

Advances and perspectives
in computational nuclear physics

Nuclear structure from no-core Monte Carlo shell model

Takashi Abe (U of Tokyo)

Kohara 3, Kona 4

Hilton Waikoloa Village

Oct. 5th – 7th, 2014

Collaborators

- U of Tokyo
 - Takaharu Otsuka (Department of Physics & CNS)
 - Noritaka Shimizu (CNS)
 - Tooru Yoshida (CNS)
 - Yusuke Tsunoda (Department of Physics)
- JAEA
 - Yutaka Utsuno
- Iowa State U
 - James P. Vary
 - Pieter Maris

Ab initio approaches

- Major challenge of nuclear physics
 - Understand the nuclear structure & reactions from *ab-initio* calculations w/ realistic **nuclear forces (potentials)**
 - *ab-initio* approaches in nuclear structure physics ($A > 4$):
 - GFMC**, **NCSM** ($A \sim 12-14$), **CC** (sub-shell closure +/- 1,2),
Green's Function theory, IM-SRG, Lattice EFT, ...

➔ demand for extensive computational resources

- ✓ *ab-initio*(-like) SM approaches (which attempt to go) beyond standard methods
 - IT-NCSM, IT-CI: R. Roth (TU Darmstadt), P. Navratil (TRIUMF), ...
 - SA-NCSM: T. Dytrych, J.P. Draayer (Louisiana State U), ...
 - **No-Core Monte Carlo Shell Model (MCSM)**

“Ab initio” in low-energy nuclear structure physics

- Solve the non-relativistic Schroedinger eq.
and obtain the eigenvalues and eigenvectors.

$$H|\Psi\rangle = E|\Psi\rangle$$

$$H = T + V_{\text{NN}} + V_{\text{3N}} + \cdots + V_{\text{Coulomb}}$$

- **Ab initio**: All nucleons are active, and Hamiltonian consists of realistic NN (+ 3N + ...) potentials.
- Two main sources of uncertainties:
 - **Nuclear forces** (interactions btw/among nucleons)
In principle, they should be obtained (directly) by QCD.
 - **Many-body methods**
CI: Finite basis space (choice of basis function and truncation),
we have to extrapolate to infinite basis dimensions

Shell model (Configuration Interaction, CI)

- Eigenvalue problem of large sparse Hamiltonian matrix

$$H|\Psi\rangle = E|\Psi\rangle$$

$$\begin{pmatrix} H_{11} & H_{12} & H_{13} & H_{14} & H_{15} & \cdots \\ H_{21} & H_{22} & H_{23} & H_{24} & & \\ H_{31} & H_{32} & H_{33} & & & \\ H_{41} & H_{33} & & \ddots & & \\ H_{51} & & & & & \\ \vdots & & & & & \end{pmatrix} \begin{pmatrix} \Psi_1 \\ \Psi_2 \\ \Psi_3 \\ \Psi_4 \\ \Psi_5 \\ \vdots \end{pmatrix} = \begin{pmatrix} E_1 & & & & & 0 \\ & E_2 & & & & \\ & & E_3 & & & \\ & & & \ddots & & \\ 0 & & & & & \end{pmatrix} \begin{pmatrix} \Psi_1 \\ \Psi_2 \\ \Psi_3 \\ \Psi_4 \\ \Psi_5 \\ \vdots \end{pmatrix}$$

Large sparse matrix (in M-scheme)

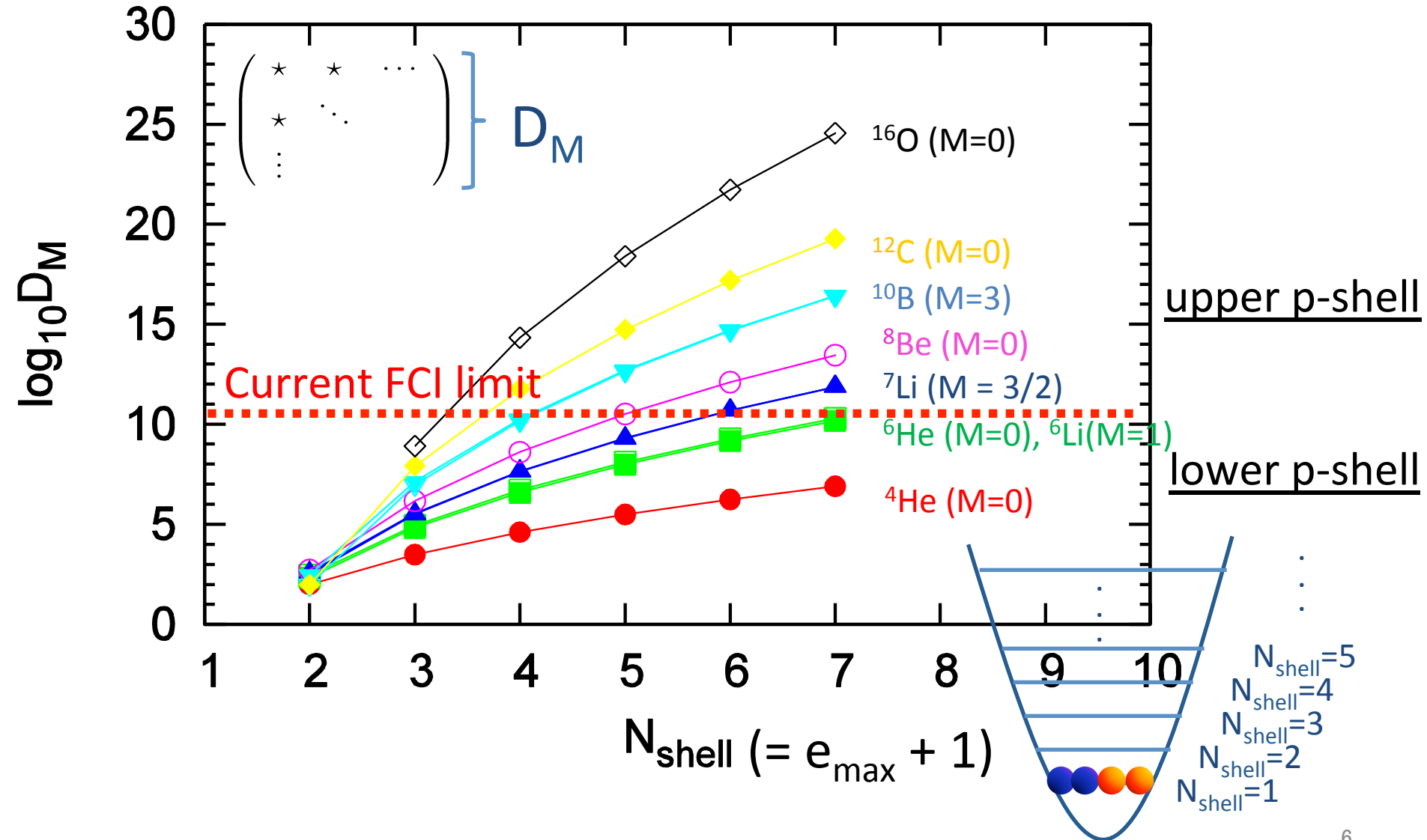
$$\sim \mathcal{O}(10^{10}) \quad \# \text{ non-zero MEs} \sim \mathcal{O}(10^{13-14})$$

Slater determinants

$$\left\{ \begin{array}{l} |\Psi_1\rangle = a_\alpha^\dagger a_\beta^\dagger a_\gamma^\dagger \cdots |-\rangle \\ |\Psi_2\rangle = a_{\alpha'}^\dagger a_{\beta'}^\dagger a_{\gamma'}^\dagger \cdots |-\rangle \\ |\Psi_3\rangle = \cdots \\ \vdots \end{array} \right.$$

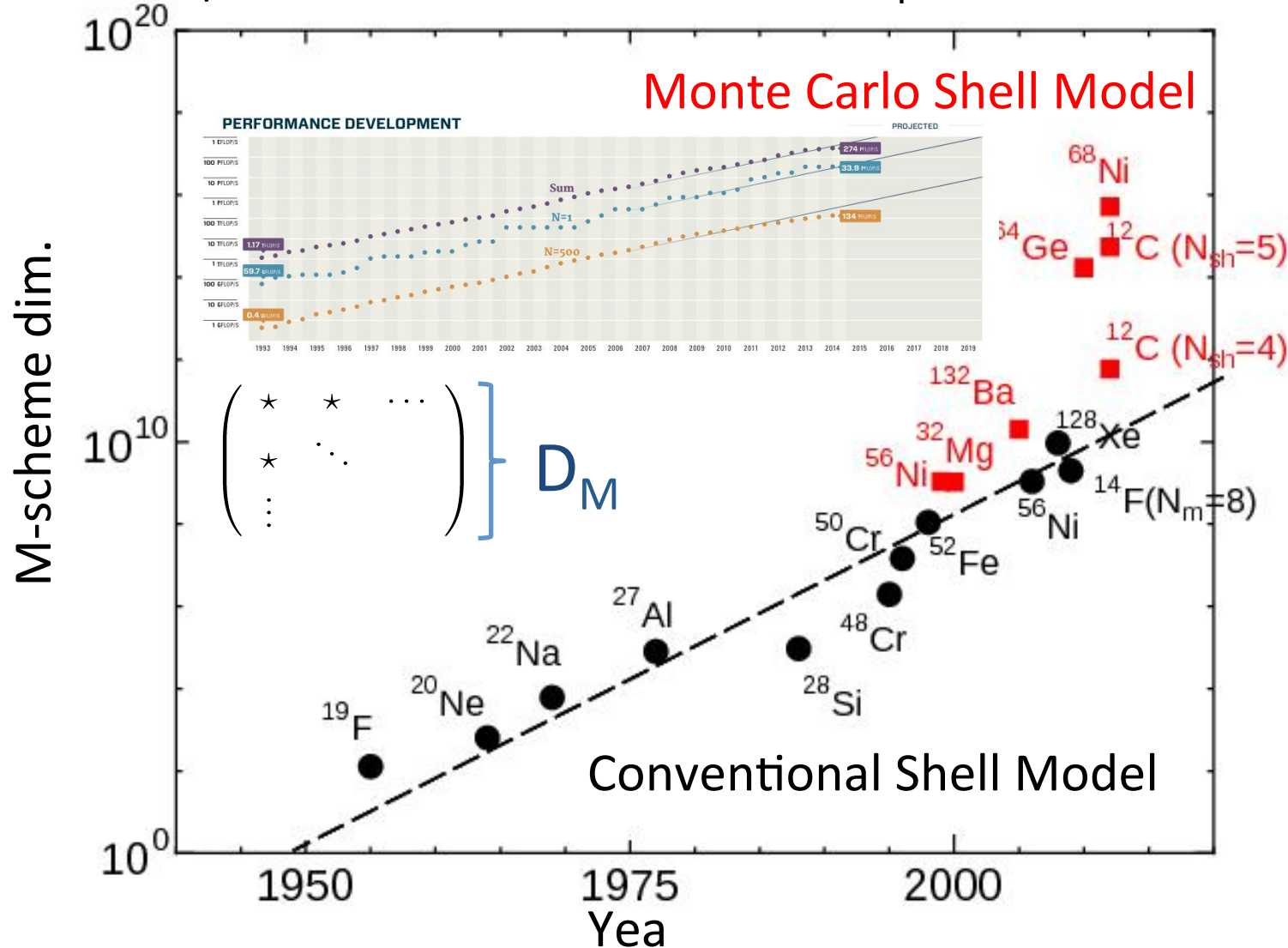
M-scheme dimension in N_{shell} truncation

No-core calculations



Historical evolution/development of the MCSM

- MCSM w/ an assumed inert core is one of the powerful shell model algorithms.

