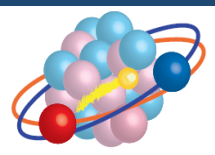


RIKEN 2koma plus seminar
Wako
July 1-2, 2015

Structure evolutions in exotic nuclei and nuclear forces - Day 2 -



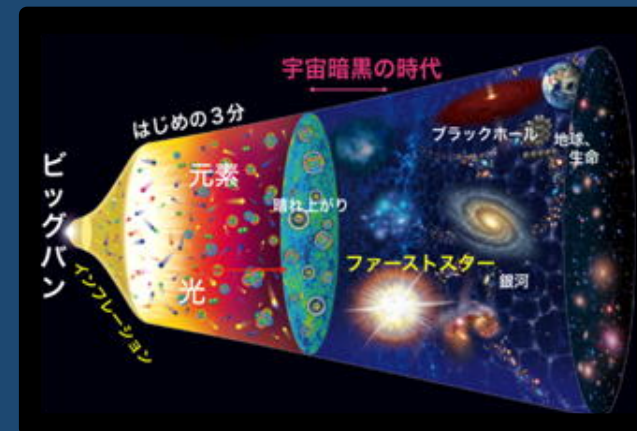
Takaharu Otsuka
University of Tokyo / MSU / KU Leuven



東京大学
THE UNIVERSITY OF TOKYO



CENTER for
NUCLEAR STUDY
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Addenda to Day 1

renormalization of tensor force
exotic silicon isotopes

Response to the renormalization of interactions

Renormalization processes

- short-range correlations
- in-medium corrections

bare interaction for free space

$$V = V_c + V_{LS} + V_T$$

effective interaction for a model space of low-momenta

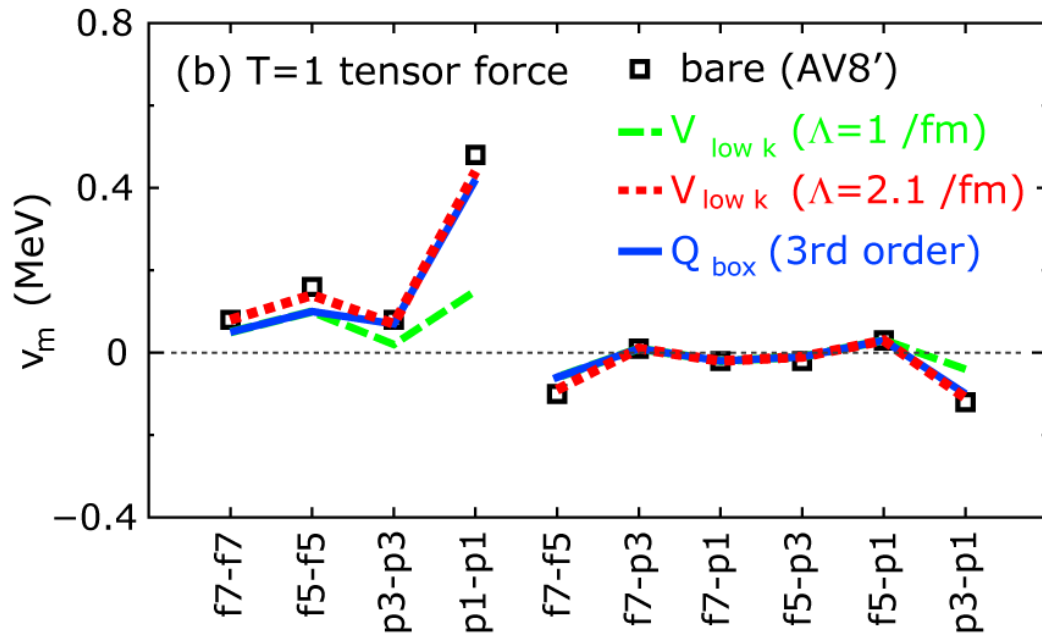
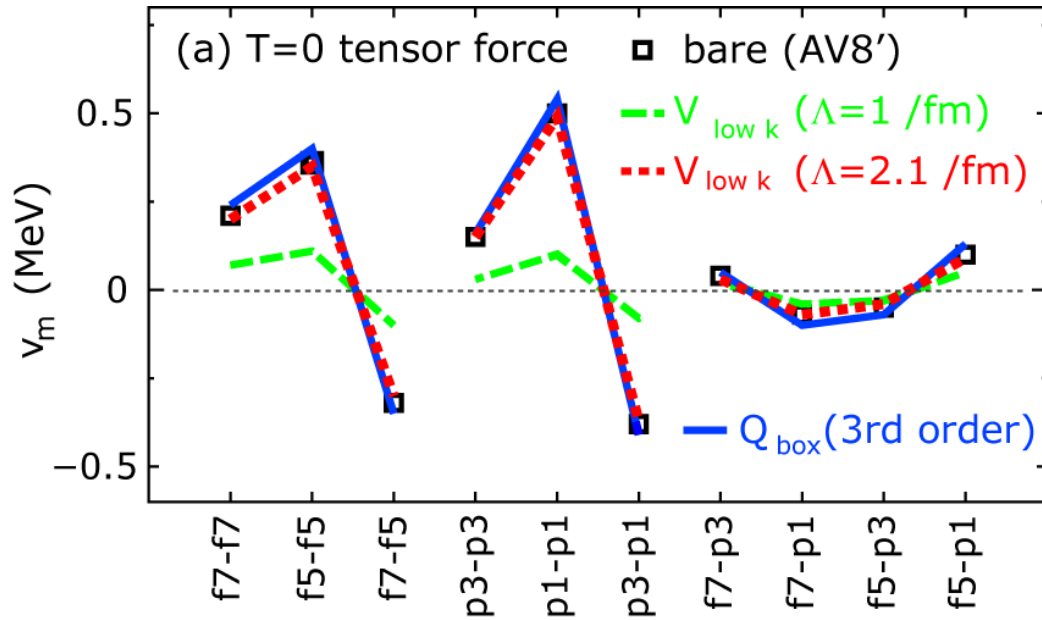
$$V' = V'_c + V'_{LS} + V'_T + V_{NNN} + \dots$$

In general, V'_x differs from V_x .

If $V_x = V'_x$, **Renormalization Persistency** holds.

- only good approx. at best, but it makes sense
- new approach to nuclear forces

Treatment of tensor force by $V_{\text{low } k}$ and Q box (3rd order)



Monopole component of **tensor** interactions in *pf* shell

□ Bare (AV8')

↓ **short-range** correlation by $V_{\text{low } k}$

--- **in-medium** correction with intermediate states (> 10 hw, 3rd order)

$V_{\text{low } k}$: Bogner, Kuo, Schwenk

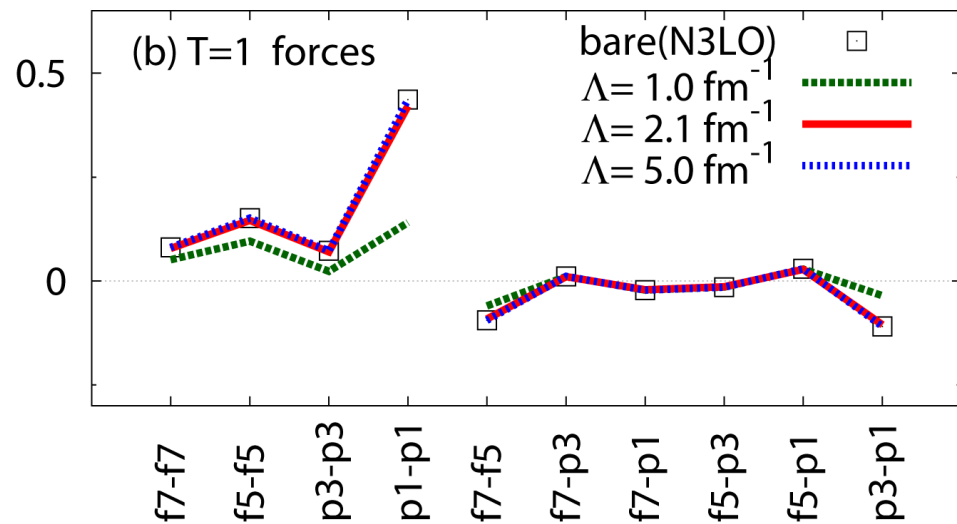
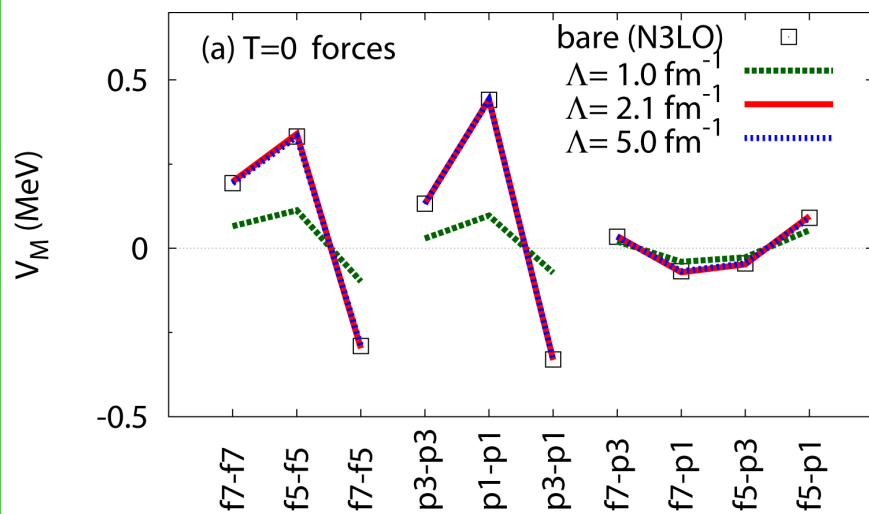
--- **only for comparison**

O, Suzuki, et al.

