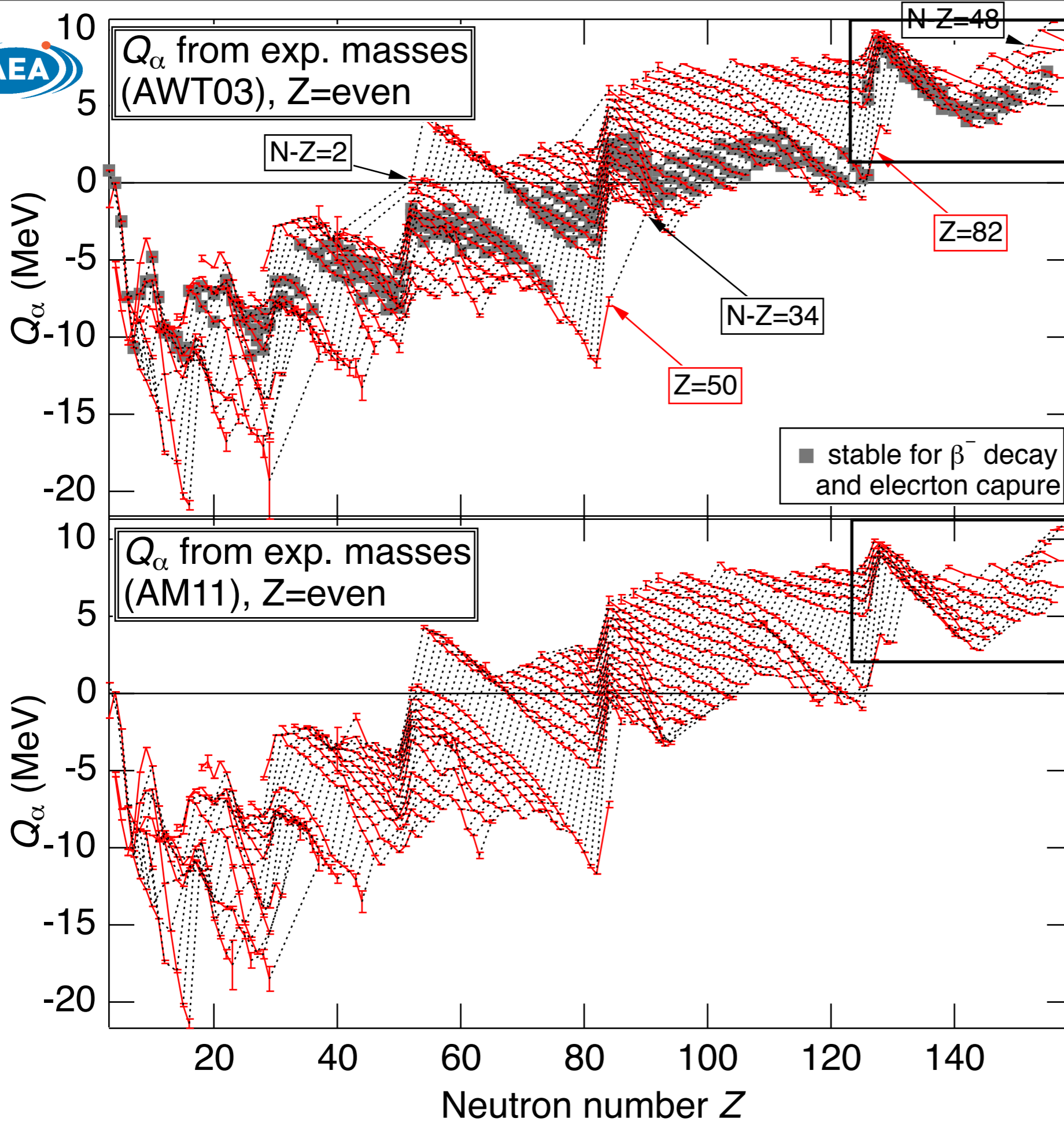
shell gaps are seen

Z=28, 50, 82 shells
as a gap

■ stable for β^- decay
and electron capture

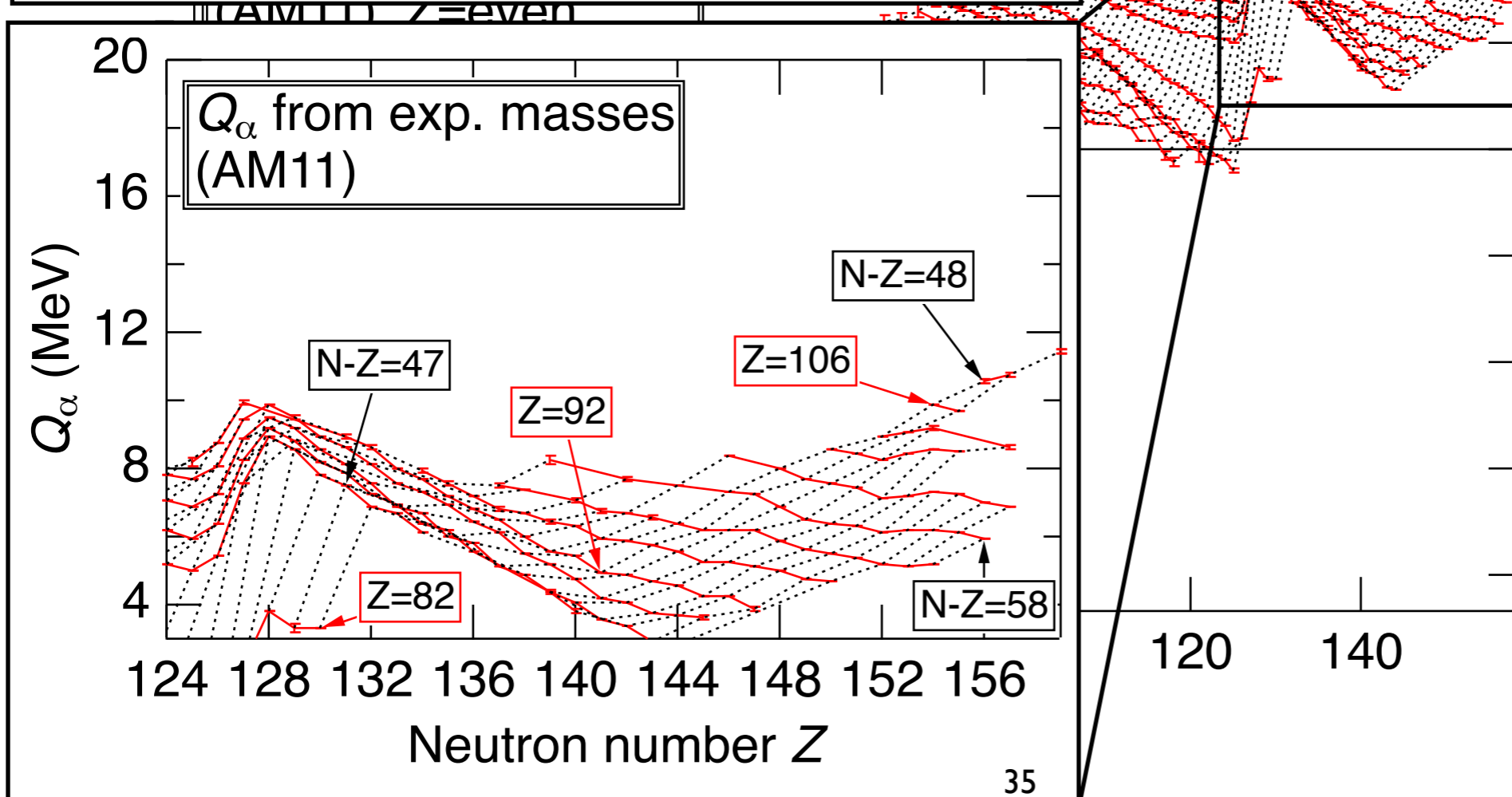
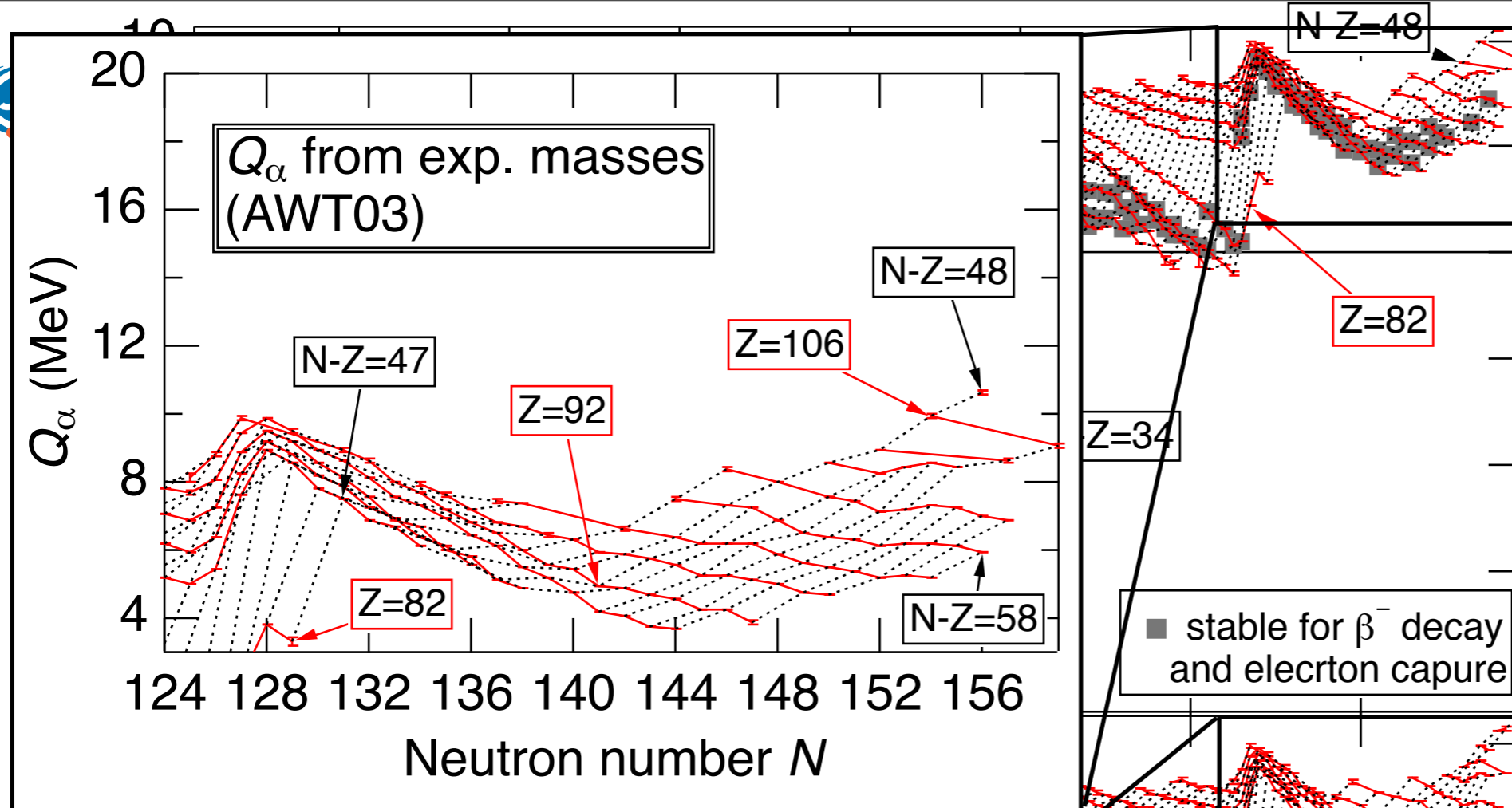


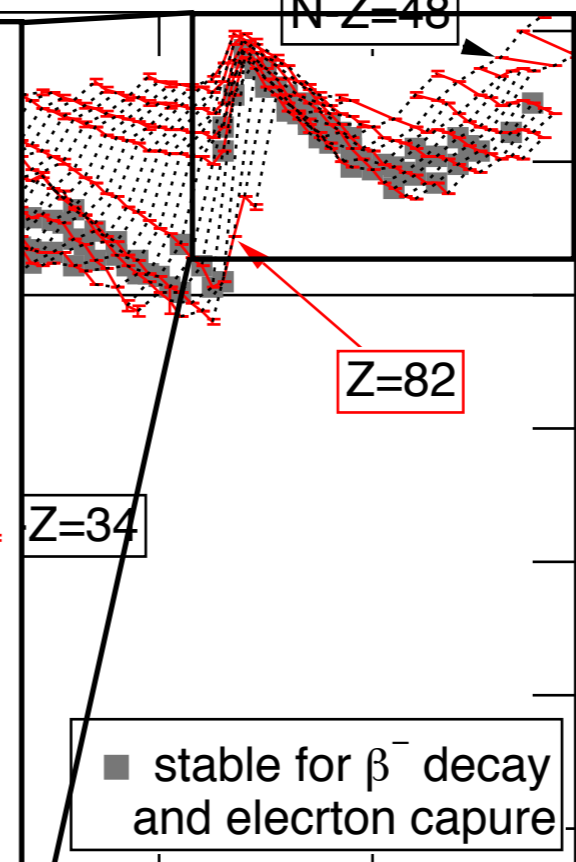
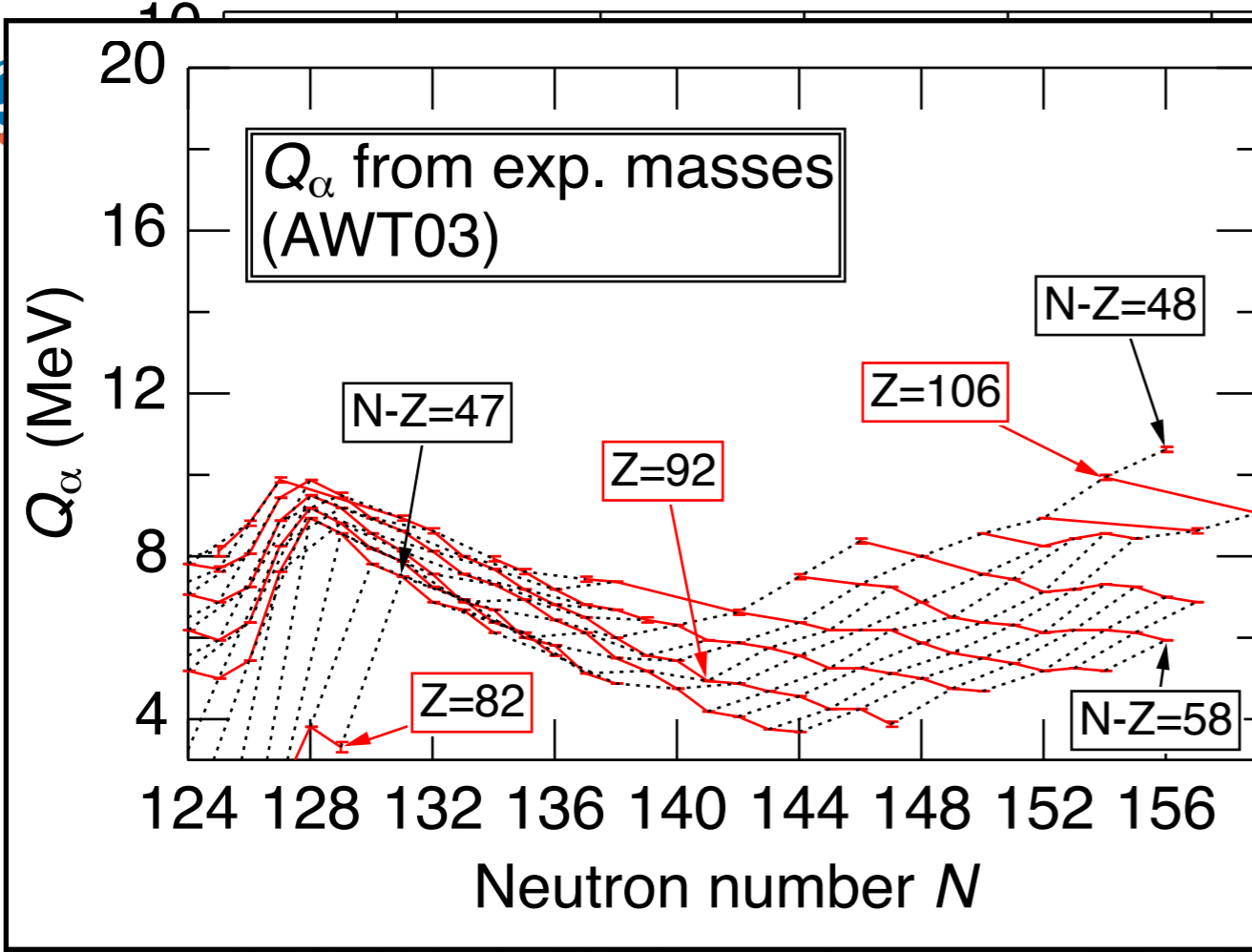
shell gaps are seen



