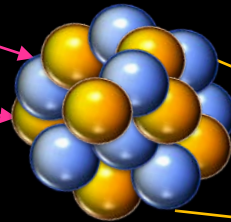


# From QCD to Compact Stars

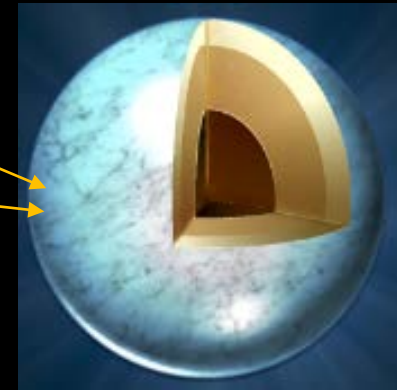
nucleon  $\sim 1$  [fm]



nucleus  $\sim 10$  [fm]



Neutron star  $\sim 10$  [km]



Advances and Perspectives in Computational Nuclear Physics  
Hilton Waikoloa Village (Hawaii, Oct. 6, 2014)  
Tetsuo Hatsuda (RIKEN)

# Plan of this Talk

1. Introduction : **phase structure** of QCD

2. **Baryon forces** from LQCD

Inoue et al. [HAL QCD Coll.], PRL 106 (2011) 162002

3. **Nuclear & Neutron Matter** from LQCD + BHF

Inoue et al. [HAL QCD Coll.], PRL 111 (2013) 112503

4. **Finite Nuclei** from LQCD + BHF

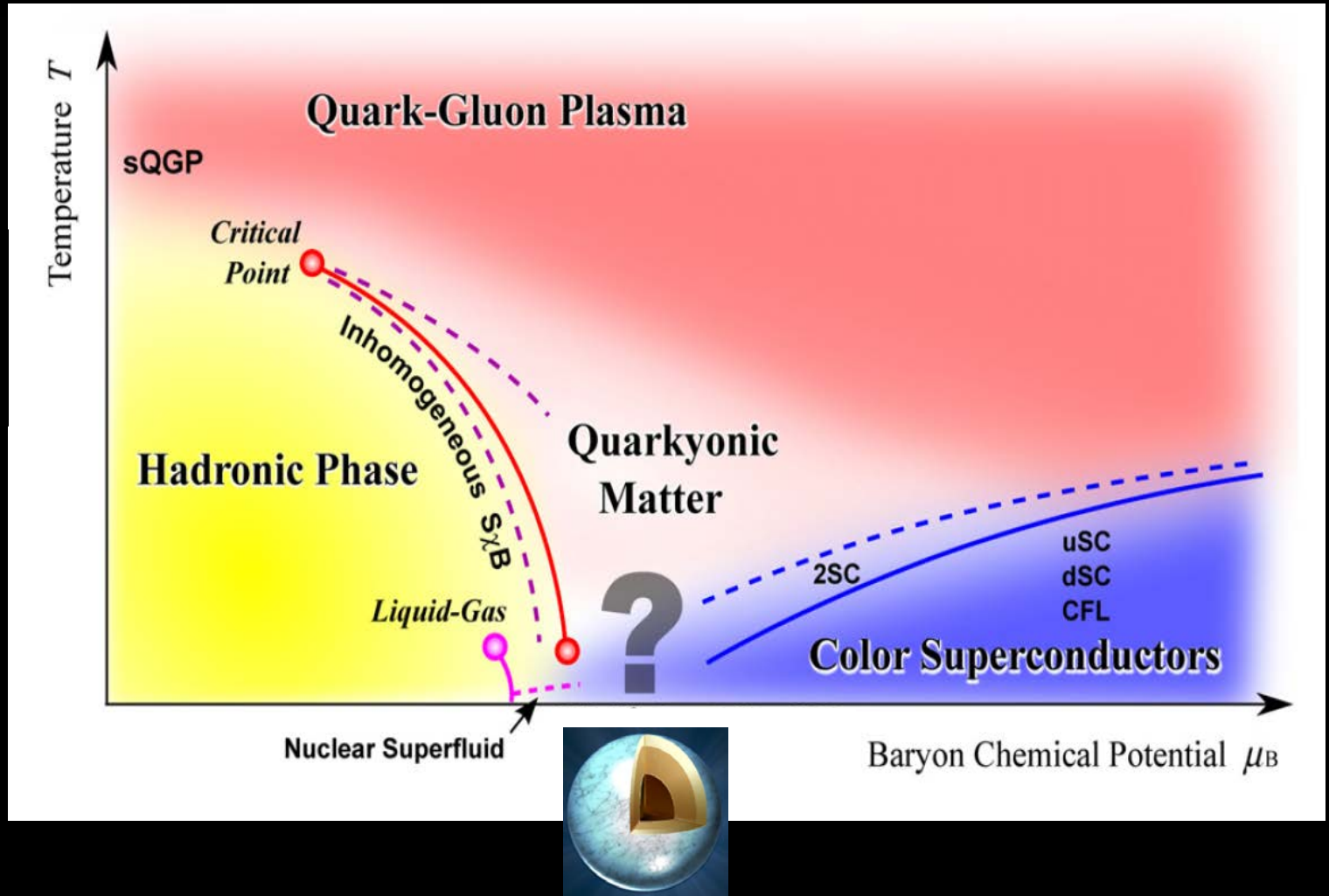
Inoue et al. [HAL QCD Coll.], arXiv: 1408.4892

5. Summary

Exotic (strange and charmed) Hadrons ( $H$ ,  $N\Omega$ ,  $T_{cc}$ ,  $T_{cs}$ ,  $Z_c$  etc)

⇒ Y. Ikeda (Oct.7, morning)

# Possible QCD Phase Structure

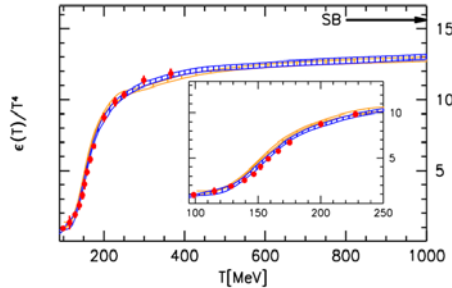


K. Fukushima and T. Hatsuda, "The Phase Diagram of Dense QCD"  
Rep. Prog. Phys. 74 (2011) 014001

# From QCD to Hot/Dense Matter

## Quantum Chromo Dynamics

Lattice gauge simulations



sign  
problem



Phenomenological  
nuclear force

Baryon interactions

Many-body  
techniques

Equation of State for Hot Matter

Relativistic  
hydrodynamics

Relativistic heavy-ion collisions

Equation of State for Dense Matter

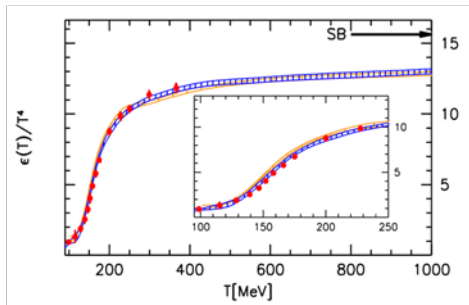
General relativity

Neutron stars

# From QCD to Hot/Dense Matter

## Quantum Chromo Dynamics

Lattice gauge simulations



sign  
problem



Lattice  
gauge theory

Baryon interactions

Many-body  
techniques

Equation of State for Hot Matter

Relativistic  
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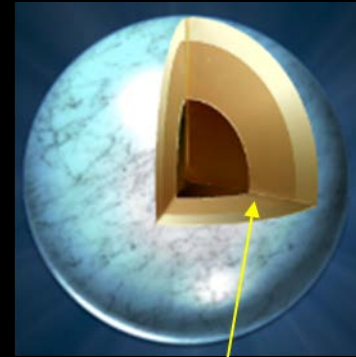
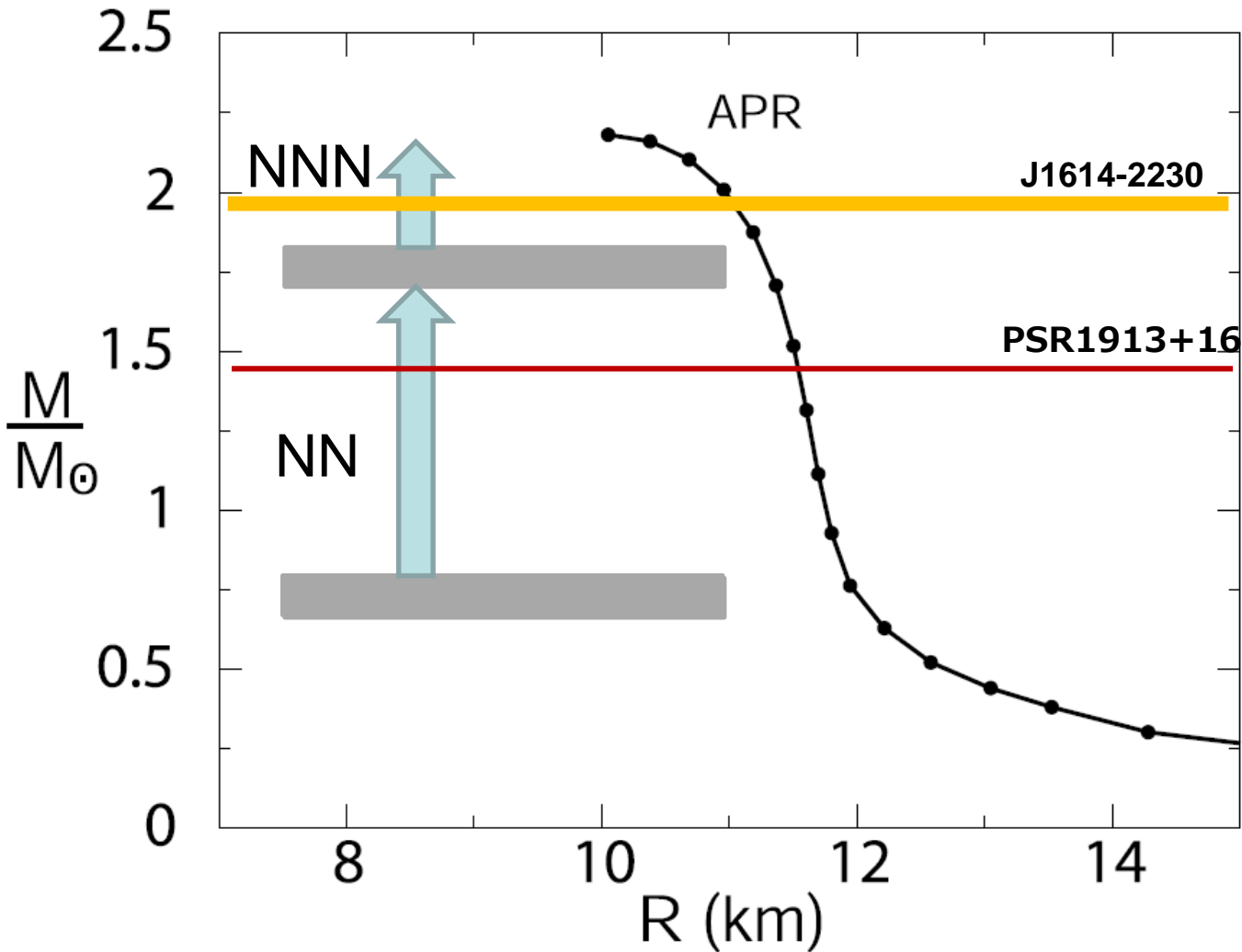
Relativistic heavy-ion collisions

