

Rhicf Simulation

Japan-Korea Phenix Joint Meeting at RIKEN

Junsang Park(Seoul National University)

Nov 27 2014

Out Line

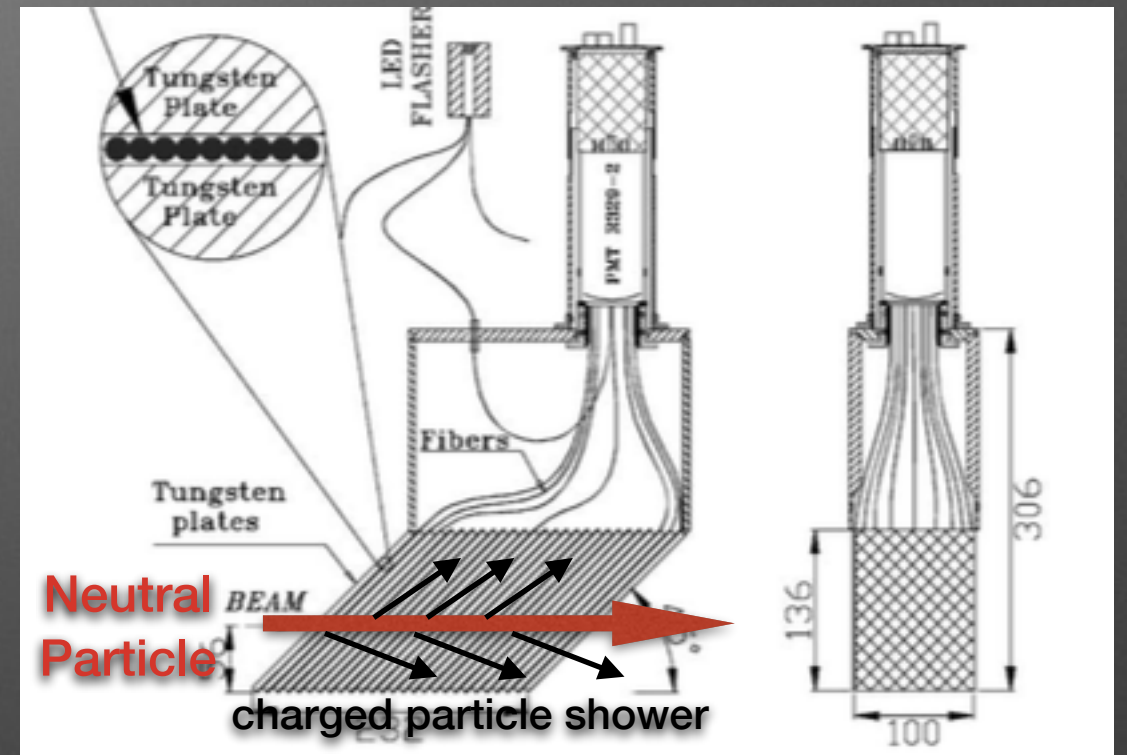
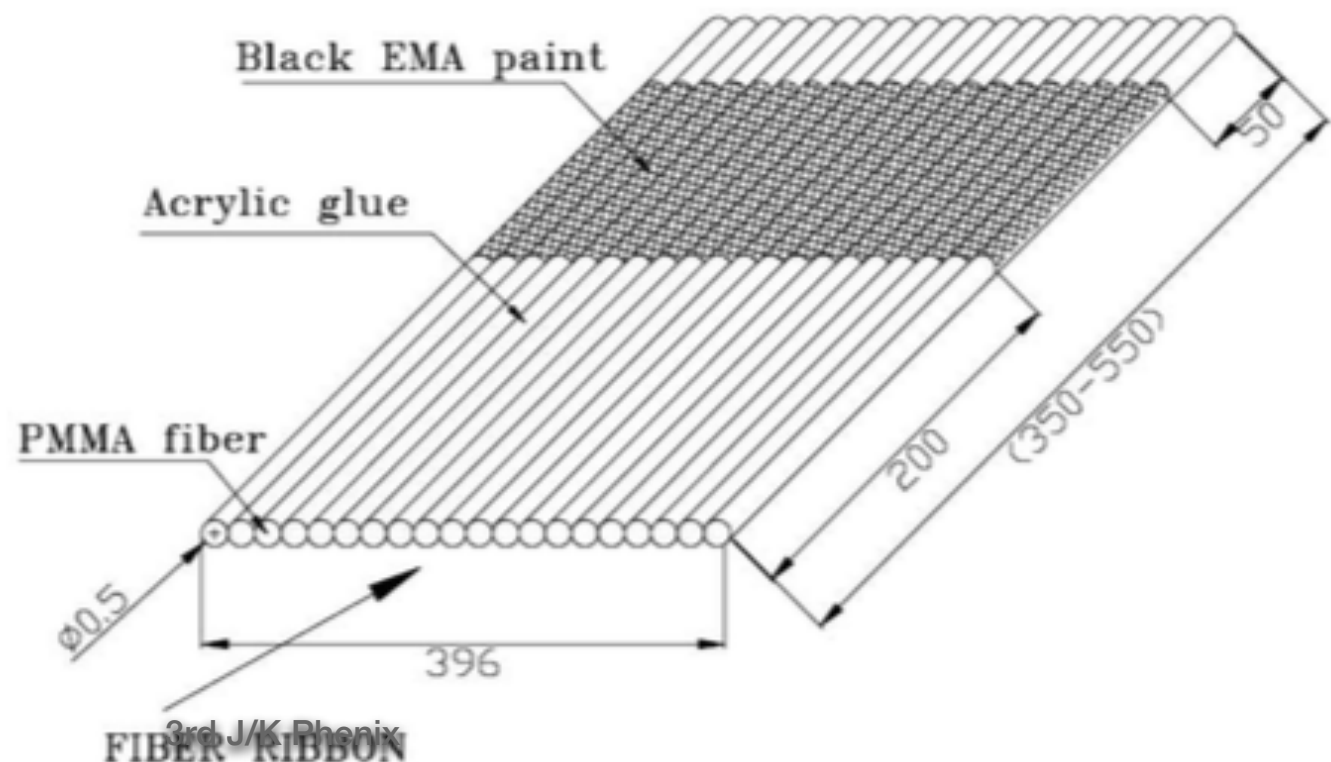
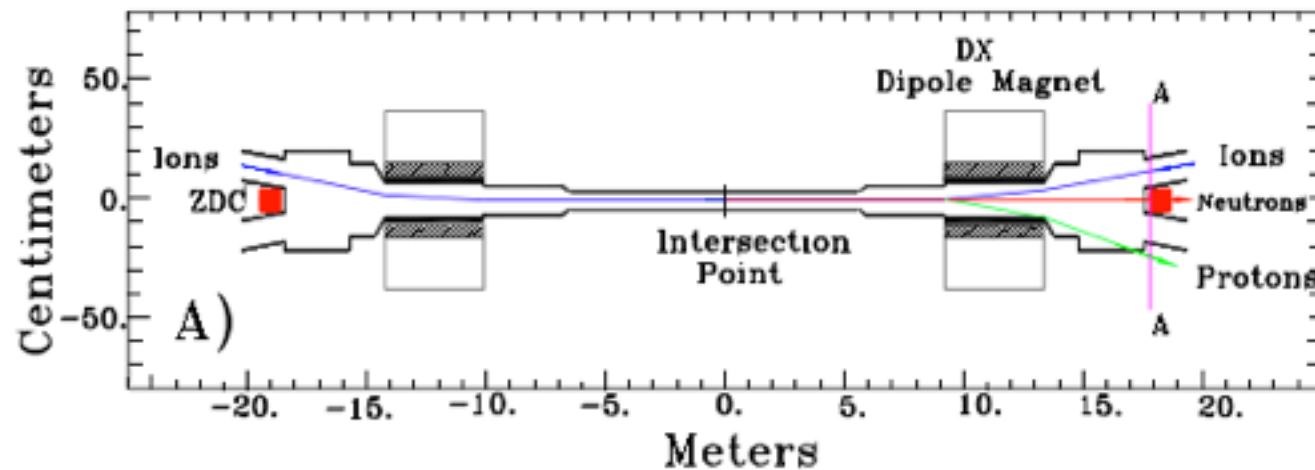
- ◎ Overview of Rhicf
 - Observing target
 - Motivation
- ◎ Simulation
- ◎ Summary and Outlook

What to observe?

- ◎ Rhic+f(forward)
→ Very forward Particle($\eta > 6$)
- ◎ Neutral Particles(π_0 , Neutron, Photon)
- ◎ Low Energy(relative to LHC)
and Low transverse momentum

How to observe?