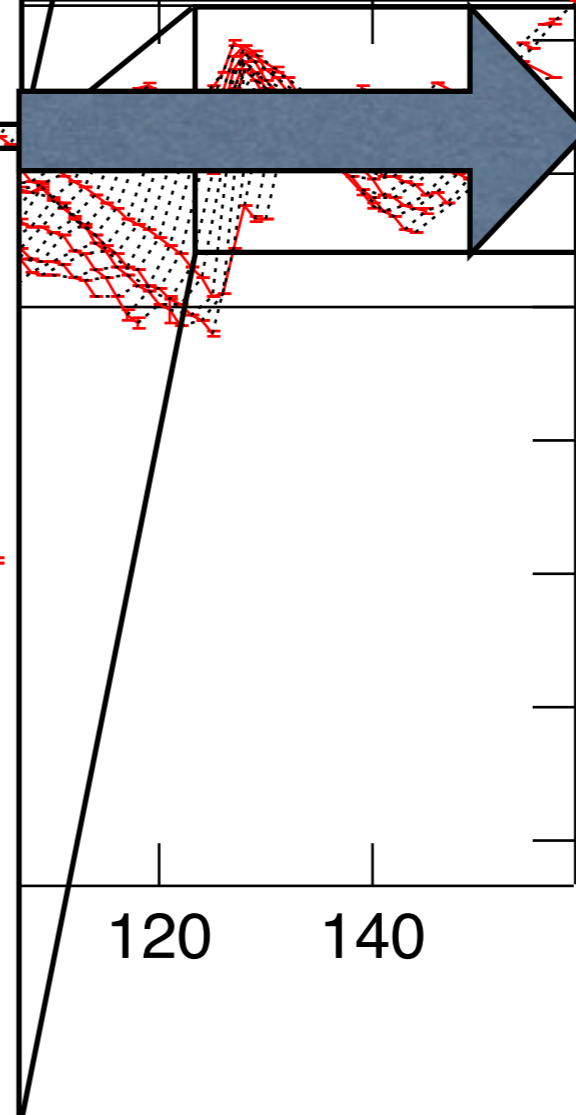
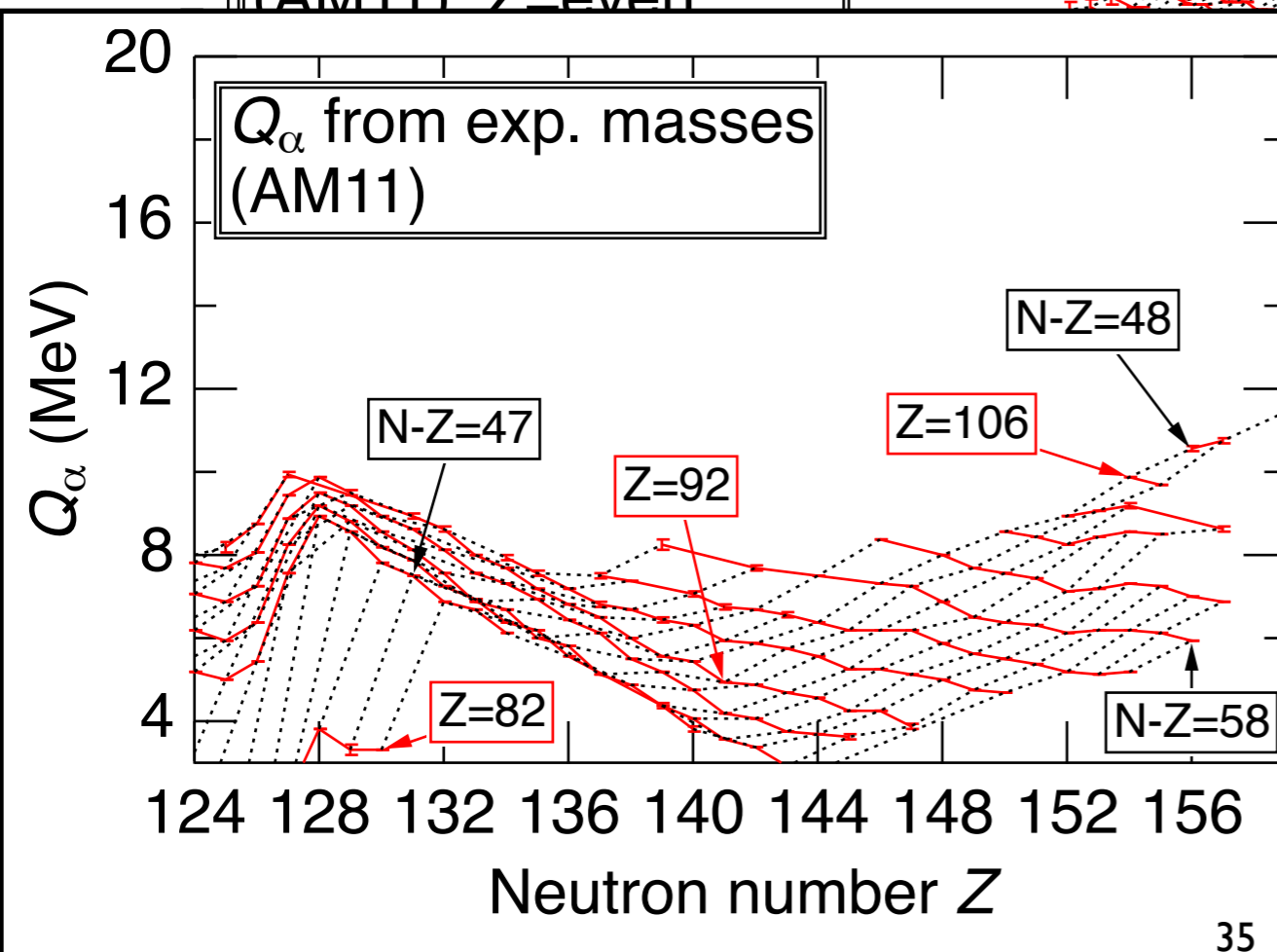
shell gaps are seen

shell gaps are seen

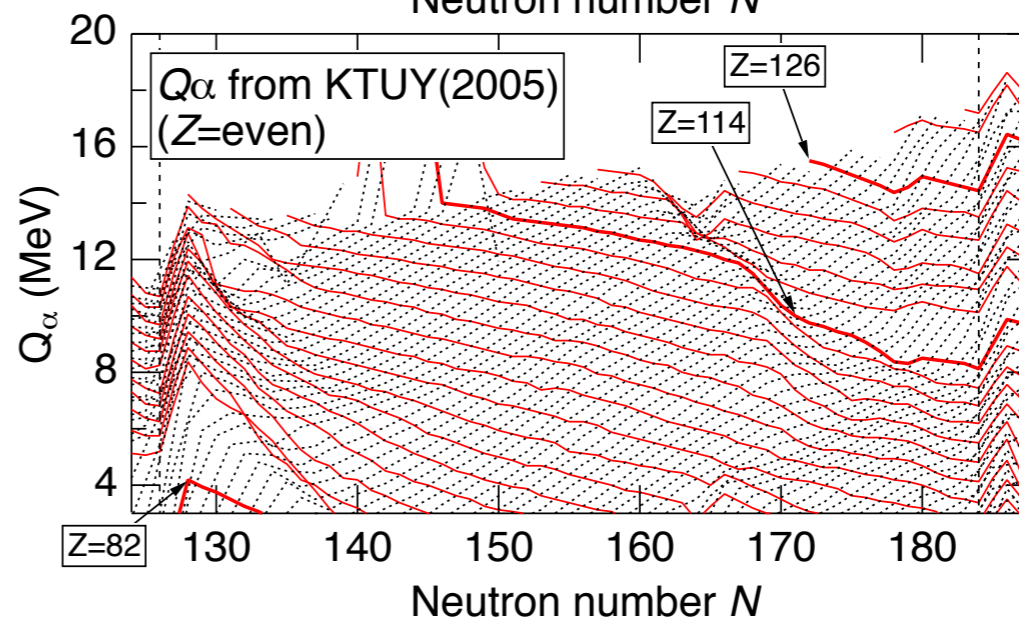
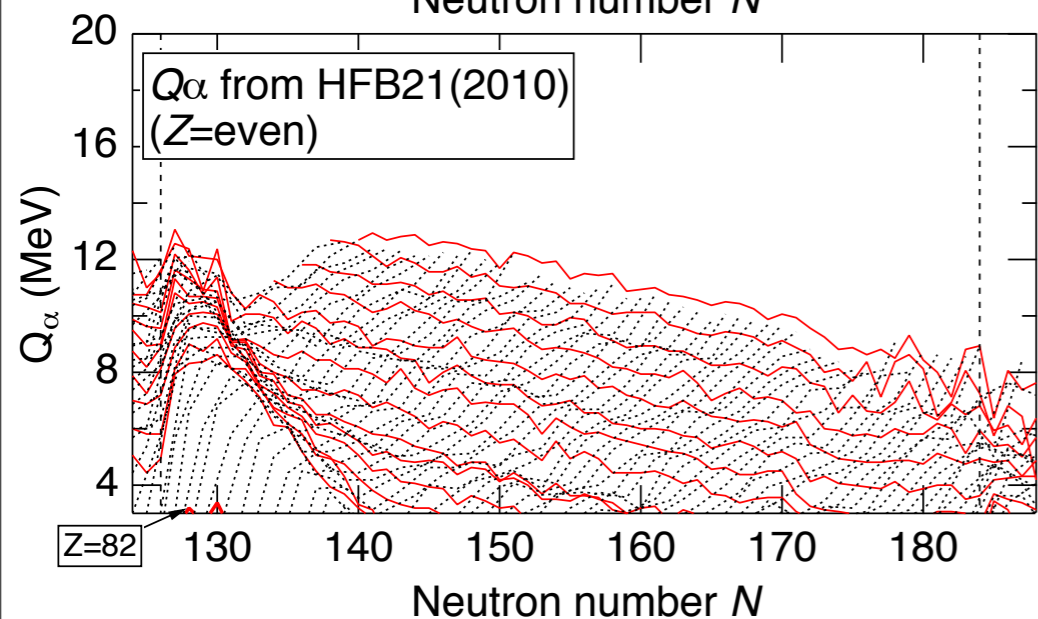
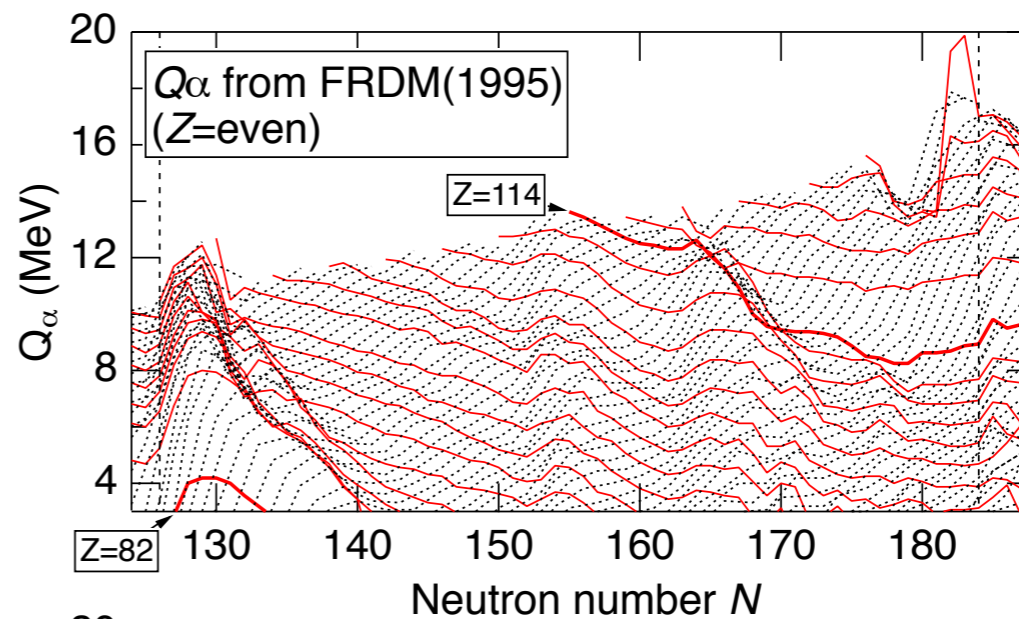
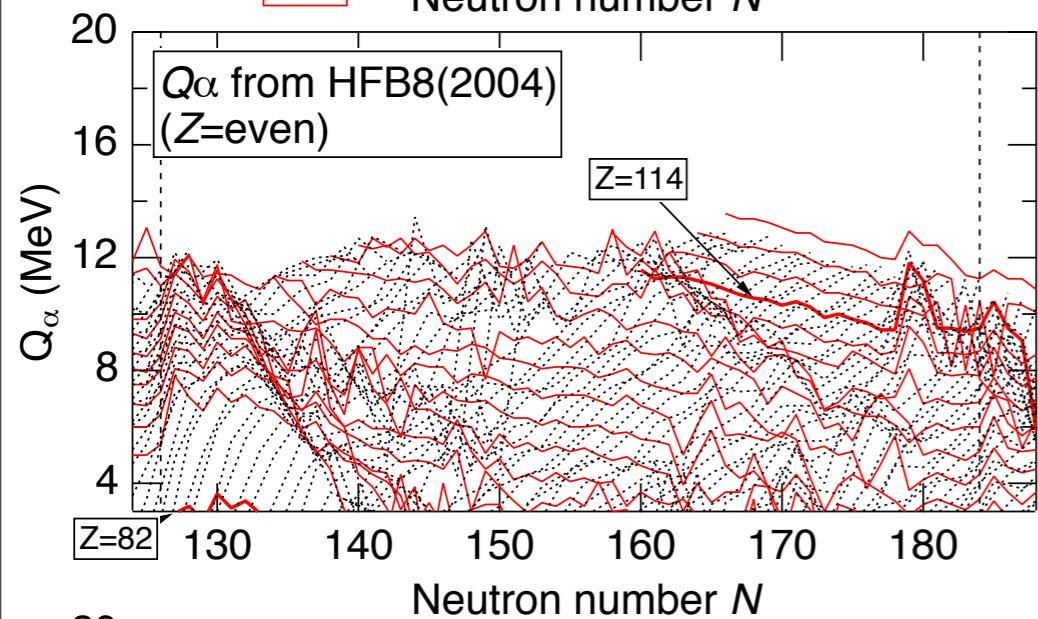
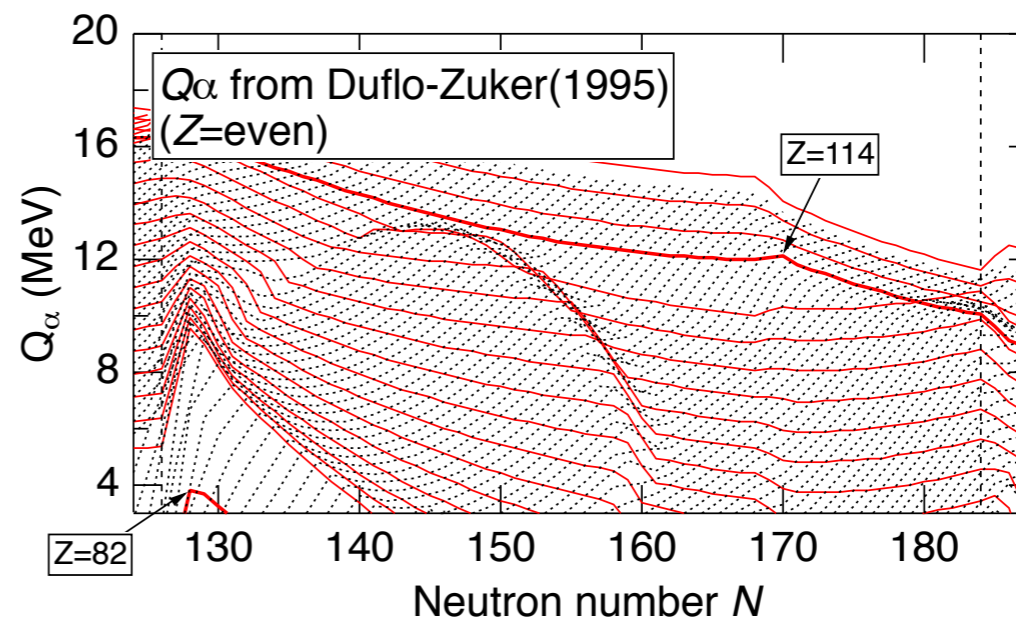
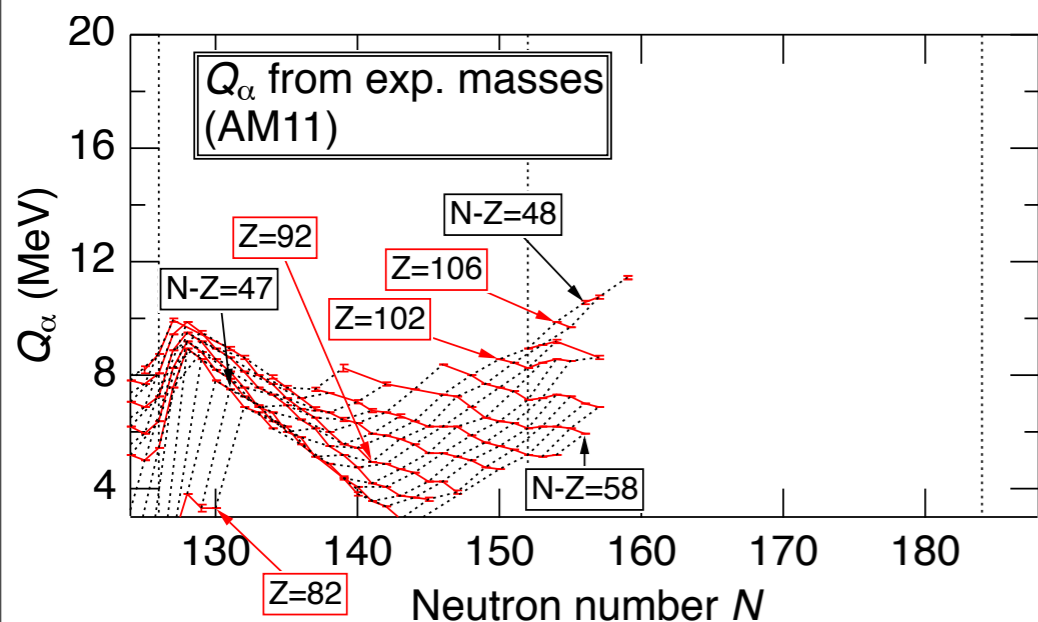