Equation of State for Dense Matter

General relativity

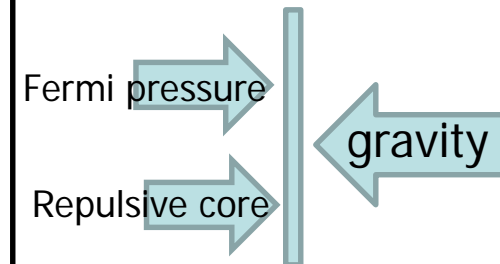
Neutron stars

# Mass-Radius relation of $N_{\star}$ (nucleons only)

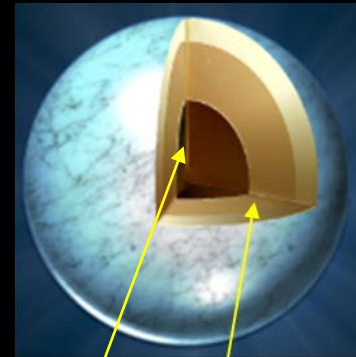
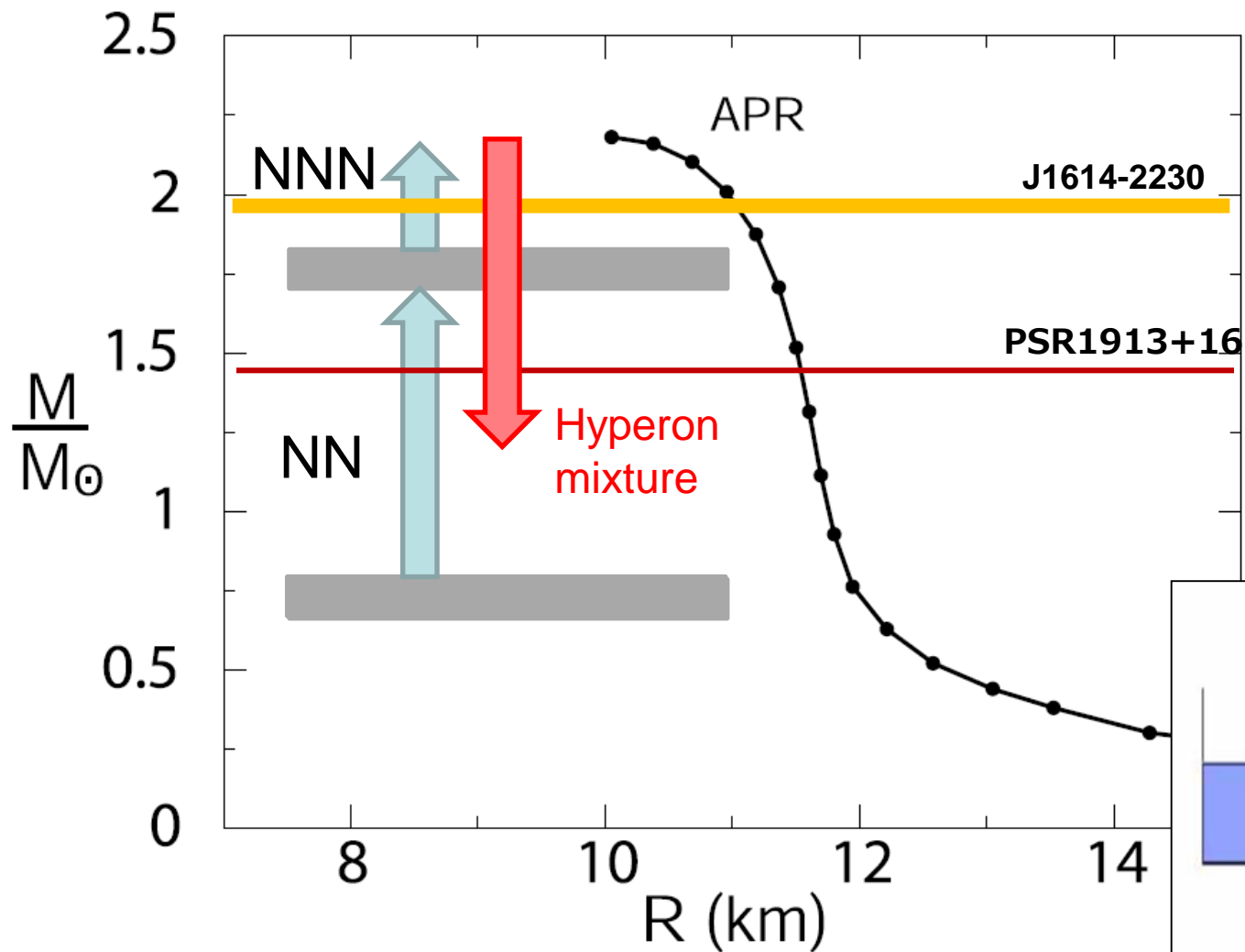


neutrons

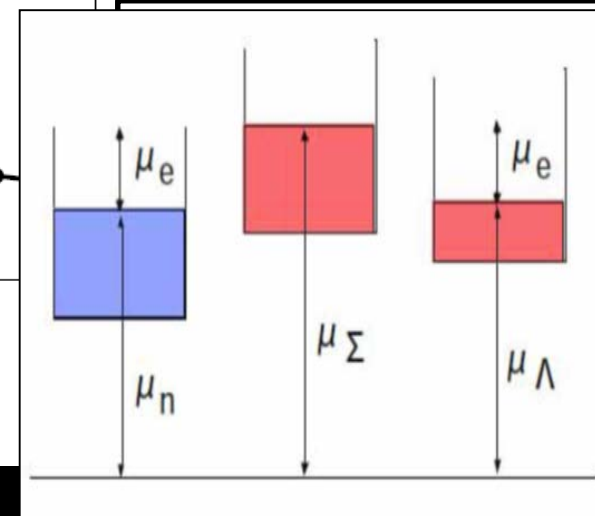
## Pressure balance



# Hyperon Crisis (Takatsuka et al., 2002)



hyperons  
neutrons



HPCI Program (FY2011-2015) Field 5: All Japan Computational Physics Collaboration

## The Origin of Matter and the Universe

(particle physics – nuclear physics – astrophysics, 11 institutions)



Lattice 2015 (July)

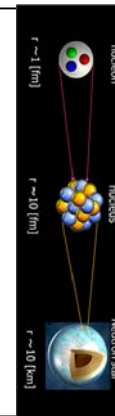
Quark Matter 2015 (Oct.)

Project 1: Baryon-Baryon interaction from lattice QCD

Project 2: Nuclear quantum many-body calculation

Project 3: Supernova explosion and black-hole formation

Project 4: First stars and galaxies





Present un-physical point simulation  
for single and multi-baryons

On-going physical point simulation  
for single and multi-baryons in K

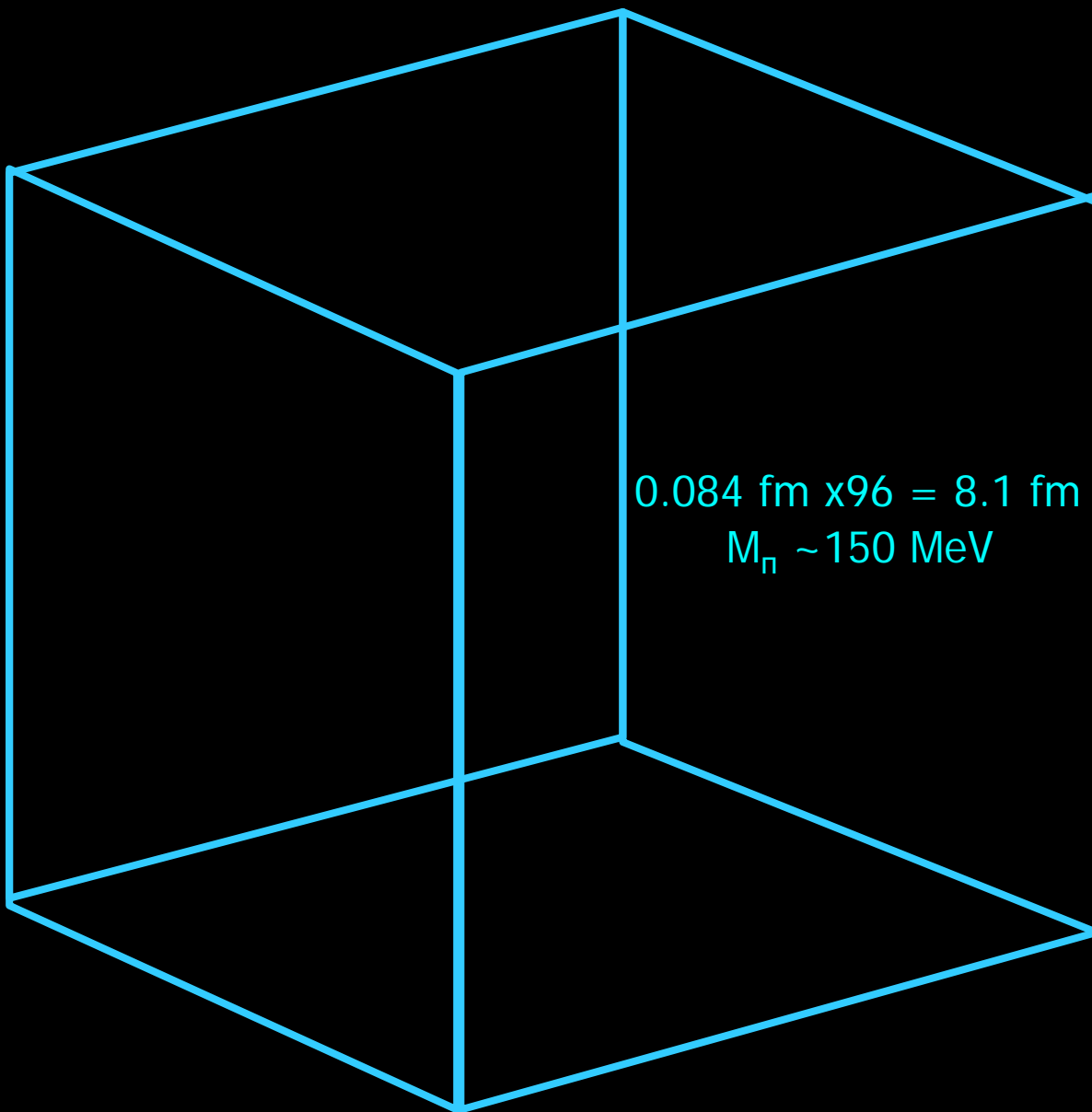
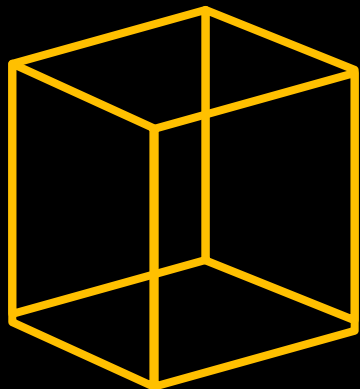
NEXT YEAR'S TALK  
Stay Tuned !!