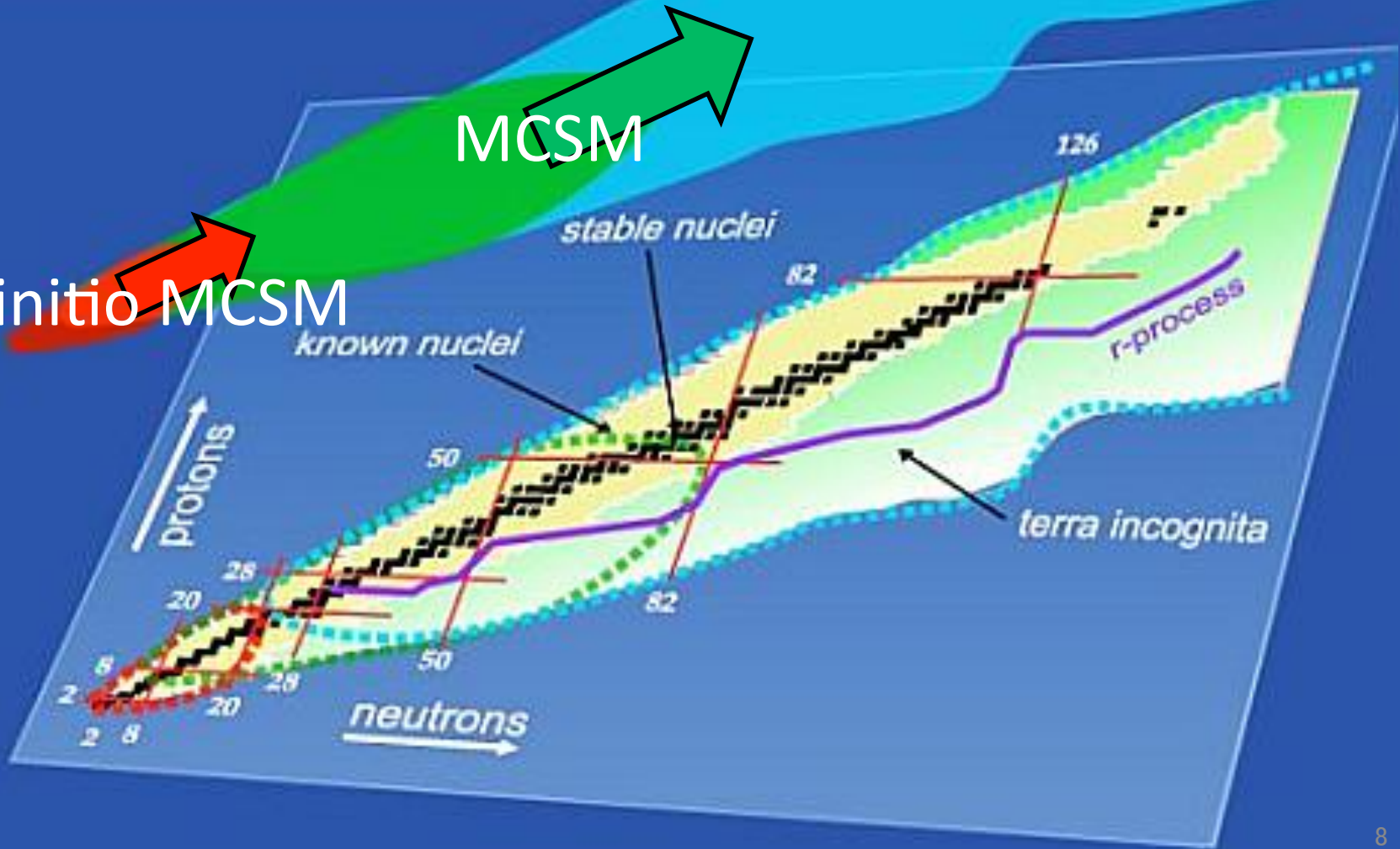


Nuclear Landscape



Ab initio MCSM

MCSM



Nuclear Landscape



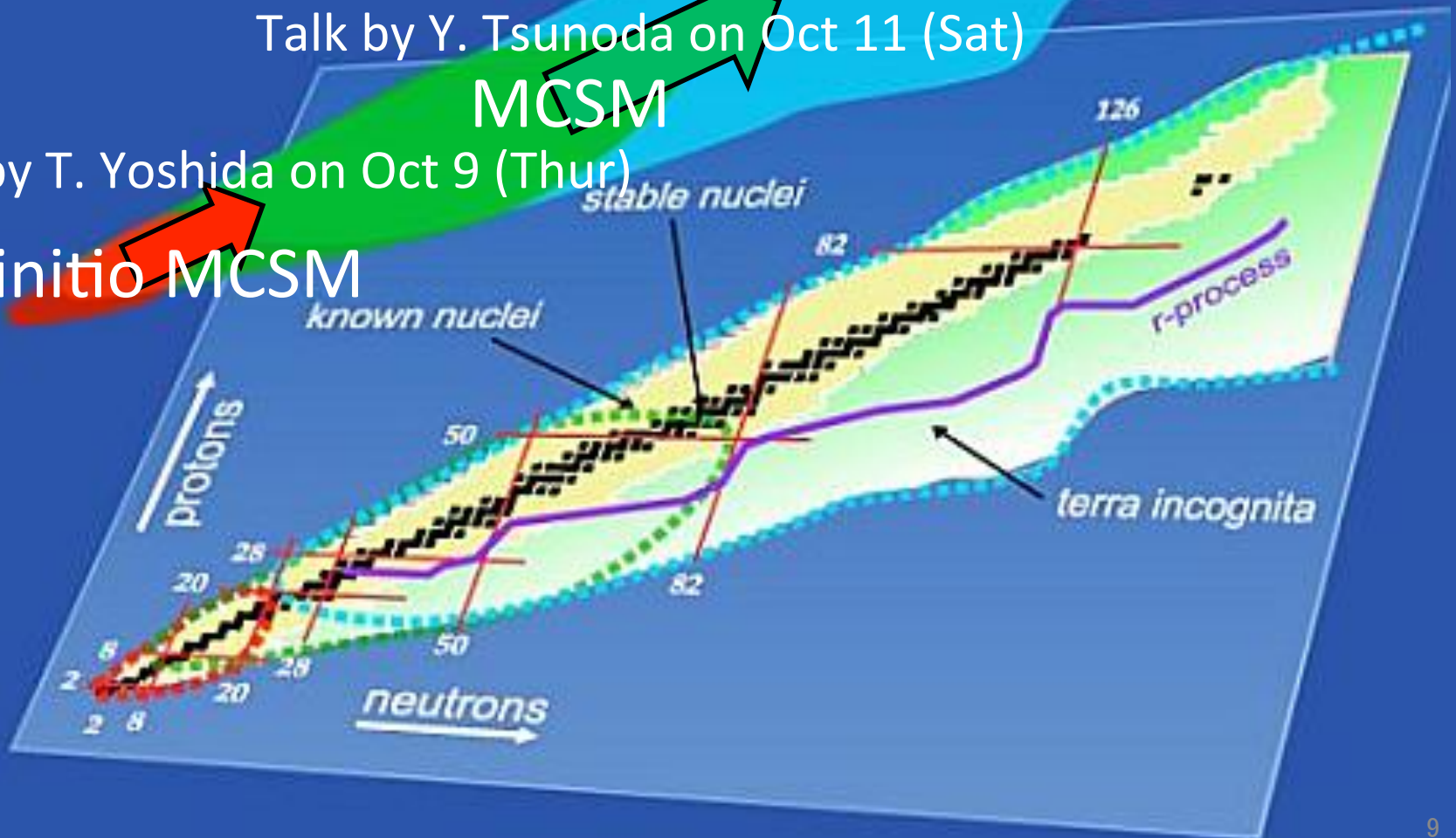
Talk by T. Otsuka on Oct 7 (Thu)

Talk by Y. Tsunoda on Oct 11 (Sat)

MCSM

Talk by T. Yoshida on Oct 9 (Thur)

Ab initio MCSM



Monte Carlo shell model (MCSM)

- Importance truncation

Talk by T. Otsuka on Oct 7 (Thu)

Standard shell model

$$H = \begin{pmatrix} * & * & * & * & * & \dots \\ * & * & * & * & & \\ * & * & * & & & \\ * & * & & \ddots & & \\ * & & & & \ddots & \\ \vdots & & & & & \ddots \end{pmatrix}$$

All Slater determinants

Diagonalization

$$\begin{pmatrix} E_0 & & & & & 0 \\ & E_1 & & & & \\ & & E_2 & & & \\ & & & \ddots & & \\ & & & & \ddots & \\ 0 & & & & & \end{pmatrix}$$

$d > O(10^{10})$

Monte Carlo shell model

$$H \sim \begin{pmatrix} * & * & \dots \\ * & \ddots & \\ \vdots & & \ddots \end{pmatrix}$$

Important bases stochastically selected

Diagonalization

$$\begin{pmatrix} E'_0 & & 0 \\ & E'_1 & \\ 0 & & \ddots \end{pmatrix}$$

$d_{\text{MCSM}} \sim O(100)$

SM Hamiltonian & MCSM many-body w.f.