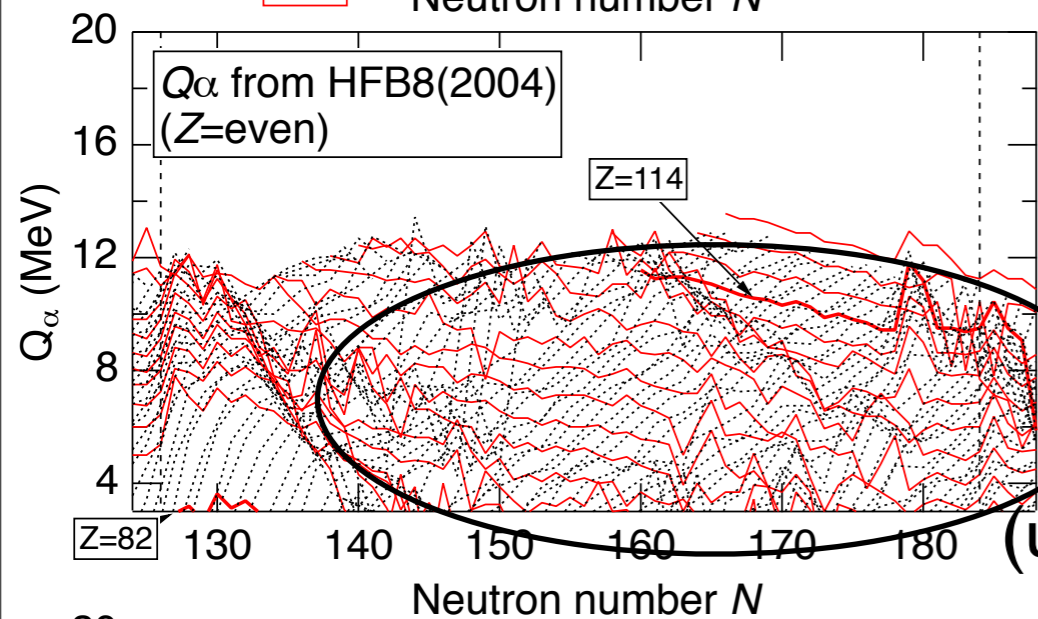
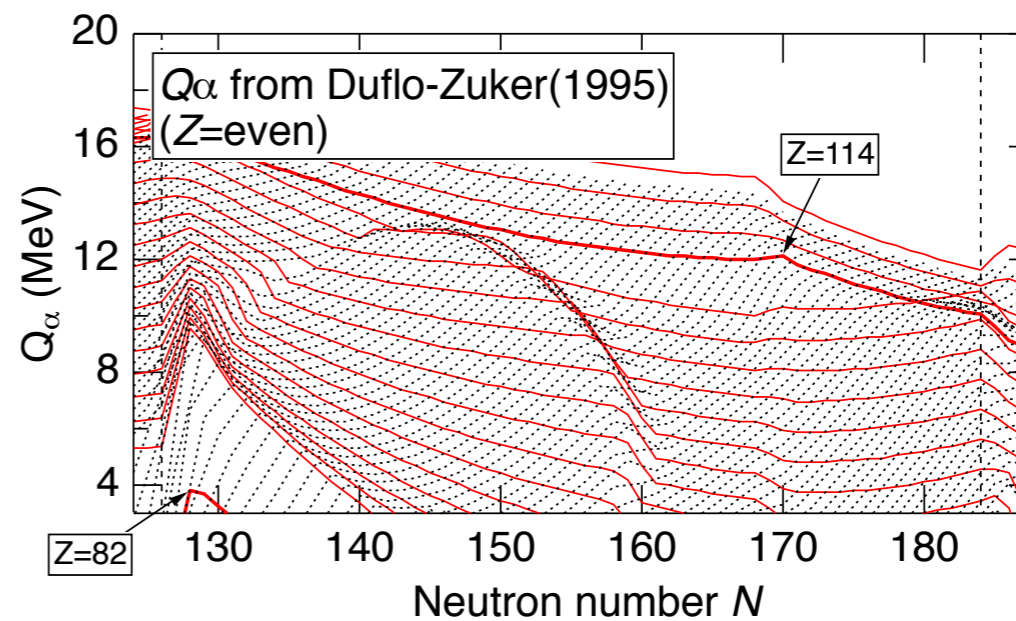
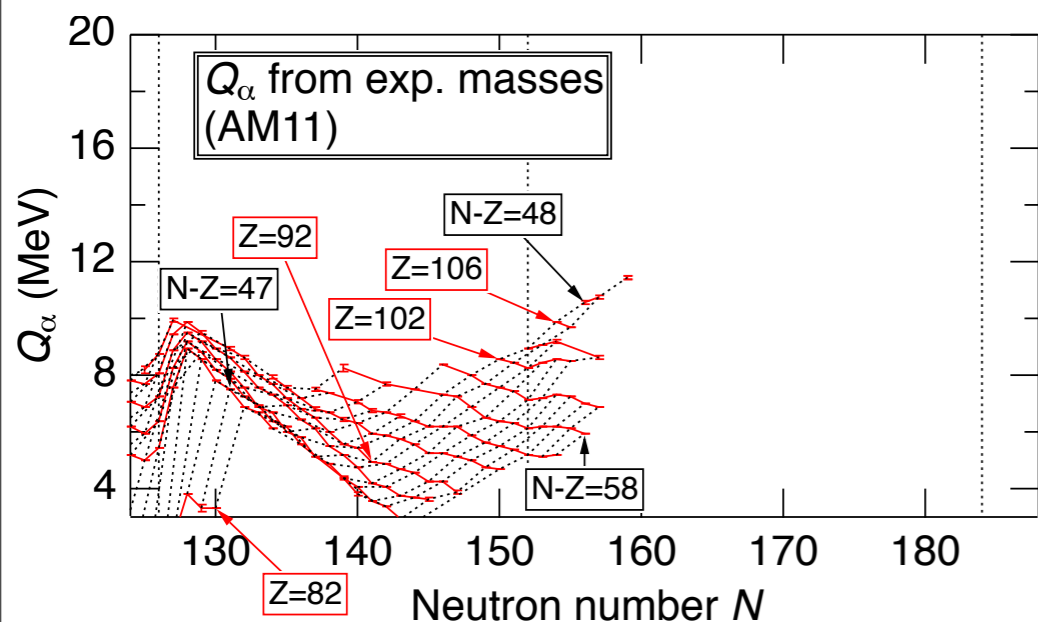


shell gaps are seen

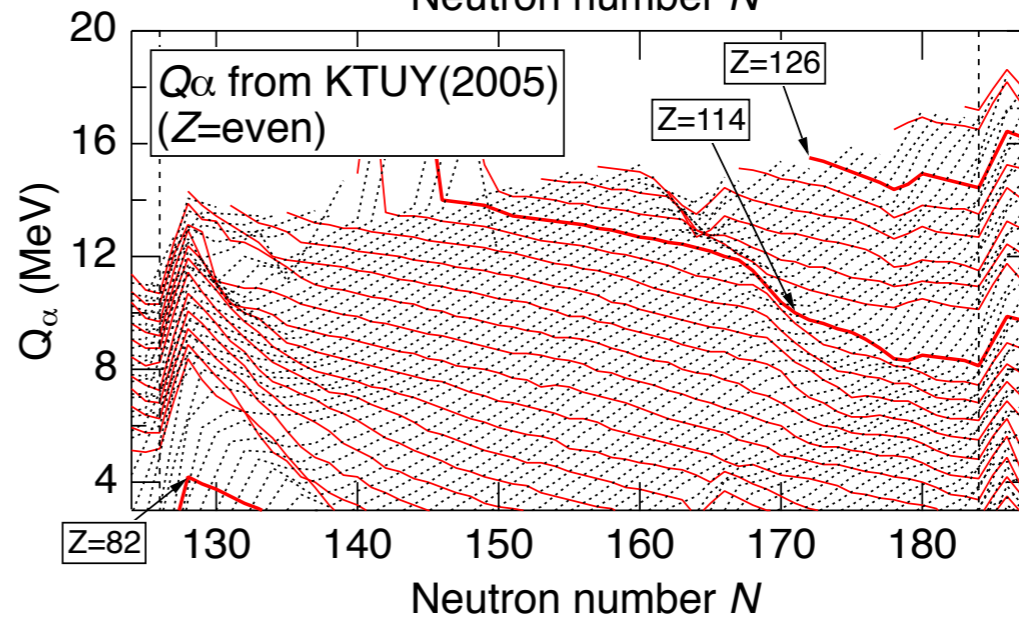
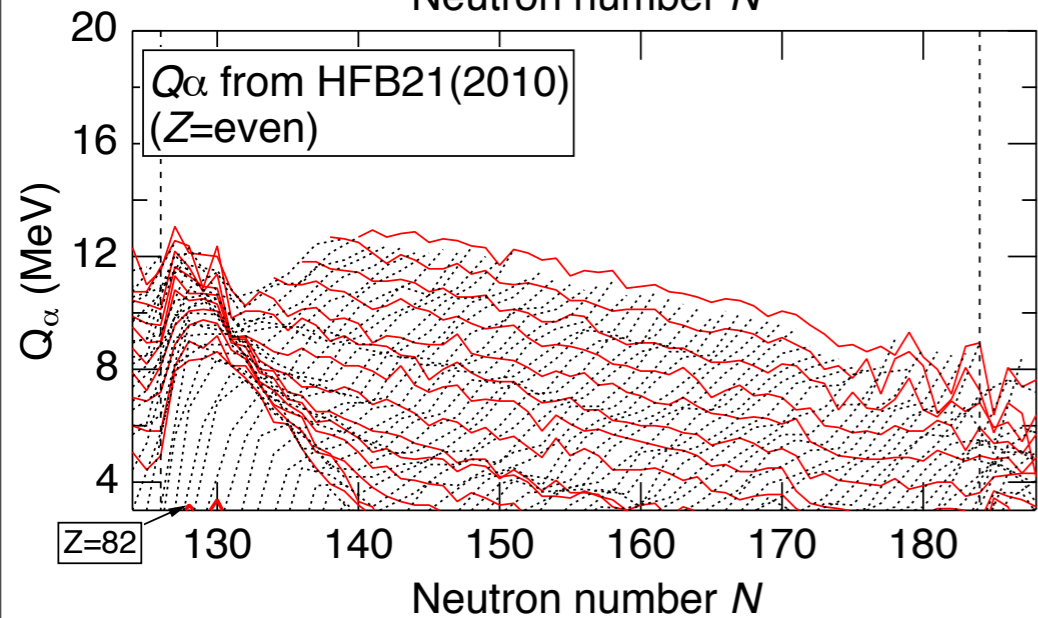
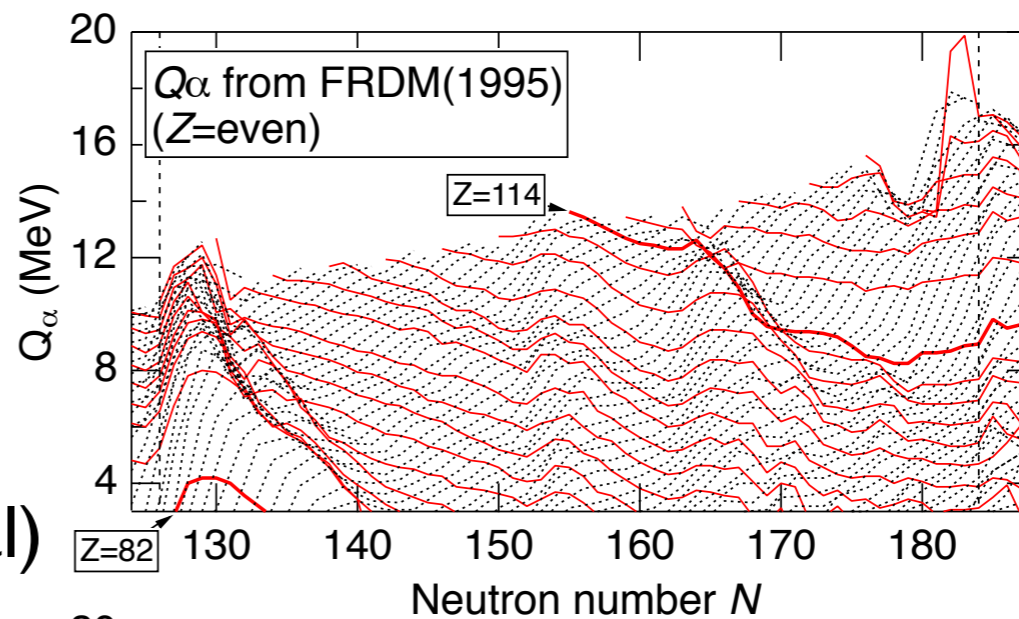