TODAY'S TALK

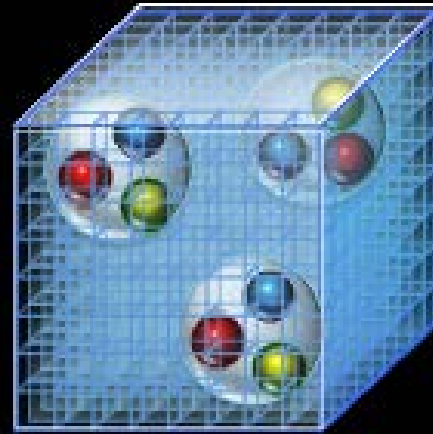
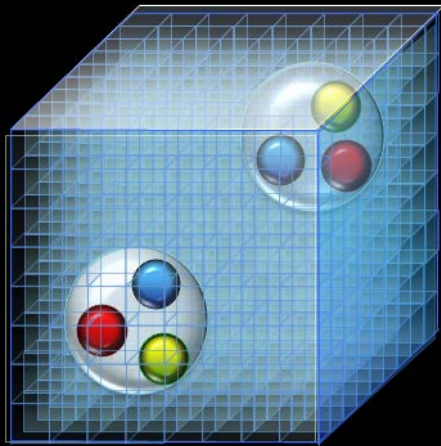


$0.121 \text{ fm} \times 32 = 3.9 \text{ fm}$   
 $m_\pi = 350\text{-}1200 \text{ MeV}$



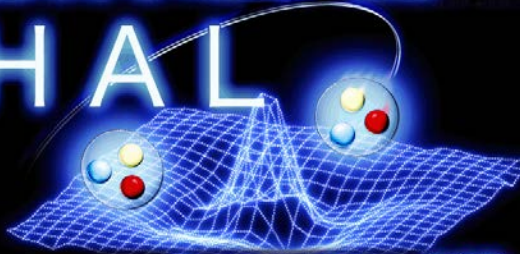
$0.084 \text{ fm} \times 96 = 8.1 \text{ fm}$   
 $M_\pi \sim 150 \text{ MeV}$

# Baryon Forces from LQCD



Hadrons to Atomic nuclei

# HAL



from Lattice QCD

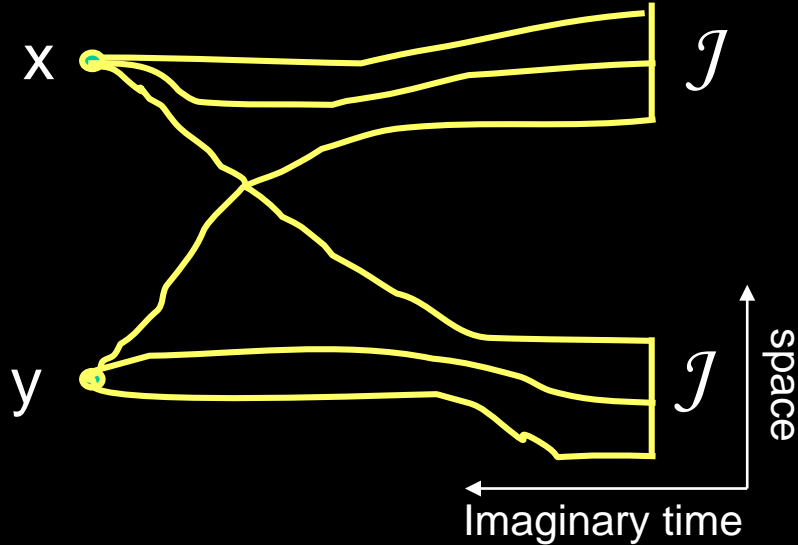
Univ. Tsukuba  
Univ. Tokyo  
RIKEN  
Nihon Univ.  
YITP (Kyoto)  
RCNP (Osaka)  
Birjand

T. Miyamoto, H. Nemura, K. Sasaki, M. Yamada  
B. Charron  
T. Doi, T. Hatsuda, Y. Ikeda, V. Krejcirik  
T. Inoue  
S. Aoki  
N. Ishii, K. Murano  
F. Etminan

Review: "Lattice QCD Approach to Nuclear Physics"

HAL QCD Collaboration, Prog. Theor. Exp. Phys. 2012 (2012) 01A105

# Hadronic correlations in LQCD



$$\begin{aligned} & \langle N_1(\mathbf{x}, t) N_2(\mathbf{y}, t) \mathcal{J}_1^\dagger(0) \mathcal{J}_2^\dagger(0) \rangle \\ &= \sum_n \langle 0 | N_1(\mathbf{x}) N_2(\mathbf{y}) | n \rangle a_n e^{-E_n t} \\ & \xrightarrow{t > t^*} \phi(\mathbf{r}, t) = \sum_{n < n^*} b_n \phi_n(\mathbf{r}) e^{-E_n t} \end{aligned}$$

## Finite Volume Method

$E_n(L) \rightarrow$  phase shift

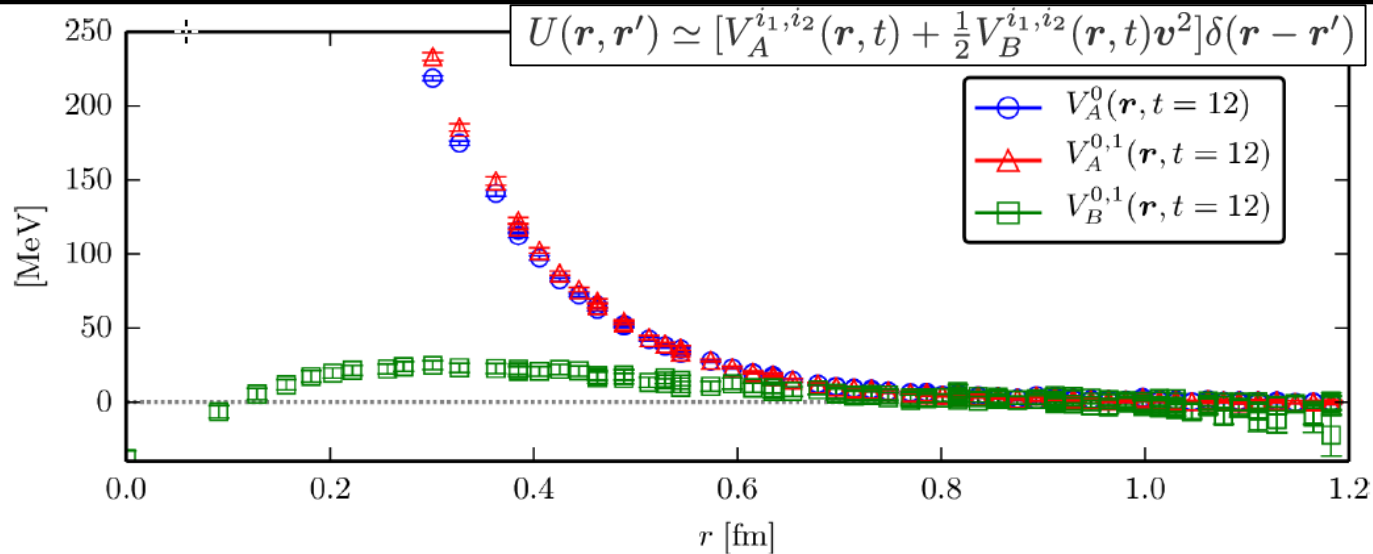
Luescher, Nucl. Phys. B354 (1991) 531

## HAL QCD Method

$\phi(r, t) \rightarrow$  kernel  $\rightarrow$  phase shift  
 $T=V+GVT$

Ishii, Aoki & Hatsuda, PRL 99 (2007) 022001  
 Ishii et al. [HAL QCD Coll.], PLB 712 (2012) 437

# Benchmark test : FV Method vs. HAL QCD Method



$N_f = 2+1$   
 $a = 0.091$  fm  
 $N_s^3 \times N_t = 32^3 \times 64$   
 $m_\pi = 700$  MeV

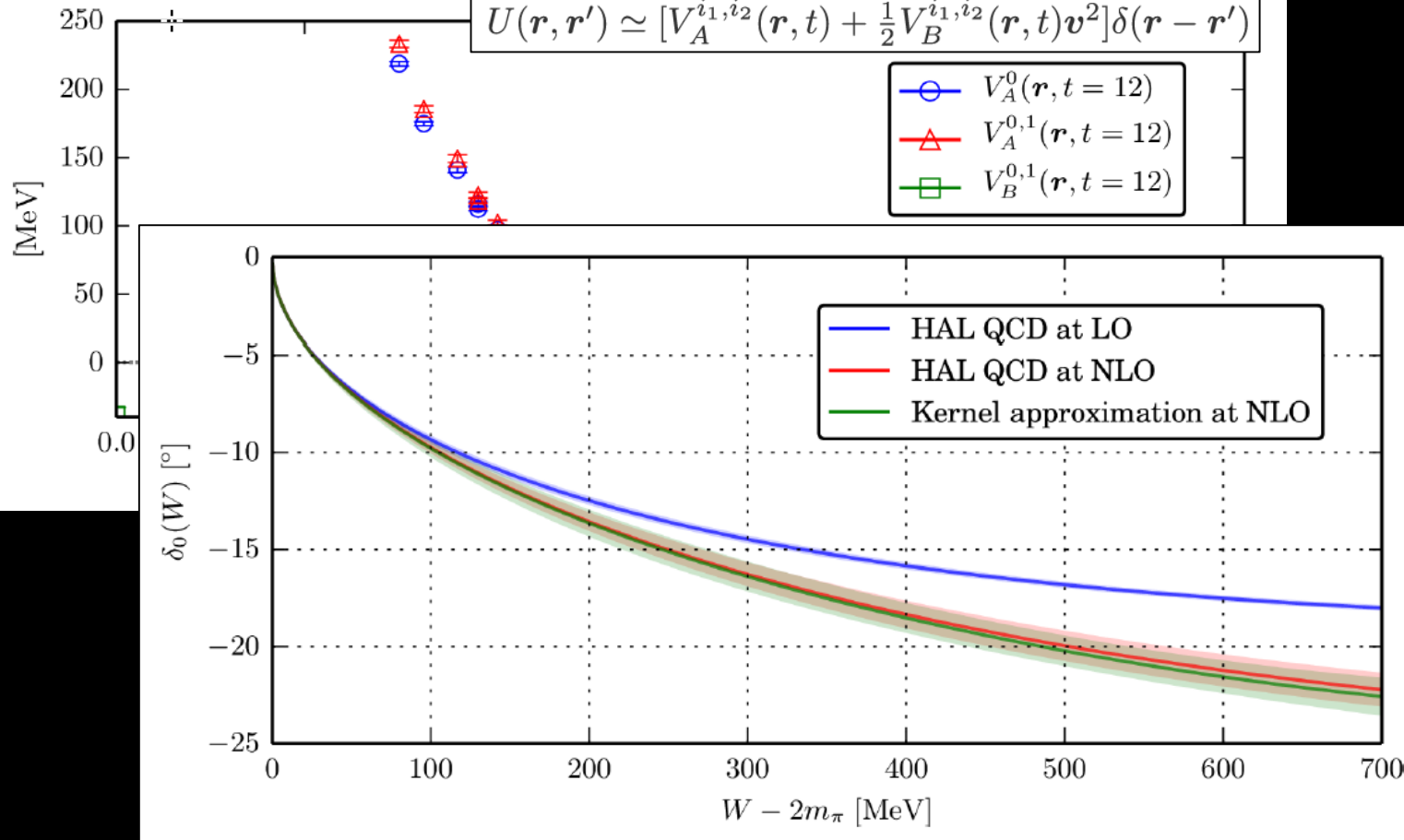
B. Charron, PhD thesis, Univ. Tokyo (2014)

See also, Kurth, Ishii, Doi, Aoki & Hatsuda, JHEP 1312 (2013) 015

# Benchmark test : FV Method vs. HAL QCD Method

$$U(\mathbf{r}, \mathbf{r}') \simeq [V_A^{i_1, i_2}(\mathbf{r}, t) + \frac{1}{2} V_B^{i_1, i_2}(\mathbf{r}, t) v^2] \delta(\mathbf{r} - \mathbf{r}')$$