PRL 104, 012501 (2010)

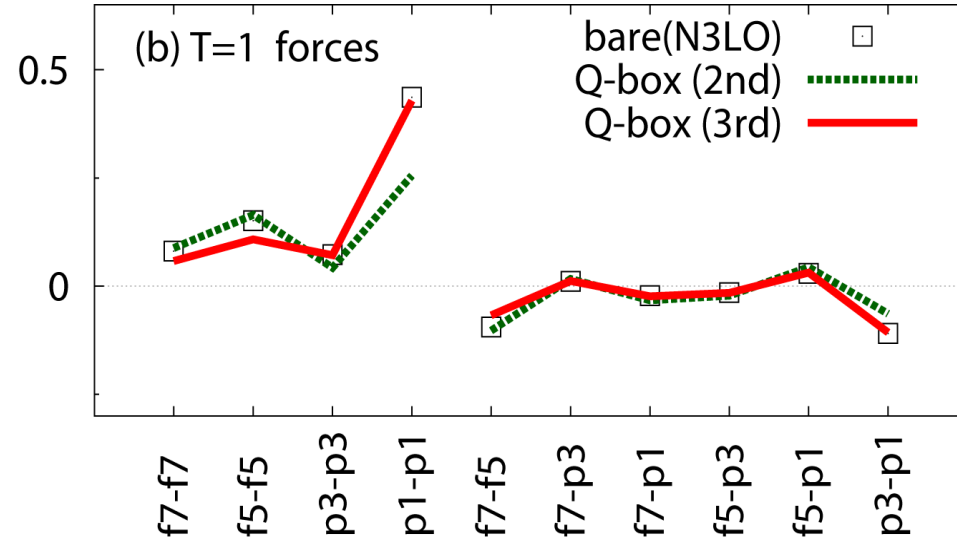
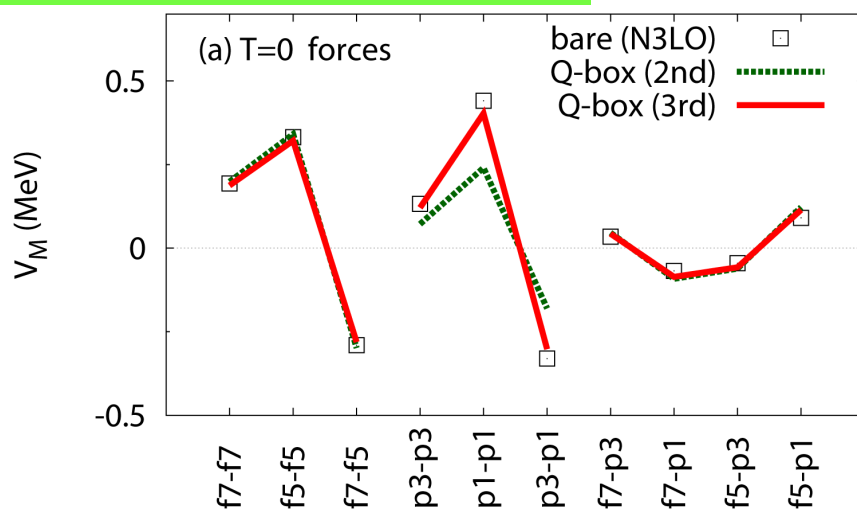
N3LO (EFT of QCD) for pf-shell

short-range correl.



in-medium correc.

$\Lambda = 2.1 \text{ fm}^{-1}$



Renormalization persistency of the tensor force in nuclei

N.Tsunoda, T.O., K.Tsukiyama, M.H.-Jensen

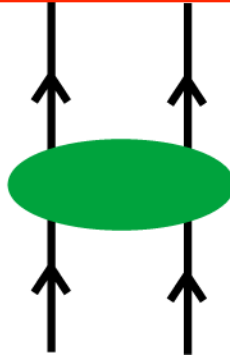
Two steps of renormalization from full space to low-momentum space relevant to many-nucleon systems

1. Treatment of short range correlations
2. Inclusion of in-medium effects

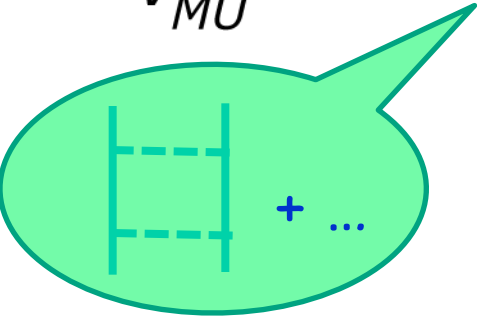
Two major components in nuclear force

(a) central force :

(strongly renormalized)



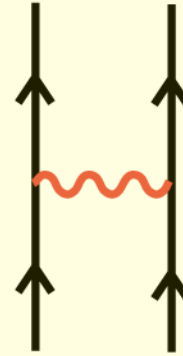
$V_{MU} =$



+

(b) tensor force :

$\pi + \rho$ meson exchange



Renormalization
Persistency

monopole component of
tensor force **in nuclear medium**

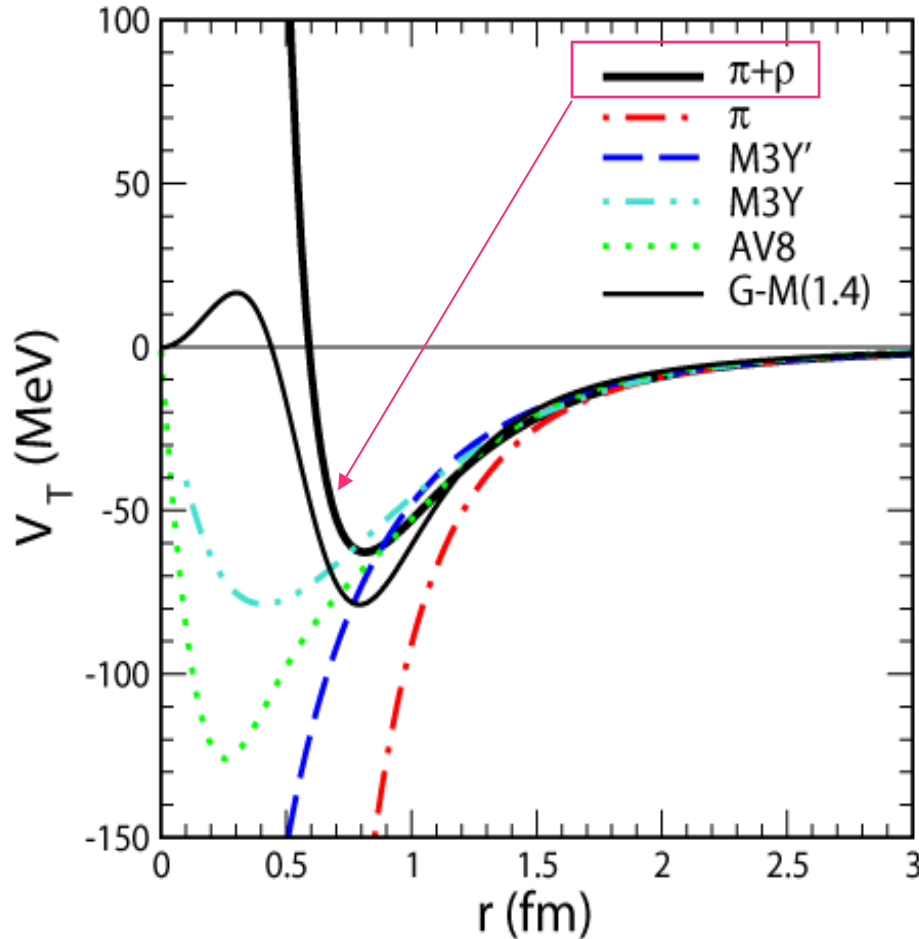
almost equal (no renormalization)



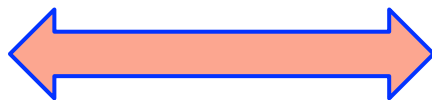
N.Tsunoda, T.O.,
K.Tsukiyama, M.H.-Jensen,
PRC84,044322 (2011)

monopole component of
tensor force **in free space**

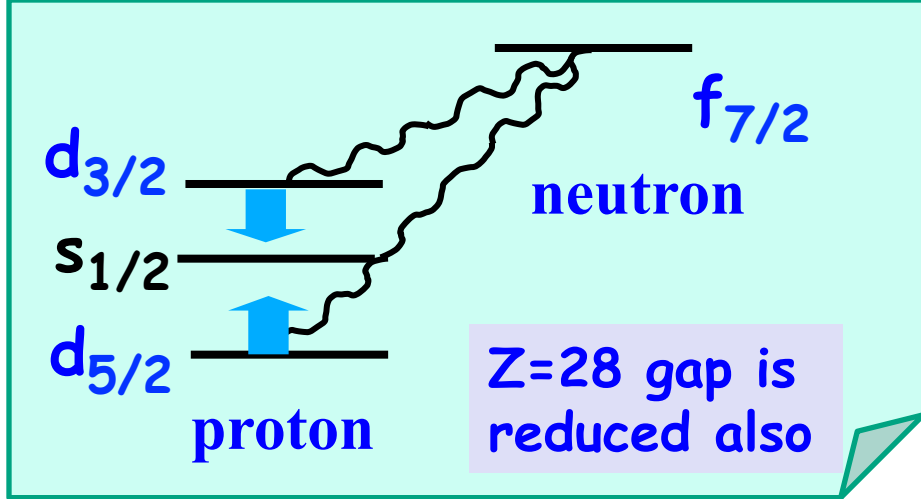
Tensor potentials



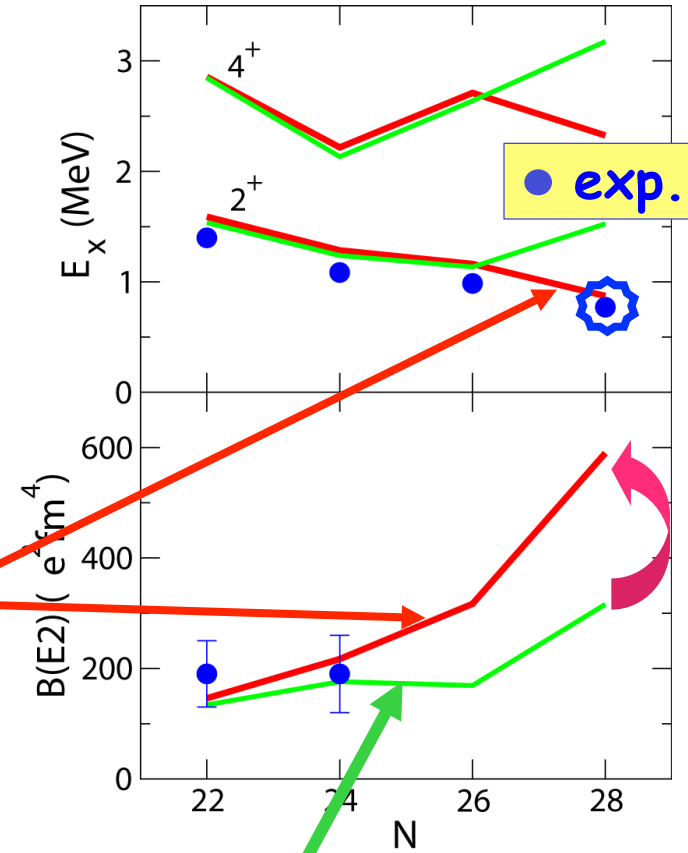
What is the *relevant part* of the tensor force to the present issue.