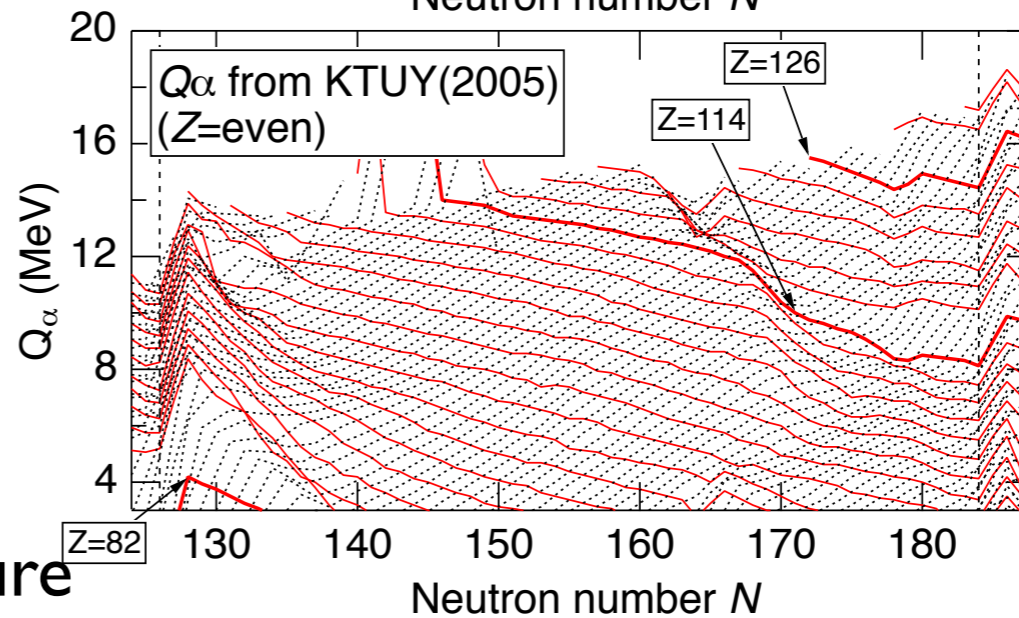
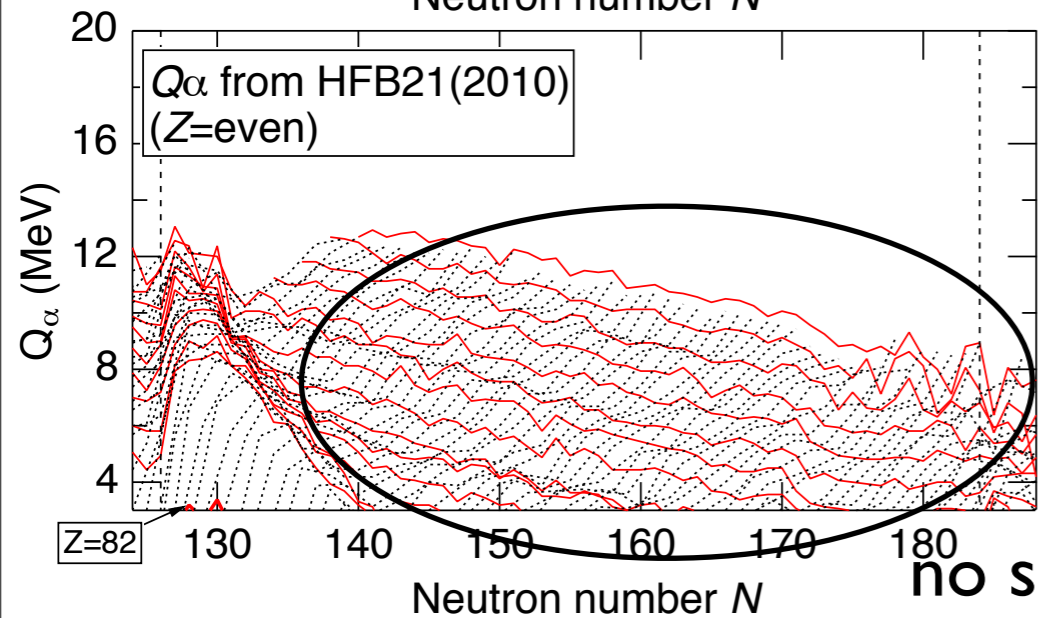
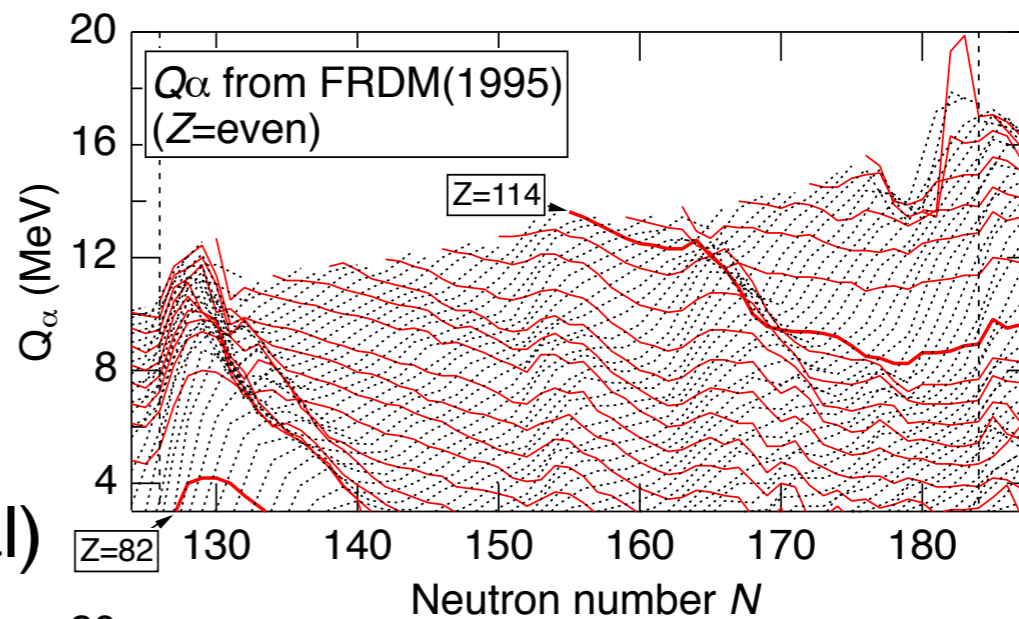
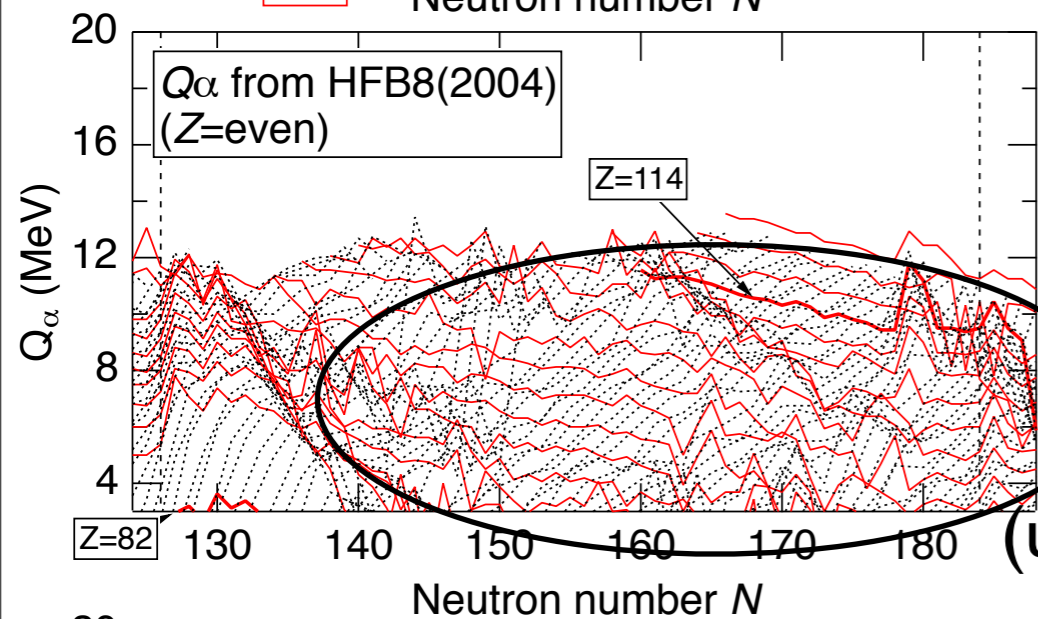
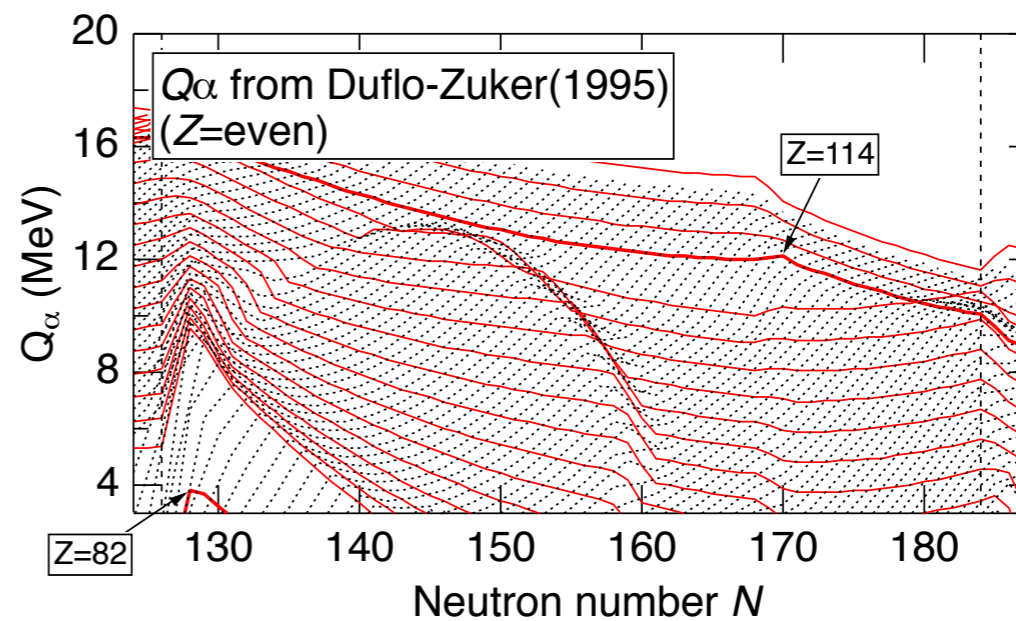
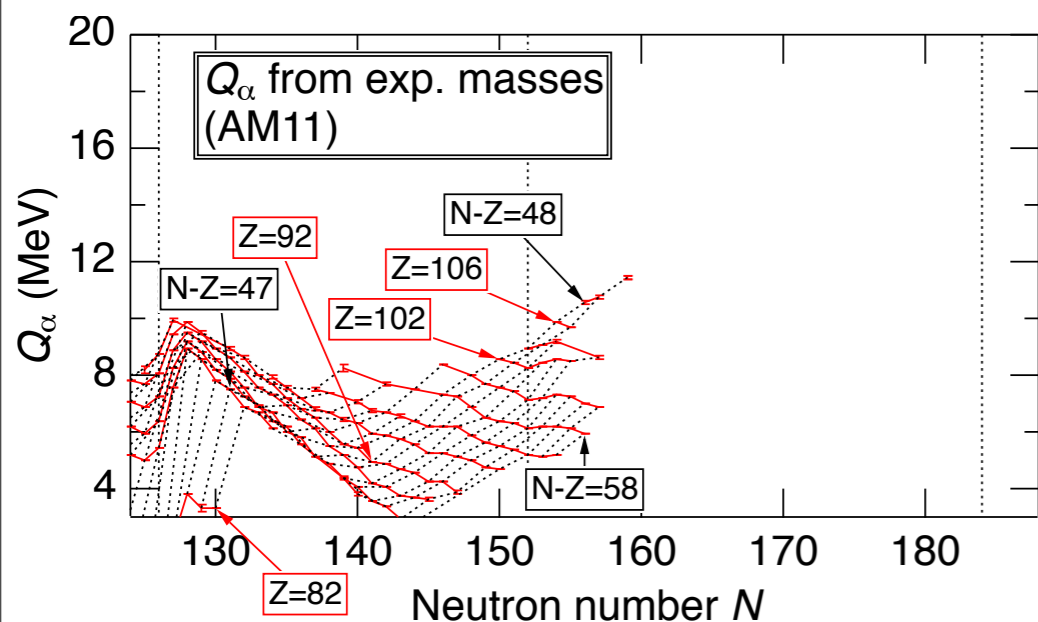
More information is required!





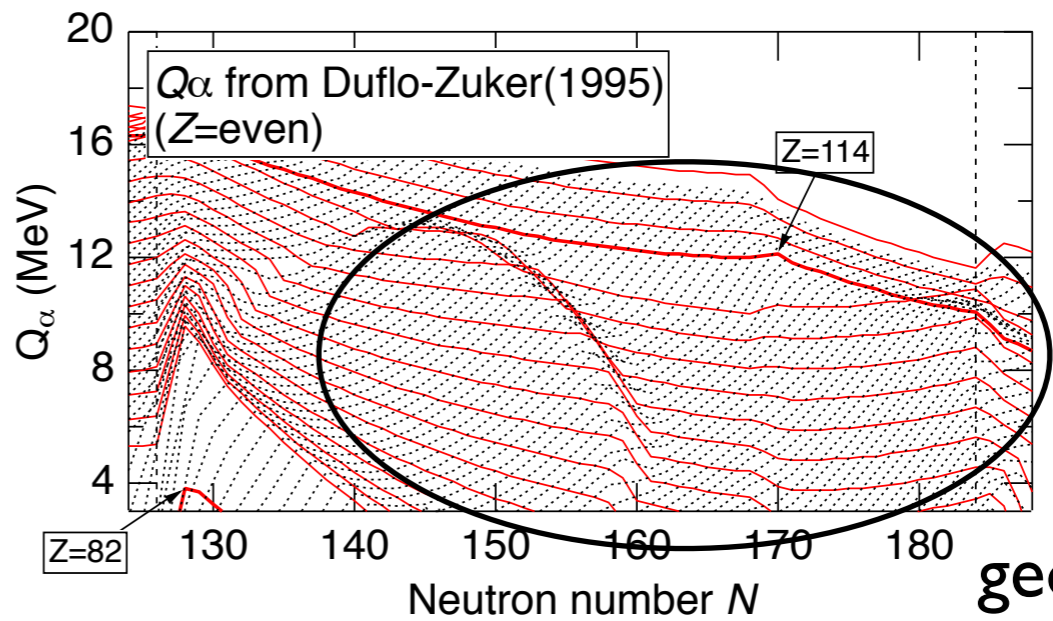
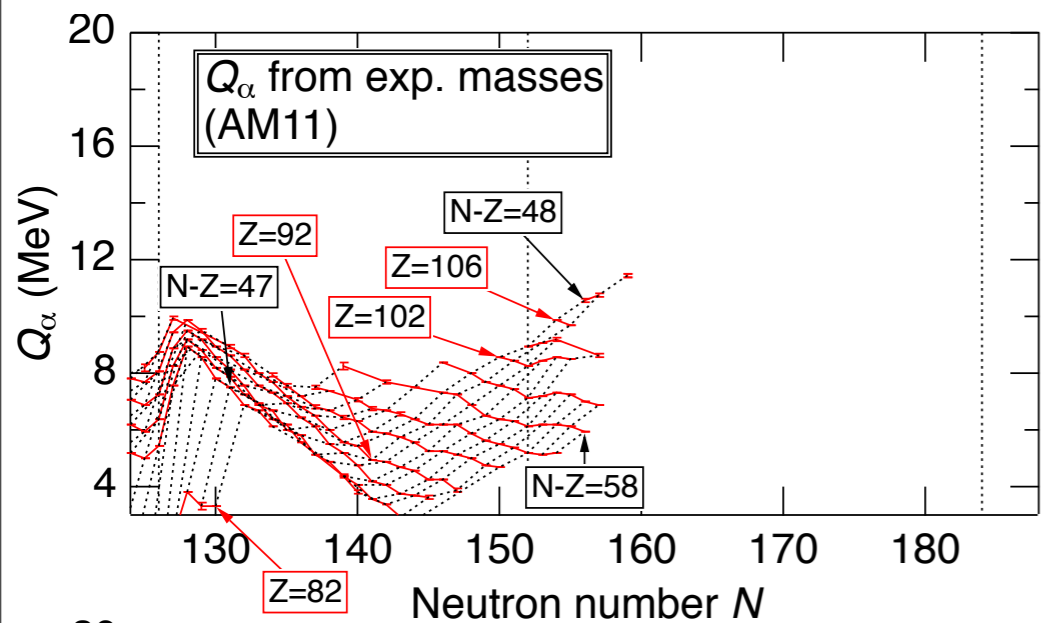
(unphysical) zigzag



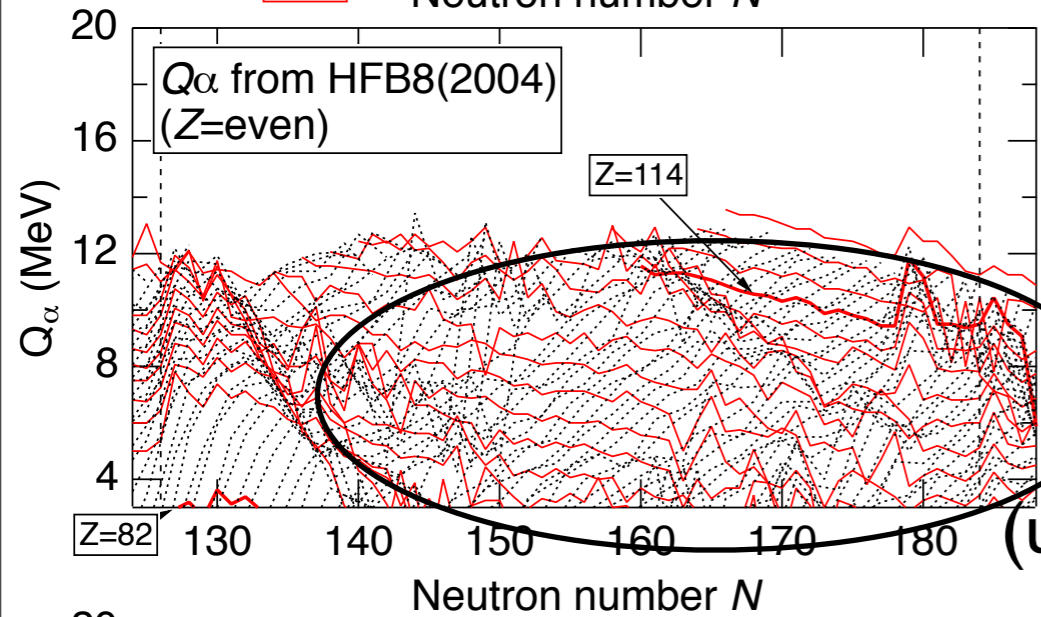


α-decay Q-value of superheavy nuclei

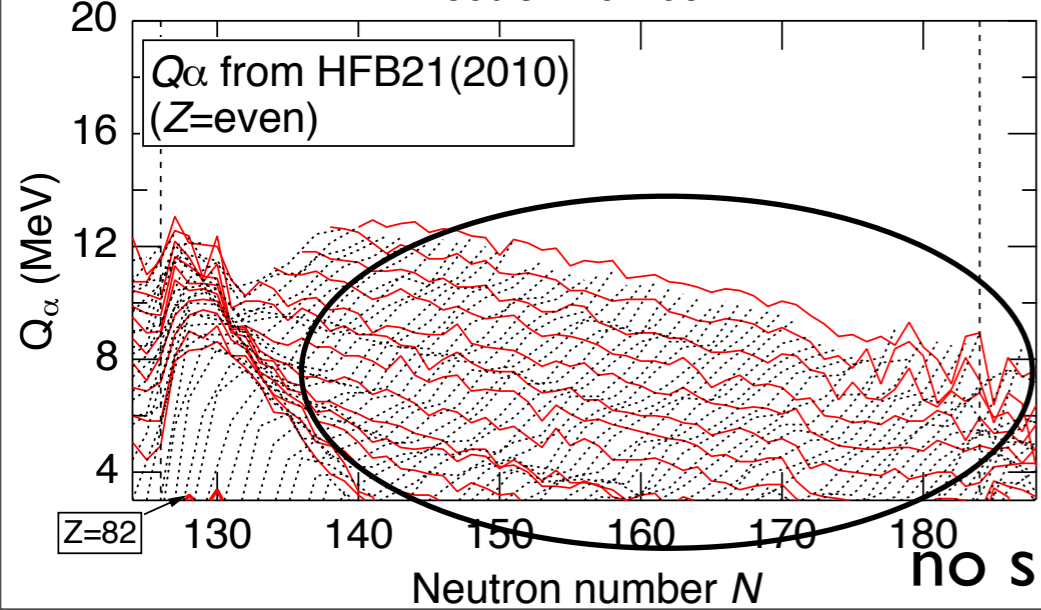
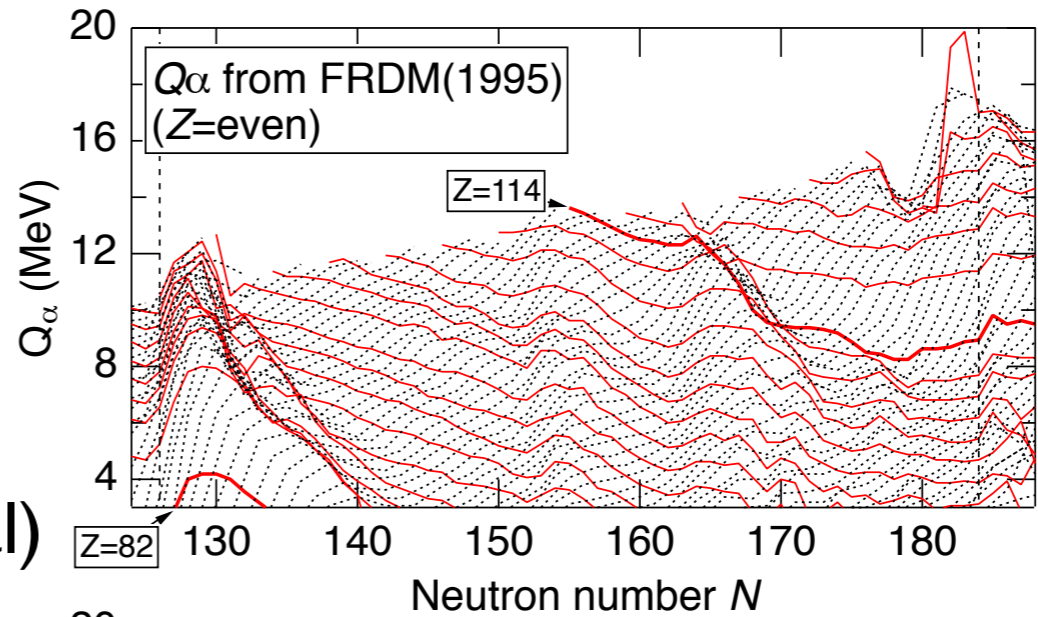
=> Prediction of structure for SHE