- 2nd-quantized non-rel. Hamiltonian (up to 2-body term, so far)

$$H = \sum_{\alpha\beta}^{N_{sps}} t_{\alpha\beta} c_{\alpha}^{\dagger} c_{\beta} + \frac{1}{4} \sum_{\alpha\beta\gamma\delta}^{N_{sps}} \bar{v}_{\alpha\beta\gamma\delta} c_{\alpha}^{\dagger} c_{\beta}^{\dagger} c_{\delta} c_{\gamma} \quad \bar{v}_{ijkl} = v_{ijkl} - v_{ijlk}$$

- Eigenvalue problem

$$H|\Psi(J, M, \pi)\rangle = E|\Psi(J, M, \pi)\rangle$$

- MCSM many-body wave function & basis function

$$|\Psi(J, M, \pi)\rangle = \sum_i^{N_{basis}} f_i |\Phi_i(J, M, \pi)\rangle \quad |\Phi(J, M, \pi)\rangle = \sum_K g_K P_{MK}^J P^{\pi} |\phi\rangle$$

These coeff. are obtained by the diagonalization.

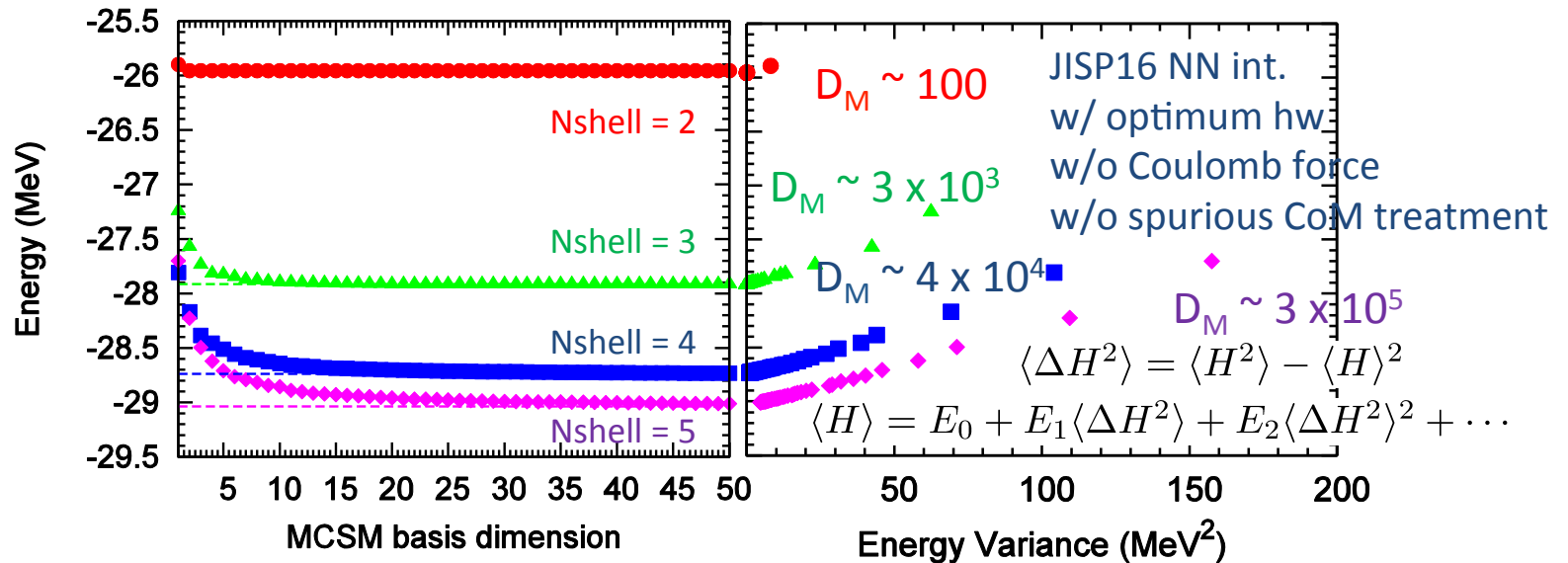
- Deformed SDs

$$|\phi\rangle = \prod_i^A a_i^{\dagger} |-\rangle \quad a_i^{\dagger} = \sum_{\alpha}^{N_{sps}} c_{\alpha}^{\dagger} D_{\alpha i} \quad (c_{\alpha}^{\dagger} \dots \text{spherical HO basis})$$

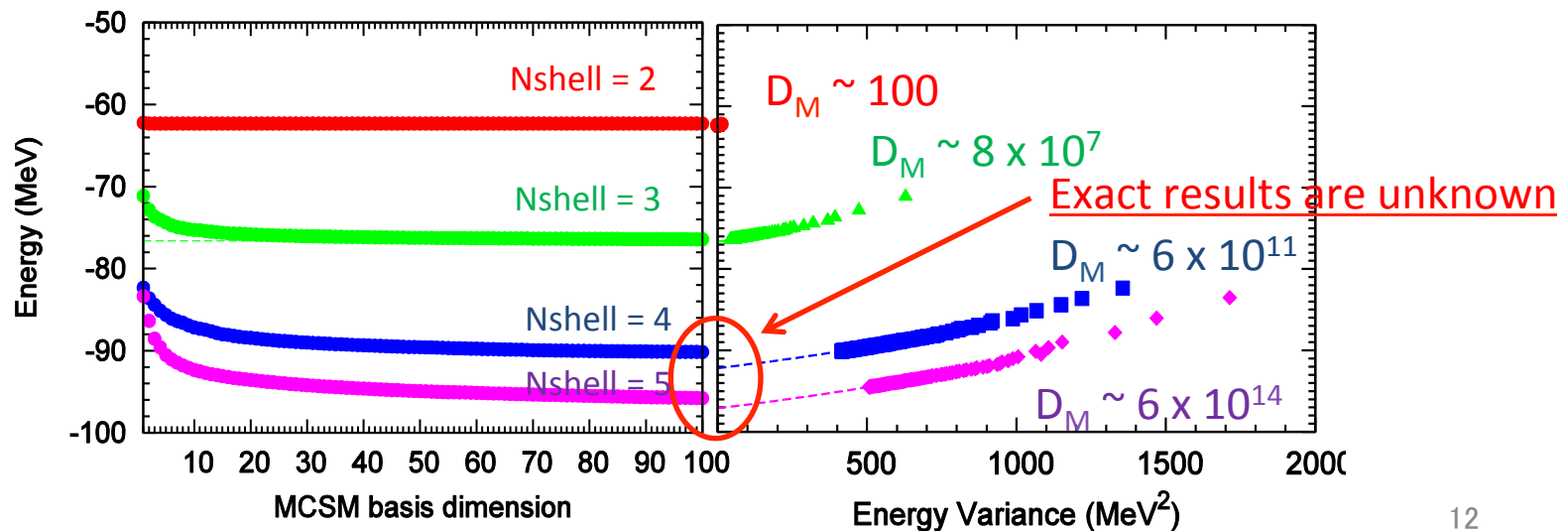
This coeff. is obtained by a stochastic sampling & CG.

Energies w.r.t. # of basis & energy variance

${}^4\text{He}(0^+; \text{gs})$



${}^{12}\text{C}(0^+; \text{gs})$



Feasibility study of MCSM for no-core calculations

PHYSICAL REVIEW C 86, 014302 (2012)

No-core Monte Carlo shell-model calculation for ^{10}Be and ^{12}Be low-lying spectra

Lang Liu (刘朗)*

Department of Physics, University of Tokyo, Hongo, Tokyo 113-0033, Japan, and State Key Laboratory of Nuclear Physics and Technology, School of Physics, Peking University, Beijing 100871, People's Republic of China

Takaharu Otsuka

Department of Physics and Center for Nuclear Study, University of Tokyo, Hongo, Tokyo 113-0033, Japan and National Superconducting Cyclotron Laboratory, Michigan State University, East Lansing, Michigan 48824, USA

Noritaka Shimizu

Department of Physics, University of Tokyo, Hongo, Tokyo 113-0033, Japan

Yutaka Utsuno

Japan Atomic Energy Agency, Tokai, Ibaraki 319-1195 Japan

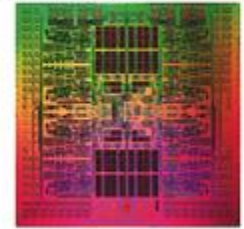
Robert Roth

Institut für Kernphysik, Technische Universität Darmstadt, D-64289 Darmstadt, Germany

(Received 24 April 2011; revised manuscript received 1 June 2012; published 3 July 2012)

Recent developments in the MCSM

- Energy minimization by the CG method
 - N. Shimizu, Y. Utsuno, T. Mizusaki, M. Honma, Y. Tsunoda & T. Otsuka, Phys. Rev. C85, 054301 (2012) ~ 30% reduction of # basis
- Efficient computation of TBMEs
 - Y. Utsuno, N. Shimizu, T. Otsuka & T. Abe, Compt. Phys. Comm. 184, 102 (2013) ~ 80% of the peak performance
- Energy variance extrapolation ($\sim 10\text{-}20\%$ in the old MCSM)
 - N. Shimizu, Y. Utsuno, T. Mizusaki, T. Otsuka, T. Abe & M. Honma, Phys. Rev. C82, 061305 (2010) Evaluation of exact eigenvalue w/ error estimate
- Summary of recent MCSM developments
 - N. Shimizu, T. Abe, Y. Tsunoda, Y. Utsuno, T. Yoshida, T. Mizusaki, M. Honma, T. Otsuka, Prog. Theor. Exp. Phys. 01A205 (2012)



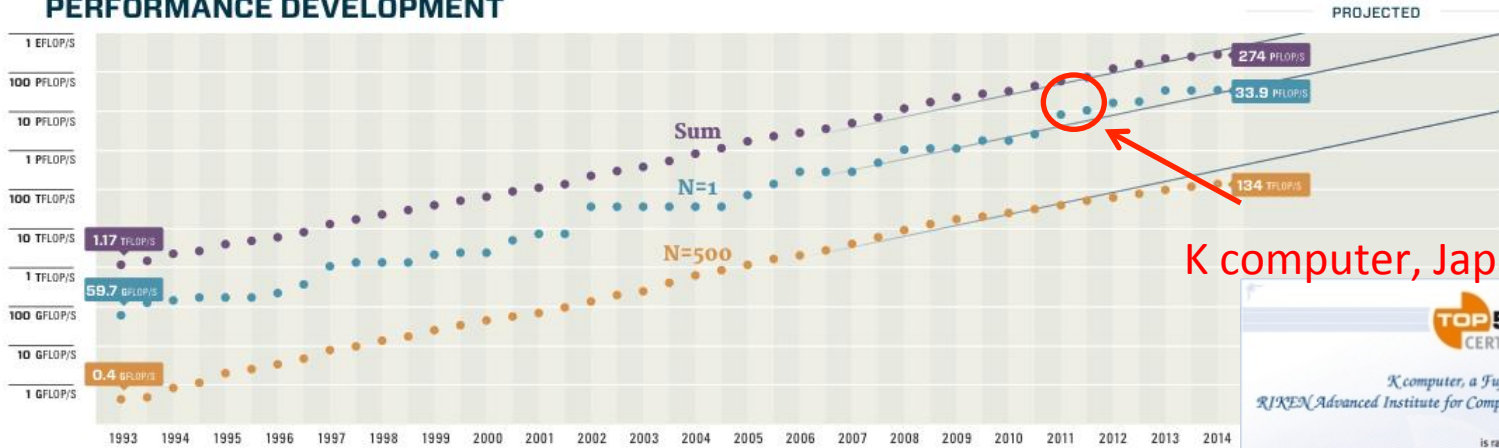
128 GFLOPS/CPU
(8 cores/CPU)