$N_f = 2+1$   
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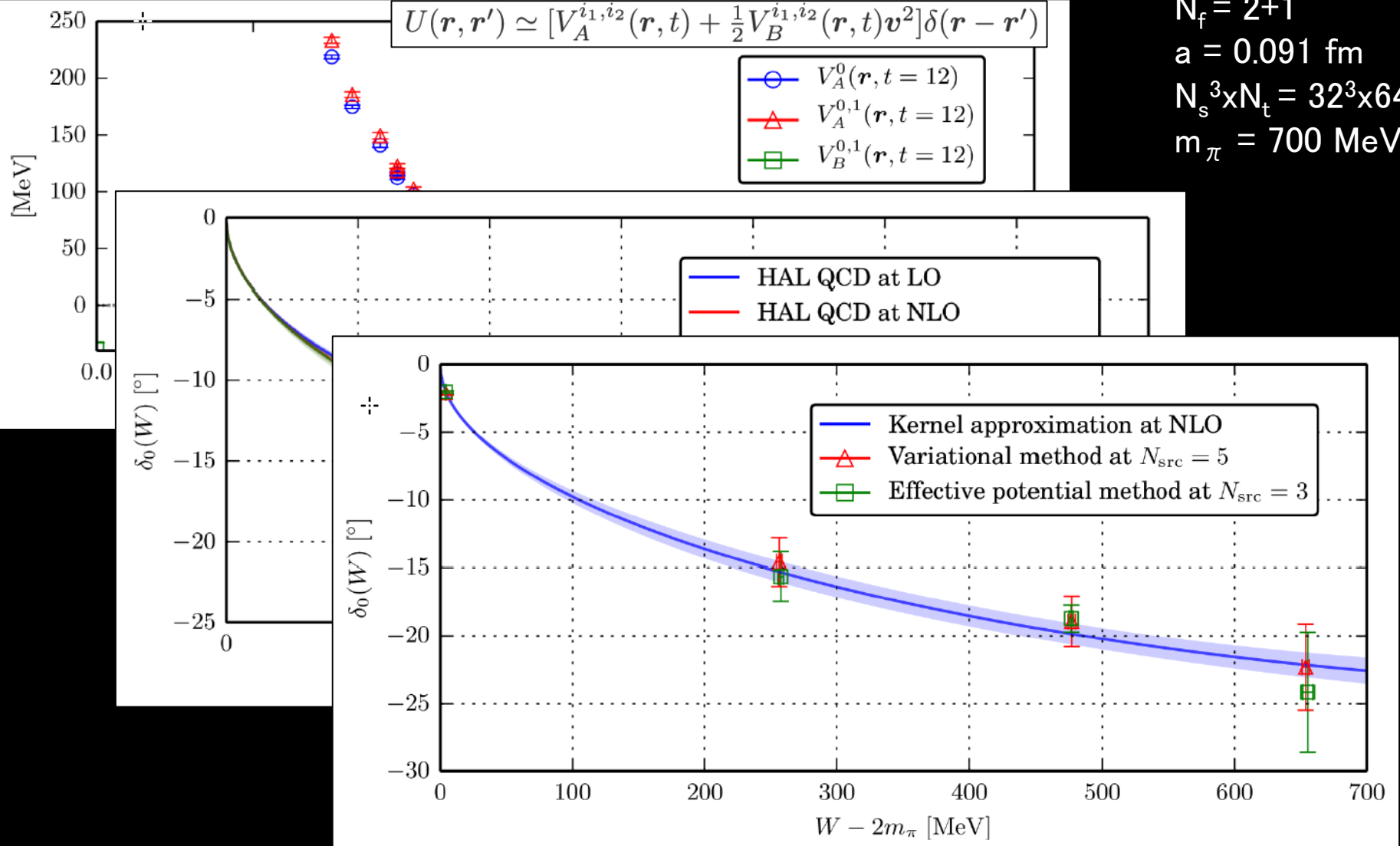
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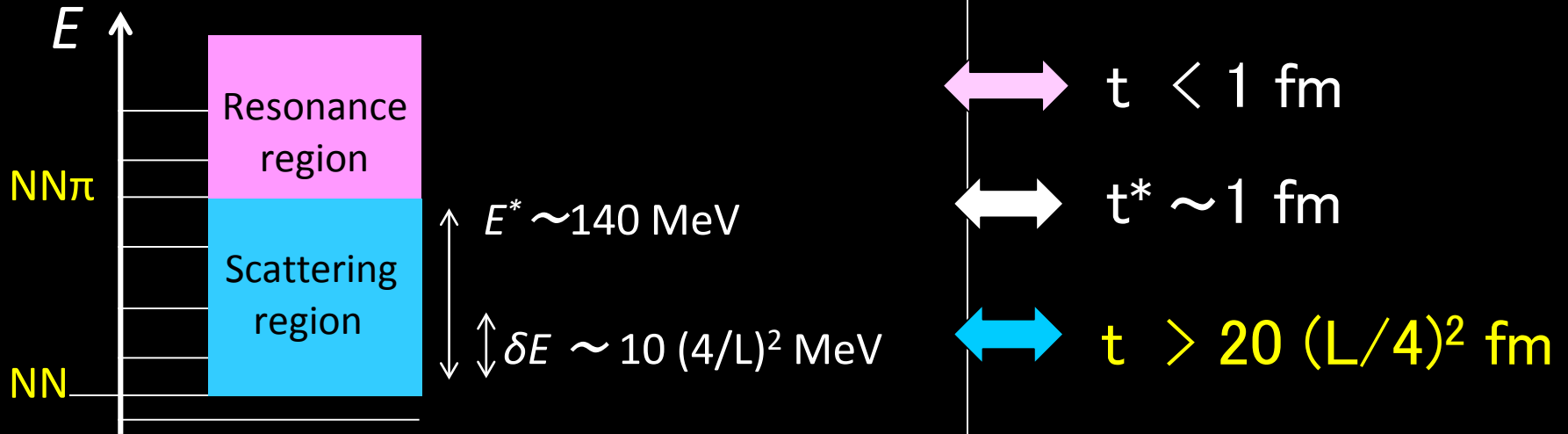
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B. Charron, PhD thesis, Univ. Tokyo (2014)

See also, Kurth, Ishii, Doi, Aoki & Hatsuda, JHEP 1312 (2013) 015

# What about NN ?



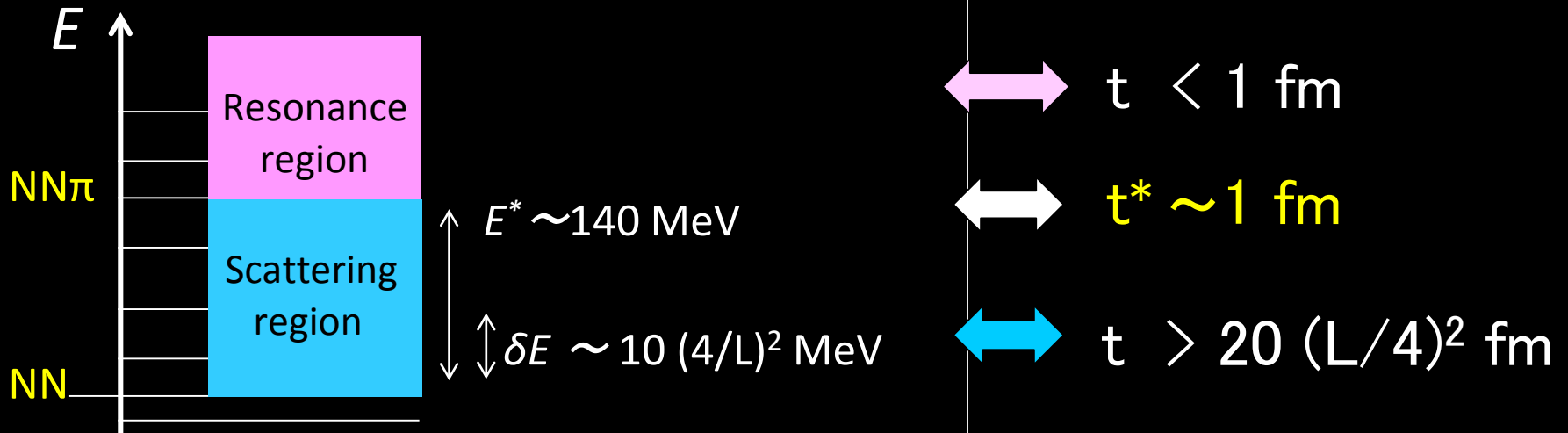
$$\left(\frac{\mathcal{S}}{\mathcal{N}}\right)_{NN} \sim \sqrt{N_{gc}} e^{-2(m_N - 3m_\pi/2)t}$$

$$\left(\frac{\mathcal{S}}{\mathcal{N}}\right)_{\pi\pi} \sim \sqrt{N_{gc}}$$

## Finite Volume Method

- Large  $t$  and huge statistics are required for NN (not for  $\pi\pi$ ) for small  $m_\pi$  and large  $L$

# What about NN ?

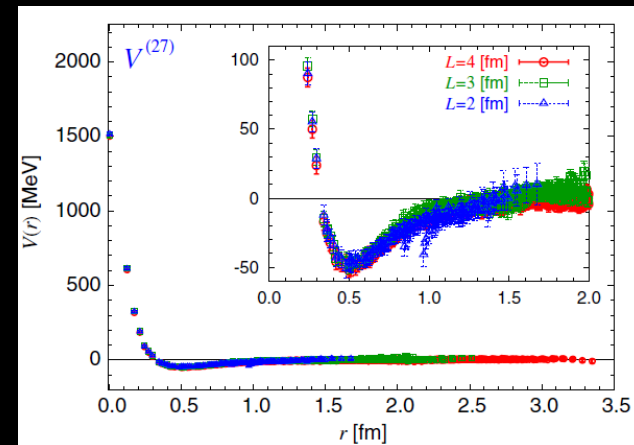


$$\left[ \left( \frac{1}{2} \frac{\partial}{\partial t} \right)^2 - \nabla^2 + m_N^2 \right] \phi(\mathbf{r}, \mathbf{t}) = m_N \int U(\mathbf{r}, \mathbf{r}') \phi(\mathbf{r}') d^3 r'$$

## HAL QCD Method

- All data for  $t > 1$  fm are signals
- $U(r, r')$  : non-local kernel  
E-independent, L-insensitive  $\rightarrow$  observables

Ishii et al. [HAL QCD Coll.], PLB 712 (2012) 437



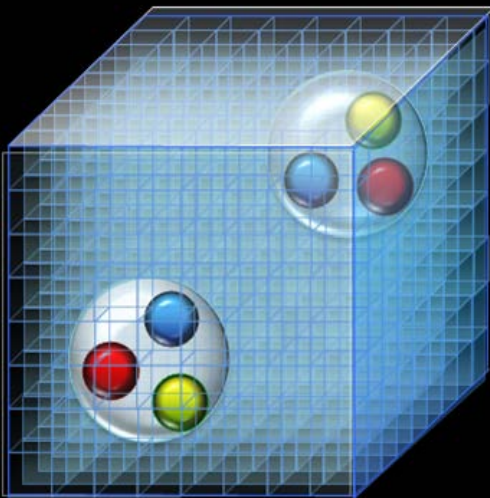


# Lattice setup

to be used in the following analyses

## 3 degenerate flavors

### Iwasaki gauge action + clover improved Wilson fermion



size	$\beta$	$C_{sw}$	$a$ [fm]	$L$ [fm]
$32^3 \times 32$	1.83	1.761	0.121(2)	3.87