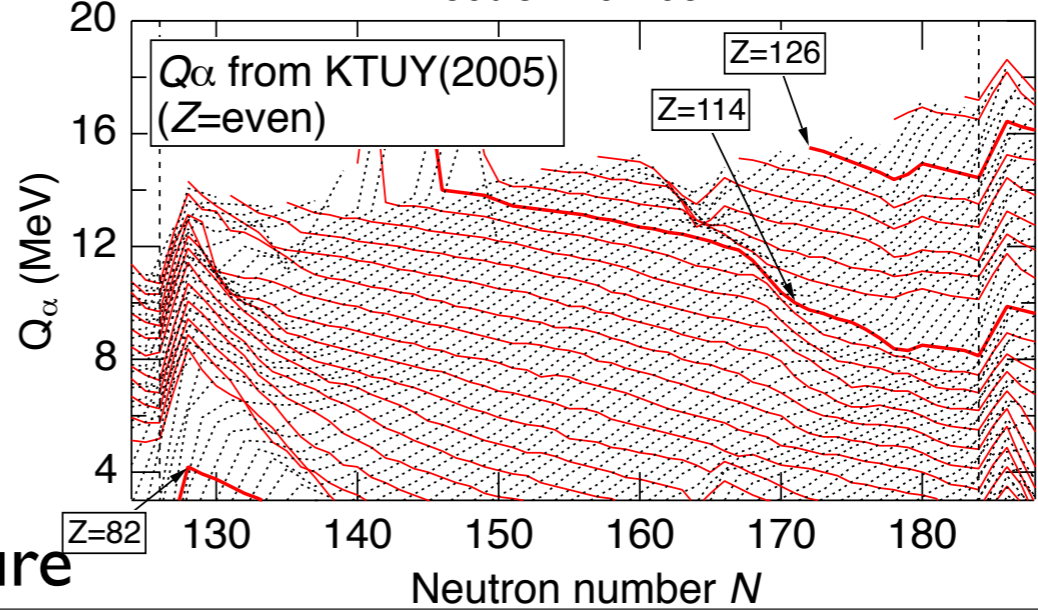
rather geometrical



(unphysical) zigzag

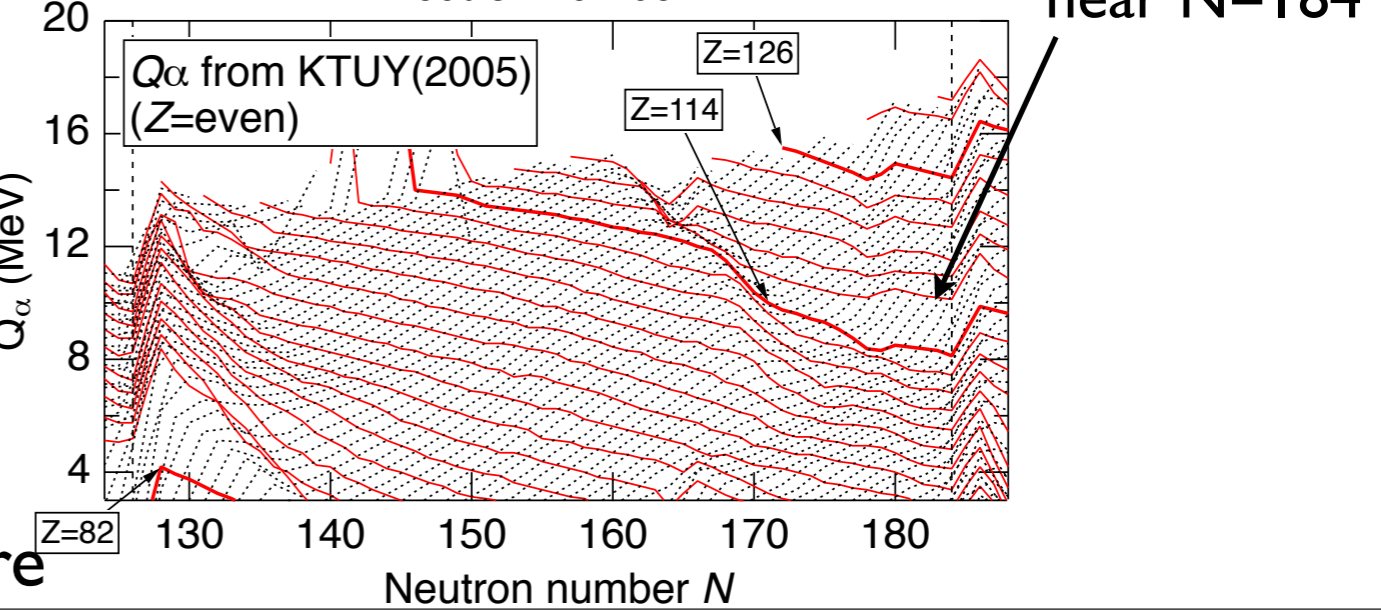
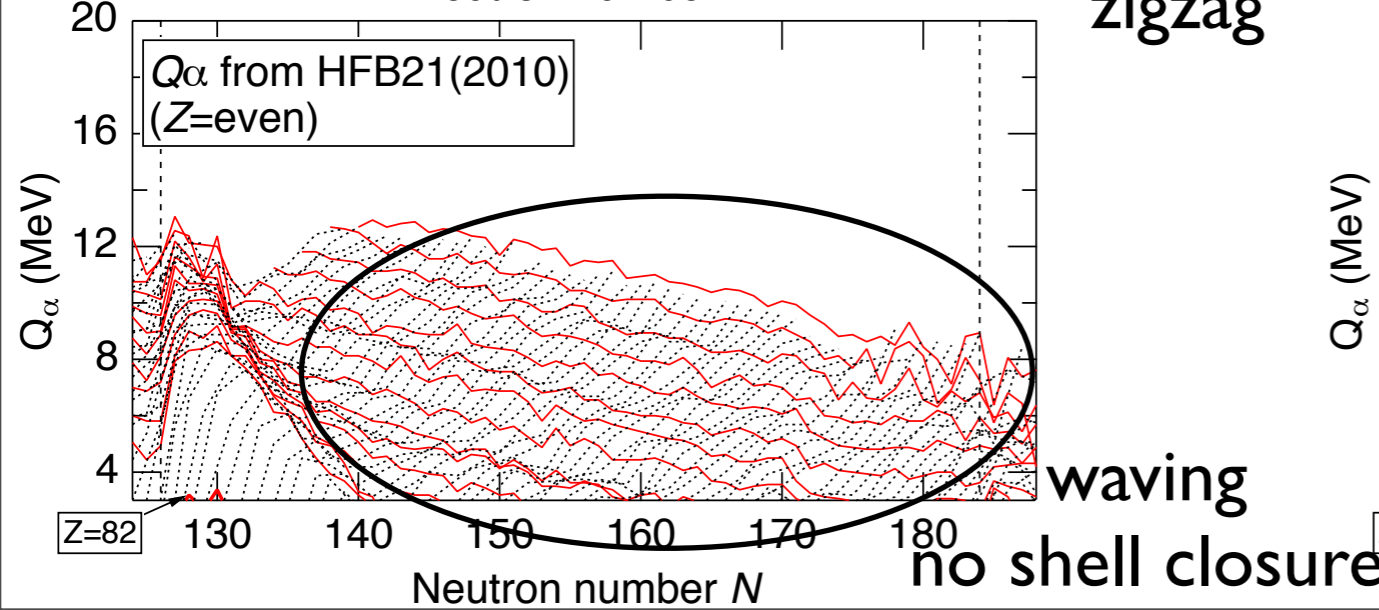
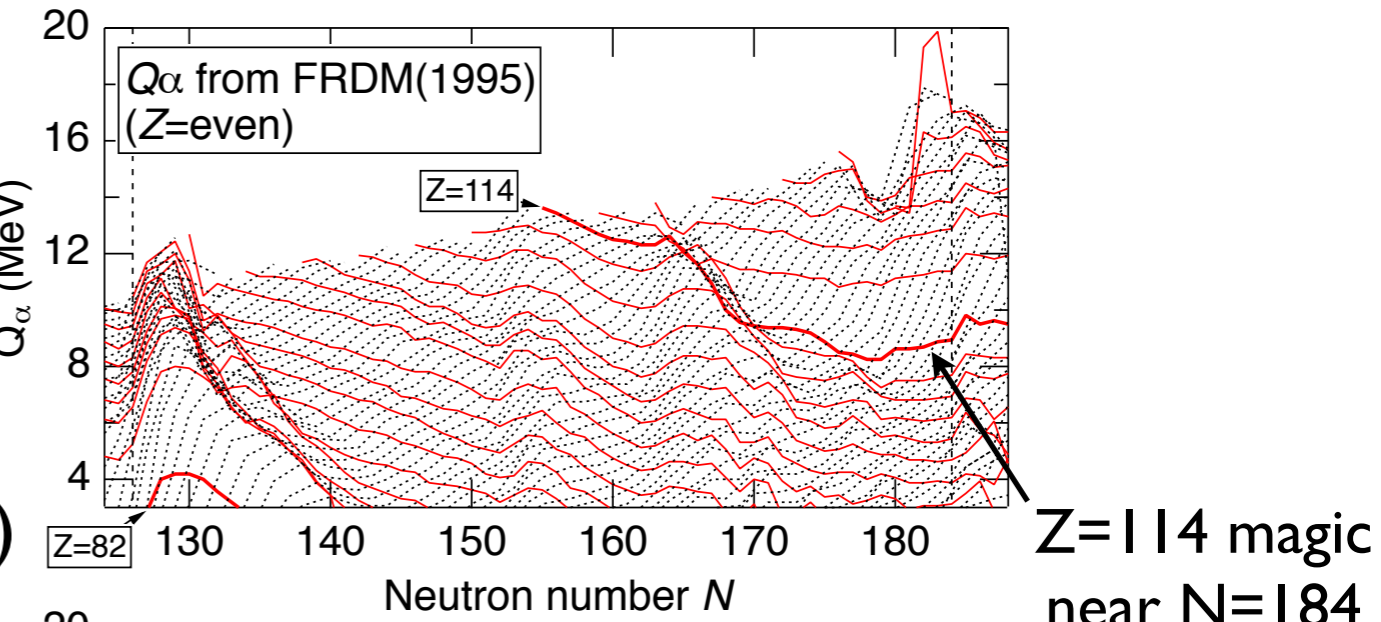
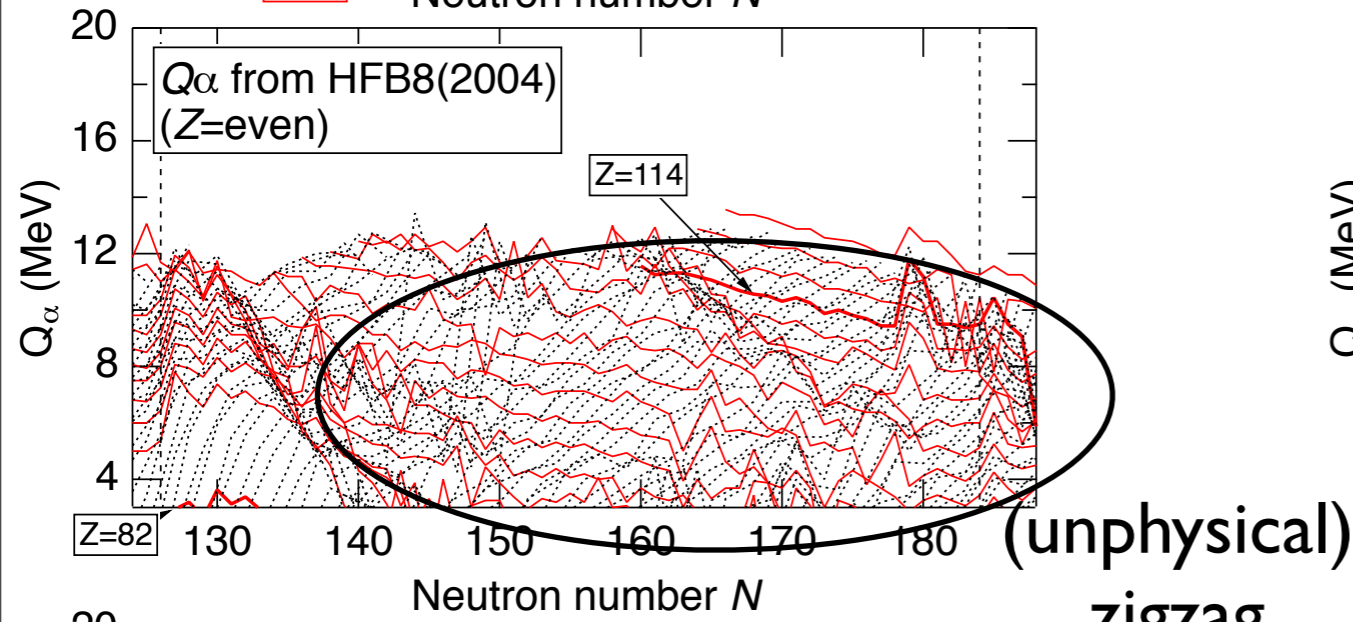
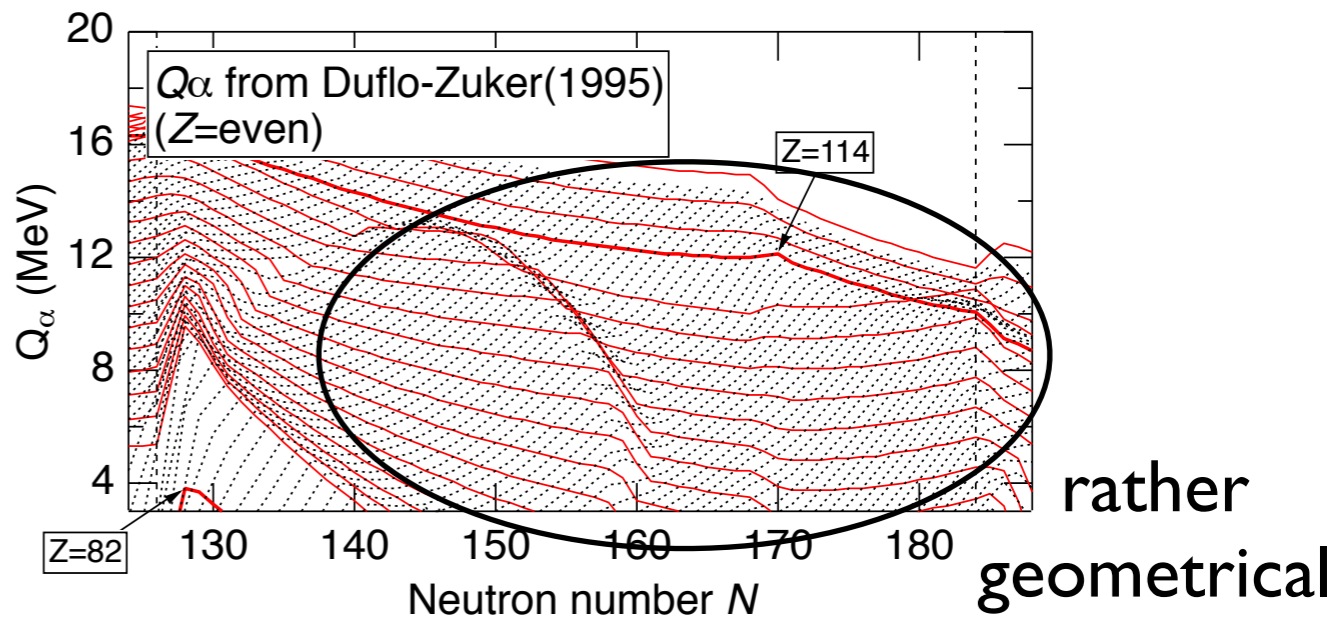
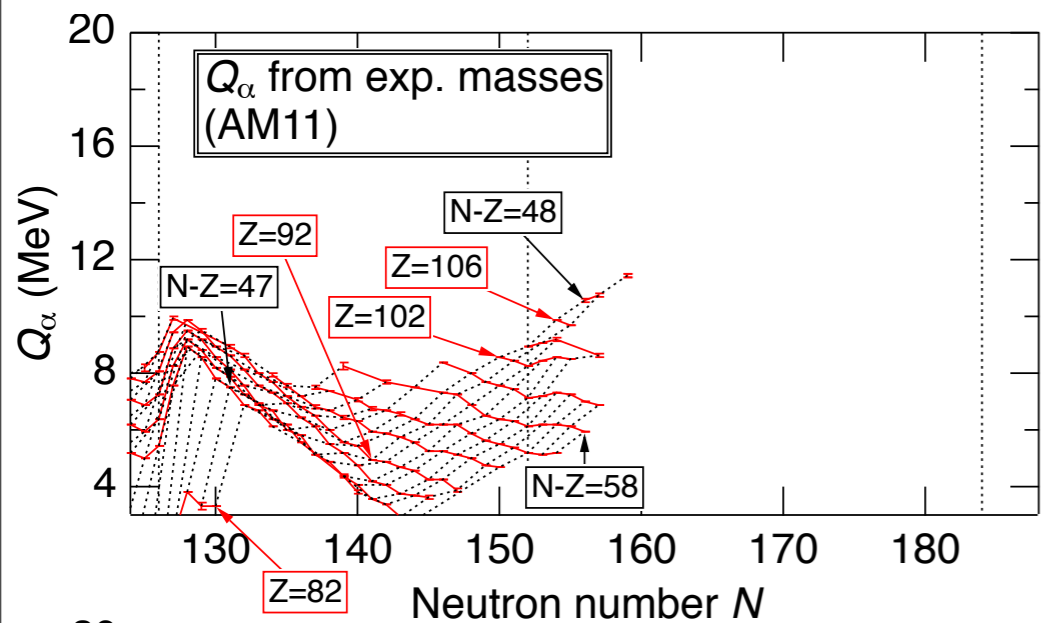