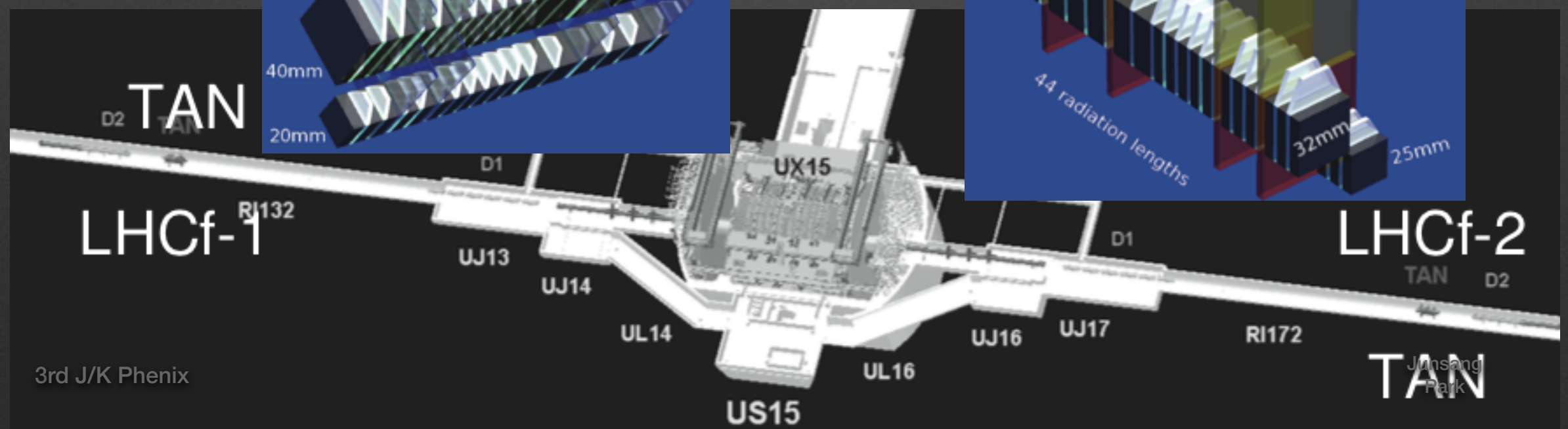
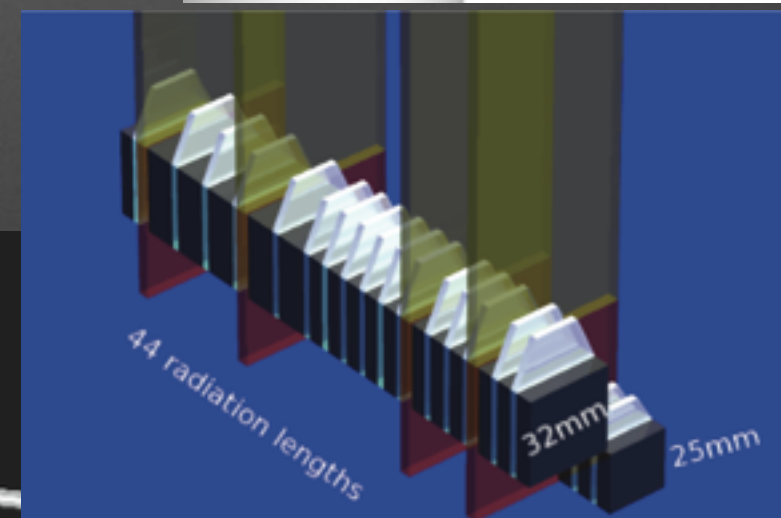
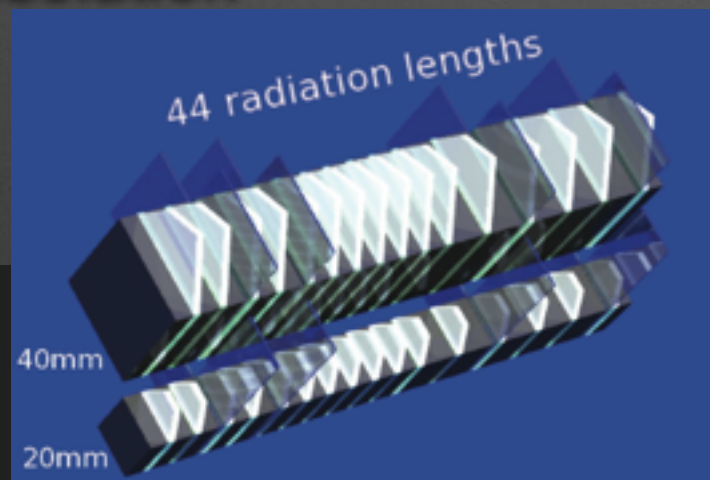
- Use ZDC and LHCf



- ZDC collects Cherenkov light created by charged particles
- Deposit energy can be calculated from PMT's signal

LHCf

- ◉ Detector for observing forward particle($\eta > 8.4$) in LHC (2009~2015)
- ◉ Having better position resolution than ZDC
- ◉ Use Arm2
- ◉ ZDC + LHCf comb make ever best resolution