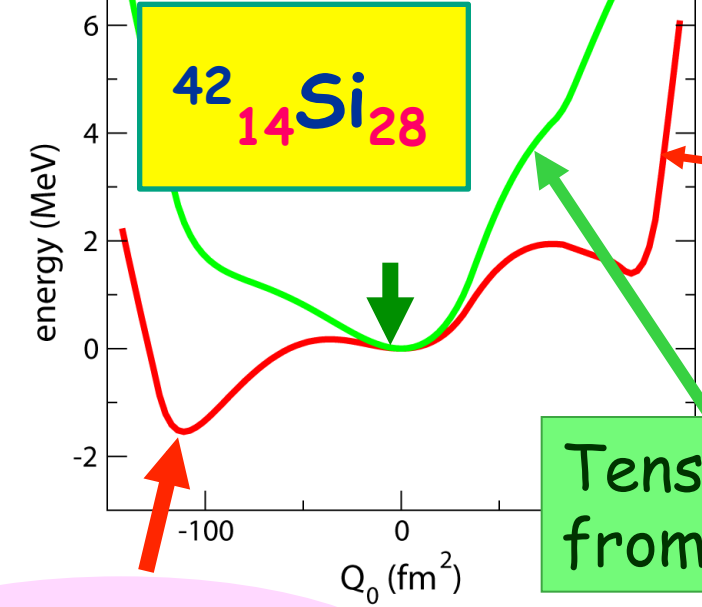
medium-long range



Si isotopes
SM calc. by Utsuno et al.



Potential Energy Surface



full

Tensor force removed
from cross-shell interaction

Strong oblate
Deformation ?

Other calculations
show a variety of shapes.

Otsuka, Suzuki and Utsuno,
Nucl. Phys. A805, 127c (2008)

^{42}Si : B. Bastin, S. Grévy et al.,
PRL 99 (2007) 022503

Tensor force effects in exotic Si isotopes

PRL 109, 182501 (2012)

PHYSICAL REVIEW LETTERS

Well Developed Deformation in ^{42}Si

S. Takeuchi,^{1,*} M. Matsushita,^{1,2,†} N. Aoi,^{1,‡} P. Doornenbal,¹ K. Li,^{1,3} T. Motobayashi,¹ H. Scheit,¹ H. Wang,^{1,3} H. Baba,¹ D. Bazin,⁴ L. C aceres,⁵ H. Crawford,⁶ P. Fallon,⁶ R. Gernh user,⁷ J. Gibeli,⁷ C. Hinke,⁷ C. R. Hoffman,¹⁰ R. Hughes,¹¹ E. Ideguchi,^{9,‡} D. Jenkins,¹² N. Kobayashi,¹³ Y. Kono,^{14,15,¶} T. Le Bleis,^{14,15,¶} J. Lee,¹ G. Lee,¹³ A. Matta,¹⁶ S. Michimasa,⁹ T. Nakamura,¹³ S. Ota,⁹ M. Petri,^{6,8} S. Shimoura,⁹ K. Steiger,⁷ K. Takahashi,¹³ M. Takechi,^{1,**} Y. Togano,^{1,**} R. Winkler,^{4,††}

¹RIKEN Nishina Center, Wako, Saitama 351-0198, Japan

PRL 99, 022503 (2007)

PHYSICAL REVIEW LETTERS

Collapse of the $N = 28$ Shell Closure in ^{42}Si

B. Bastin,² S. Gr vy,^{1,*} D. Sohler,³ O. Sorlin,^{1,4} Zs. Dombr adi,³ N.L. Achouri,² J.C. Ang elique,⁵ D. Baiborodin,⁵ R. Borcea,⁶ C. Bourgeois,⁴ A. Buta,⁶ A. B urger,^{7,8} R. Chapman,⁹ J.C. Dalouzy,¹ Z. Dlouh y,³ Z. Elekes,³ S. Franchoo,⁴ S. Iacob,⁶ B. Laurent,² M. Lazar,⁶ X. Liang,⁹ E. Li nard,² J. Mrazek,⁵ L. Nalpas,³ N. A. Orr,² Y. Penionzhkevich,¹⁰ Zs. Podoly ak,¹¹ F. Pougheon,⁴ P. Roussel-Chomaz,¹ M.G. Saint-Laurent,¹ M. Stanoiu,^{4,6} and I. Stefan¹
F. Nowacki¹² and A. Poves¹³

The energies of the excited states in very neutron-rich ^{42}Si and $^{41,43}\text{P}$ have been measured using in-beam γ -ray spectroscopy from the fragmentation of secondary beams of $^{42,44}\text{S}$ at 39A MeV. The low 2^+ energy of ^{42}Si , 770(19) keV, together with the level schemes of $^{41,43}\text{P}$, provides evidence for the disappearance of the $Z = 14$ and $N = 28$ spherical shell closures, which is ascribed mainly to the action of proton-neutron tensor forces. New shell model calculations indicate that ^{42}Si is best described as a well-deformed oblate rotor.

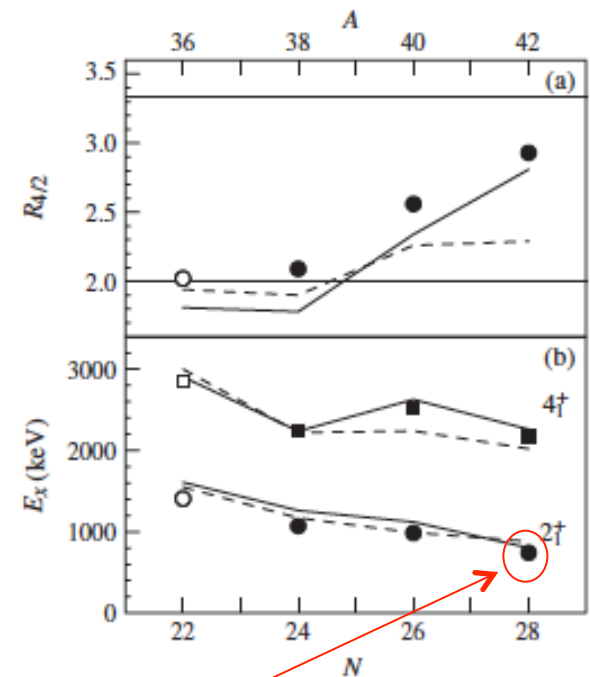


FIG. 3. (a) Ratio between the energies of the 2_1^+ and 4_1^+ states ($R_{4/2}$) for Si isotopes. The horizontal lines at 2.0 and 3.3 indicate the vibrational and rotational limits, respectively. (b) Excitation energies for 2_1^+ and 4_1^+ states, which are indicated by circles and squares, respectively. Filled symbols are results of the present study, and solid and dashed lines represent predictions of the SM with SDPF-MU [17] and SM with SDPF-U-MIX [33], respectively (see text for details). The 2^+ energies of the $N = 24, 26, 28$ Si isotopes have been measured in previous works [12,14,16].

Tensor-force driven Jahn-Teller effect -> shape evolution

PHYSICAL REVIEW C **86**, 051301(R) (2012)

Utsuno, TO, Brown, Honma, Mizusaki, Shimizu

C 86, 051301(R) (2012).