K_uds	M_P.S. [MeV]	M_B <sup>8</sup> [MeV]
0.13660	1170.9(7)	2274(2)
0.13710	1015.2(6)	2031(2)
0.13760	836.8(5)	1749(1)
0.13800	672.3(6)	1484(2)
0.13840	468.9(8)	1161(2)

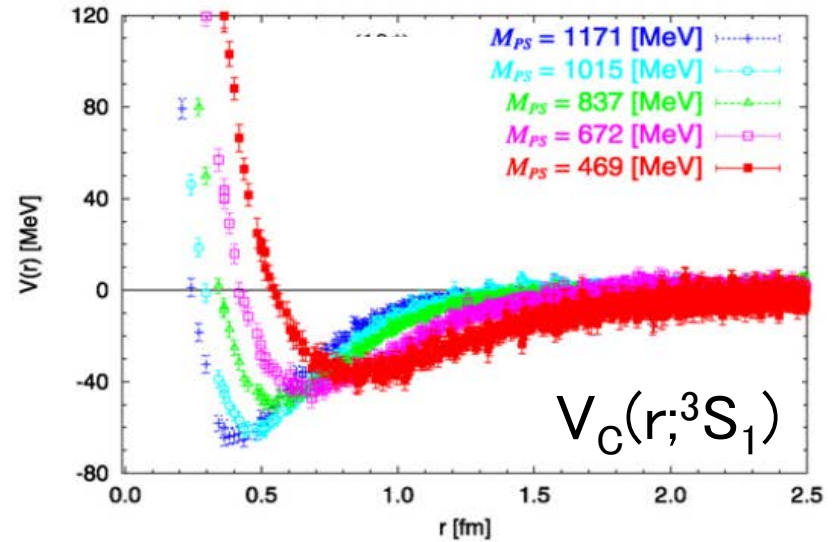
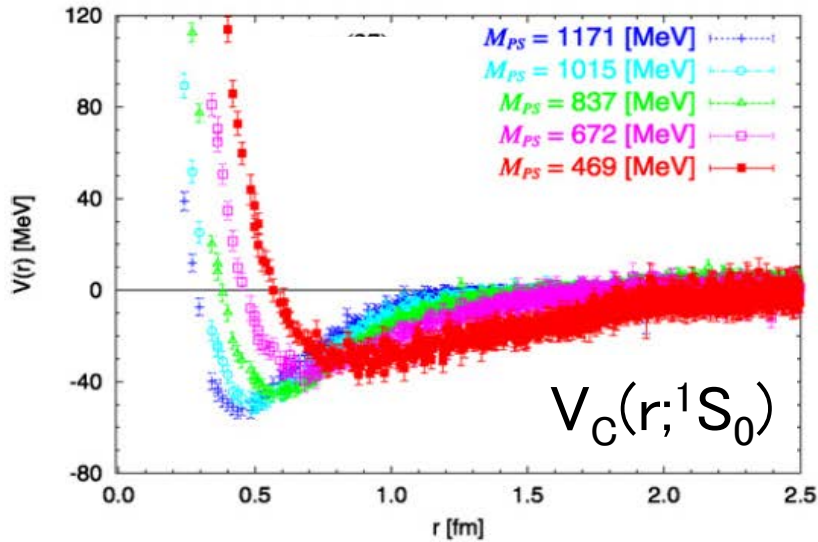
$$V(\vec{r}, \nabla) = V_C(r) + S_{12}V_T(r) + \vec{L} \cdot \vec{S} V_{LS}(r) + \{V_D(r), \nabla^2\} + \dots$$

LO

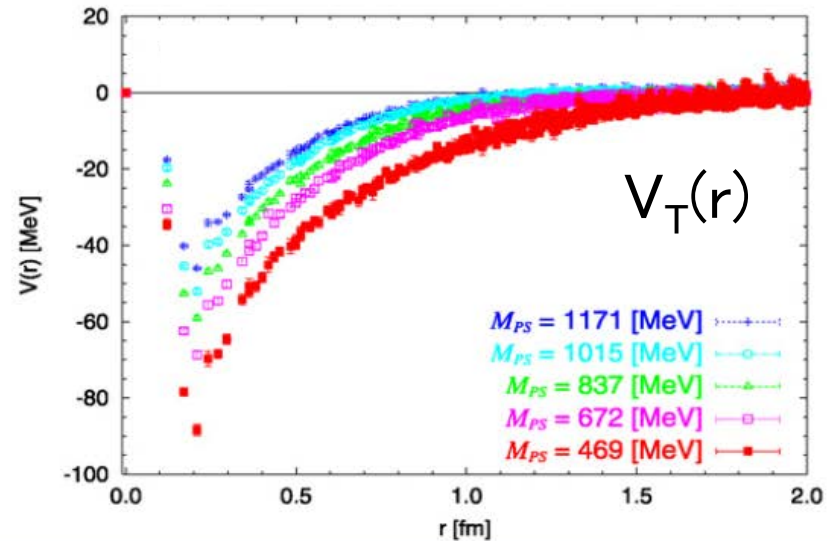
LO

NLO

NNLO

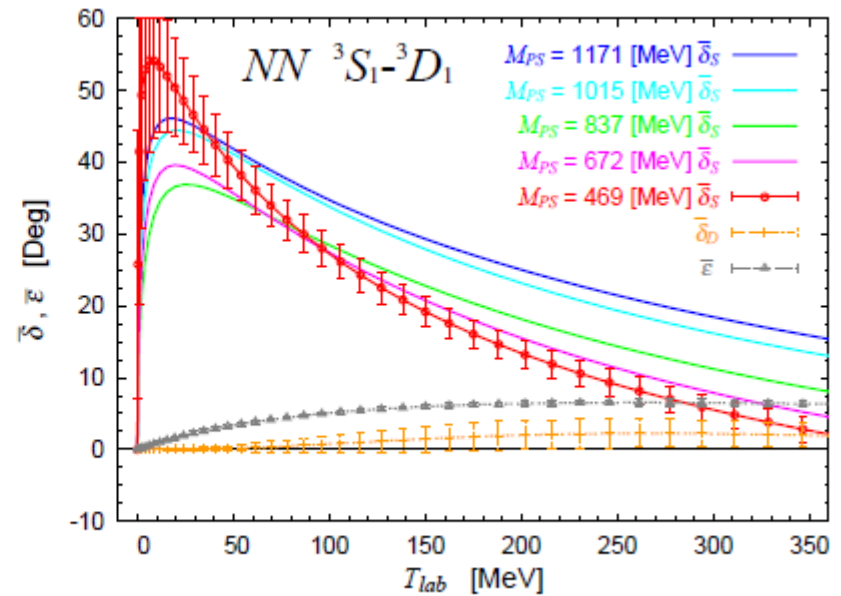
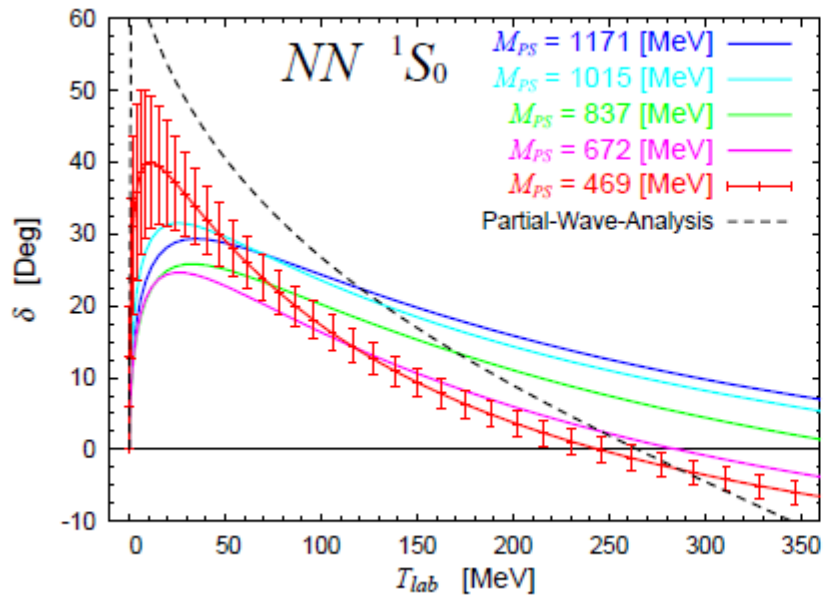


NN  
Central & Tensor Forces  
in 3-flavor QCD



HAL QCD Coll.  
Phys. Rev. Lett. 106 (2011) 162002  
Nucl. Phys. A881 (2012) 28

# NN phase shifts in 3-flavor QCD



Stronger attraction in the deuteron channel

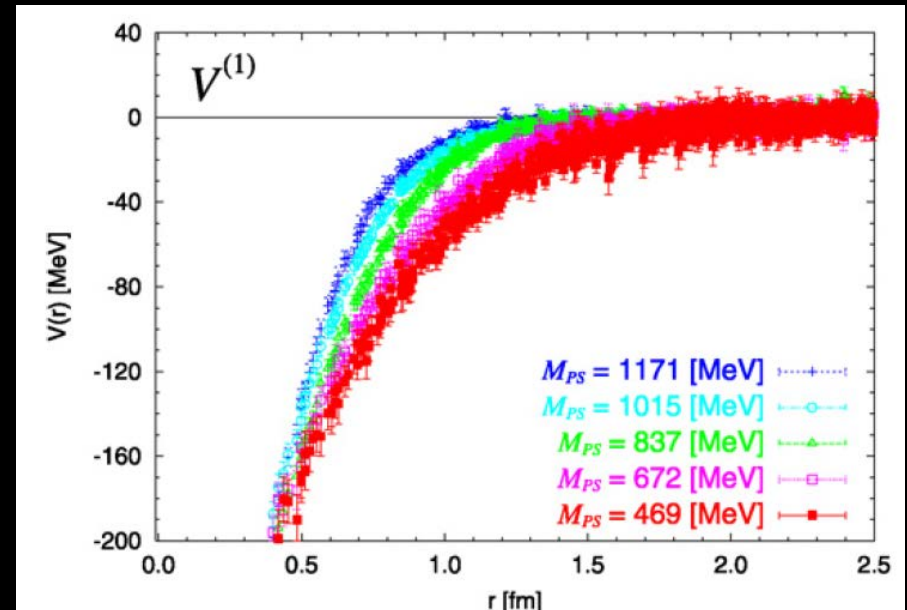
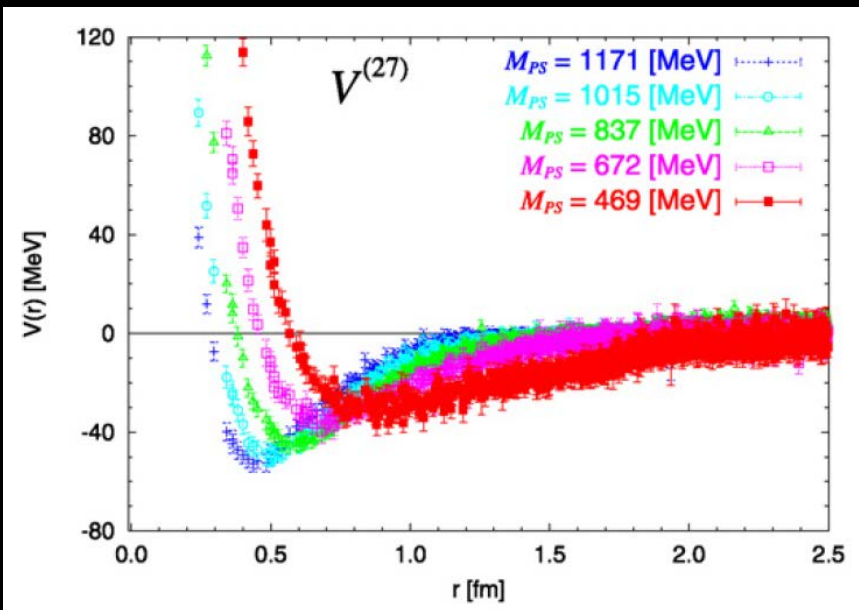
HAL QCD Coll.,  
Phys. Rev. Lett. 106 (2011) 162002  
Nucl. Phys. A881 (2012) 28

# BB Forces in 3-flavor QCD

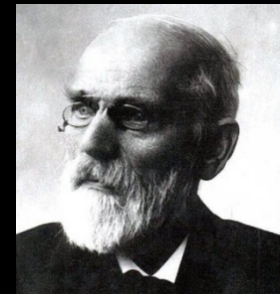
HAL QCD Coll.  
Phys. Rev. Lett. 106 (2011) 162002  
Nucl. Phys. A881 (2012) 28

PP (uud-uud) channel  
(partial) Pauli blocking

H (uds-uds) channel  
No Pauli blocking



Pauli and van der Waarls  
at work !