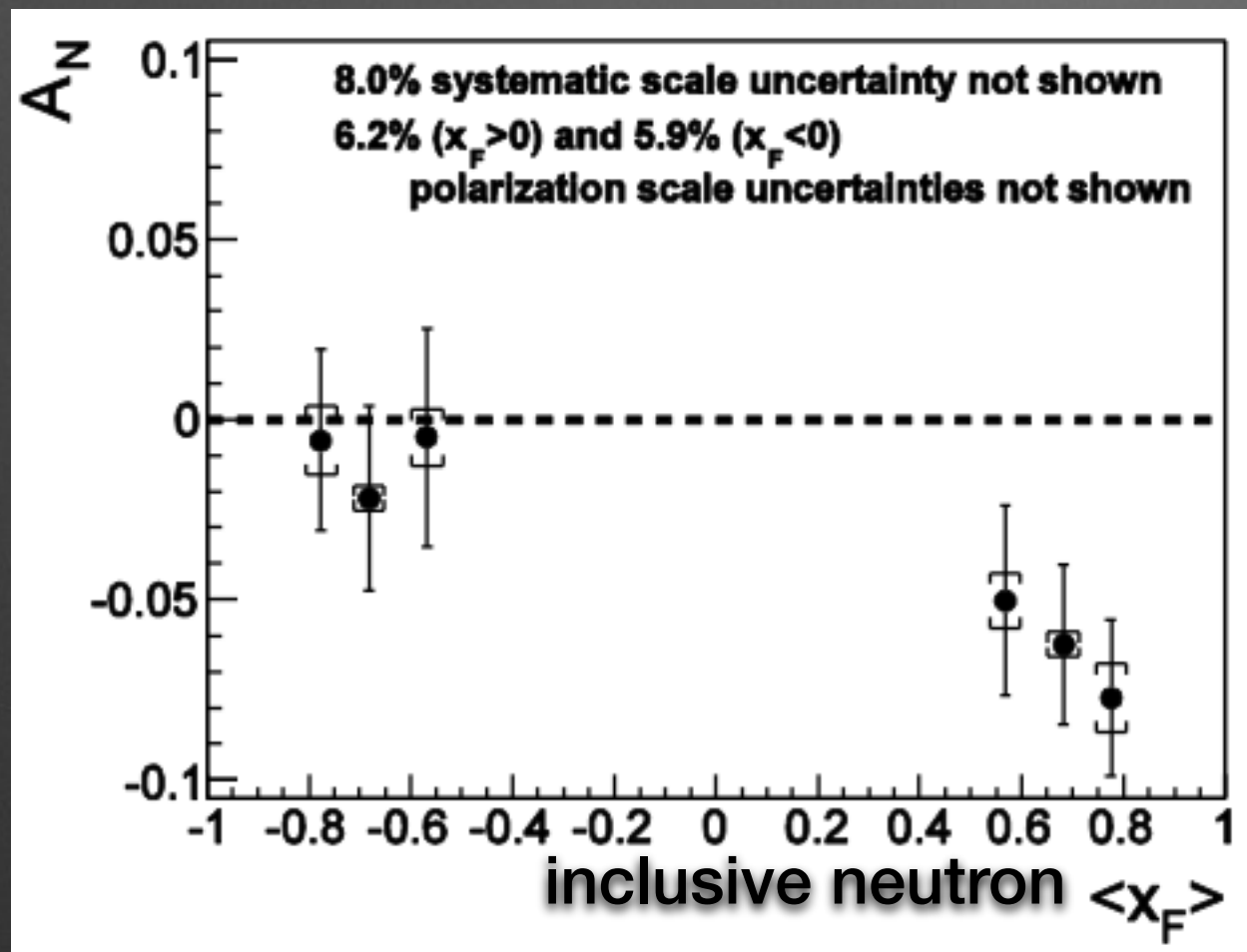


Physics Motivation

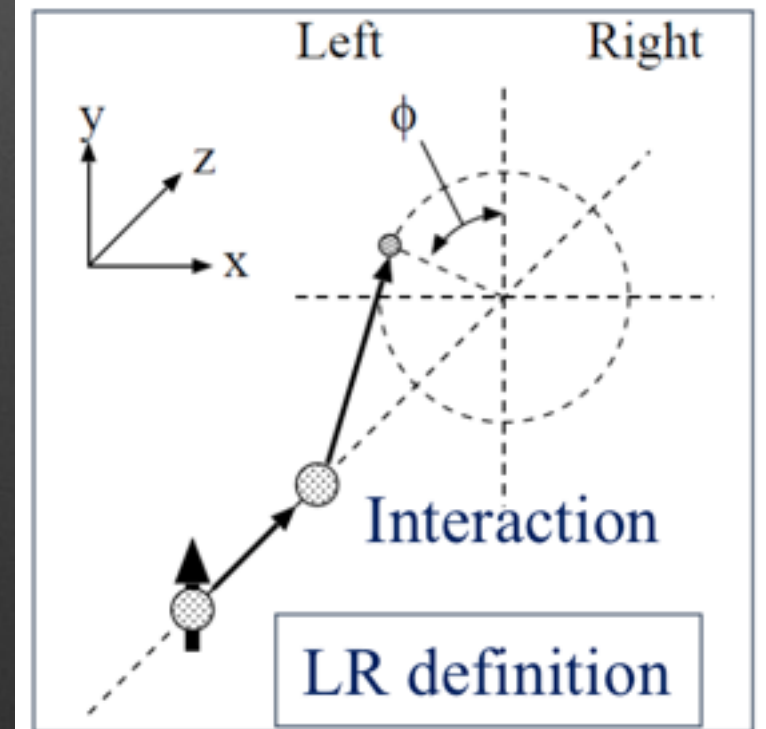
- ◉ Spin Physics
→ Forward neutron's A_N
- ◉ Verify hadronic interaction model
→ This connects to development of Cosmic Ray Physics

Forward neutron's A_N

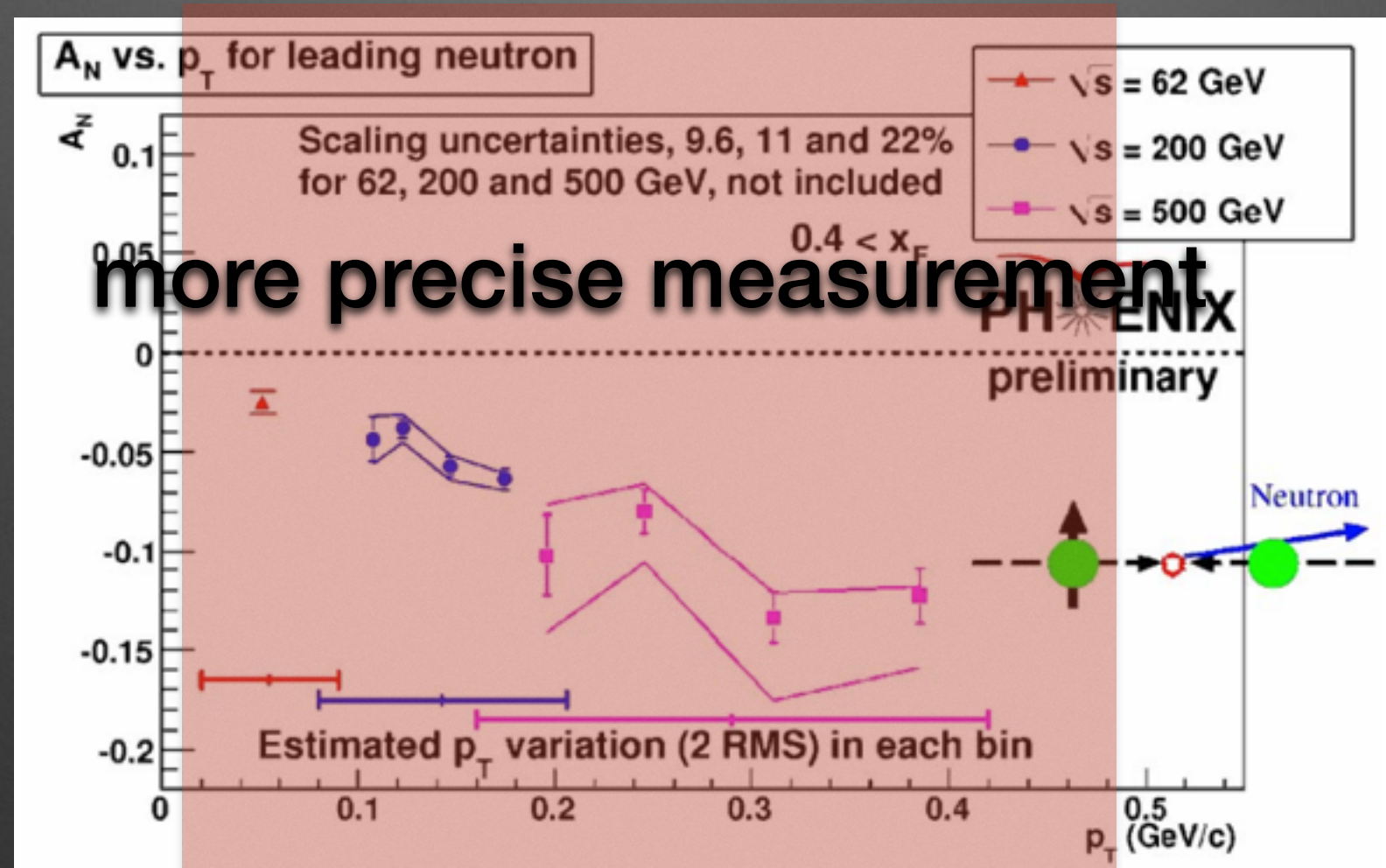
- Run5 data shows relation between X_F and A_N



$$A_N \equiv \frac{d\sigma^{\uparrow} - d\sigma^{\downarrow}}{d\sigma^{\uparrow} + d\sigma^{\downarrow}} = \frac{\sigma_L^{\uparrow} - \sigma_R^{\uparrow}}{\sigma_L^{\uparrow} + \sigma_R^{\uparrow}}$$



AN vs \sqrt{s}



Hadronic interaction models

- ◉ Theoretical basis
 - pQCD
 - Regge Theory
 - energy conservation
- ◉ Phenomenology(Models)
 - hadronization
 - Pomeron interaction
 - etc..
- ◉ Tuning parameters with data