RIBF data close to

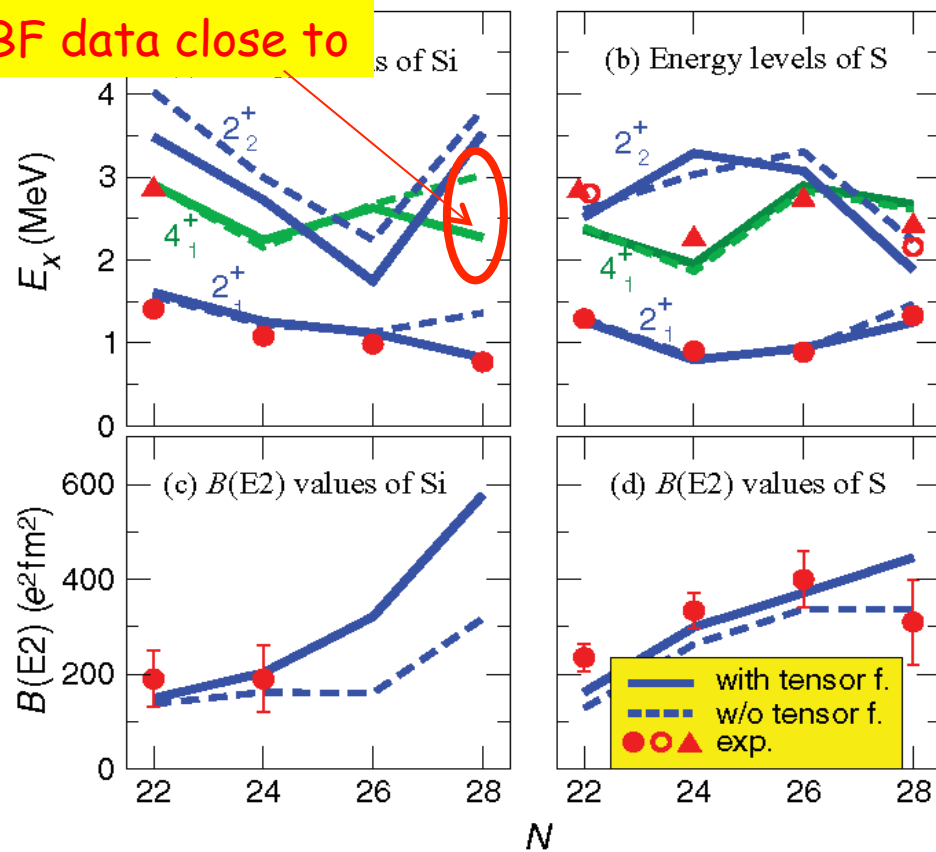
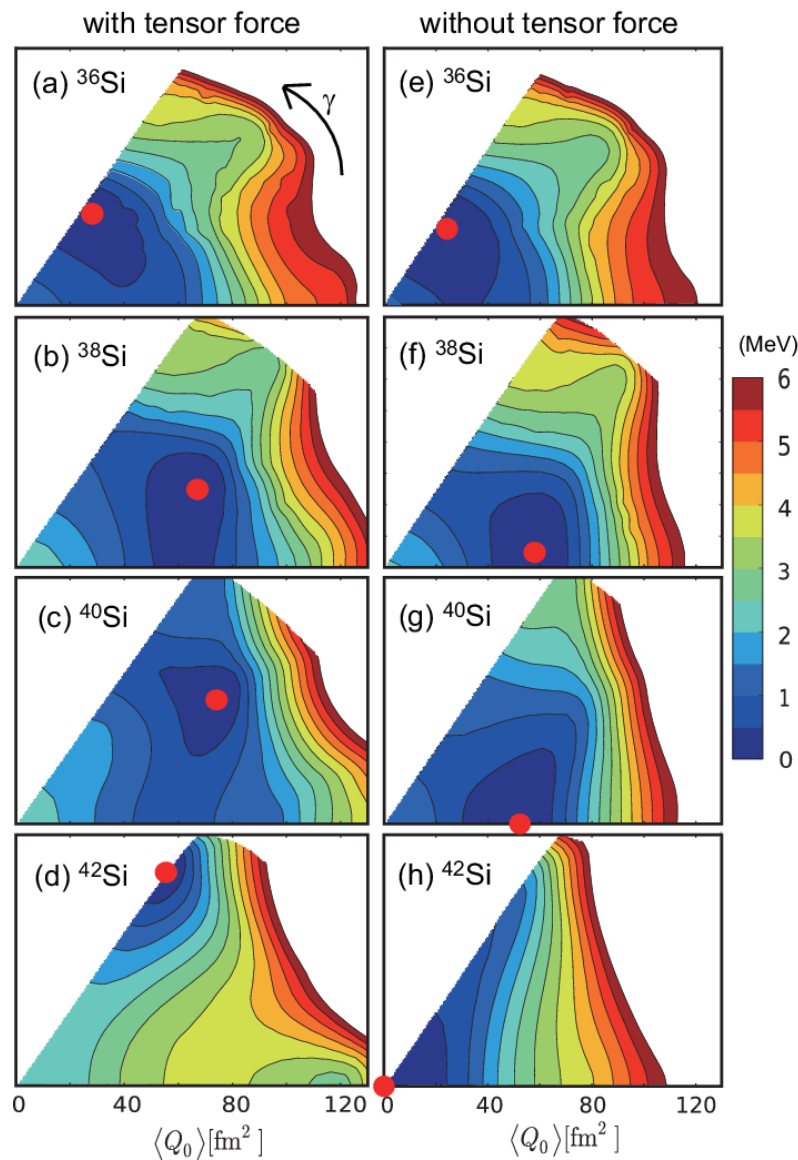


FIG. 3. (Color online) (a) and (b) $2_{1,2}^+$ (blue lines and red circles) and 4_1^+ (green lines and red triangles) energy levels and (c) and (d) $B(E2; 0_1^+ \rightarrow 2_1^+)$ values of Si and S isotopes for $N = 22-28$. Symbols are experimental data [13–19]. Solid (dashed) lines are calculations with (without) the cross-shell tensor force.



Tensor force effects on binding energy (two-neutron separation energy)

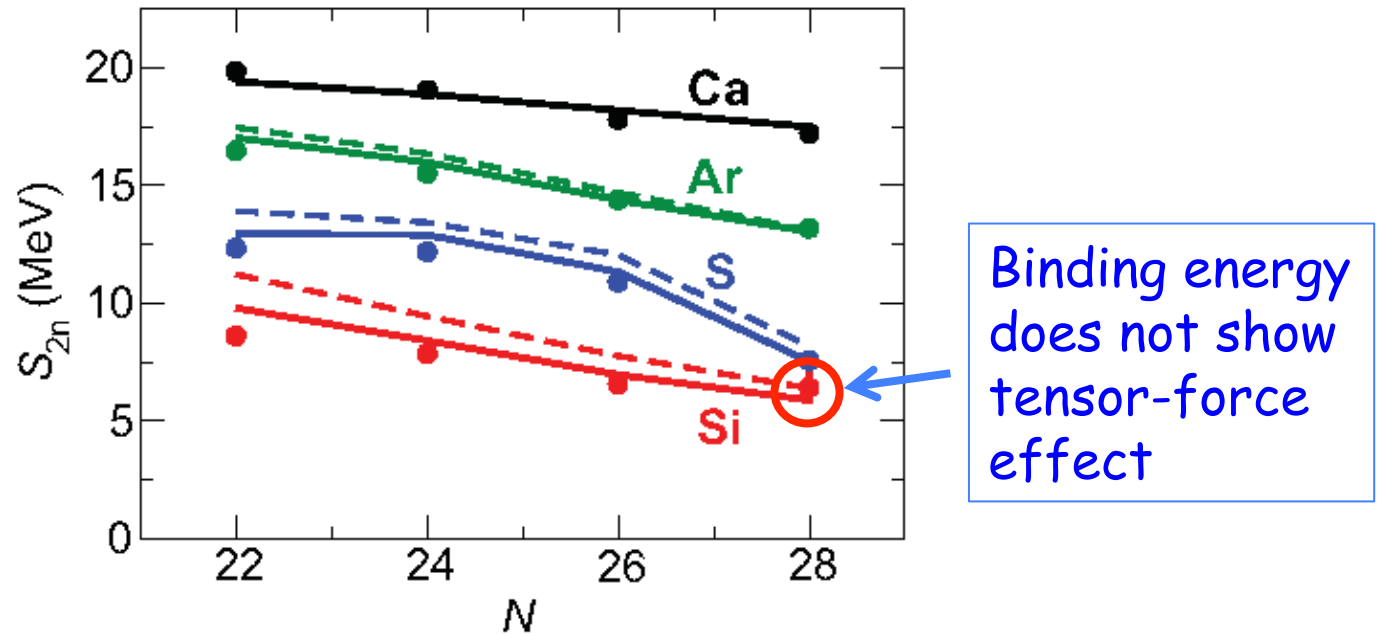


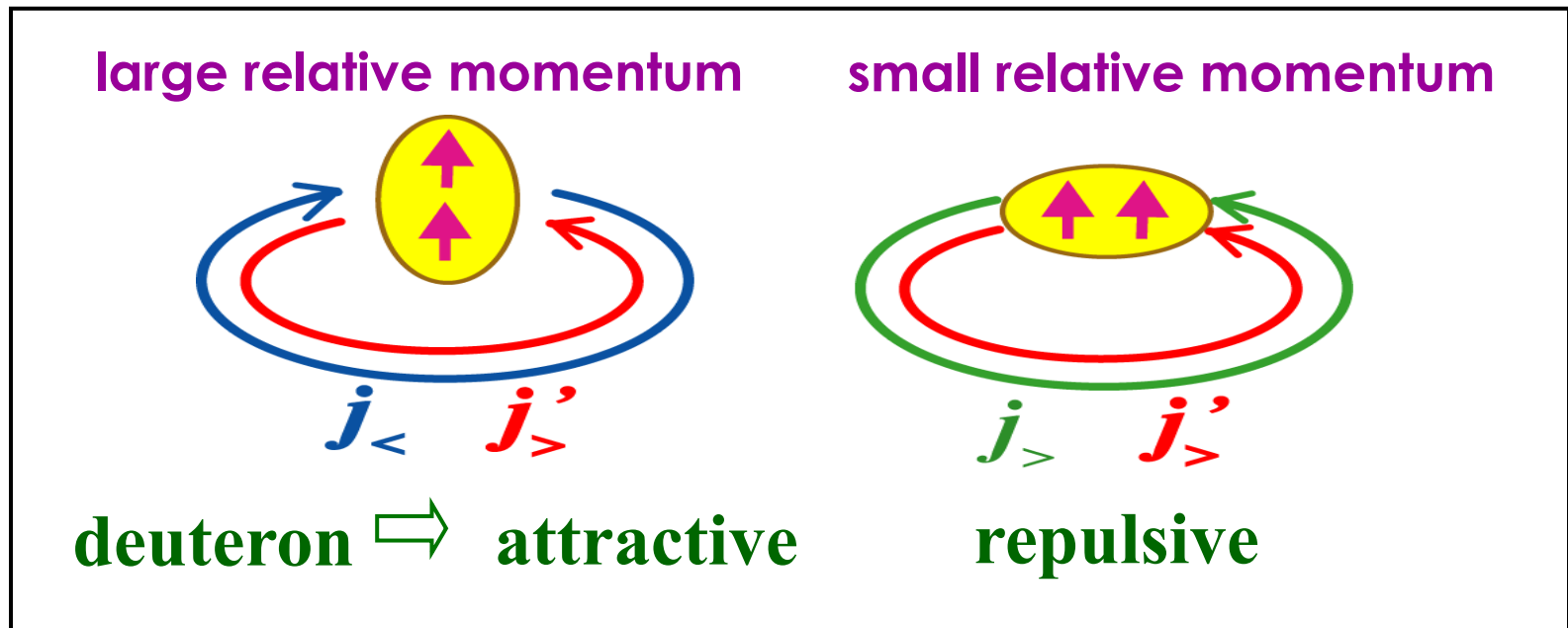
FIG. 6. (Color online) Two-neutron separation energies of Si and S isotopes from $N = 22$ to 28. Solid (dashed) lines are calculations with (without) the cross-shell tensor force. Points are experimental data [40,41].