Tofu inter-connection
6D Mesh/Torus



NAME	SPECS	SITE	COUNTRY	CORES	R _{MAX} PFLDP/S	POWER MW
1 Tianhe-2 (Milkyway-2)	NUDT, Intel Ivy Bridge (12C, 2.2 GHz) & Xeon Phi (57C, 1.1 GHz), Custom interconnect	NSCC Guangzhou	China	3,120,000	33.9	17.8
2 Titan	Cray XK7, Operon 6274 (16C 2.2 GHz) + Nvidia Kepler GPU, Custom interconnect	DOE/SC/ORNL	USA	560,640	17.6	8.2
3 Sequoia	IBM BlueGene/Q, Power BQC (16C 1.60 GHz), Custom interconnect	DOE/NNSA/LLNL	USA	1,572,864	17.2	7.9
4 K computer	Fujitsu SPARC64 VIIIifx (8C, 2.0GHz), Custom interconnect	RIKEN AICS	Japan	705,024	10.5	12.7
5 Mira	IBM BlueGene/Q, Power BQC (16C, 1.60 GHz), Custom interconnect	DOE/SC/ANL	USA	786,432	8.59	3.95

PERFORMANCE DEVELOPMENT

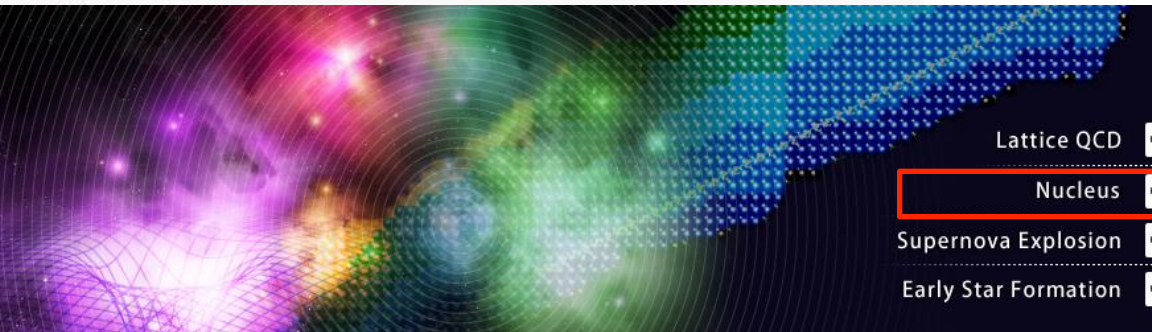


K computer, Japan



HPCI Strategic Program Field 5

"The origin of matter and the universe"



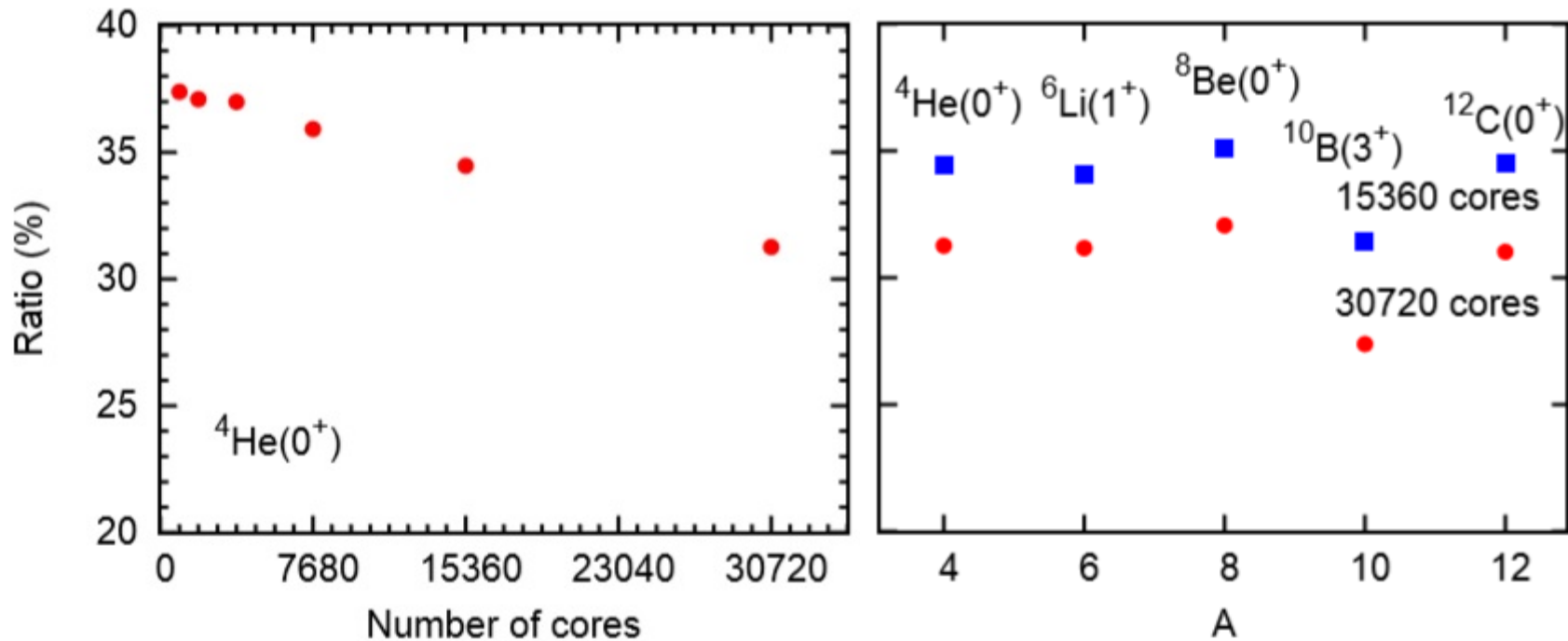
- Lattice QCD
- Nucleus
- Supernova Explosion
- Early Star Formation



Peak performance on the K computer

Peak performance

- Optimization of 15th basis dim. of the w.f. in $N_{\text{shell}} = 5$ w/ 100 CG iterations (MPI/OpenMP, 8 threads)

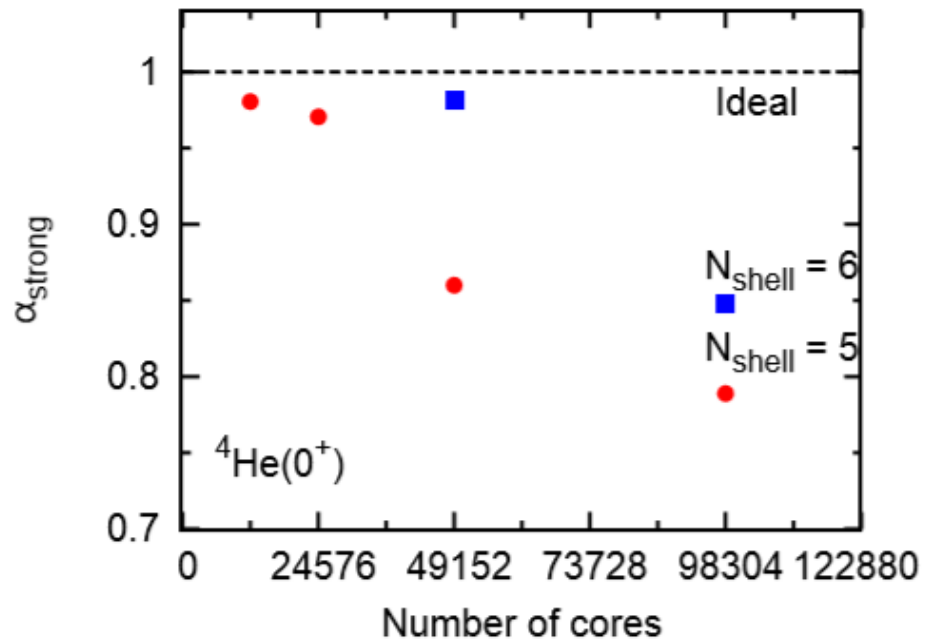
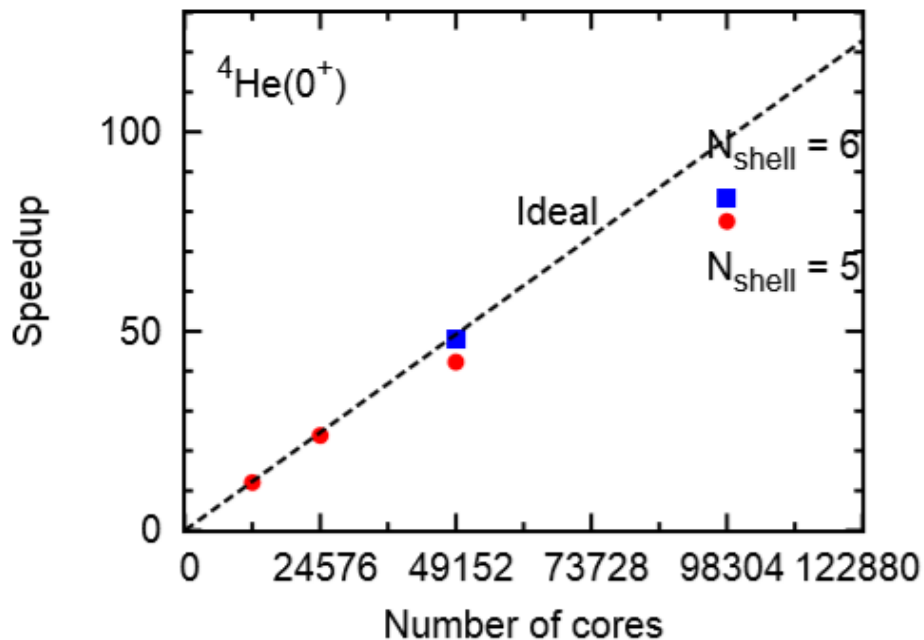