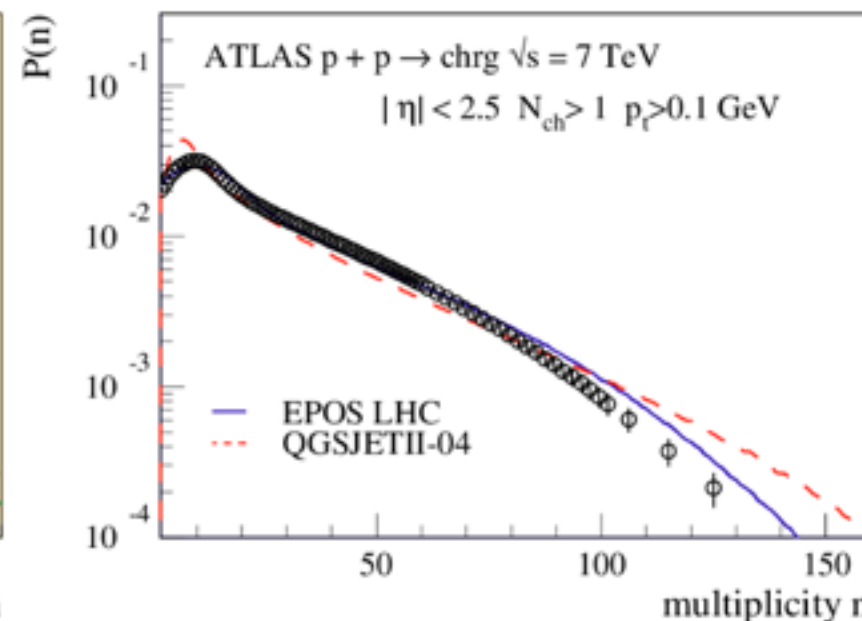
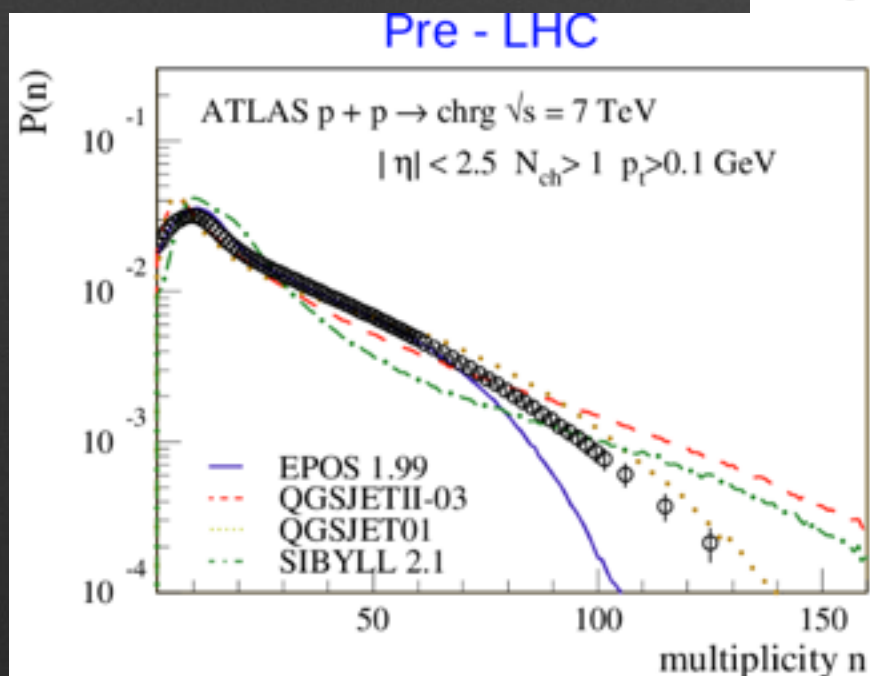
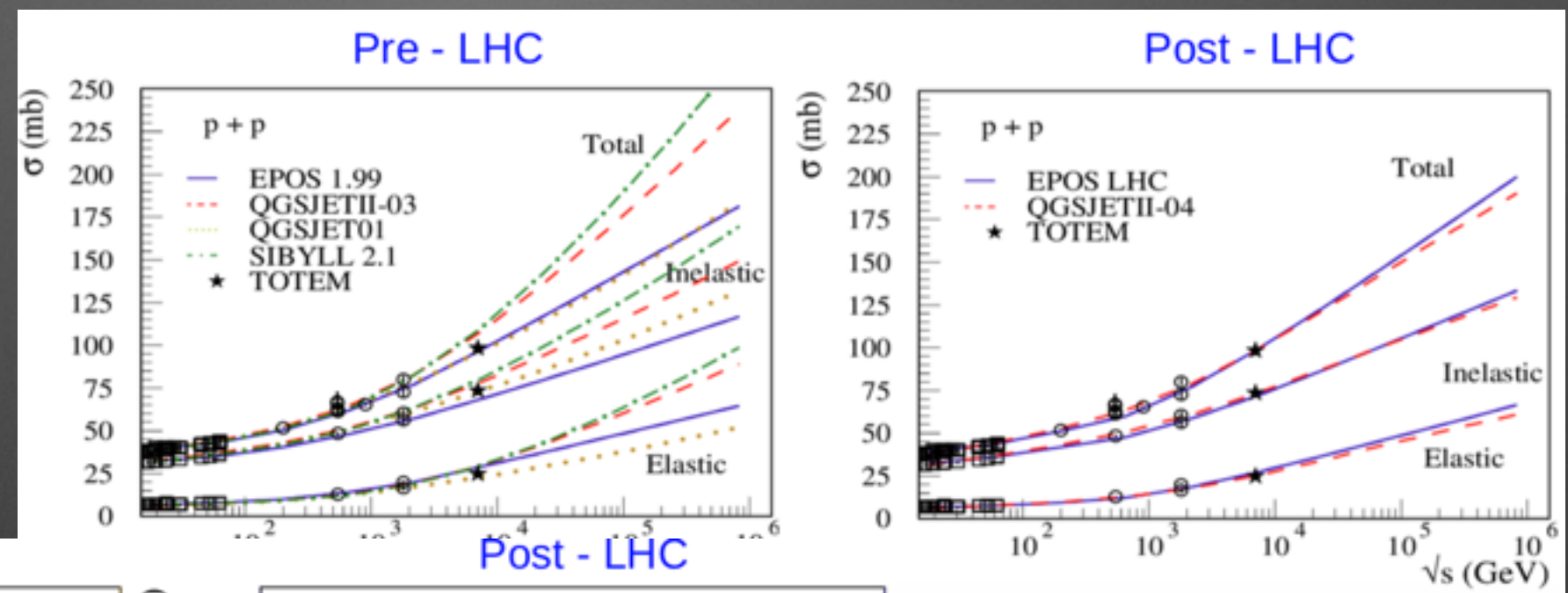
Low P_T region!

Various Models

- ◎ QGSJET series : Regge theory + String dynamics + semi-hadprocess + Pomeron interaction..
- ◎ SIBYLL series : mini-jet model + Eikonal approximation
+ Dual parton model +
- ◎ DPMJET series, EPOS series and so on....

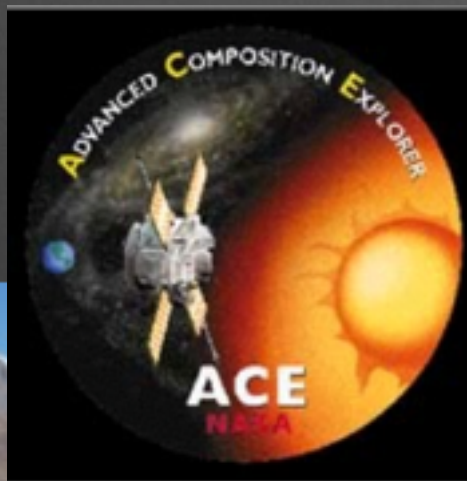
Model verification & tuning

- Verifying or tuning model by experiment data



Why are these models important to Cosmic Ray Physics?

- ◉ Observation of CR(cosmic ray)
 - direct method(satellite)→waste space, time and \$
 - indirect method(Surface Detector, Radio telescope)