Stancu, Brink and Flocard, Phys. Lett. 68B, 108 (1977)

Zero-range spin-momentum tensor coupling term

$$\begin{aligned}
 v_T = & \frac{1}{2} T \{ [(\boldsymbol{\sigma}_1 \cdot \mathbf{k}')(\boldsymbol{\sigma}_2 \cdot \mathbf{k}') - \frac{1}{3}(\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2)k'^2] \delta(r_1 - r_2) \\
 & + \delta(r_1 - r_2) [(\boldsymbol{\sigma}_1 \cdot \mathbf{k})(\boldsymbol{\sigma}_2 \cdot \mathbf{k}) - \frac{1}{3}(\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2)k^2] \} \\
 & + U \{ (\boldsymbol{\sigma}_1 \cdot \mathbf{k}') \delta(r_1 - r_2) (\boldsymbol{\sigma}_1 \cdot \mathbf{k}) \\
 & - \frac{1}{3}(\boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2) [k' \cdot \delta(r_1 - r_2) \mathbf{k}] \}, \quad (1)
 \end{aligned}$$

This is not be a good approximation to the tensor force itself, but may simulate the monopole effect of the tensor shown below, picking up differences in relative momenta.

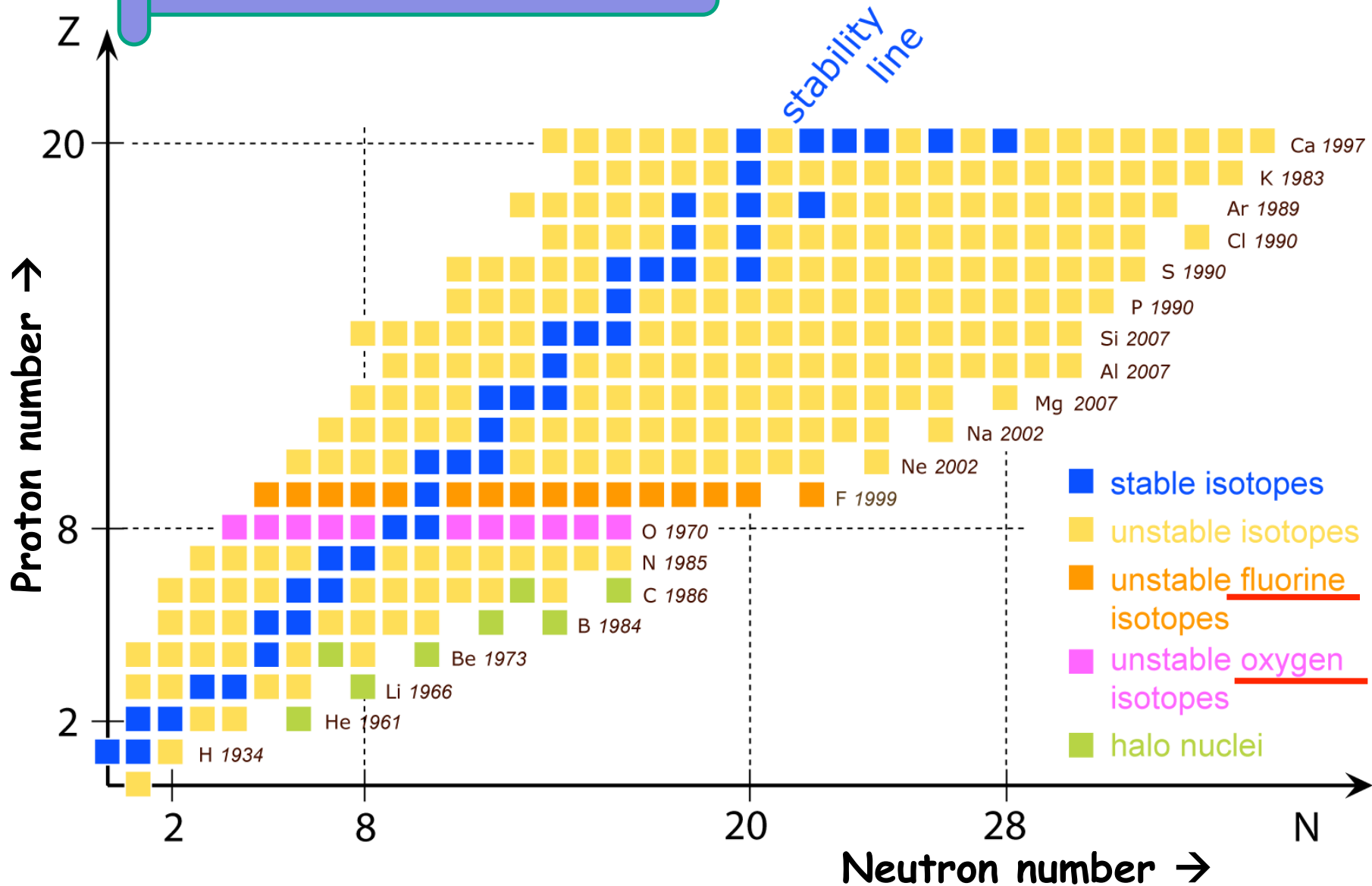


Outline

1. Introduction
2. Shell model and monopole interaction
3. Shell evolution and tensor force
4. Multiple quantum liquid in exotic nuclei
5. Shell evolution and three-nucleon force
6. Monte Carlo Shell Model
7. Summary

Nuclear Chart - Left Lower Part -

Why is the drip line of Oxygen so near ?



Single-Particle Energy for Oxygen isotopes

by microscopic eff. int.

G-matrix+ core-pol. : Kuo, Brown

$V_{\text{low-k}}$: Bogner, Kuo, Schwenk

by phenomenological eff. int.

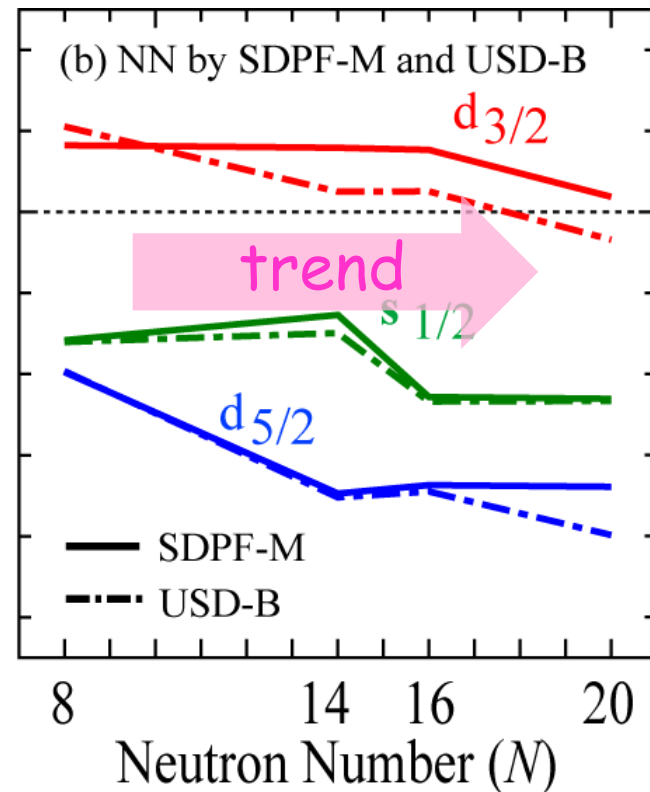
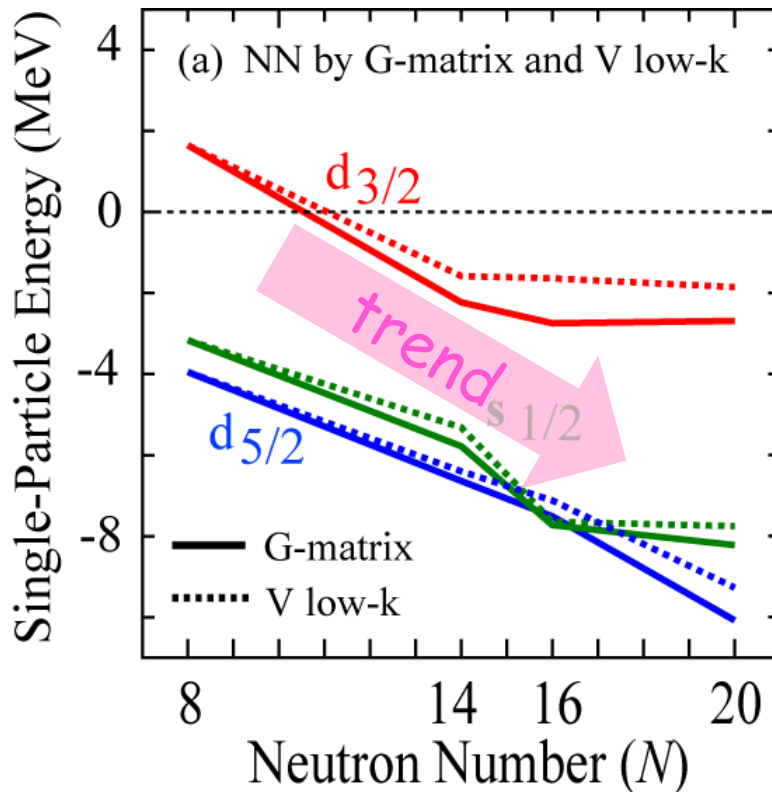
- G-matrix + fit -

SDPF-M

Utsuno, O., Mizusaki, Honma,
Phys. Rev. C **60**, 054315 (1999)

USD-B

Brown and Richter,
Phys. Rev. C **74**, 034315 (2006)



What is the origin of
the *repulsive modification* of
 $T=1$ monopole matrix elements ?