~30 % thru p-shell nuclei

Speed-up & strong scaling on the K computer

Speed-up (strong scaling)

- Optimization of 48th basis dim. of the ${}^4\text{He} (0^+)$ w.f. in $N_{\text{shell}} = 6$ w/ 100 CG iterations



Scaling up to ~ 100,000

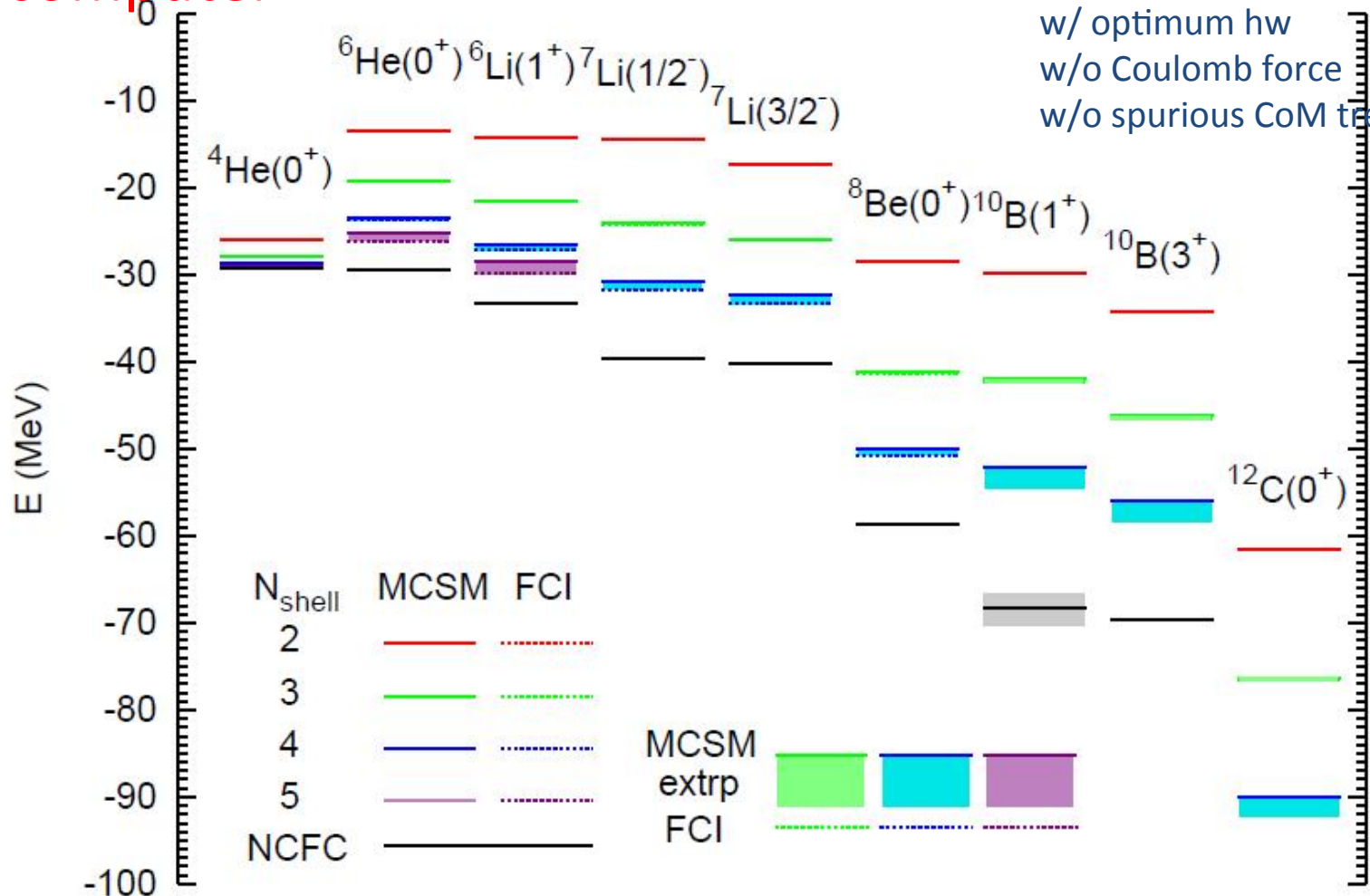
cores

Energies of the Light Nuclei

T. Abe, P. Maris, T. Otsuka, N. Shimizu, Y. Utsuno, J. P. Vary, Phys Rev C86, 054301 (2012)

Pre K computer

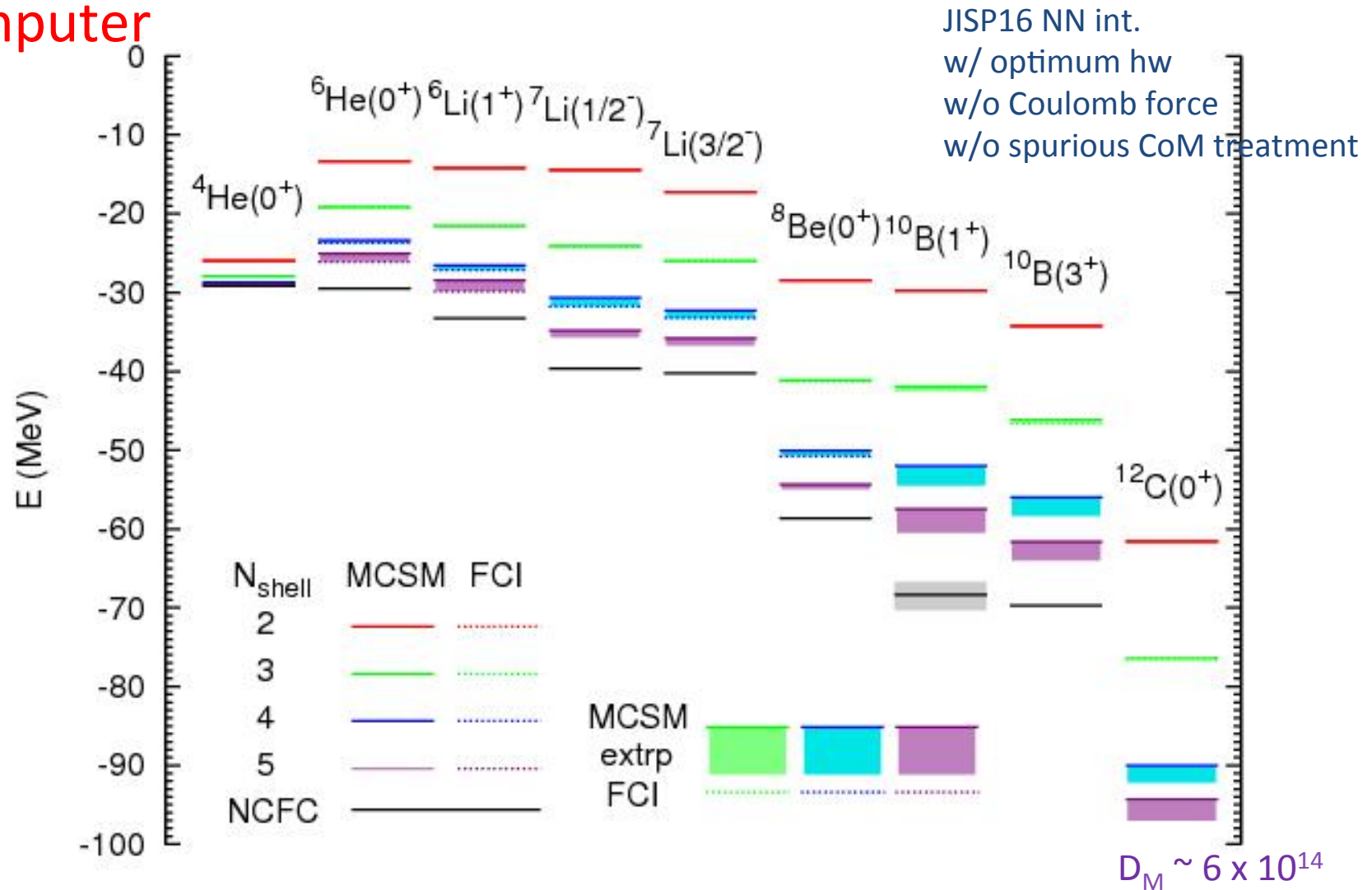
JISP16 NN int.
w/ optimum hw
w/o Coulomb force
w/o spurious CoM treatment



MCSM results w/ E-var extrp are consistent w/ FCI results

Energies of the Light Nuclei

K computer



Some MCSM results are not reachable in the current FCI

CPU time

core * hours

	$N_{\text{shell}} = 2$	$N_{\text{shell}} = 3$	$N_{\text{shell}} = 4$	$N_{\text{shell}} = 5$	$N_{\text{shell}} = 6$	$N_{\text{shell}} = 7$
${}^4\text{He} (0^+)$	1,300	2,000	2,400	10,000	70,000	400,000
${}^8\text{Be} (0^+)$	1,500	5,000	10,000	40,000	200,000	1,000,000
${}^{12}\text{C} (0^+)$	1,400	6,000	17,000	50,000	250,000	1,300,000
${}^{16}\text{O} (0^+)$	-----	6,000	15,000	70,000	280,000	1,400,000

FX10 @ U of Tokyo

K computer

For 100 bases

CPU time does not explode exponentially w.r.t. A & N_{shell} in the MCSM

“Ab initio” in low-energy nuclear structure physics

- Solve the non-relativistic Schroedinger eq.
and obtain the eigenvalues and eigenvectors.

$$H|\Psi\rangle = E|\Psi\rangle$$

$$H = T + V_{\text{NN}} + V_{\text{3N}} + \cdots + V_{\text{Coulomb}}$$

- **Ab initio**: All nucleons are active, and Hamiltonian consists of realistic NN (+ 3N) potentials.
- Two main sources of uncertainties:
 - **Nuclear forces** (interactions btw/among nucleons)
In principle, they should be obtained (directly) by QCD.
 - **Many-body methods**
CI: **Finite basis space** (choice of basis function and truncation),
we have to extrapolate to infinite basis dimensions

Extrapolations in the MCSM

- Two steps of the extrapolation
 1. Extrapolation of our MCSM (approx.) results to the FCI (exact) results in fixed model space

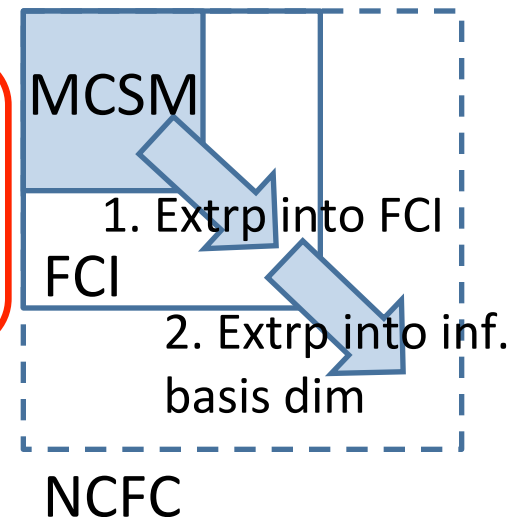
Energy-variance extrapolation

N. Shimizu, Y. Utsuno, T. Mizusaki, T. Otsuka, T. Abe, & M. Honma, Phys. Rev. C82, 061305(R) (2010)

2. Extrapolation into the infinite model space

- Exponential fit w.r.t. N_{\max} in the NCFC
- IR- & UV-cutoff extrapolations

Not applied on the MCSM, so far...