# BB Forces in 3-flavor QCD

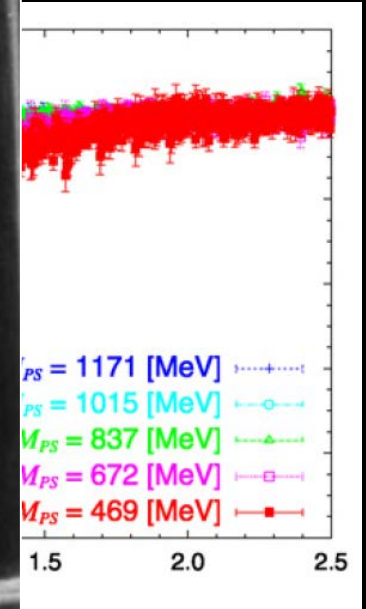
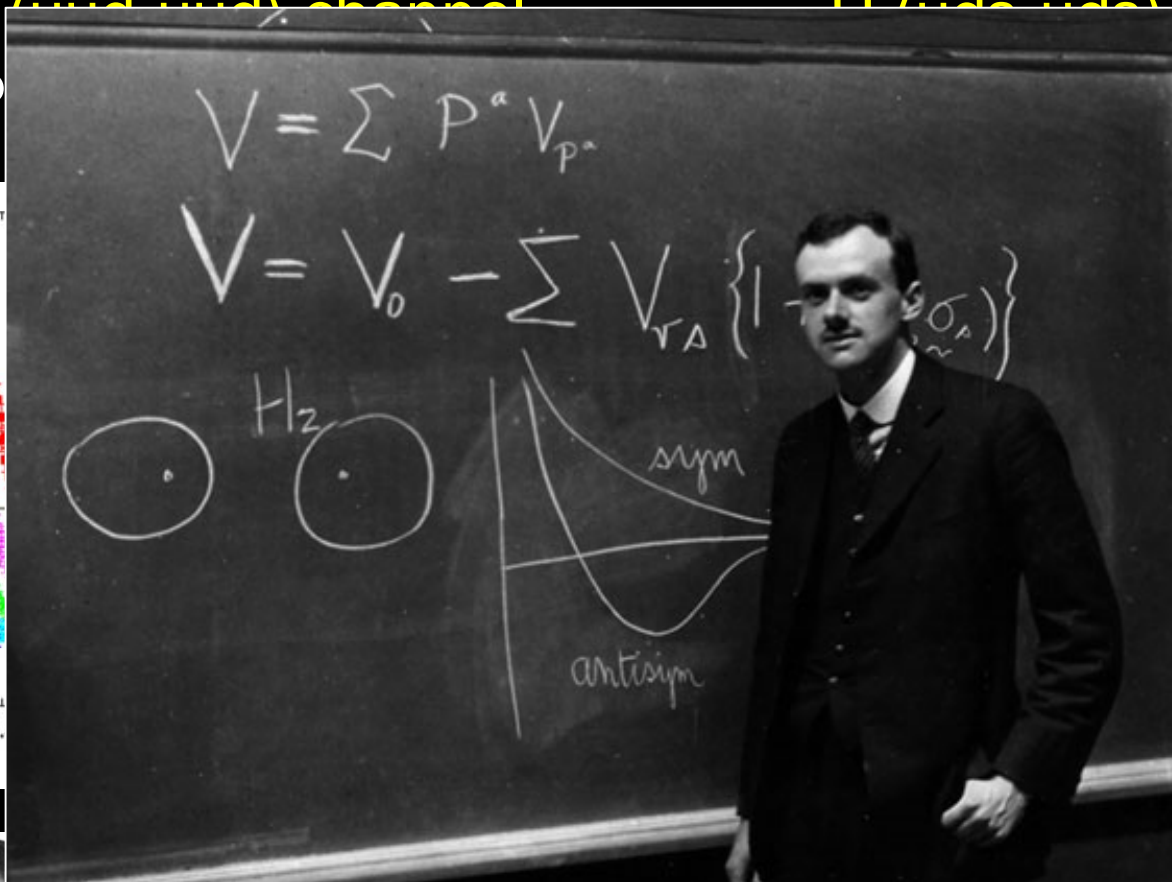
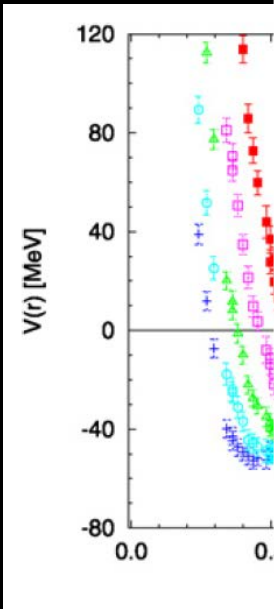
HAL QCD Coll.  
 Phys. Rev. Lett. 106 (2011) 162002  
 Nucl. Phys. A881 (2012) 28

PP (quark-quark) channel

HH (quark-quark) channel

(p)

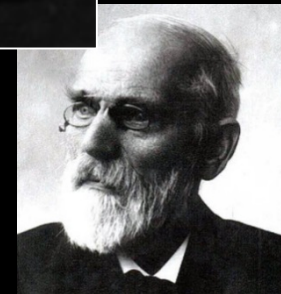
locking



$M_{PS} = 1171$  [MeV] ●  
 $M_{PS} = 1015$  [MeV] ○  
 $M_{PS} = 837$  [MeV] ▲  
 $M_{PS} = 672$  [MeV] □  
 $M_{PS} = 469$  [MeV] ■



Pauli and van der Waarls  
 at work !

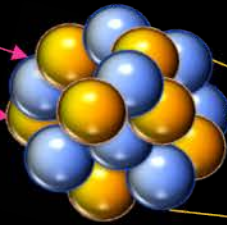


# Nuclear Matter, Neutron Matter & Finite Nuclei from LQCD + BHF

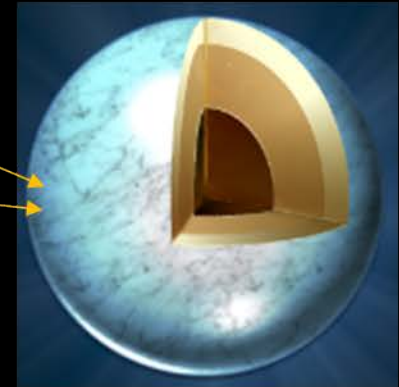
nucleon  $\sim 1$  [fm]



nucleus  $\sim 10$  [fm]



Neutron star  $\sim 10$  [km]

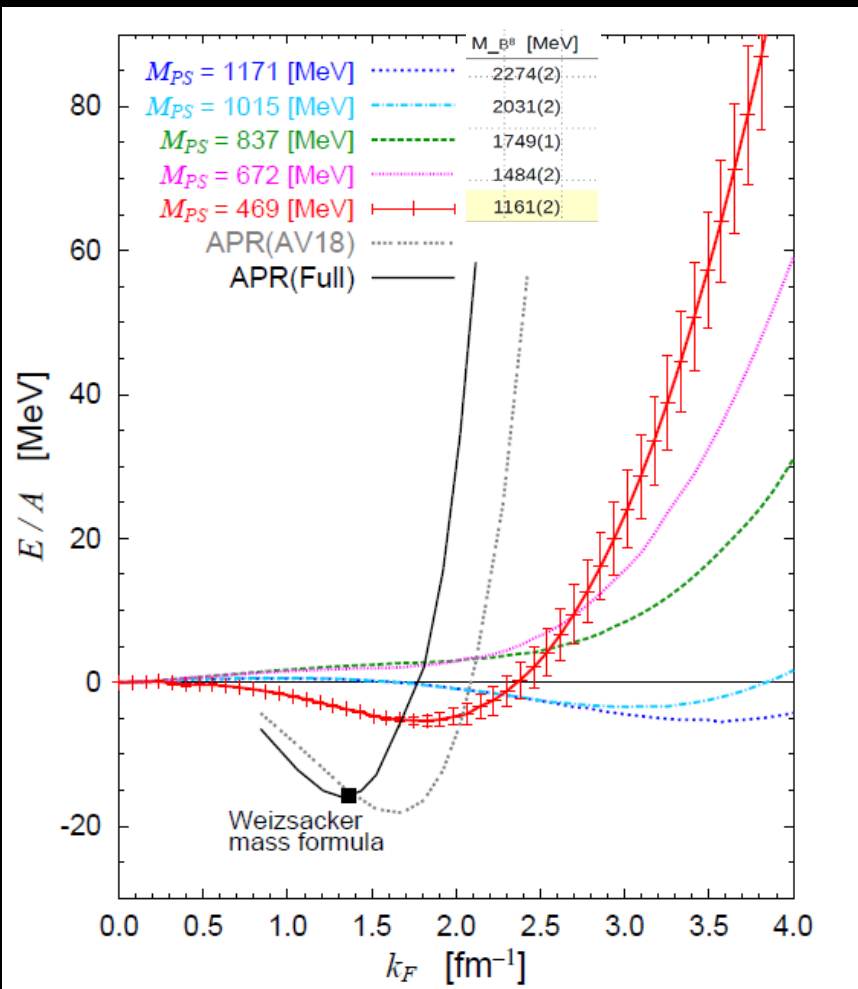


# Nuclear EOS from Lattice NN force + BHF calculation

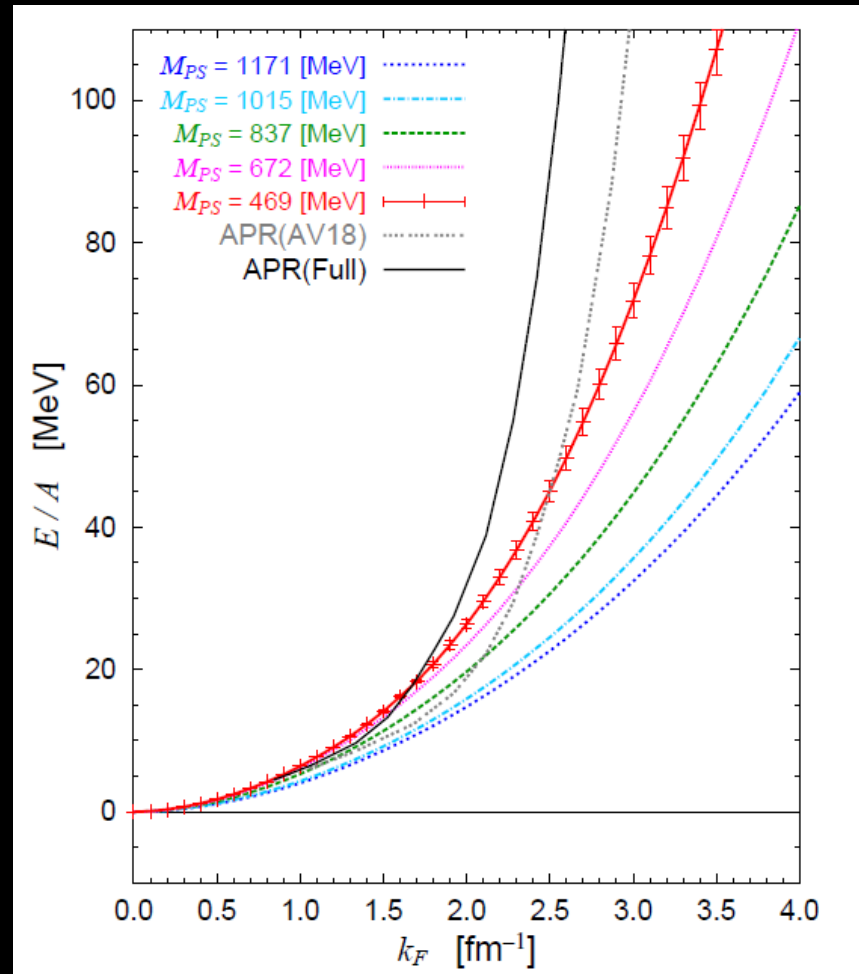
(NN force:  $^1S_0$ ,  $^3S_1$ ,  $^3D_1$  channels only)