The same puzzle as in the pf shell

A solution within *bare* 2-body interaction
is very unlikely
(considering efforts made so far)

Zuker, Phys. Rev. Lett. 90, 042502 (2003)

→ 3-body interaction

The clue : Fujita-Miyazawa 3N mechanism (Δ -hole excitation)

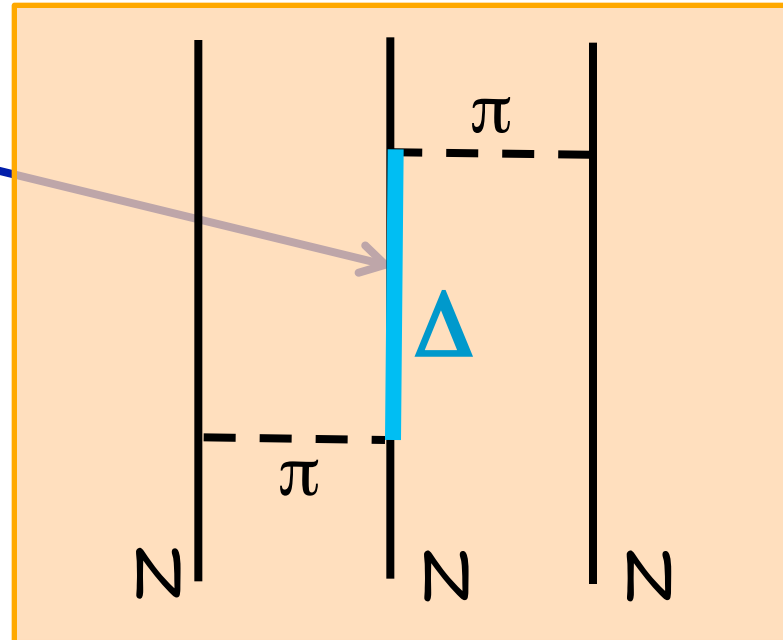
Progress of Theoretical Physics, Vol. 17, No. 3, March 1957

Pion Theory of Three-Body Forces

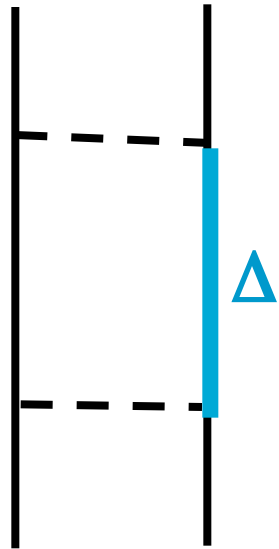
Jun-ichi FUJITA and Hironari MIYAZAWA

Δ particle
 $m=1232$ MeV
 $S=3/2, I=3/2$

Oset, Toki and Weise
Pionic modes of excitation
Phys. Rep. 83, 281 (1982)



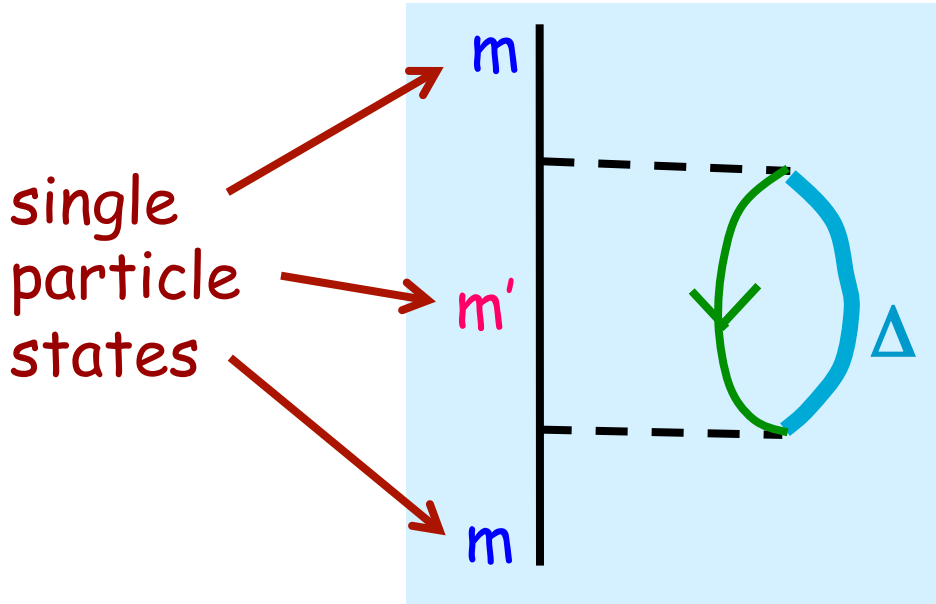
Renormalization of NN interaction due to Δ excitation in the intermediate state



Modification to
bare NN interaction
(for NN scattering)

$T=1$
attraction
between NN
effectively

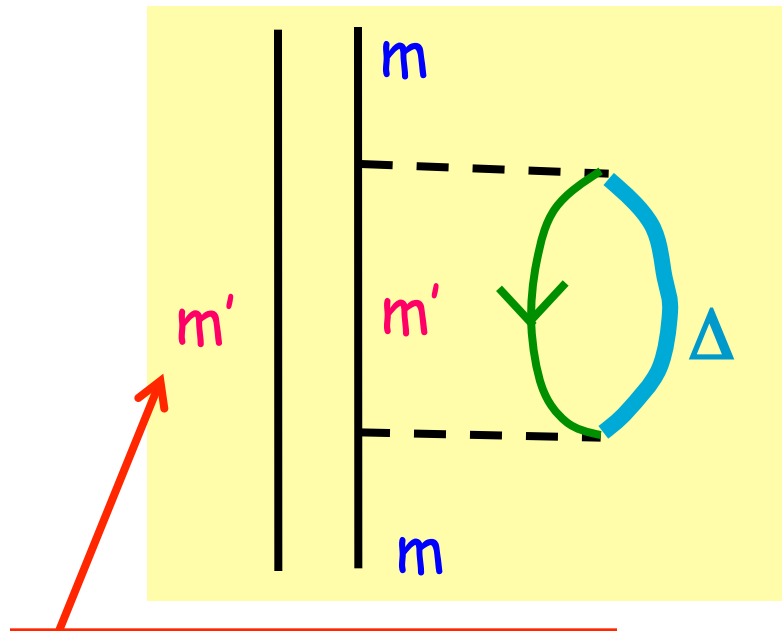
Pauli blocking effect on the renormalization of single-particle energy



Renormalization of single particle energy due to

Δ -hole excitation

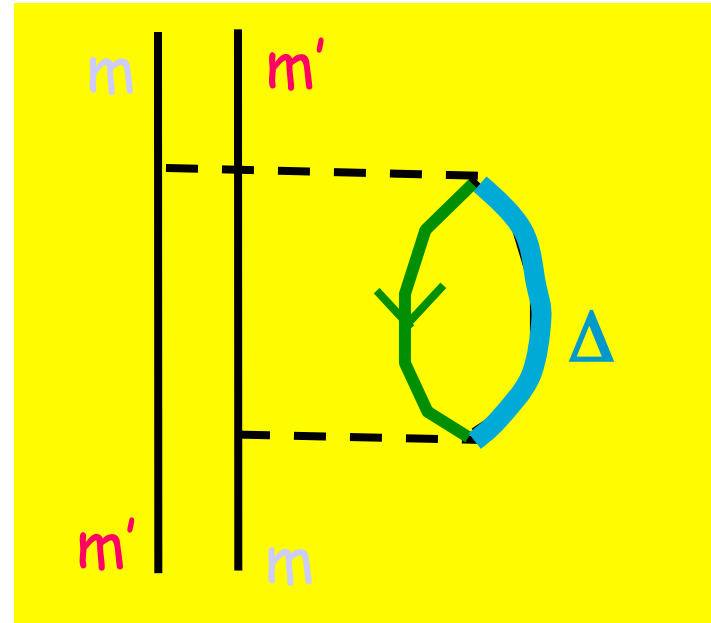
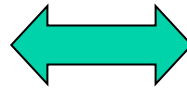
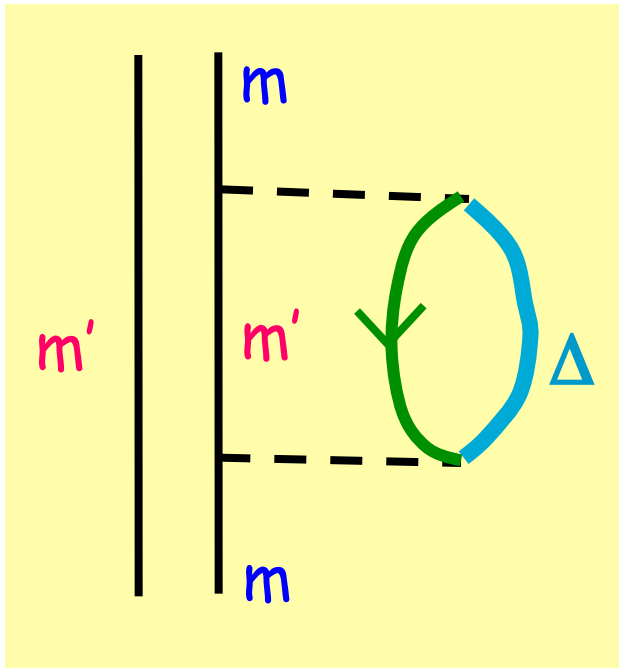
→ more binding (attractive)



Another valence particle in state m'

Pauli Forbidden
→ *The effect is suppressed*

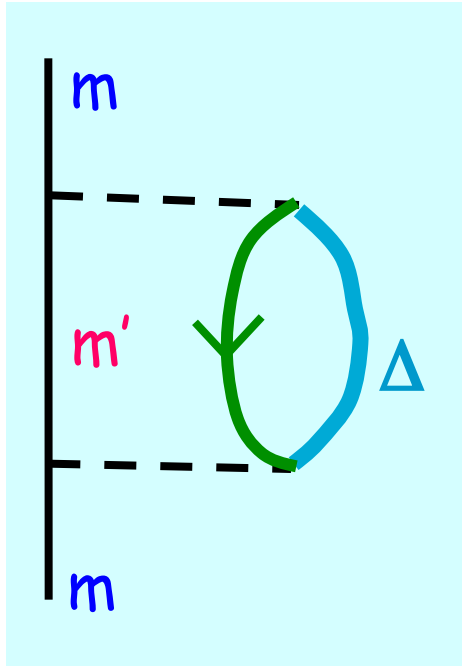
Inclusion of Pauli blocking



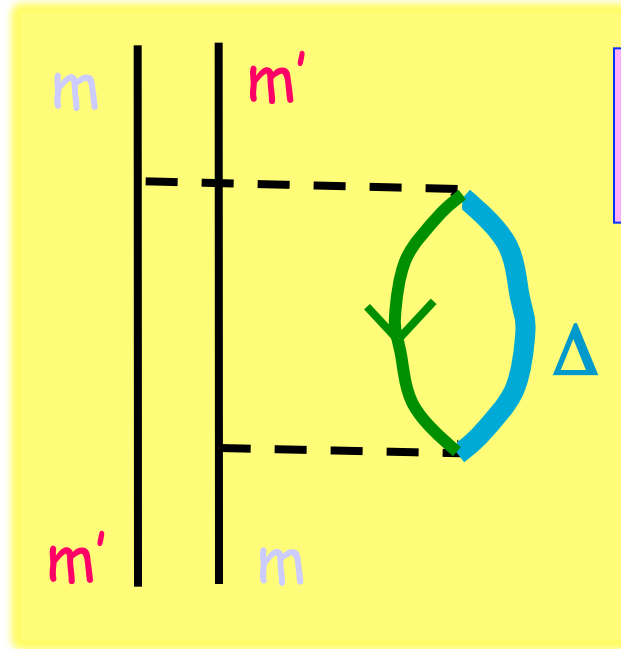
Pauli forbidden
(from previous page)

This Pauli effect is
included automatically
by the exchange term.