waving no shell closure



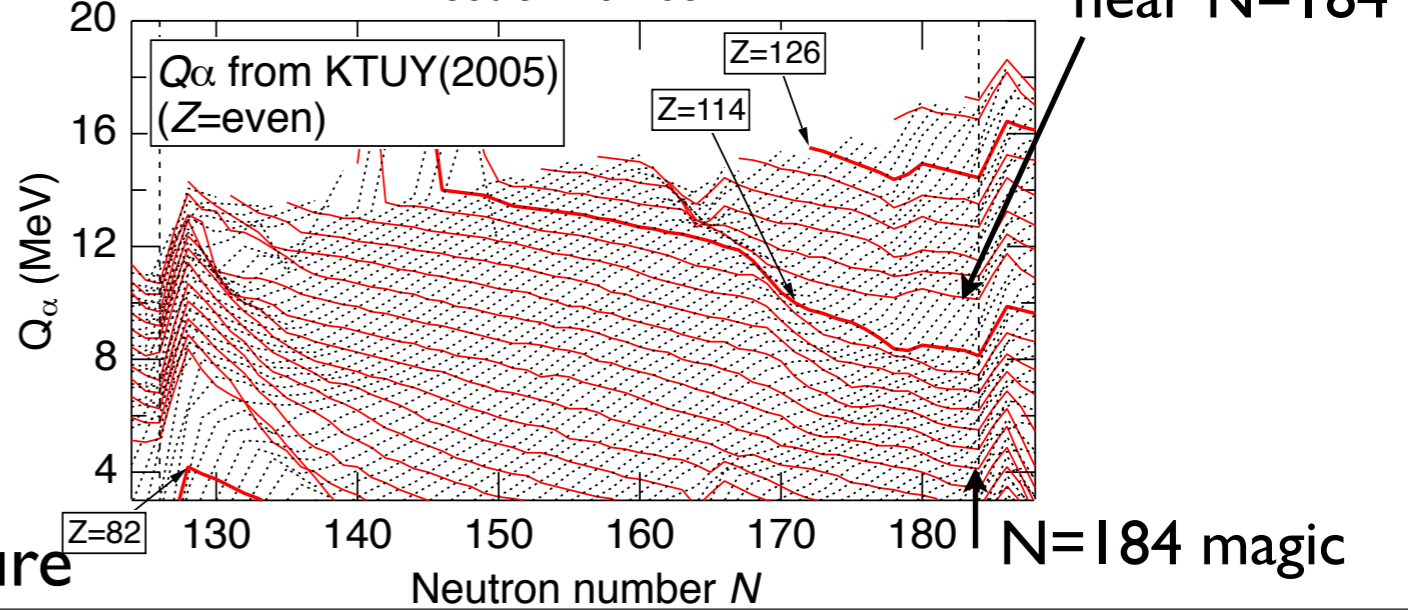
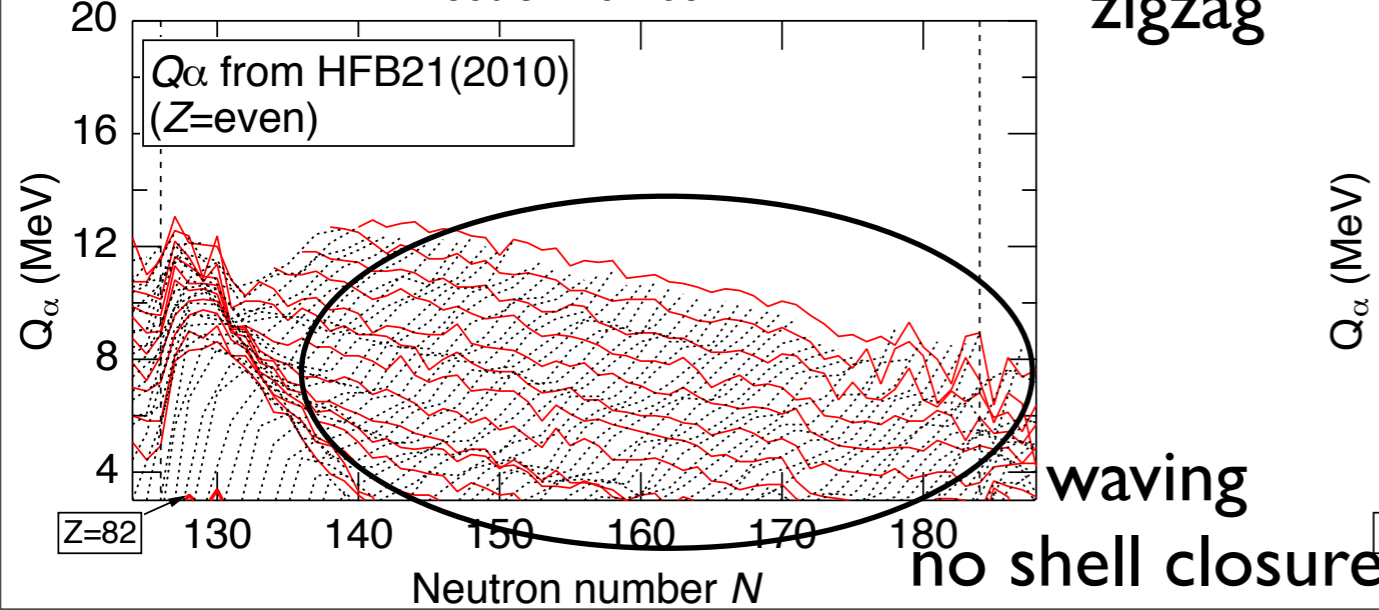
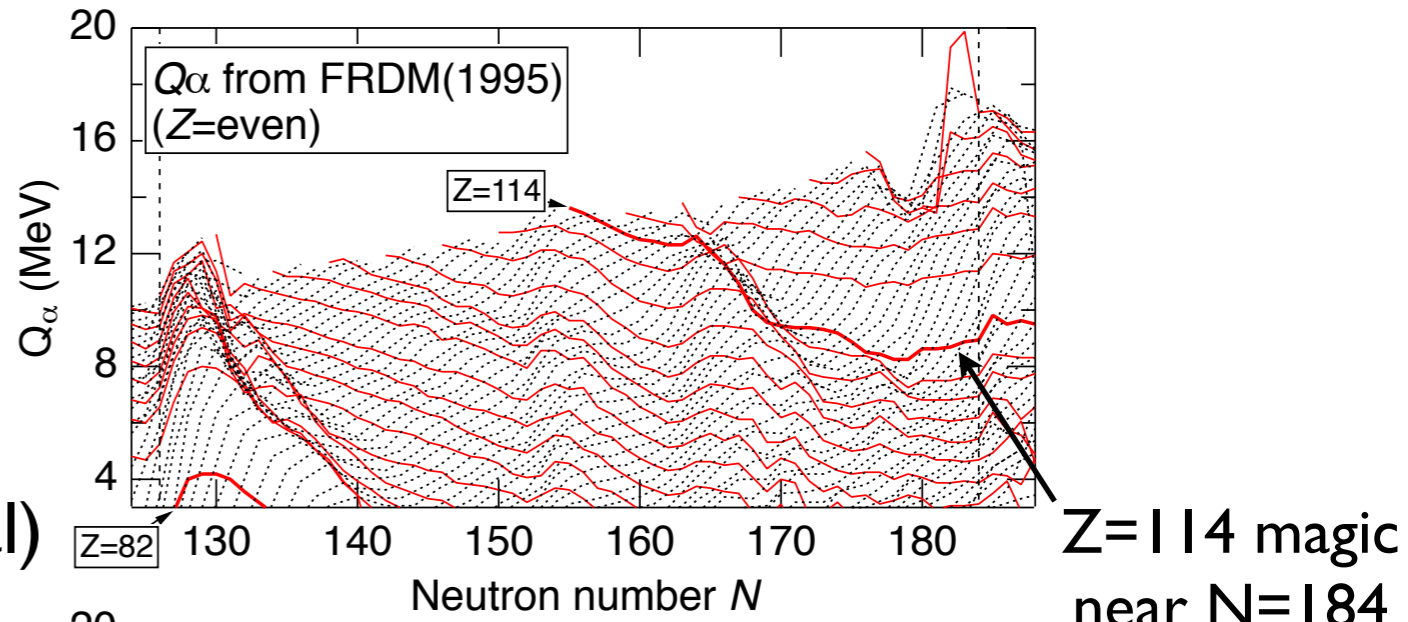
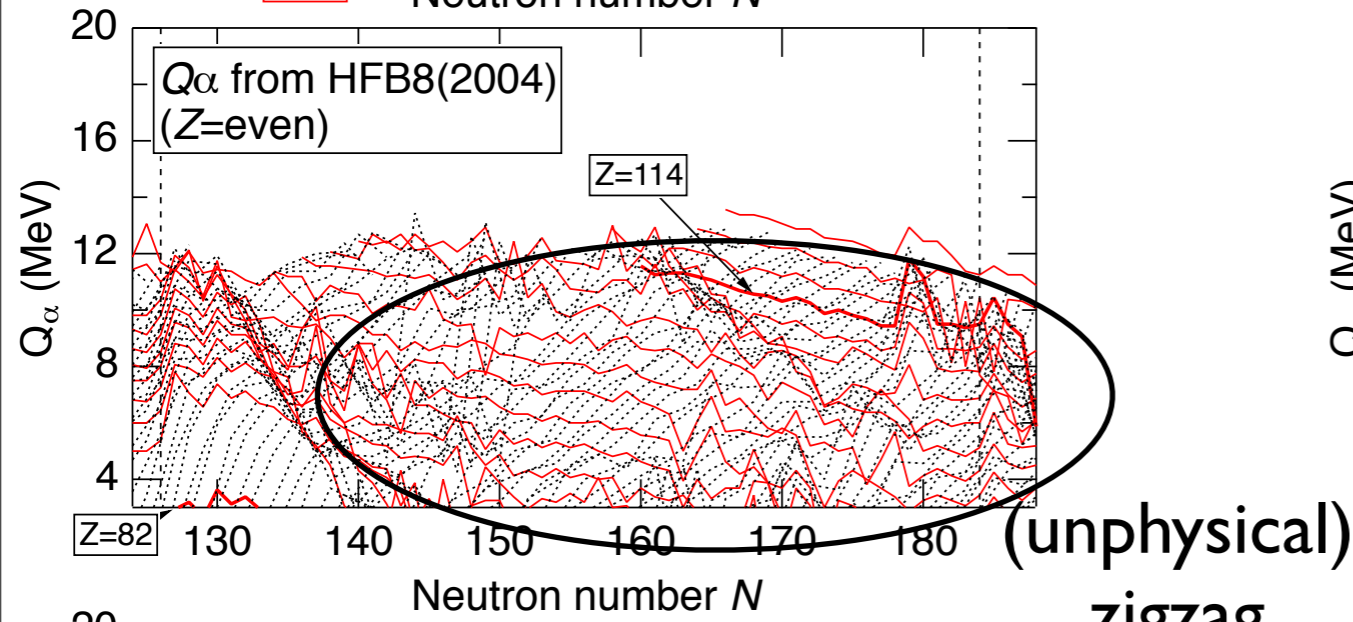
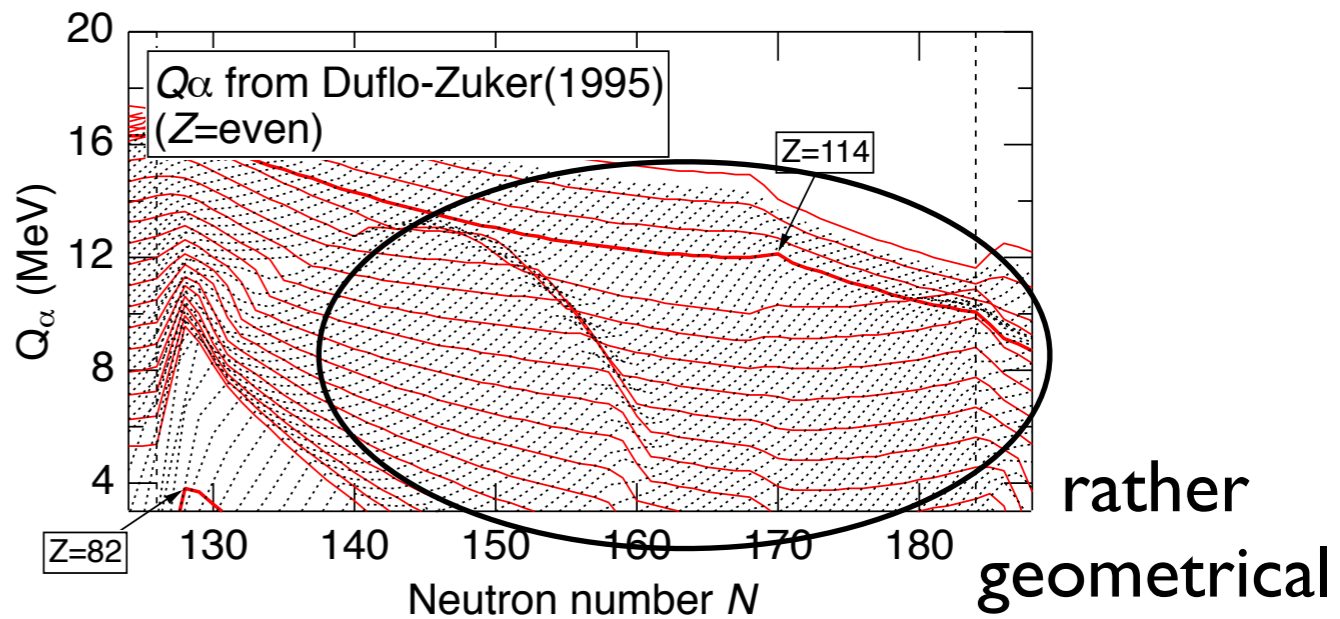
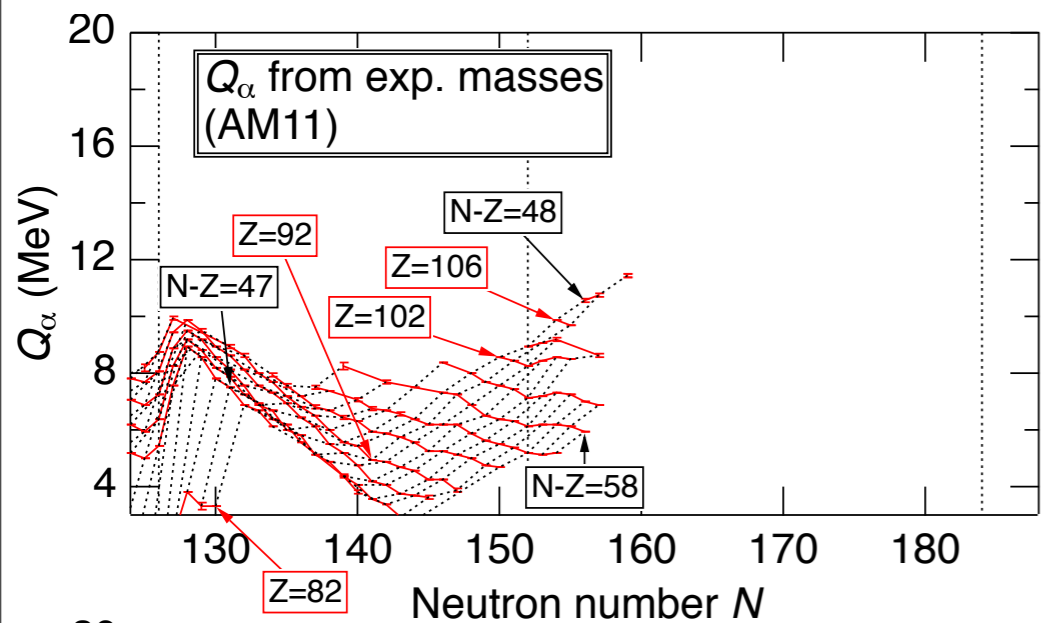
α-decay Q-value of superheavy nuclei