HAL QCD Coll., Phys. Rev. Lett. 111 (2013) 112503

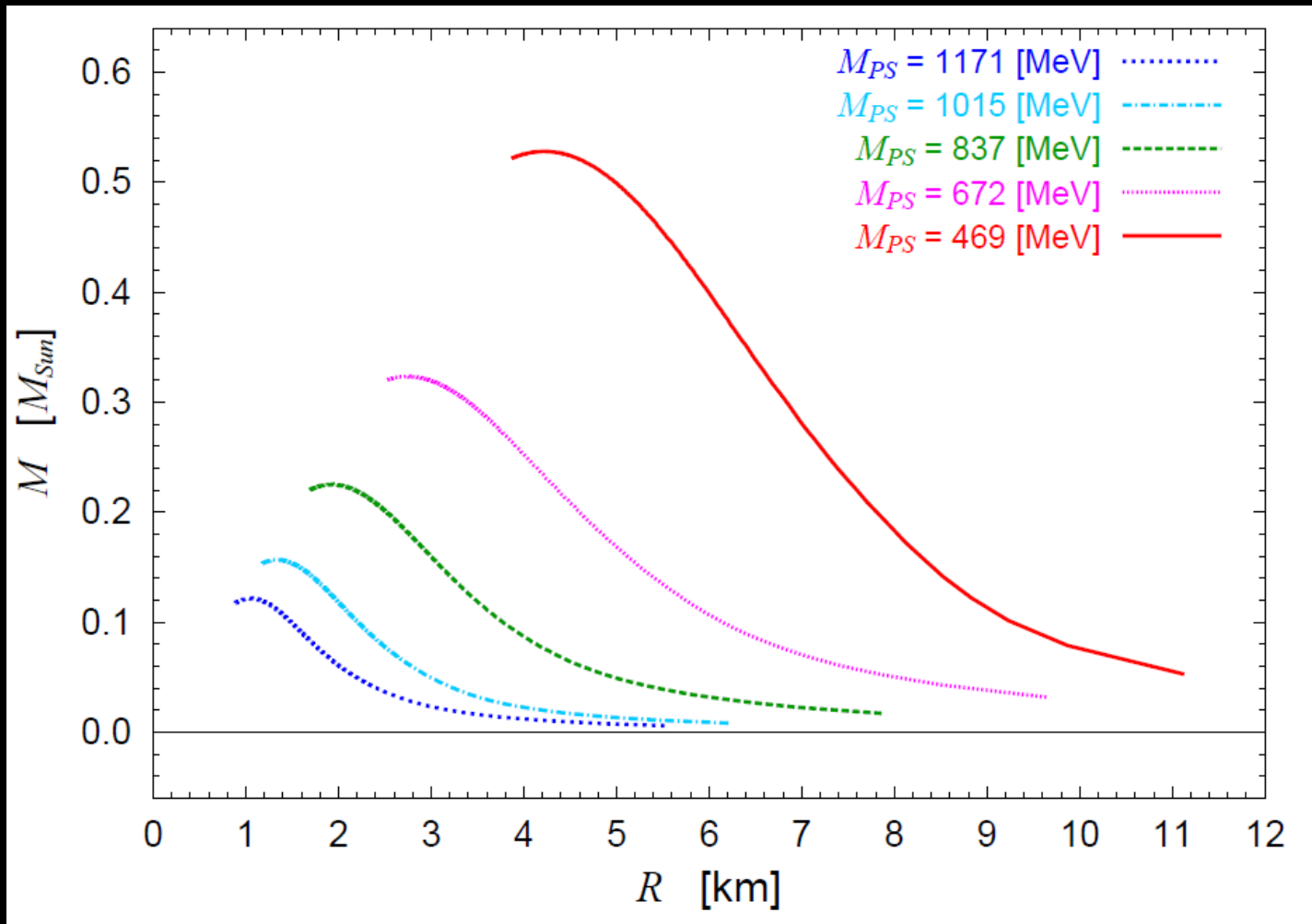
## Nuclear Matter



## Neutron Matter

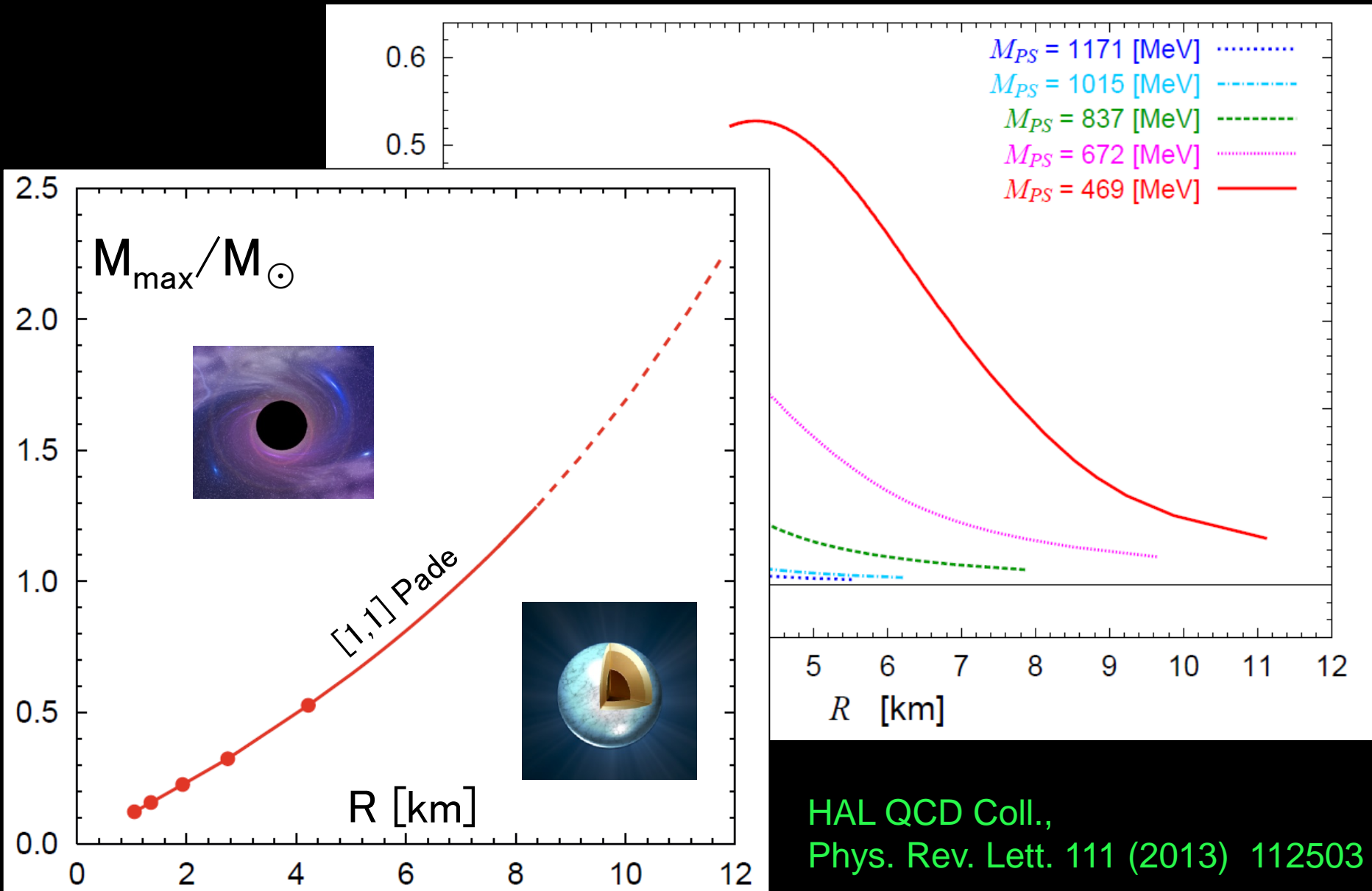


# Neutron Star from “Lattice EOS”





# Neutron Star from "Lattice EOS"



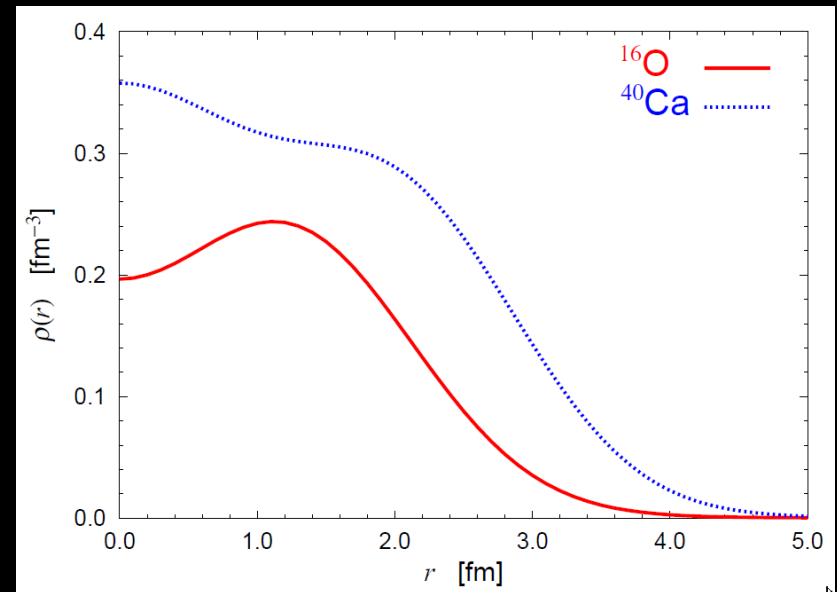
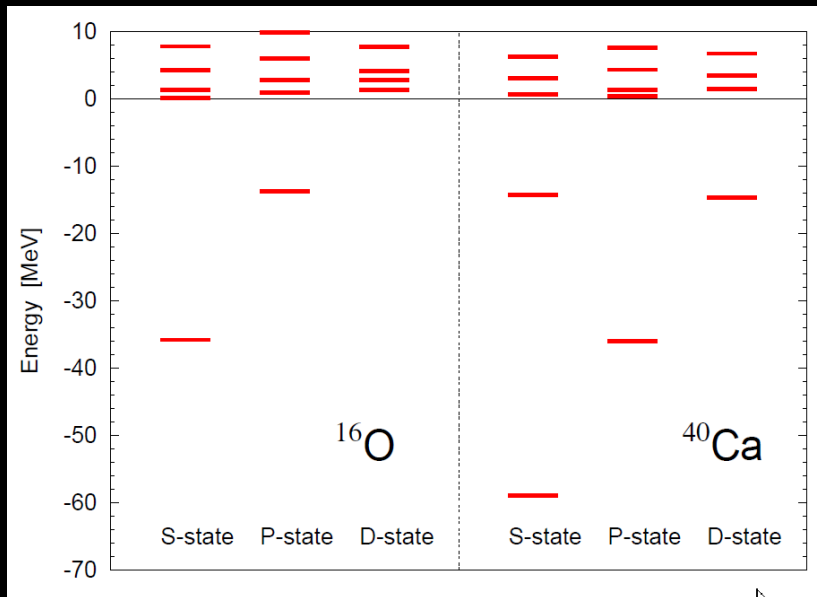
HAL QCD Coll.,  
Phys. Rev. Lett. 111 (2013) 112503