Most important message with Fujita-Miyazawa 3NF



+



Pauli blocking

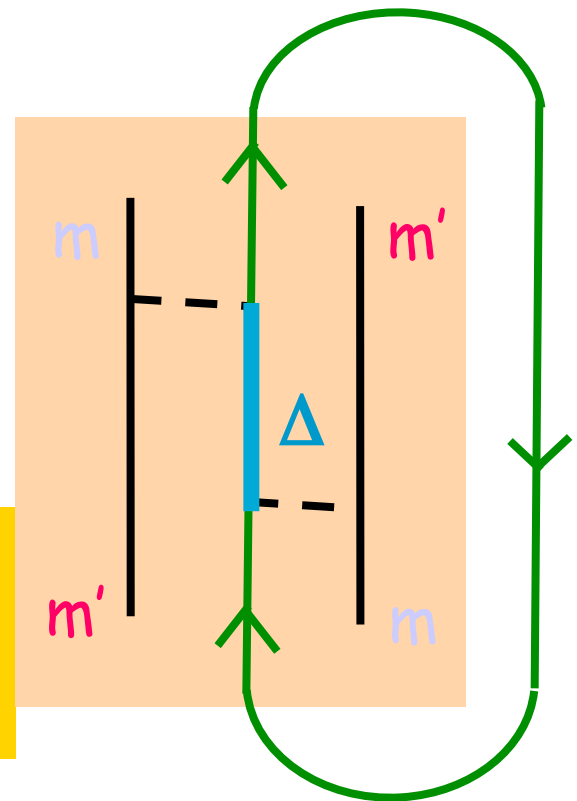
Effective monopole repulsive interaction



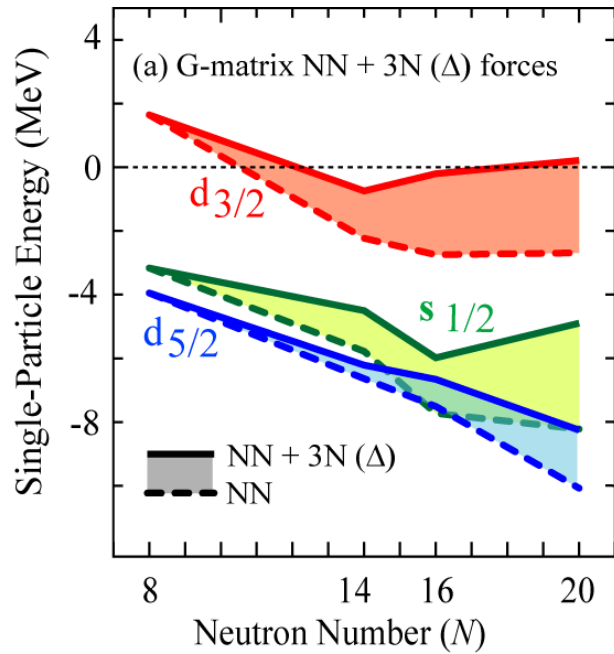
Renormalization of single particle energy

same

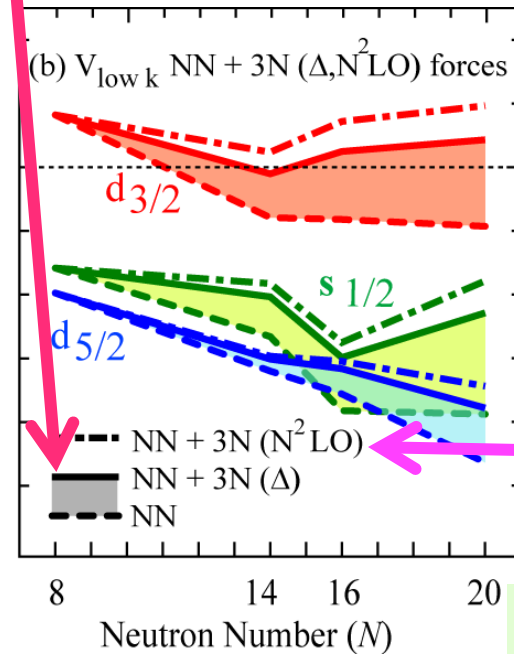
Monopole part of Fujita-Miyazawa 3-body force



(i) Δ -hole excitation in a conventional way



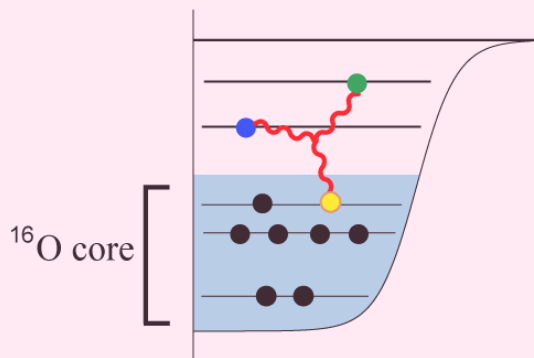
(ii) EFT with Δ



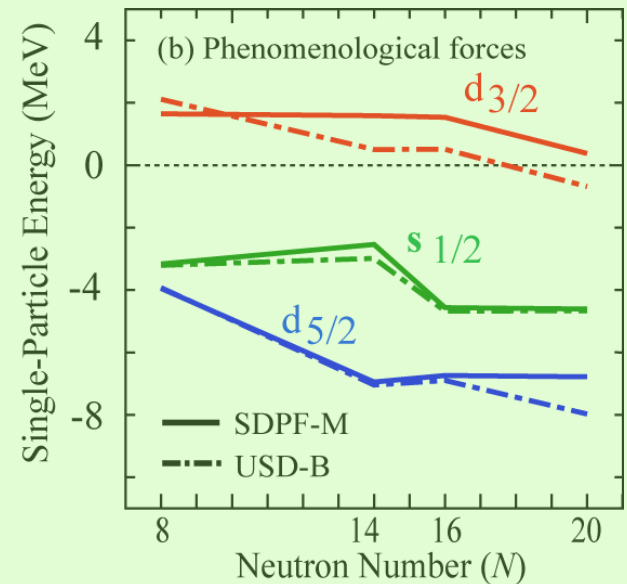
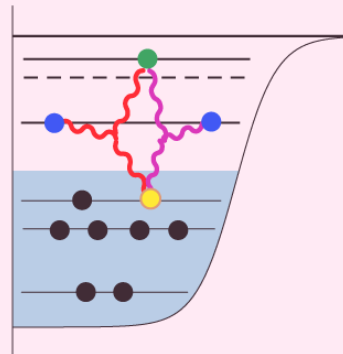
Δ -hole dominant role in determining oxygen drip line

(iii) EFT incl. contact terms (N^2LO)

(c) 3-body interaction

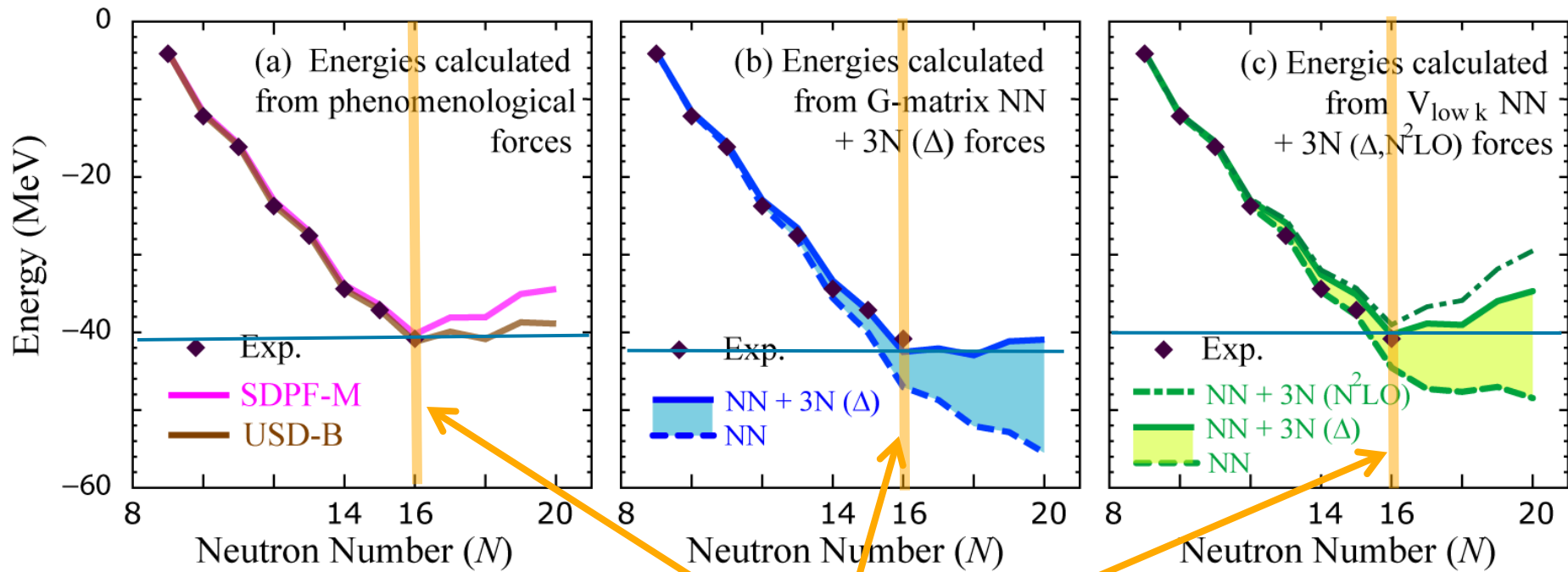
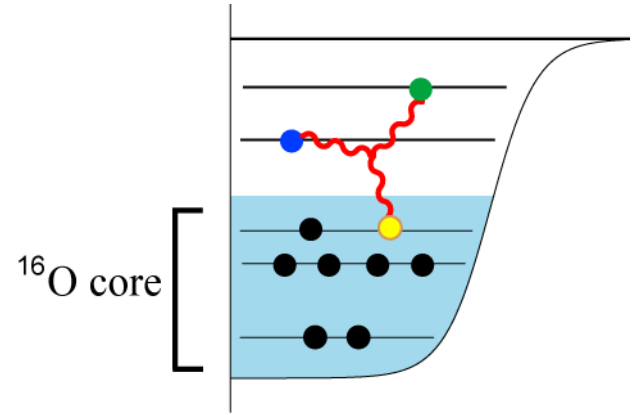