Yangbajing International Cosmic Ray Observatory

3rd J/K Phenix

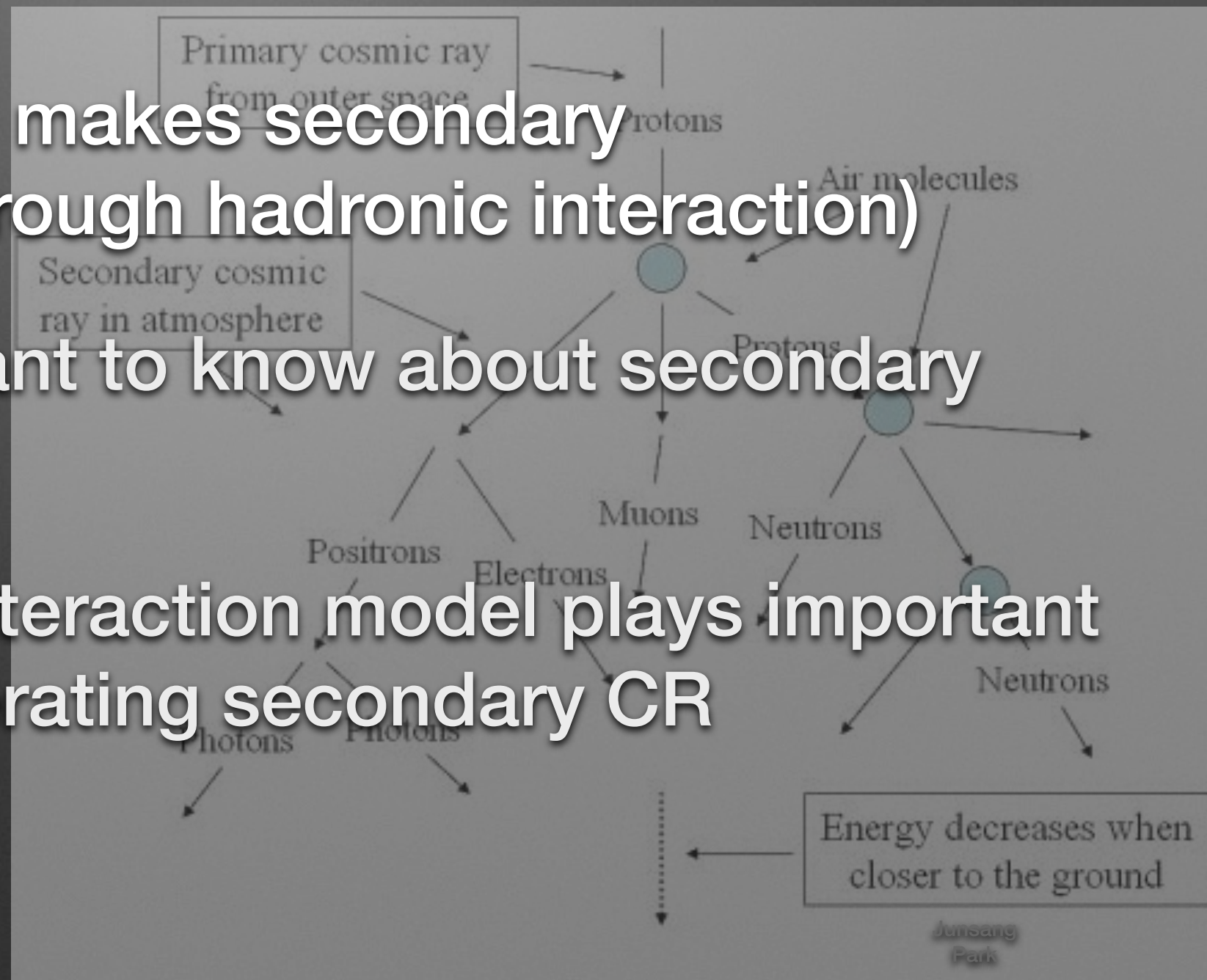


Very Energetic Radiation Telescope Array, Arizona

Junsang
Park

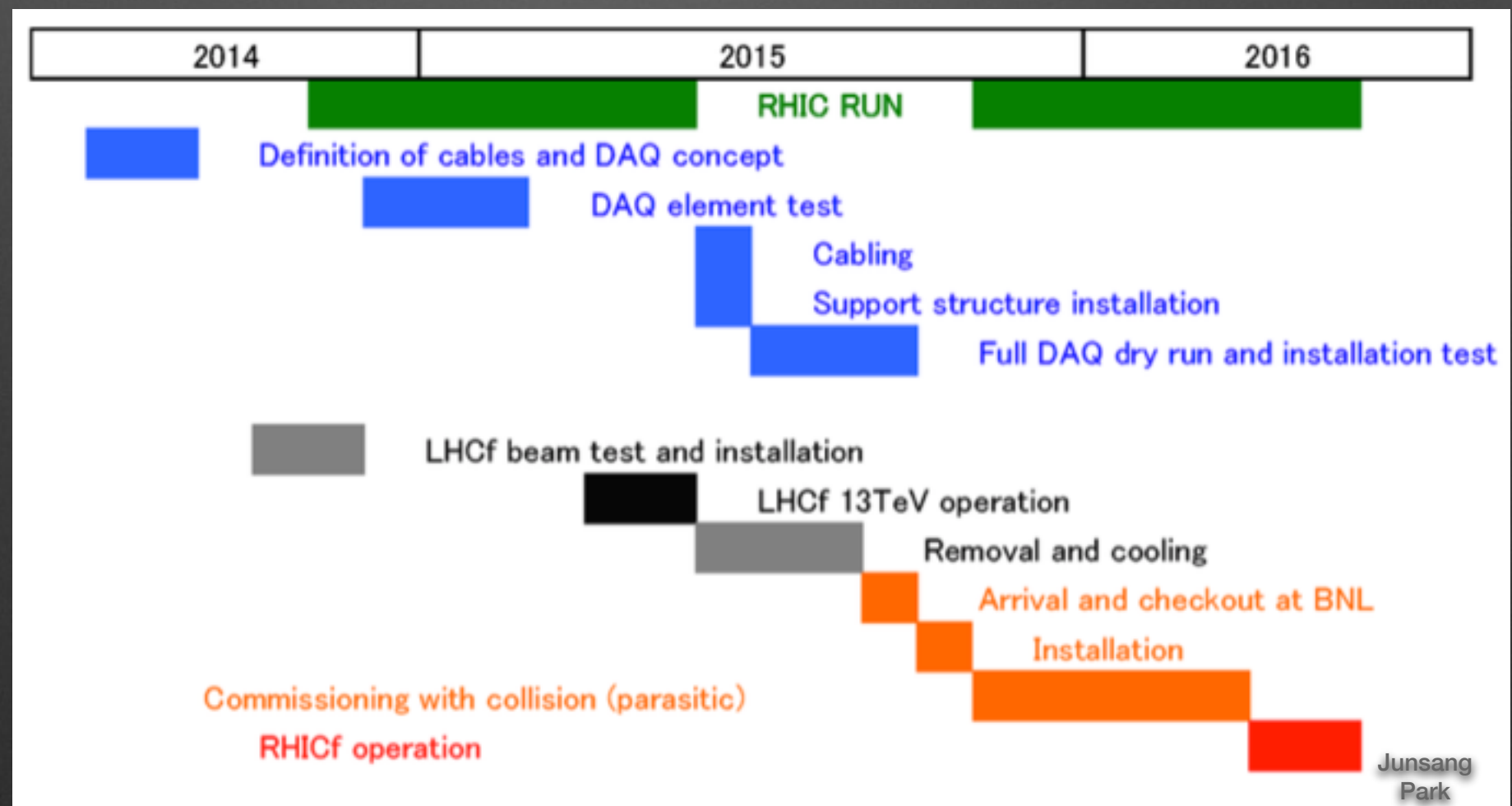
Why are these models important to Cosmic Ray Physics?

- Primary CR makes secondary particles(through hadronic interaction)
- It is important to know about secondary particles
- Hadronic interaction model plays important role in generating secondary CR



Schedules

- ◉ 2015 - Run in LHC
- ◉ Installation starts from end of 2015
- ◉ Mainrun in 2016



Rhicf summary

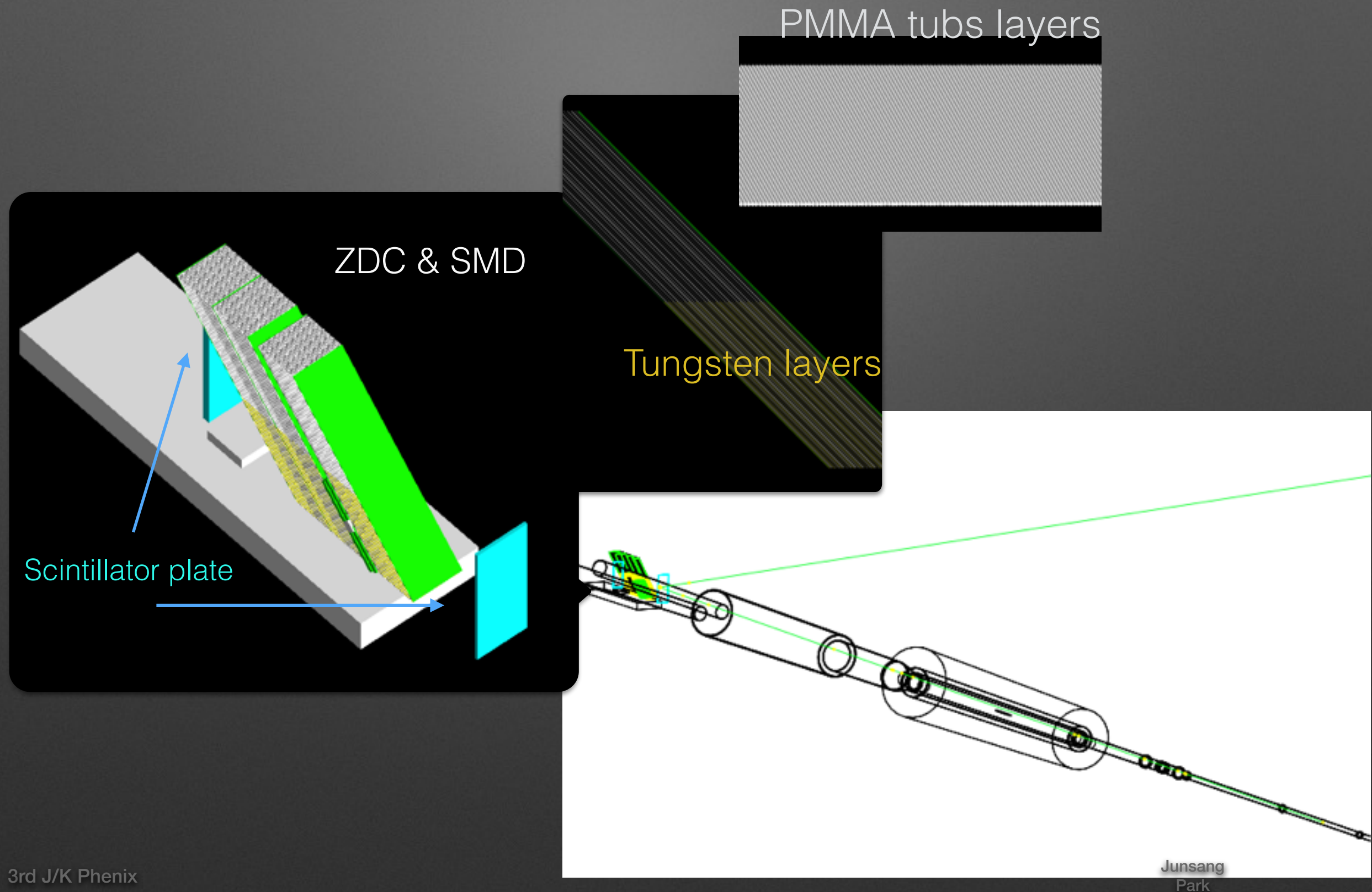
- ◉ Observe neutral particles with Low energy, small P_T and $\eta > 6$
- ◉ Data from Rhicf contributes to spin physics, cosmic ray physics and particle physics.
- ◉ Operation in 2016