Extrapolation to the infinite basis space

- Two ways of the extrapolation to the infinite basis space

1. Empirical exponential form (w/ fixed hw)

$$E(N) = E(N = \infty) + a \exp(-bN)$$

P. Maris, A. M. Shirokov, & J. P. Vary, Phys. Rev. C79, 014308 (2009)

2. Cutoff extrapolations

- IR-cutoff extrapolation (w/ UV-saturated data)

$$E(\lambda) = E(\lambda = 0) + a \exp(-b/\lambda)$$

- IR- & UV-cutoff extrapolations (w/ any data, ideally)

$$E(\lambda, \Lambda) = E(\lambda = 0, \Lambda = \infty) + a \exp(-b/\lambda) + c \exp(-\Lambda^2/d^2)$$

S. A. Coon, M. I. Avetian, M. K. G. Kruse, U. van Kolck, P. Maris, J. P. Vary, Phys. Rev. C86, 054002 (2012)

S. A. Coon, arXiv:1303.6358

S. A. Coon, arXiv:1408.0738

R. J. Furnstahl, G. Hagen, T. Papenbrock, Phys. Rev. C86, 031301(R) (2012)

S. N. More, A. Ekstrom, R. J. Furnstahl, G. Hagen, T. Papenbrock, Phys. Rev. C87, 044326 (2013)

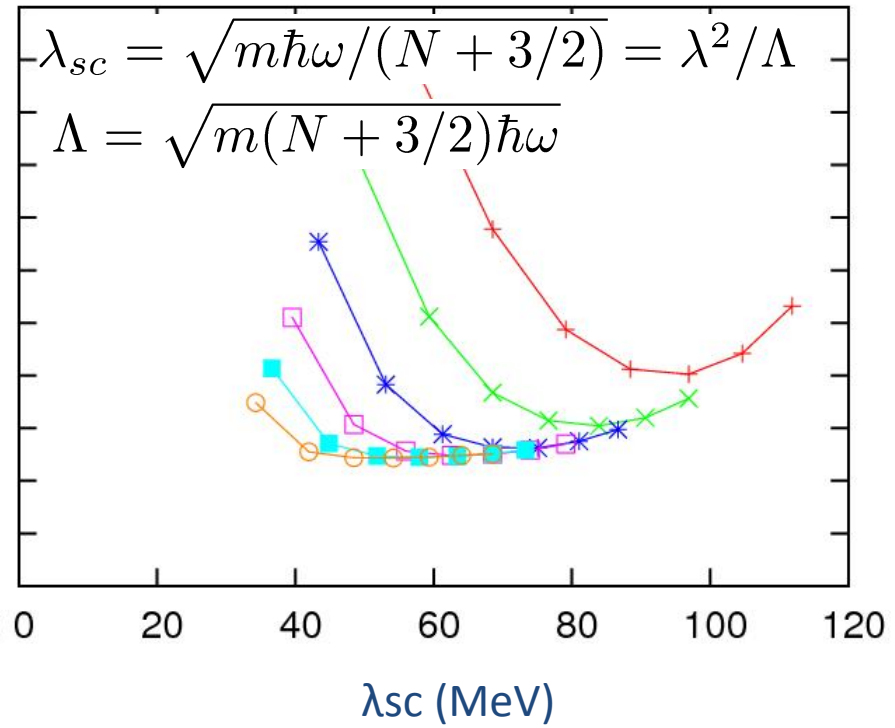
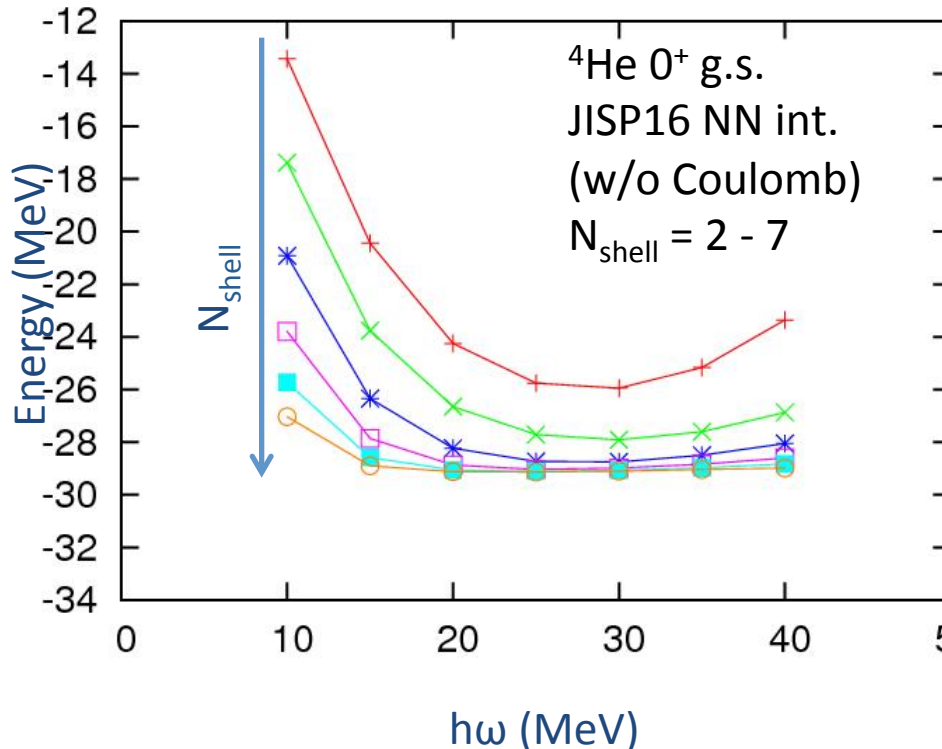
R. J. Furnstahl, S. N. More, T. Papenbrock, arXiv:1312.6876

R. J. Furnstahl, G. Hagen, T. Papenbrock, K. A. Wendt, arXiv:1408.0252

E. D. Jurgenson, P. Maris, R. J. Furnstahl, W. E. Ormand & J. P. Vary, Phys. Rev. C87, 054312

(2013)

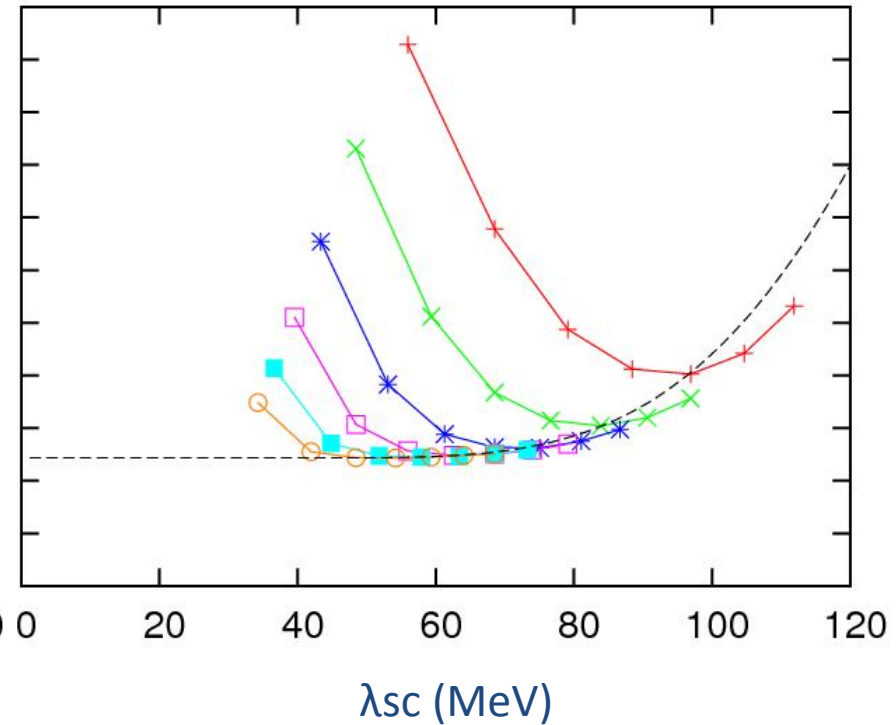
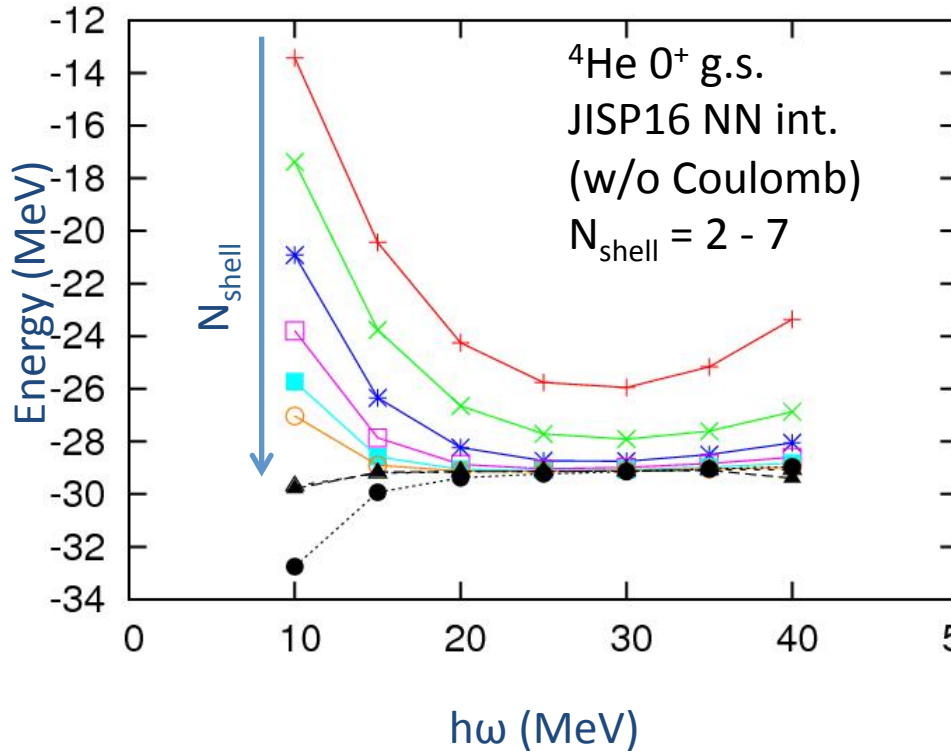
Empirical & IR-cutoff extrapolations