=> Prediction of structure for SHE



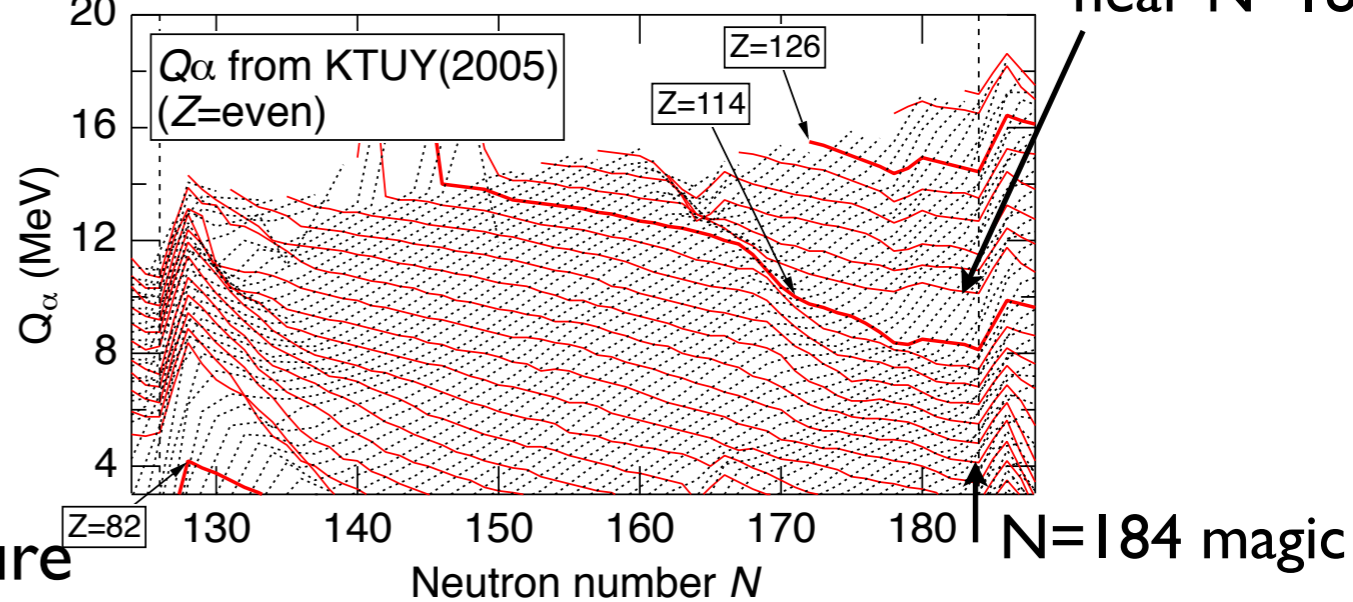
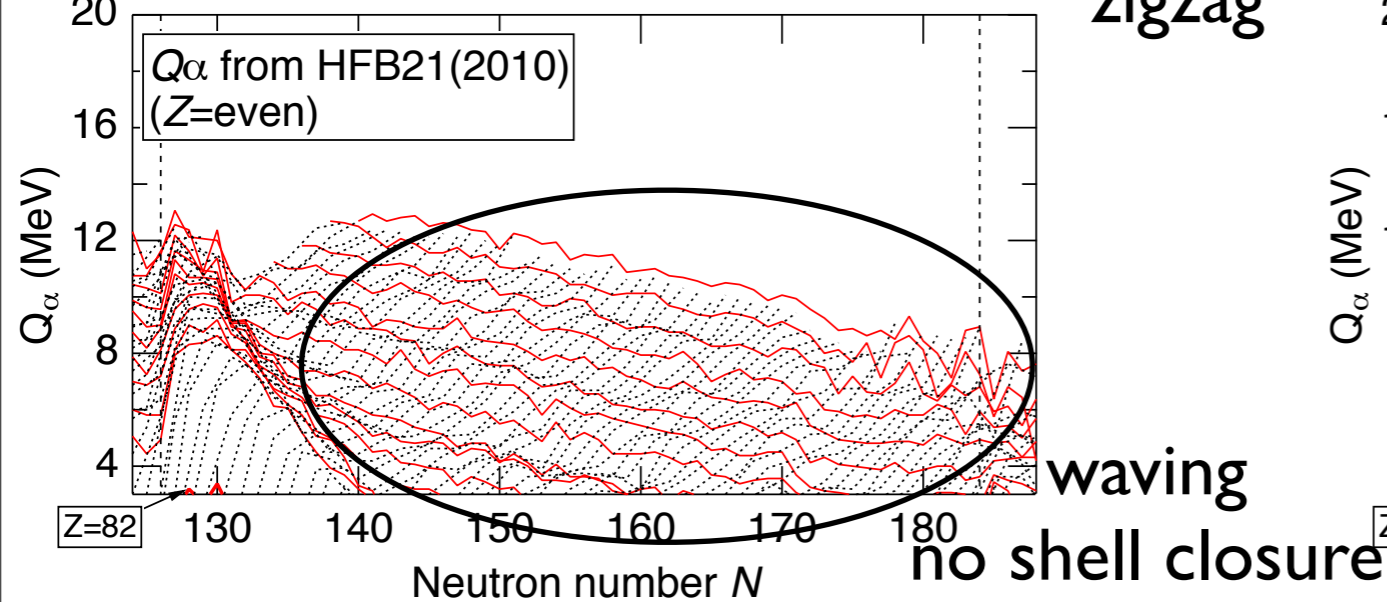
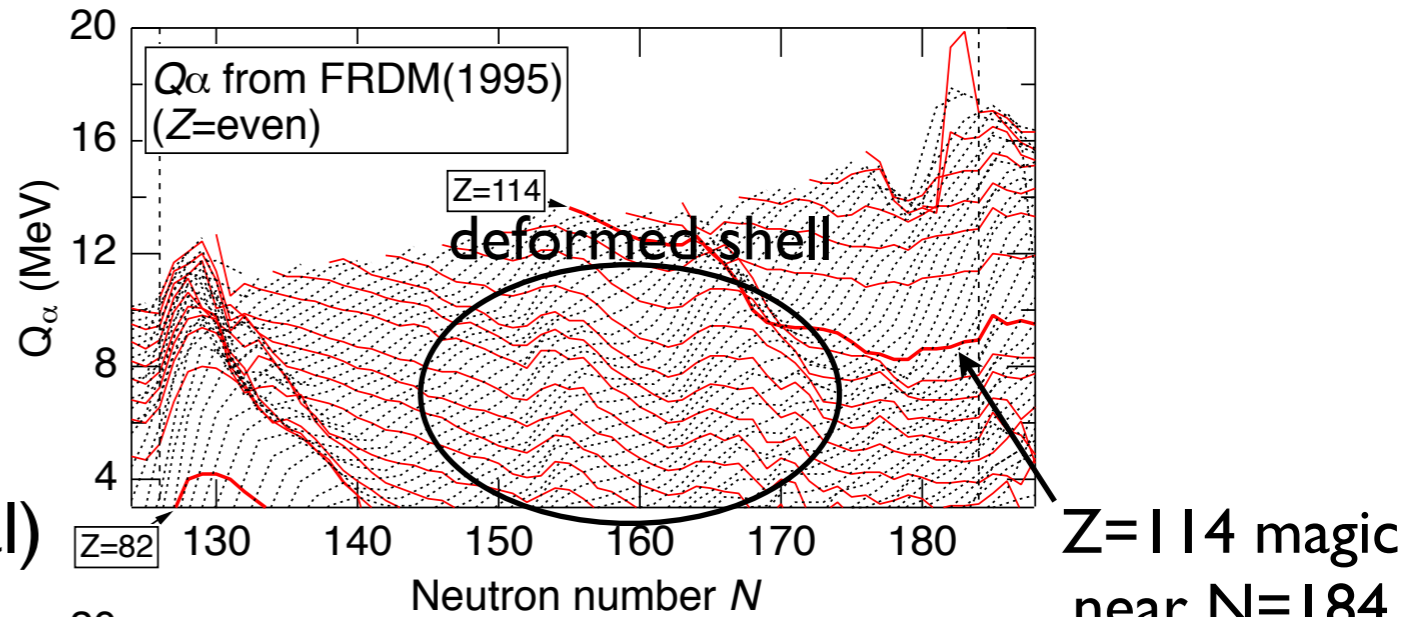
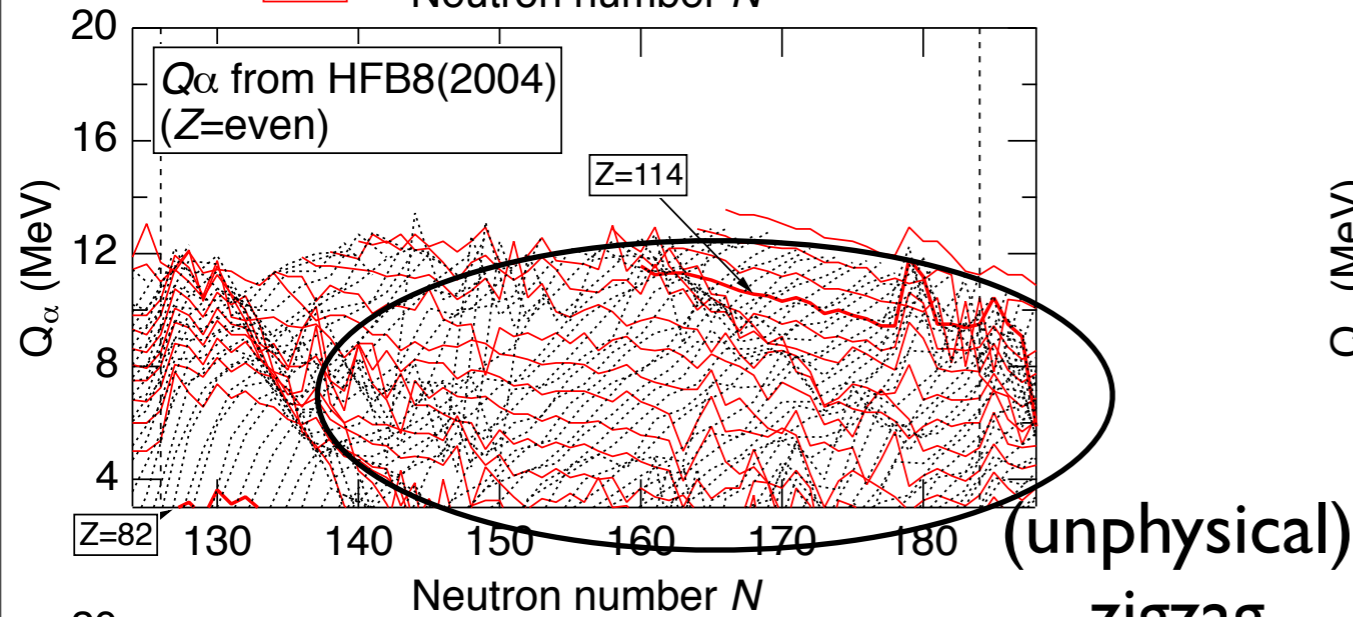
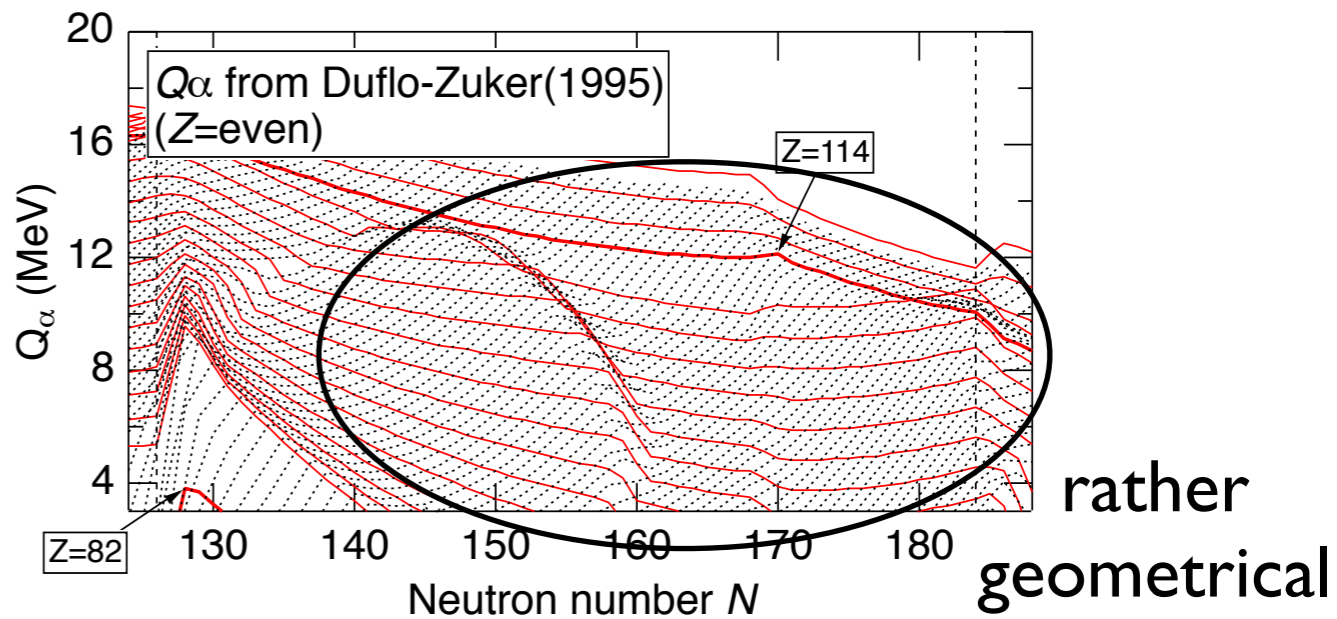
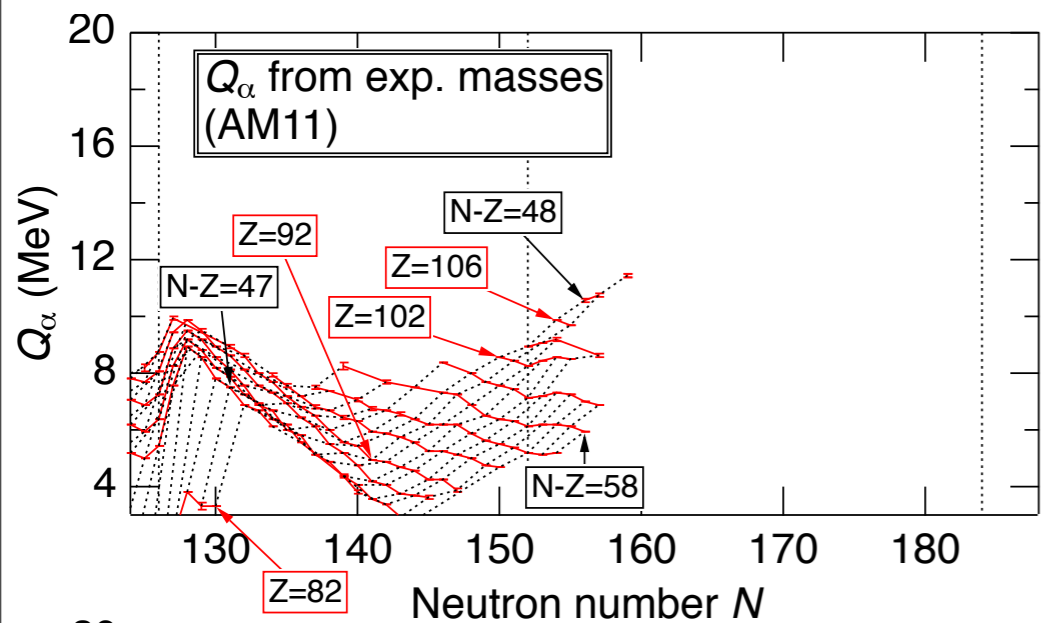
α-decay Q-value of superheavy nuclei

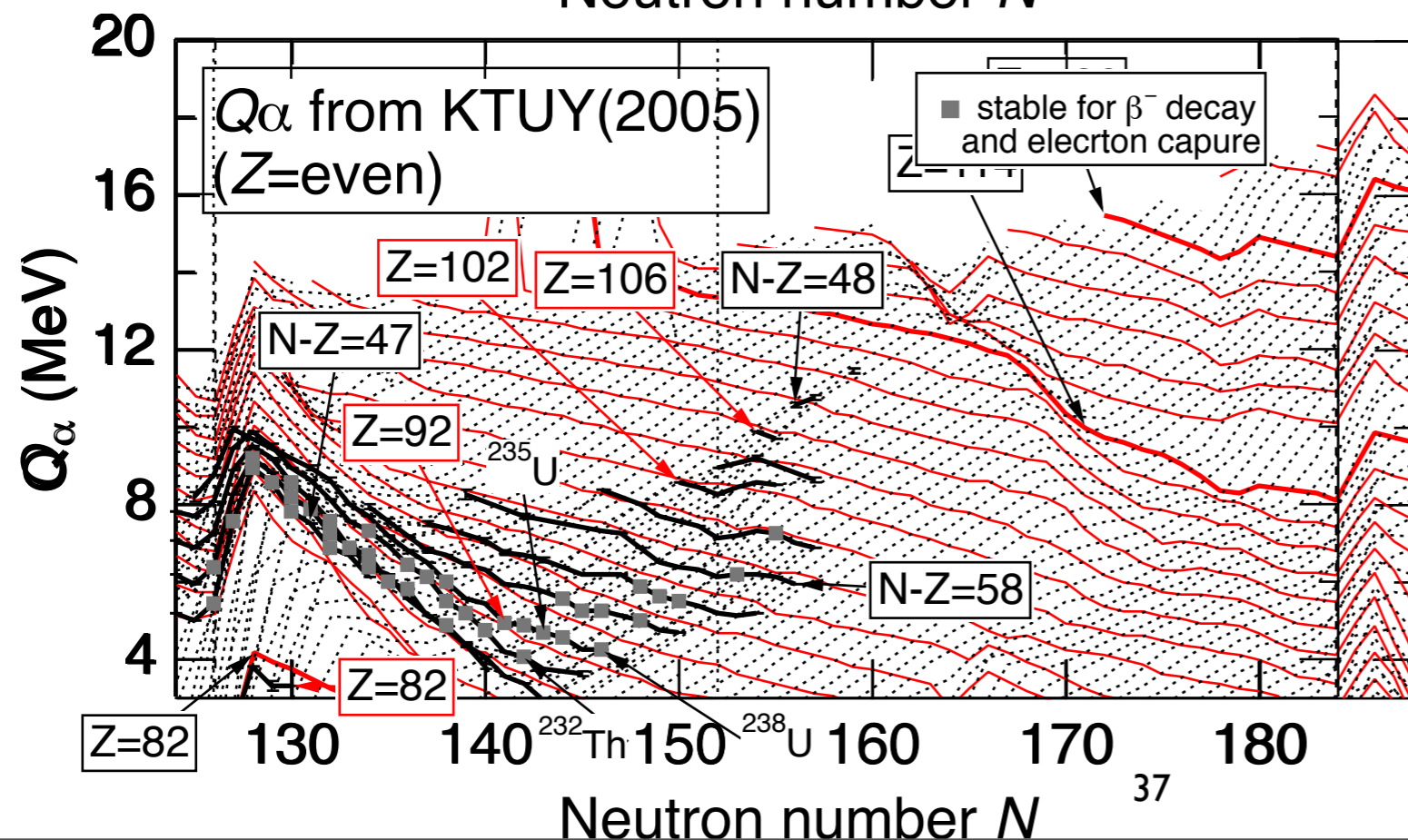
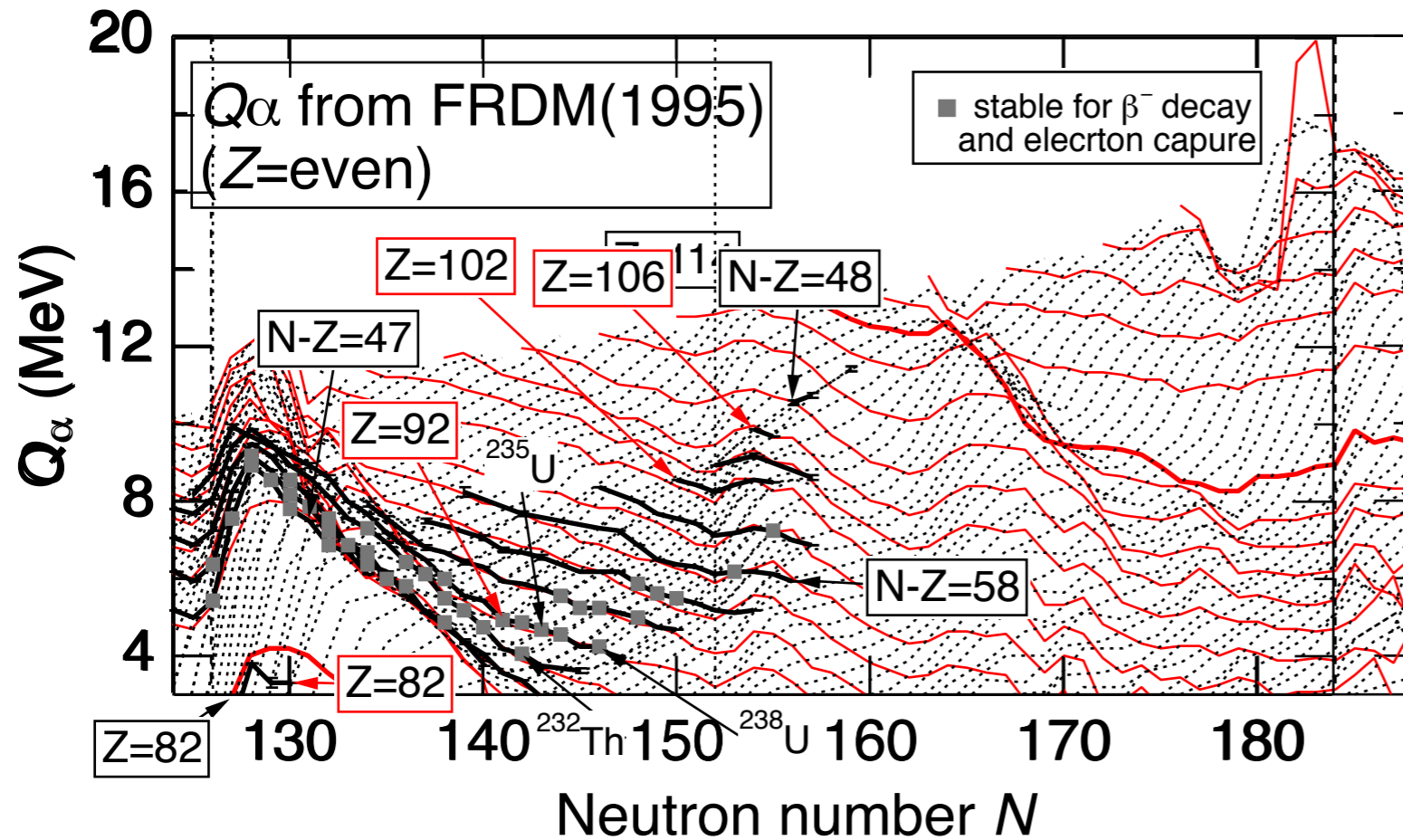
=> Prediction of structure for SHE