Simulation

- ◉ Generally before experiment starts, Need to estimate results of experiment for determination of parameter
(e.g:structure of detector, matter and so on)
- ◉ Use Geant4 simulator for consistency between Phenix detectors and LHCf
(present code for Phenix is based on geant3)
- ◉ geometry code is completed

Rhicf

Simulation



Summary & outlook

- ◉ Localpol's geometry code is completed
- ◉ ZDC's energy resolution calculation is on going
- ◉ Make more proper PhysicsList to Rhicf
- ◉ Need to make Multithread activate
- ◉ Calculate efficiency, position resolution and energy resolution after implementation of LHCf

Back Up

Definition of Variables

$$x = \frac{p_L}{\sqrt{s}/2} = \frac{2p_L}{\sqrt{s}}$$

Feynman- x

$$y = \ln \frac{E+p_L}{m_T} = \frac{1}{2} \ln \frac{E+p_L}{E-p_L}$$

Rapidity

$$m_T^2 = p_T^2 + m^2$$

Transverse mass squared

$$d^3p/E = d^2p_T \frac{dp_L}{E} = d^2p_T dy$$

Phase space element

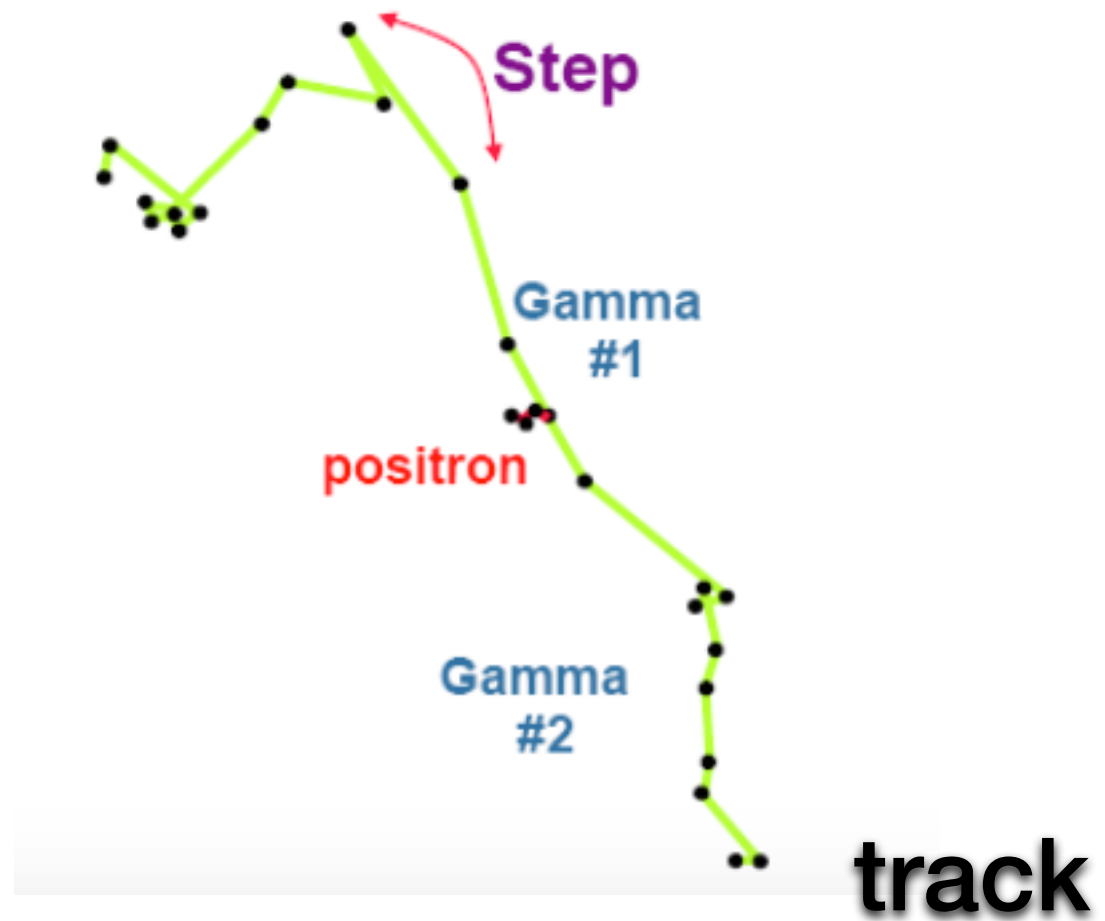
$$\eta \equiv y_{m=0} = -\ln \tan(\theta/2)$$