Λ : UV cutoff λ_{sc} : IR cutoff

IR cutoff scaling w/ UV saturated data

Empirical & IR-cutoff extrapolations



MCSM(empirical): $-29.389 \sim -29.077$ MeV
 ($N_{\text{shell}} = 3 - 7$, $h\omega = 15 - 35$ MeV)

$O(100)$ keV error?

c.f.) NCFC: $-29.164(2)$ MeV

Extrapolated results to infinite N_{max}

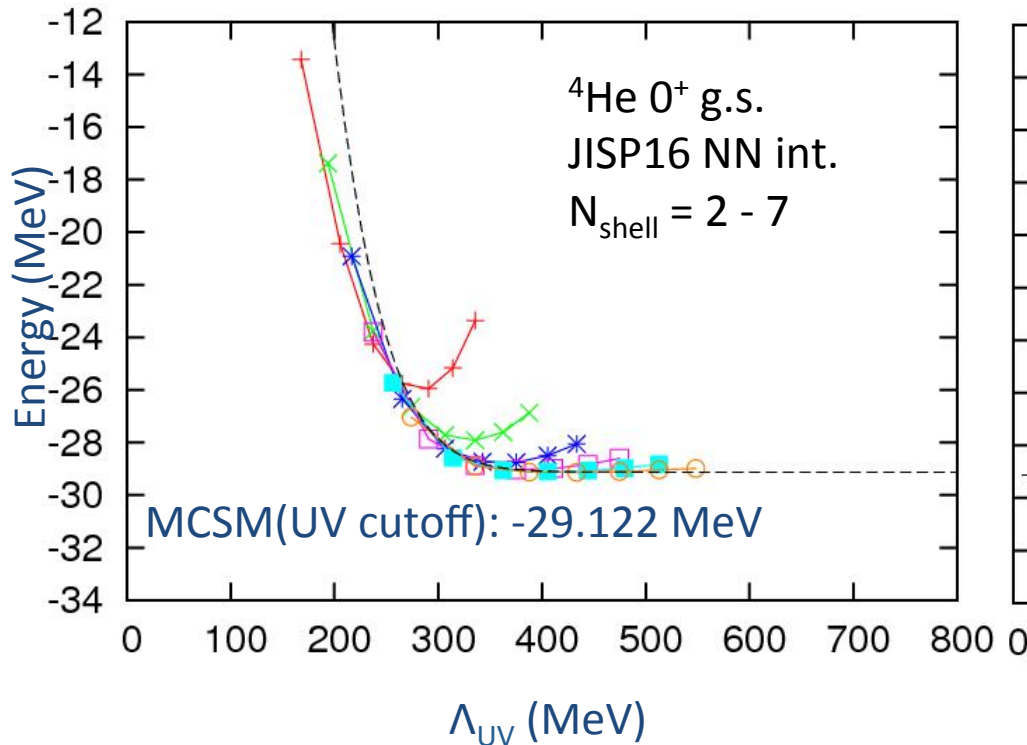
MCSM(IR cutoff): ~ -29.142 MeV
 (w/ UV-saturated data)

$O(10)$ keV error?

Error estimates are needed.

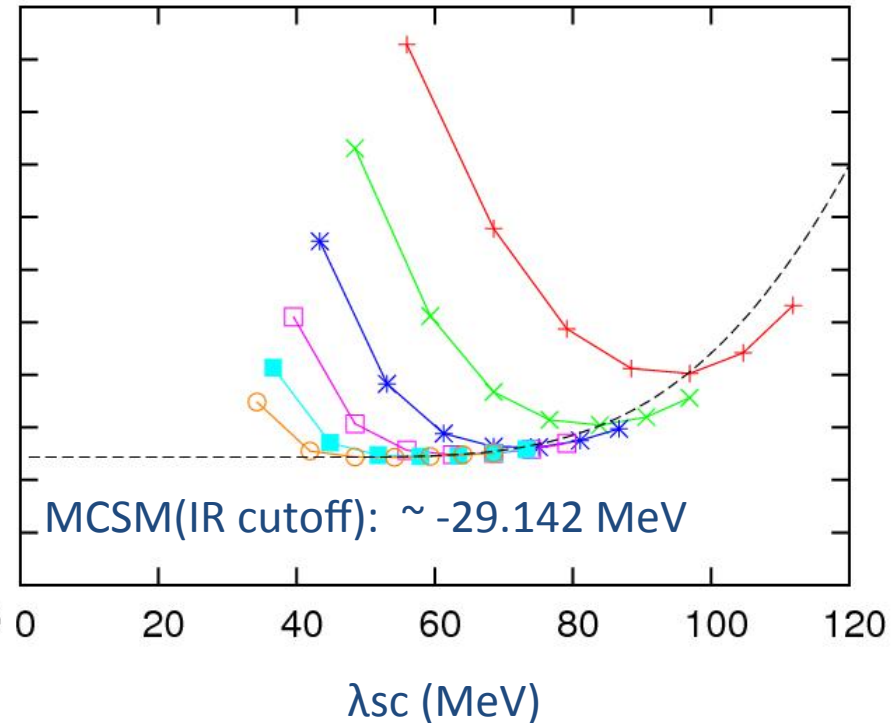
UV-cutoff extrapolation

$$E(\Lambda) = E(\Lambda = \infty) + c \exp(-\Lambda^2/d^2)$$



IR-cutoff extrapolation

$$E(\lambda) = E(\lambda = 0) + a \exp(-b/\lambda)$$



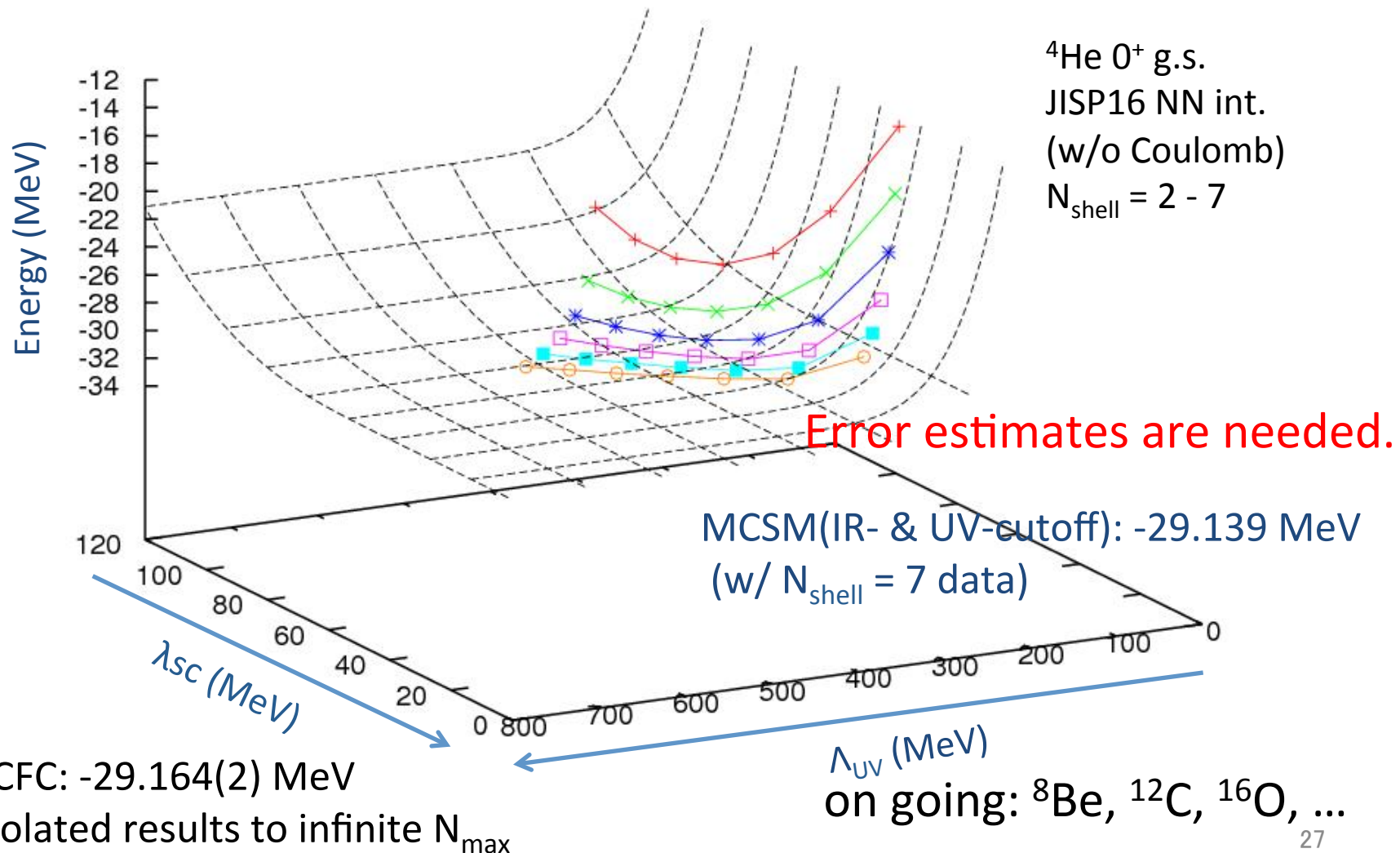
c.f.) NCFC: -29.164(2) MeV
 Extrapolated results to infinite N_{max}

on going: ${}^8\text{Be}$, ${}^{12}\text{C}$, ${}^{16}\text{O}$, ...