# Finite Nuclei from Lattice NN force + BHF calculation

(NN force:  $^1S_0$ ,  $^3S_1$ ,  $^3D_1$  channels only)

Inoue et al. [HAL QCD Coll.], arXive 1408.4892

Bound nuclei start to appear from  $m_\pi = 470$  MeV

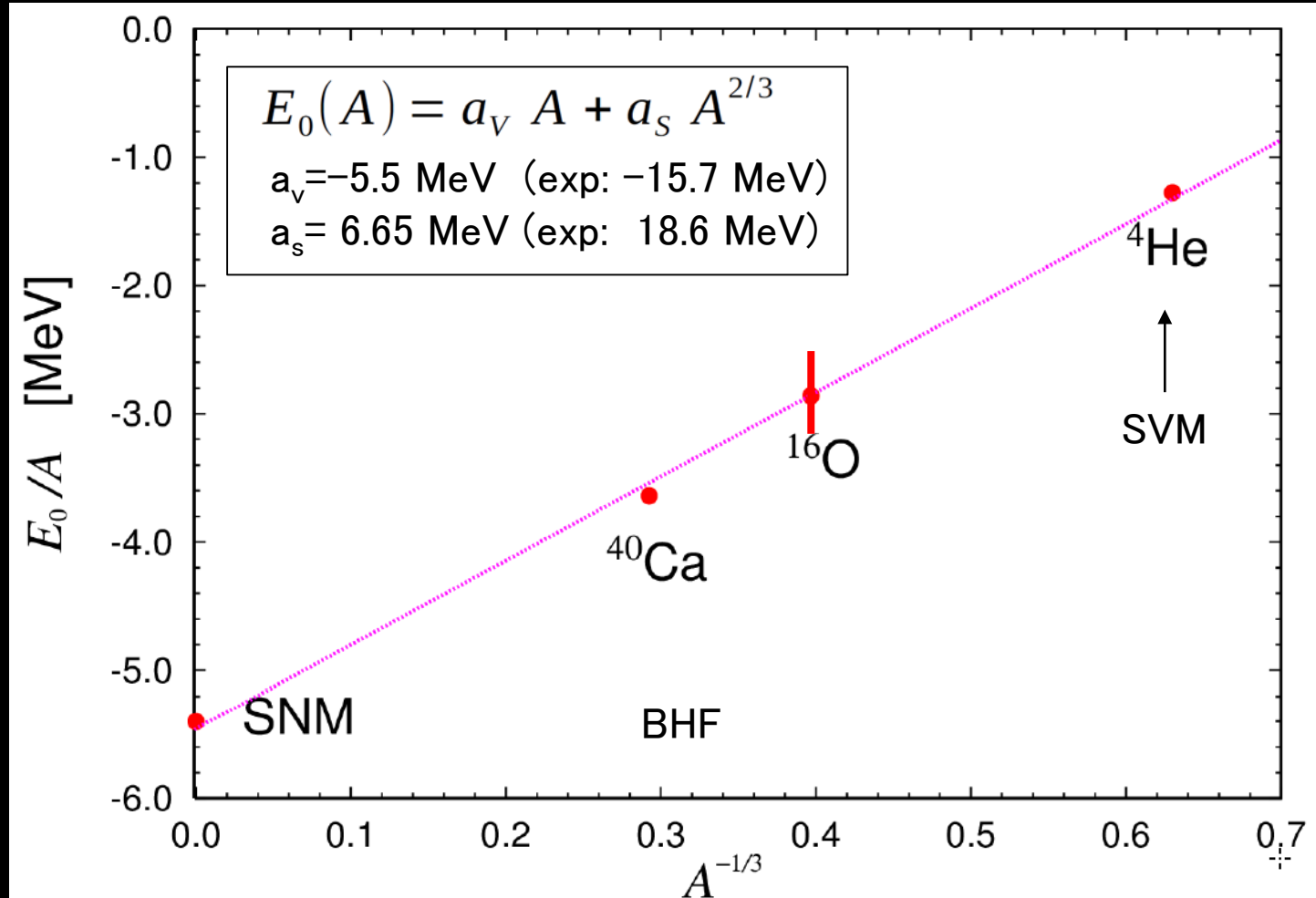


	Single particle level				Total energy		Radius
	1S	1P	2S	1D	$E_0$	$E_0/A$	$\sqrt{\langle r^2 \rangle}$
$^{16}\text{O}$	-35.8	-13.8		-14.3	-34.7	-2.17	2.35
$^{40}\text{Ca}$	-59.0	-36.0	-14.7	-14.3	-112.7	-2.82	2.78

# Nuclear Binding Energy from Lattice NN Force

Inoue et al. [HAL QCD Coll.], arXiv 1408.4892

## Bethe-Weizacker behavior at $m_\pi = 470$ MeV



# Summary

## 1. From LQCD to Compact Stars

- The best but hardest strategy : solve the sign problem
- Second and doable strategy (HAL QCD) : derive the BB and BBB forces  
→ EoS (with hyperons and 3-body forces) based on LQCD

## 2. BB forces from LQCD

- $T=V+GV$  (HAL QCD method) is lattice friendly:  
     $V$  is L-insensitive, all the data for  $t > 1$  fm are signals, etc.
- LQCD at unphysical points :  $m_\pi=470-1170$ MeV,  $L \sim 4$ fm  
    → quark-mass dependence of the BB forces  
    origin of the short-range repulsion (Pauli at work)
- Physical-point simulations are on-going by K-computer.

## 3. Nuclear/Neutron matter and Finite Nuclei from LQCD

- Neutron star becomes heavier as quark mass decreases
- Finite nuclei start to be bound at (and possible below)  $m_\pi=470$ MeV.

Present un-physical point simulation  
for single and multi-baryons

On-going physical point simulation  
for single and multi-baryons in K

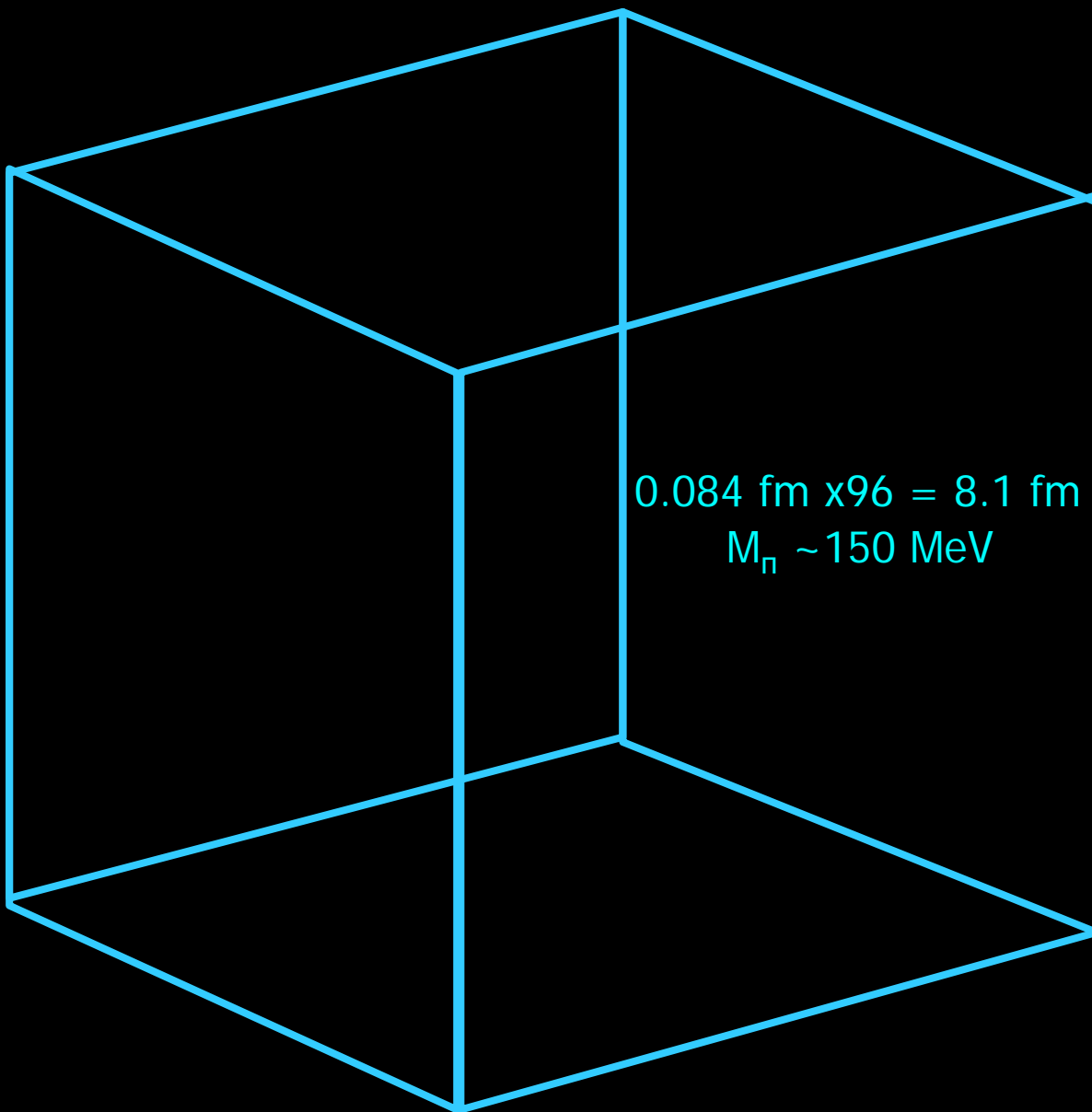
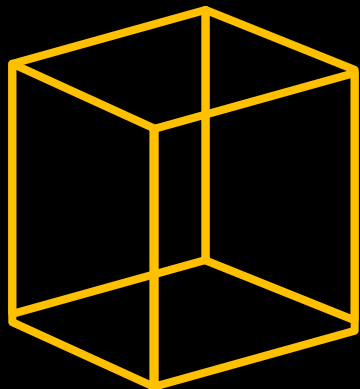
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$0.121 \text{ fm} \times 32 = 3.9 \text{ fm}$   
 $m_\pi = 350\text{-}1200 \text{ MeV}$



$0.084 \text{ fm} \times 96 = 8.1 \text{ fm}$   
 $M_\pi \sim 150 \text{ MeV}$

END