Pseudorapidity

$$p_L = m_T \sinh y$$

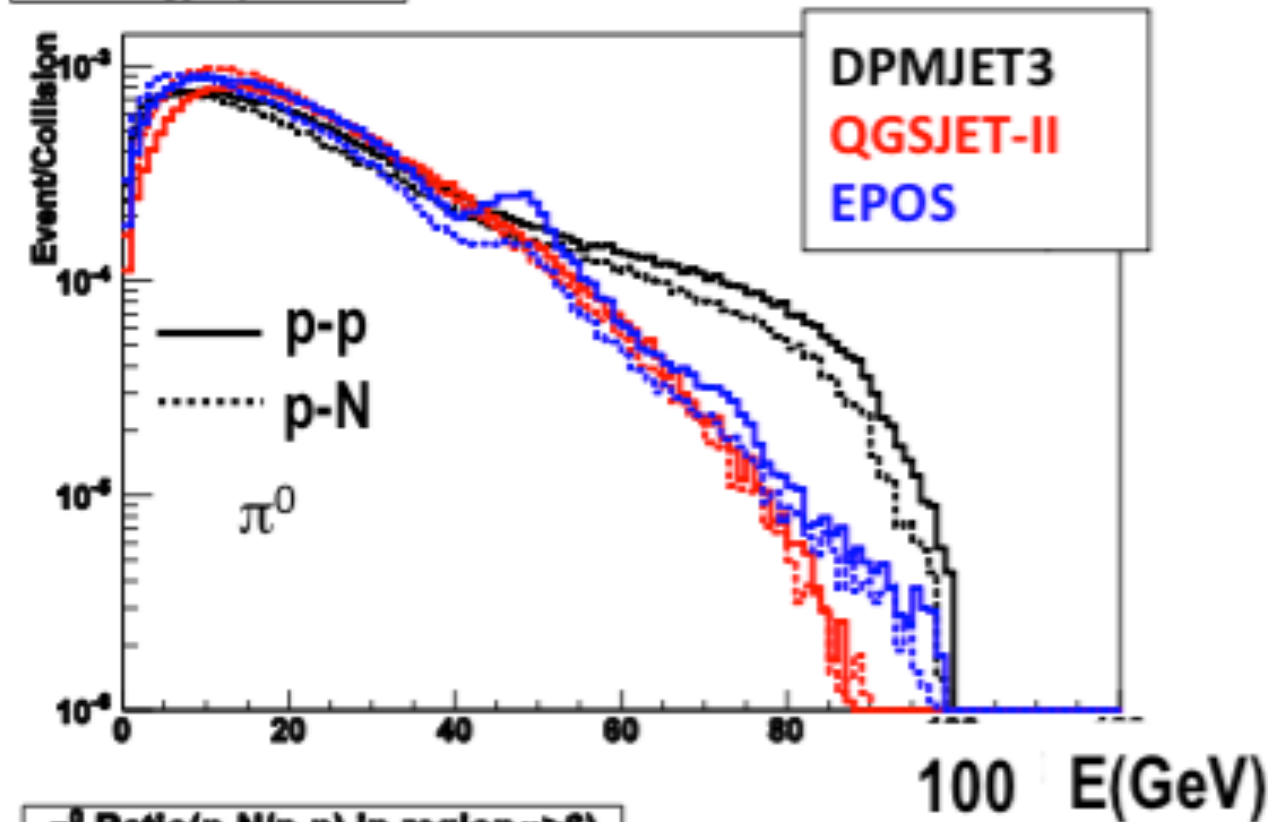
$$E = m_T \cosh y$$

Transportation in geant4

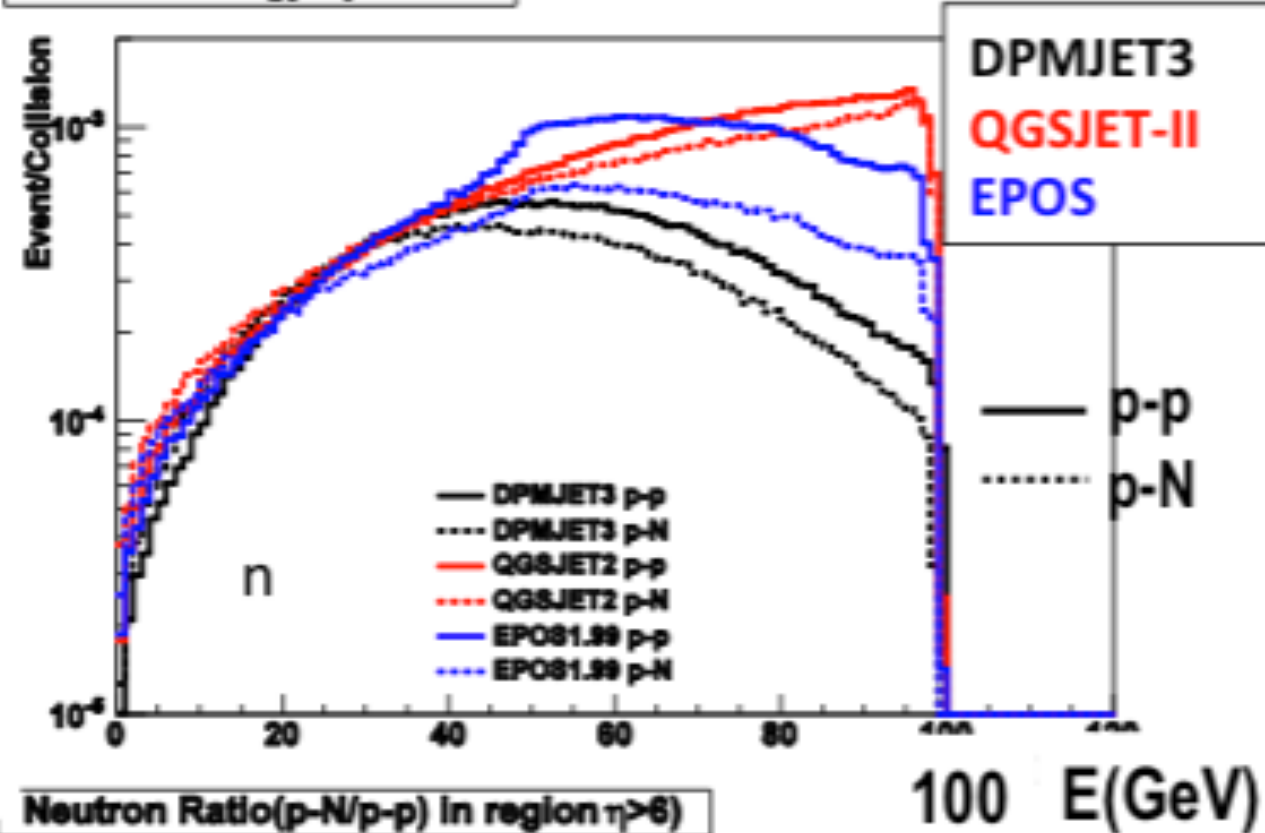


$E\pi^0$ and En (200 GeV p-p / p-N)