Preliminary

IR- & UV-cutoff extrapolation

$$E(\lambda, \Lambda) = E(\lambda = 0, \Lambda = \infty) + a \exp(-b/\lambda) + c \exp(-\Lambda^2/d^2)$$



Density Plots from ab initio calc.

- Green's function Monte Carlo (GFMC)

- “Intrinsic” density is constructed by aligning the moment of inertia among samples

R. B. Wiringa, S. C. Pieper, J. Carlson, & V. R. Pandharipande, Phys. Rev. C62, 014001 (2000)

- No-core full configuration (NCFC)

- Translationally-invariant density is obtained by deconvoluting the intrinsic & CM w.f.

C. Cockrell J. P. Vary & P. Maris, Phys. Rev. C86, 034325 (2012)

- Lattice EFT

- Triangle structure in carbon-12

E. Epelbaum, H. Krebs, T. A. Lahde, D. Lee, & U.-G. Meissner, Phys. Rev. Lett. 109, 252501 (2012)

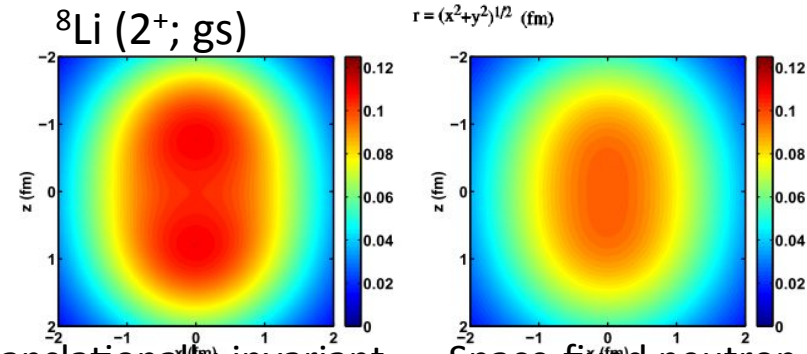
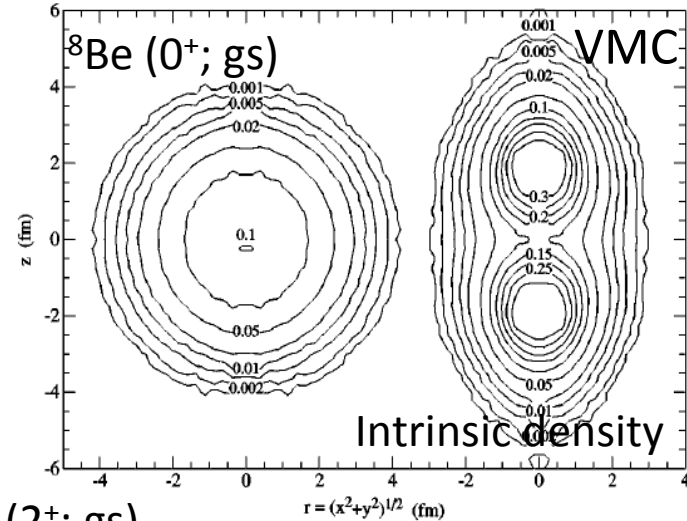
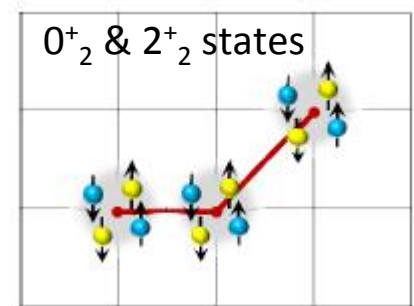
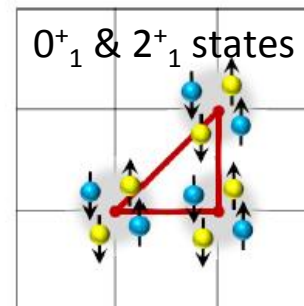


FIG. 12: (Color online) The $y = 0$ slice of the translationally-invariant neutron density (left) of the 2^+ gs of ${}^8\text{Li}$. The space-fixed neutron density (right) is shown in the right. These densities were calculated with $\beta = 10$ and $\hbar\Omega = 12.5$ MeV.



Density plots in MCSM

Talk by T. Yoshida on Oct 9

(Thur)

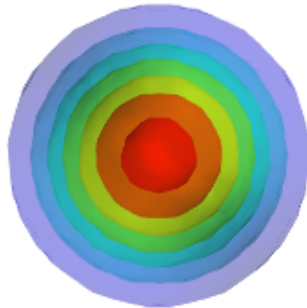
$$|\Phi\rangle = \sum_{i=1}^{N_{basis}} c_i |\Phi_i\rangle = c_1 \text{img}_1 + c_2 \text{img}_2 + c_3 \text{img}_3 + c_4 \text{img}_4 + \dots$$

Angular-momentum projection

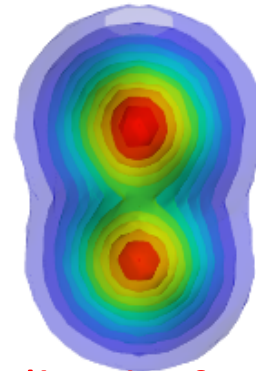
$$|\Psi\rangle = \sum_{i=1}^{N_{basis}} c_i P^J P^\pi |\Phi_i\rangle$$

Rotation of each basis by diagonalizing Q-moment

$$|\Phi'\rangle = \sum_{i=1}^{N_{basis}} c_i R(\Omega_i) |\Phi_i\rangle$$



$^8\text{Be } 0^+$ ground state



Laboratory frame

“Intrinsic” (body-fixed) frame

Densities in lab. & body-fixed frames can be constructed by MCSM

Preliminary Density plots of Be isotopes (0_1^+)

($N_{\text{shell}} = 4$, $hw = 25$ MeV, $\beta = 0$, JISP16 NN w/o Coulomb)

Talk by T. Yoshida on Oct 9

A = 8

A = 9

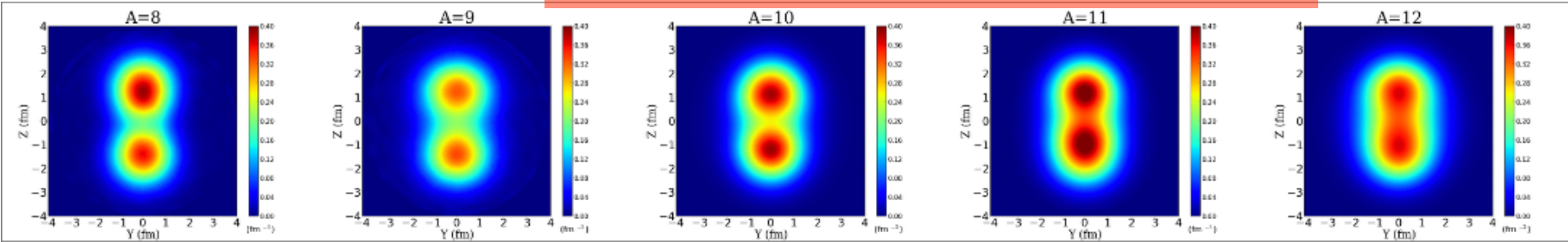
A = 10

(Thur) A = 11

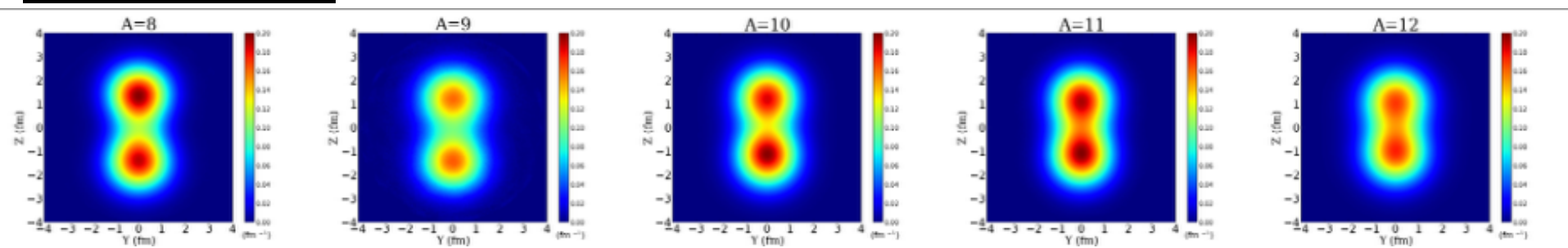
A = 12

Matter density

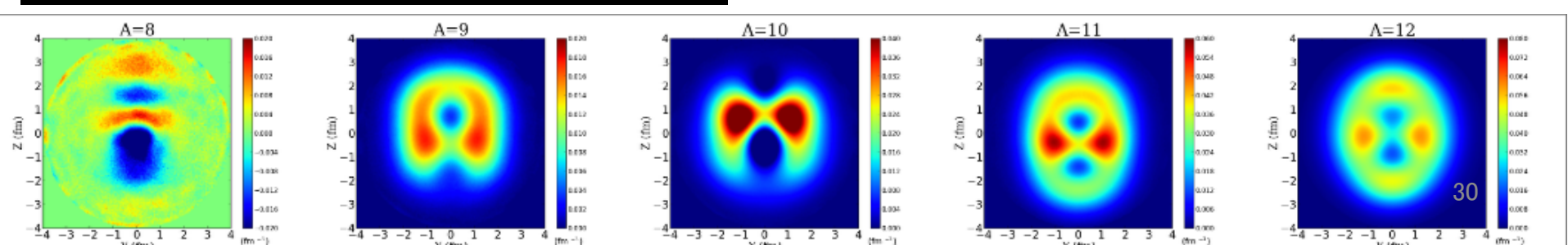
2- α structure is vanishing as A increases



Proton density



Neutron density – Proton density



Summary

- MCSM can be applied to no-core calculations of the p-shell nuclei.
 - Benchmarks for the p-shell nuclei have been performed and gave good agreements w/ FCI results. Some results are obtained only by MCSM.
 - Extension to larger model spaces ($N_{\text{shell}} = 6, 7, \dots$), extrapolation to infinite basis space, & comparison with the another truncation (N_{max})

Perspective

- MCSM algorithm/computation
 - Error estimates of the extrapolations
 - Inclusion of the 3-body force (thru. effective 2-body force)
 - GPGPU
- Physics
 - Cluster states, non-yrast states, unnatural parity states, ...
 - sd-shell nuclei

END