α-decay Q-value of superheavy nuclei

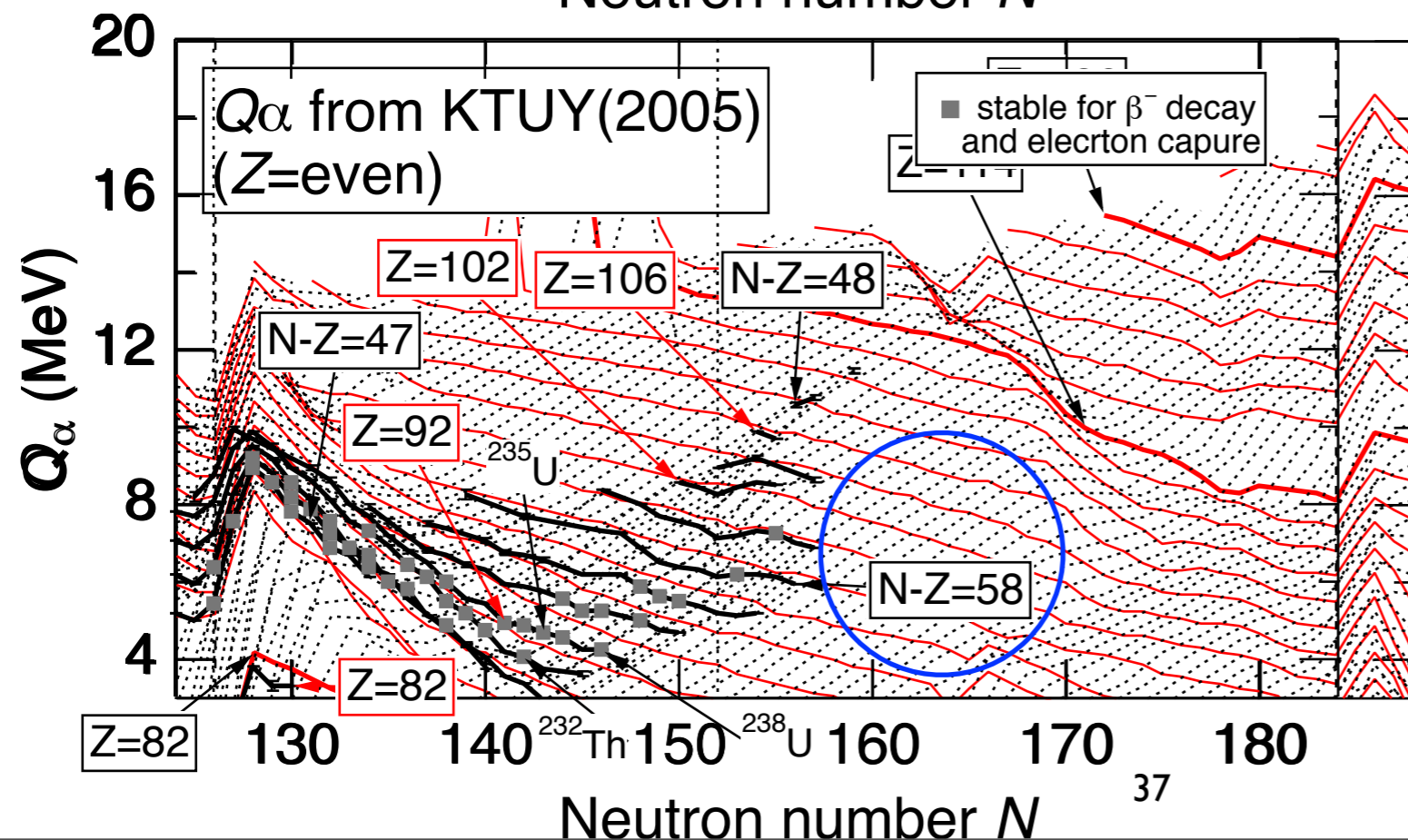
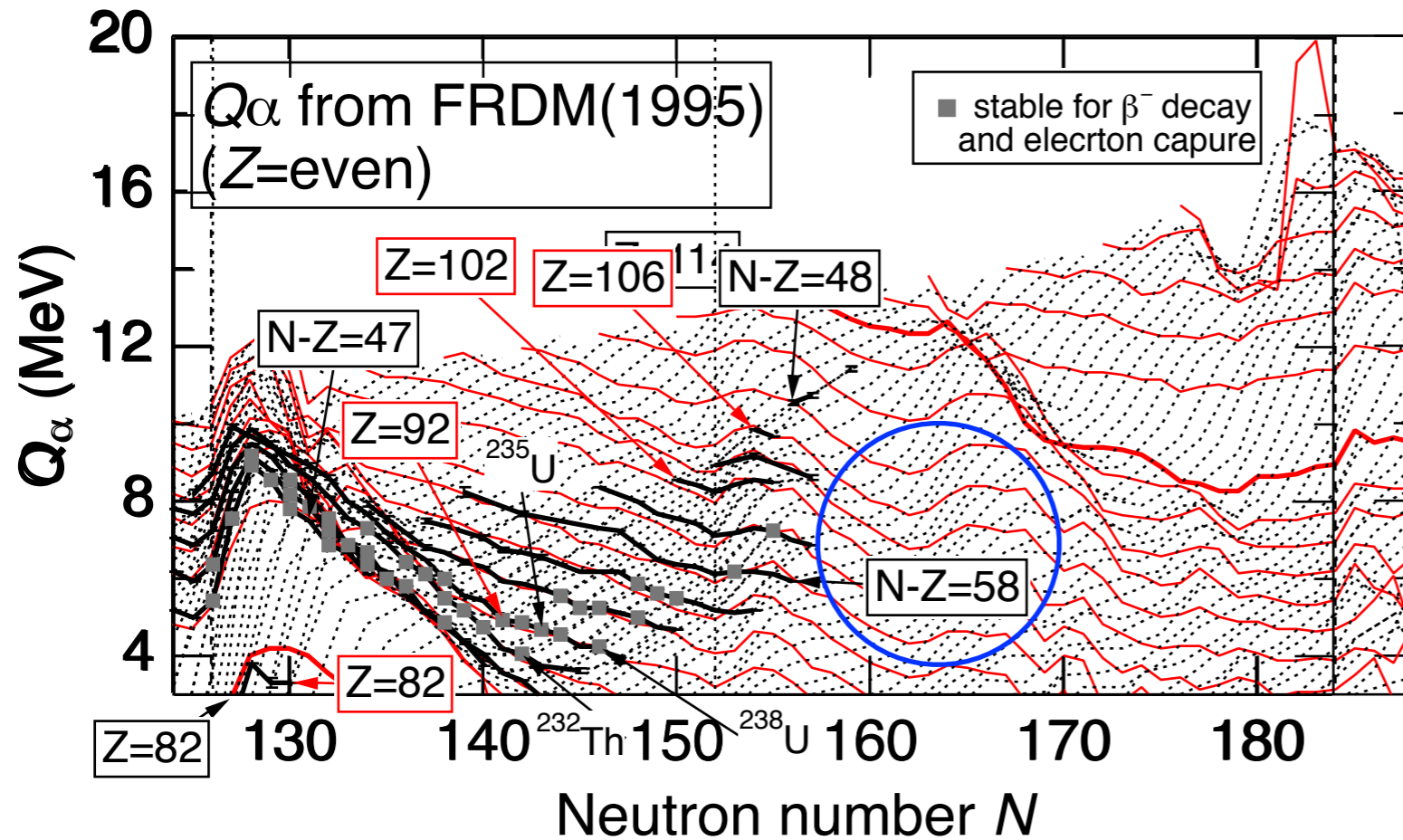
=> Prediction of structure for SHE





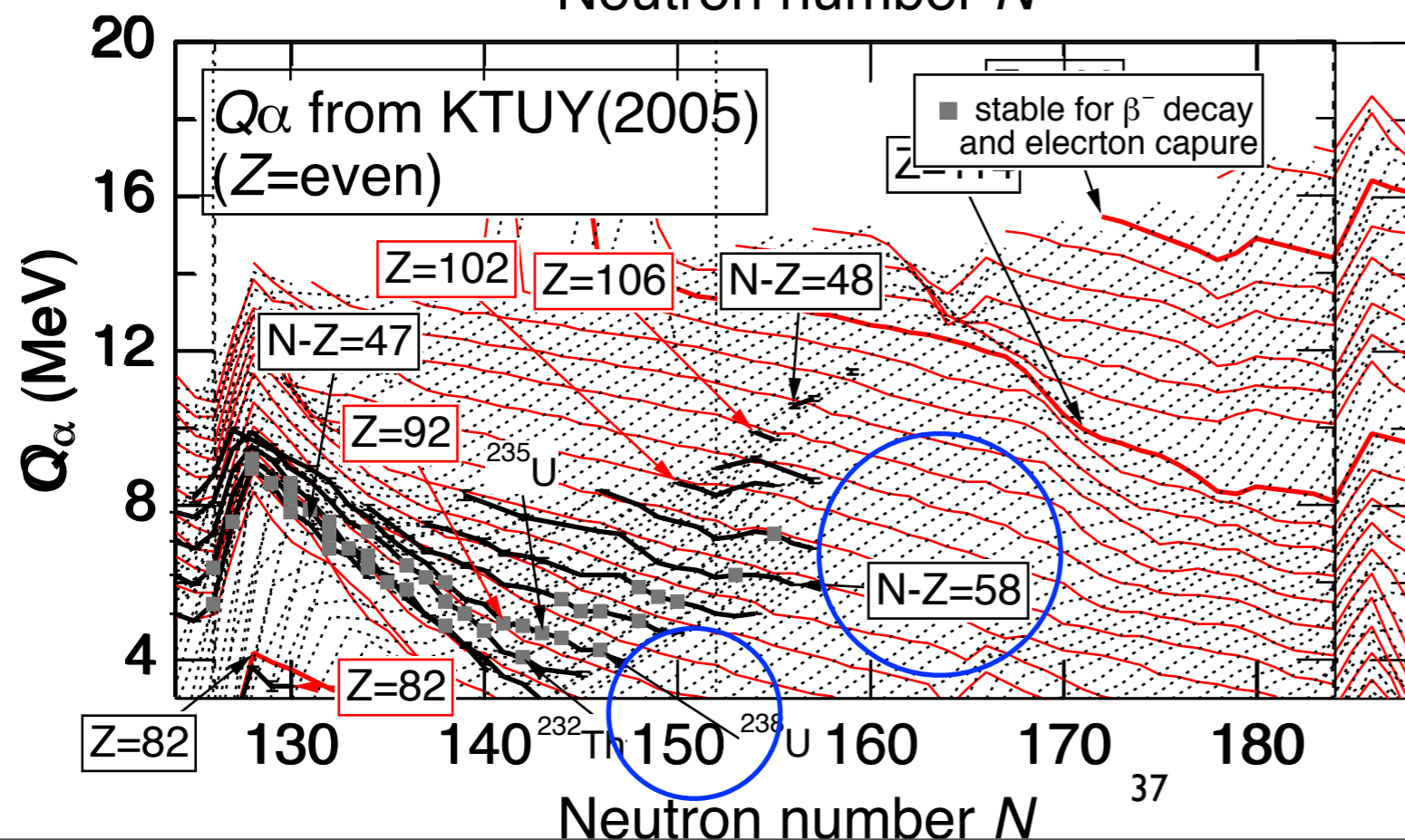
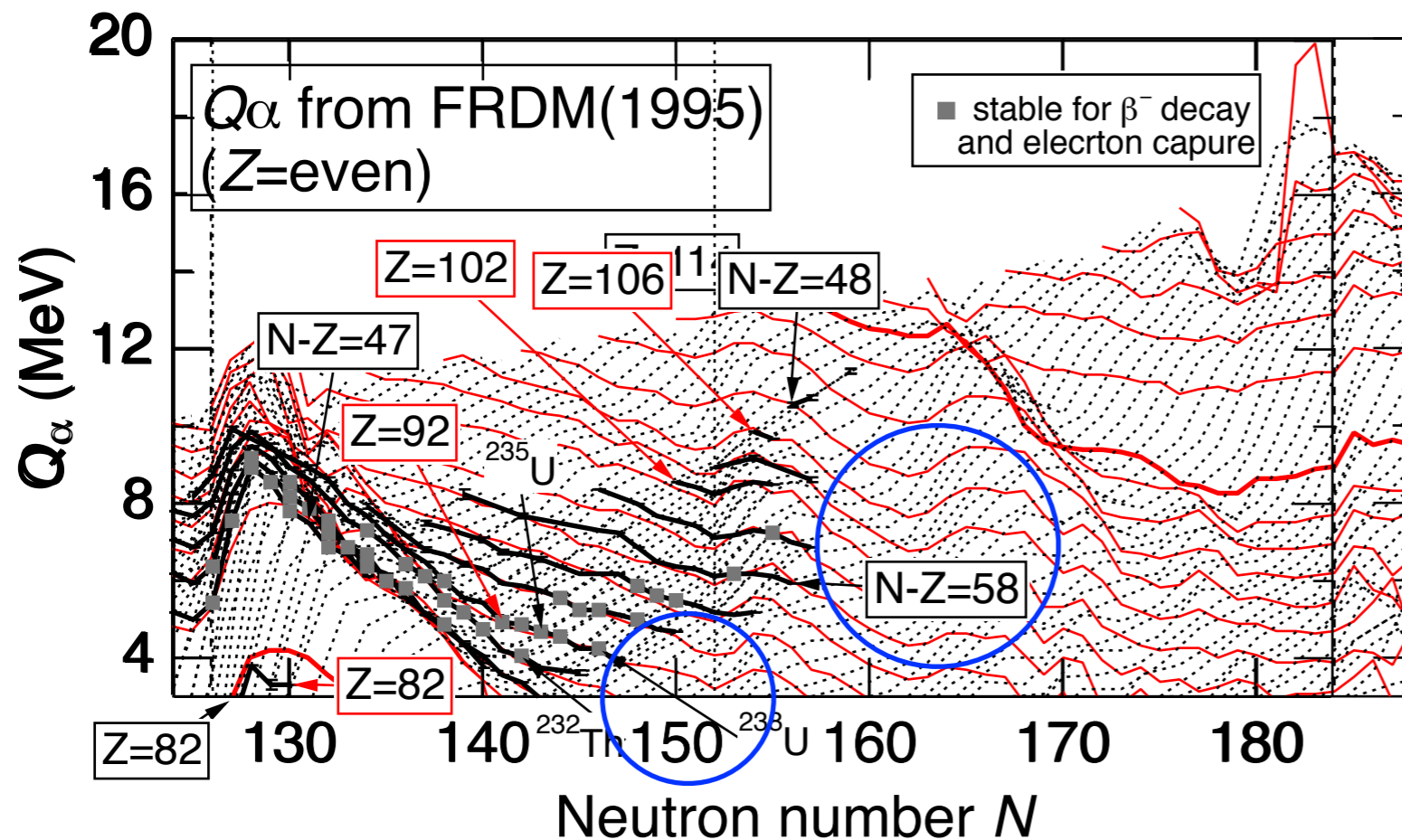
Search for N=162 deformed shell **in the ground-state** is required.

(Also for N=152)



Search for N=162 deformed shell **in the ground-state** is required.

(Also for N=152)



Search for $N=162$ deformed shell **in the ground-state** is required.

(Also for $N=152$)

- We give a short review of systematical properties of experimental nuclear masses.
 - Mass-systematics like G-K is a good tool to check mass values.
- We survey various mass formulae:
 - Old-parametrized mass formulae (in 1976, 88) generally fail to extrapolation.
 - HFB type mass formulae sometimes give anomaly on GK-sys or alpha-chain sys.
- At the n-rich, $A=130$ and 195 related to the r-process, there is poor exp. mass data.
- In the neutron-rich heavy mass region and superheavy mass region, mass measurements is required for ground-state information.