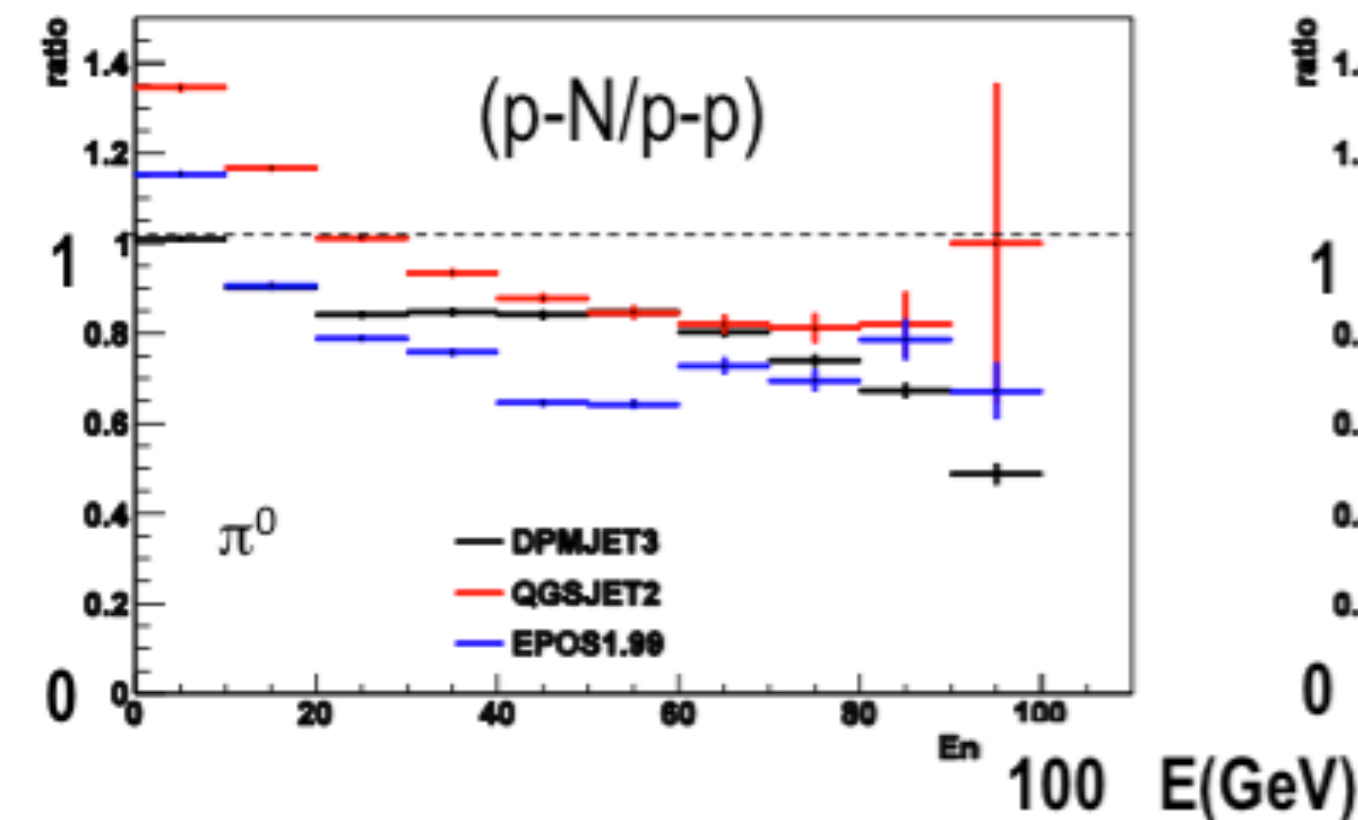
π^0 Energy Spectrum



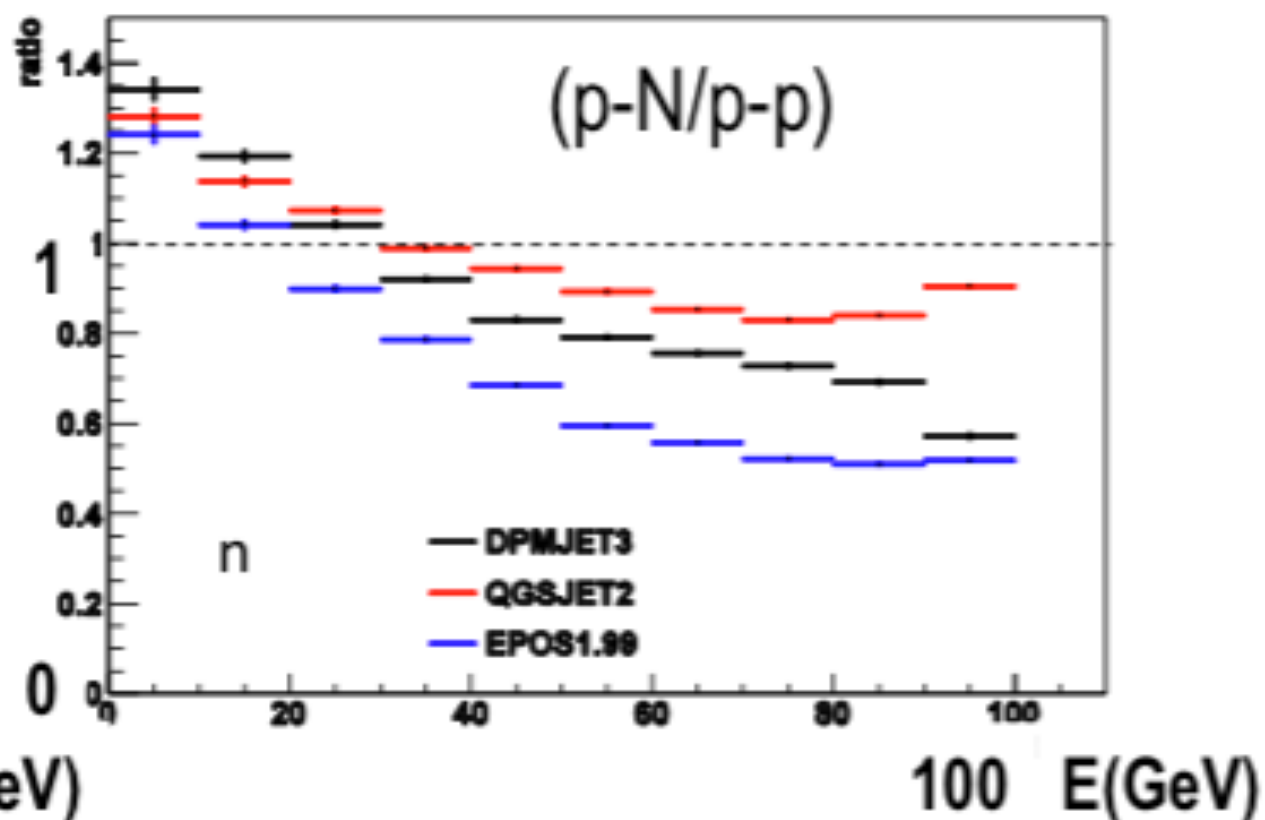
Neutron Energy Spectrum



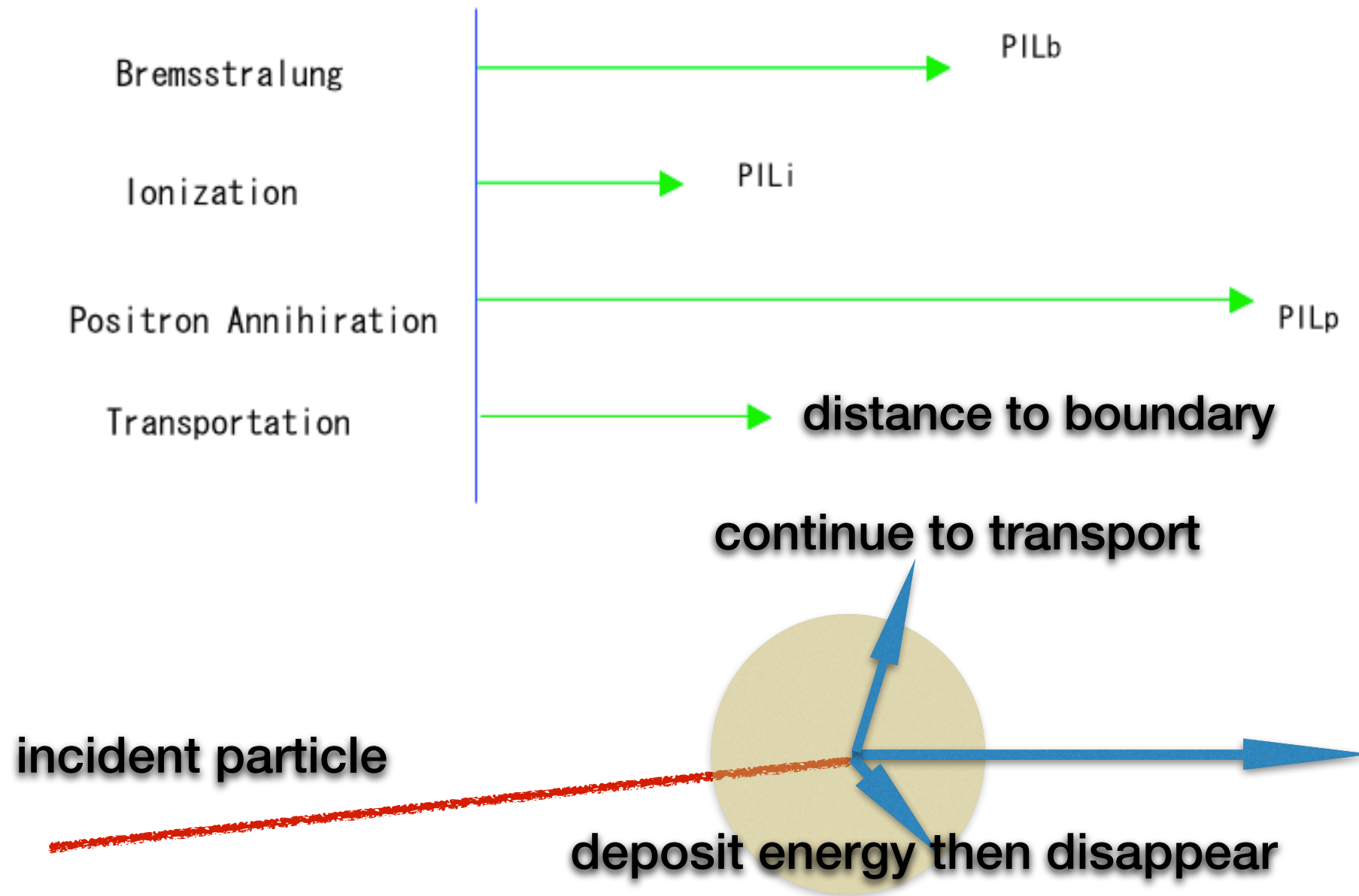
π^0 Ratio(p-N/p-p) in region $\eta > 6$



Neutron Ratio(p-N/p-p) in region $\eta > 6$



conversion NLL to distance



Hadronic interaction models

Old generation : QGSJET01 SIBYLL 2.1 DPMJET 2.55 VENUS (<1999)

All Glauber based

But differences in hard, remnants, diffraction ...

semi-hard

soft

NEXUS
3.97

Attempt to get everything described in a consistent way (energy sharing)

New generation : (QGSJET II-03) (DPMJET III) (EPOS 1.99) (2005-2012)

LHC tuned :

QGSJET II-04

EPOS LHC (2013-)

Theory ++ :

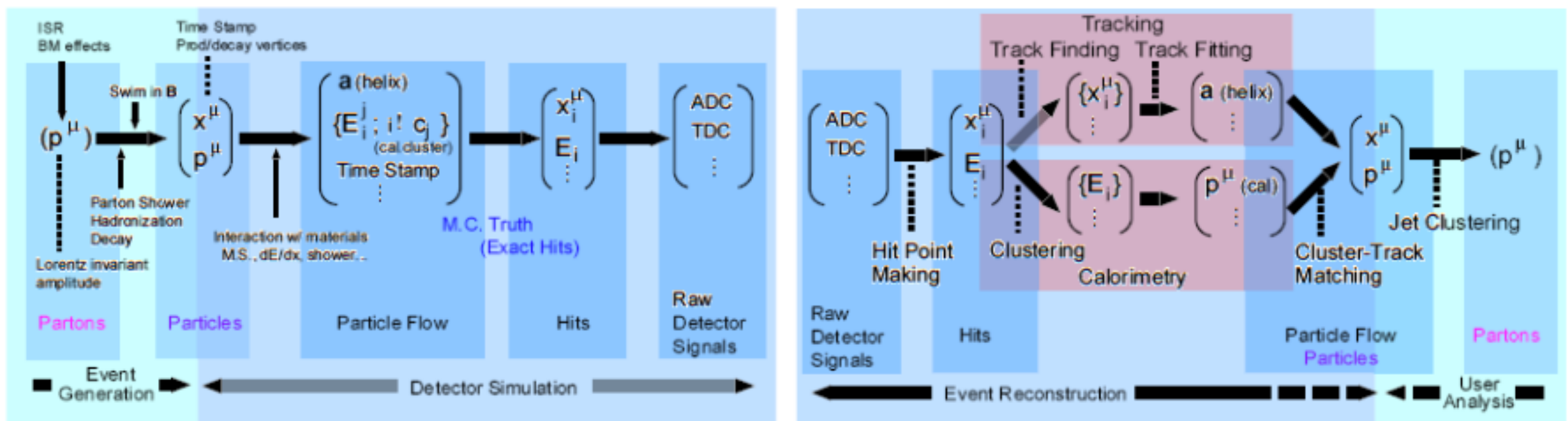
- Loop diagrams
- rho0 resonance
- optimized for CR

Phenomenology ++ :

- Nuclear effect
- High density effect (QGP)
- all type of data studied

Only model used in HEP (SPS, RHIC, LHC)

Flow of simulation



QuickSim (CLIB)

skipping...

JLCSM (Geant3)

skipping...

JUPITER&Satellites (Geant4)