(d) 3-body interaction with one more neutron added to (c)



Ground-state energies of oxygen isotopes

NN force + $3N$ -induced NN force
(Fujita-Miyazawa force)



Drip line

Conventional calculation with $\pi N\Delta$ coupling

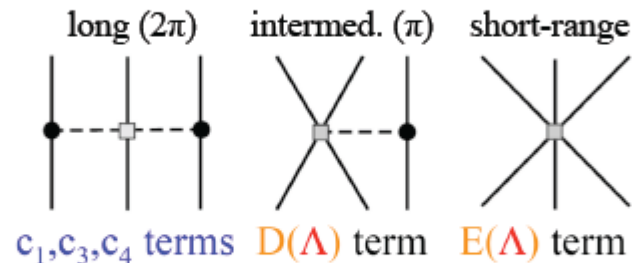
π exchange with radial cut-off at 0.5 fm , $\Delta E = 293$ MeV

$$f_{\{\pi N\Delta\}}/f_{\{\pi NN\}} = \sqrt{9/2}$$

A.M. Green, Rep. Prog. Phys. 39, 1109 (1976)

Low-momentum 3N interactions

from leading N^2 LO chiral EFT $\sim (Q/\Lambda)^3$ van Kolck (1994), Epelbaum et al. (2002)



c_1 from πN , consistent with NN

Meissner (2007)

$$c_1 = -0.9^{+0.2}_{-0.5}, \quad c_3 = -4.7^{+1.2}_{-1.0}, \quad c_4 = 3.5^{+0.5}_{-0.2}$$

c_3, c_4 important for structure, large uncertainties at present

NN for smooth cutoff $V_{low k}$ ($n_{exp}=4$) from N^3 LO(500)

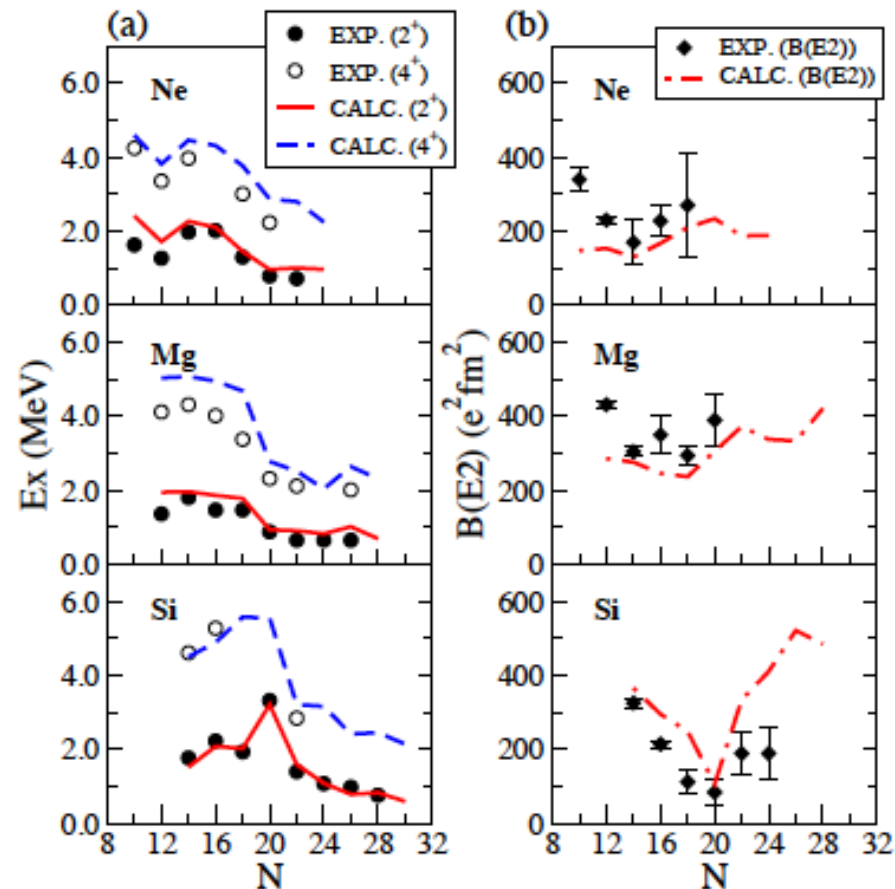
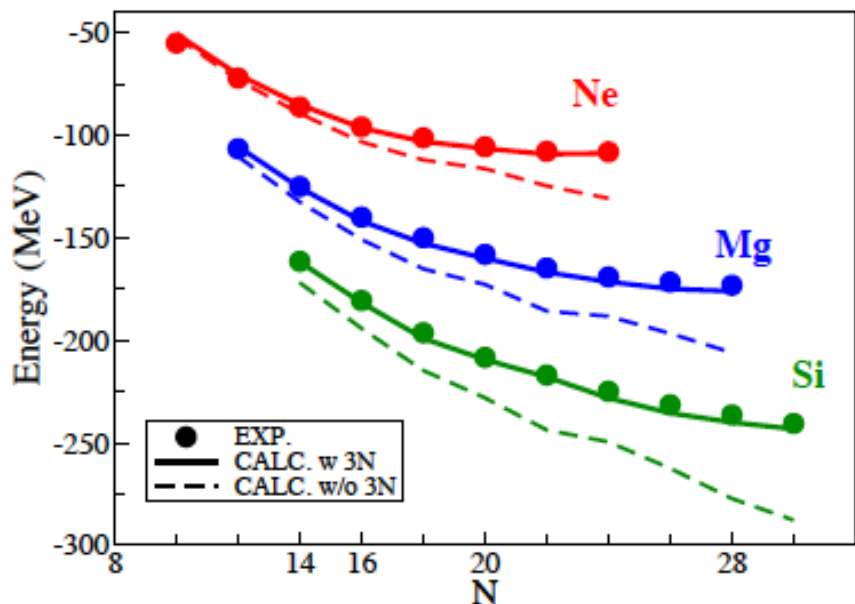
D, E terms fitted to E(3H) and radius(4He)

Ne-Mg-Si in the sd-pf shell

Prototype of future shell-model calculation

New & preliminary

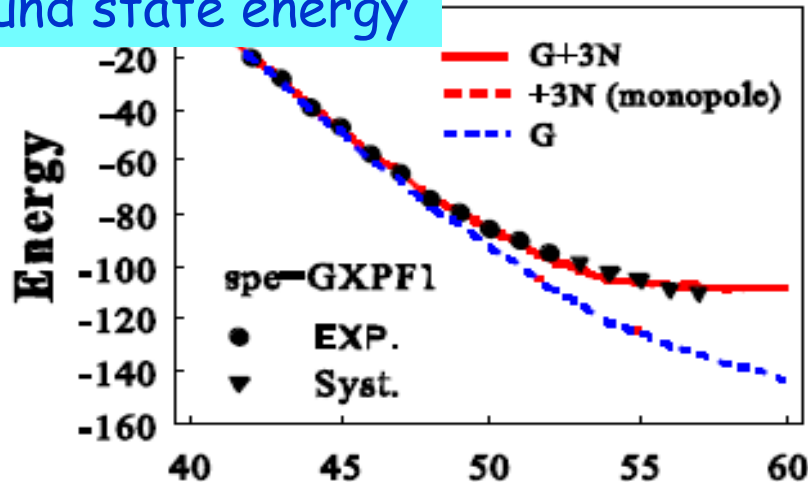
Chiral N3LO NN interaction
+ EKK (Extended Kuo-Krenciglowa) method
+ three-body Fujita-Miyazawa force



In preparation, Tsunoda, Shimizu, Otsuka,
H.-Jensen, Takayanagi, and Suzuki

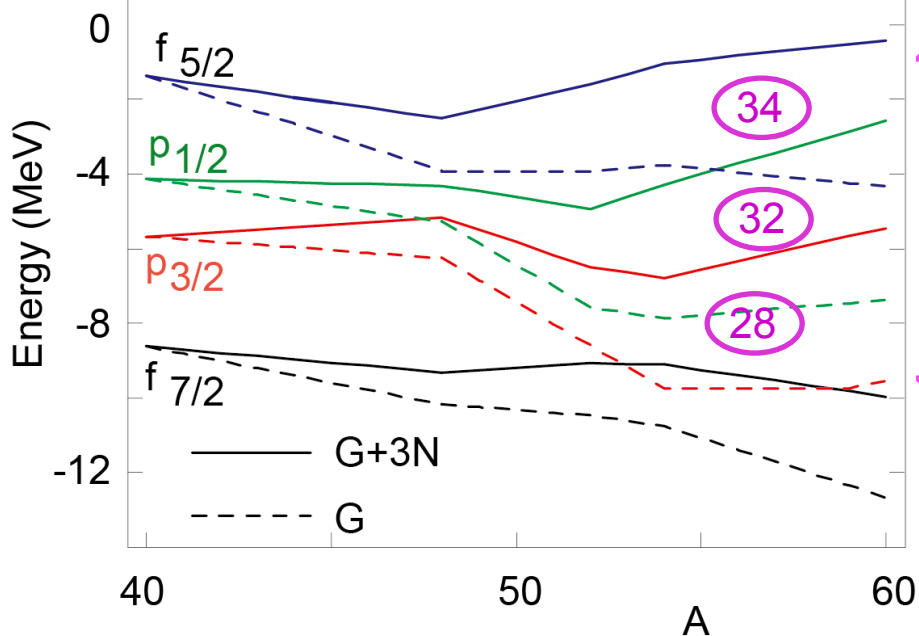
Ca isotopes

ground state energy



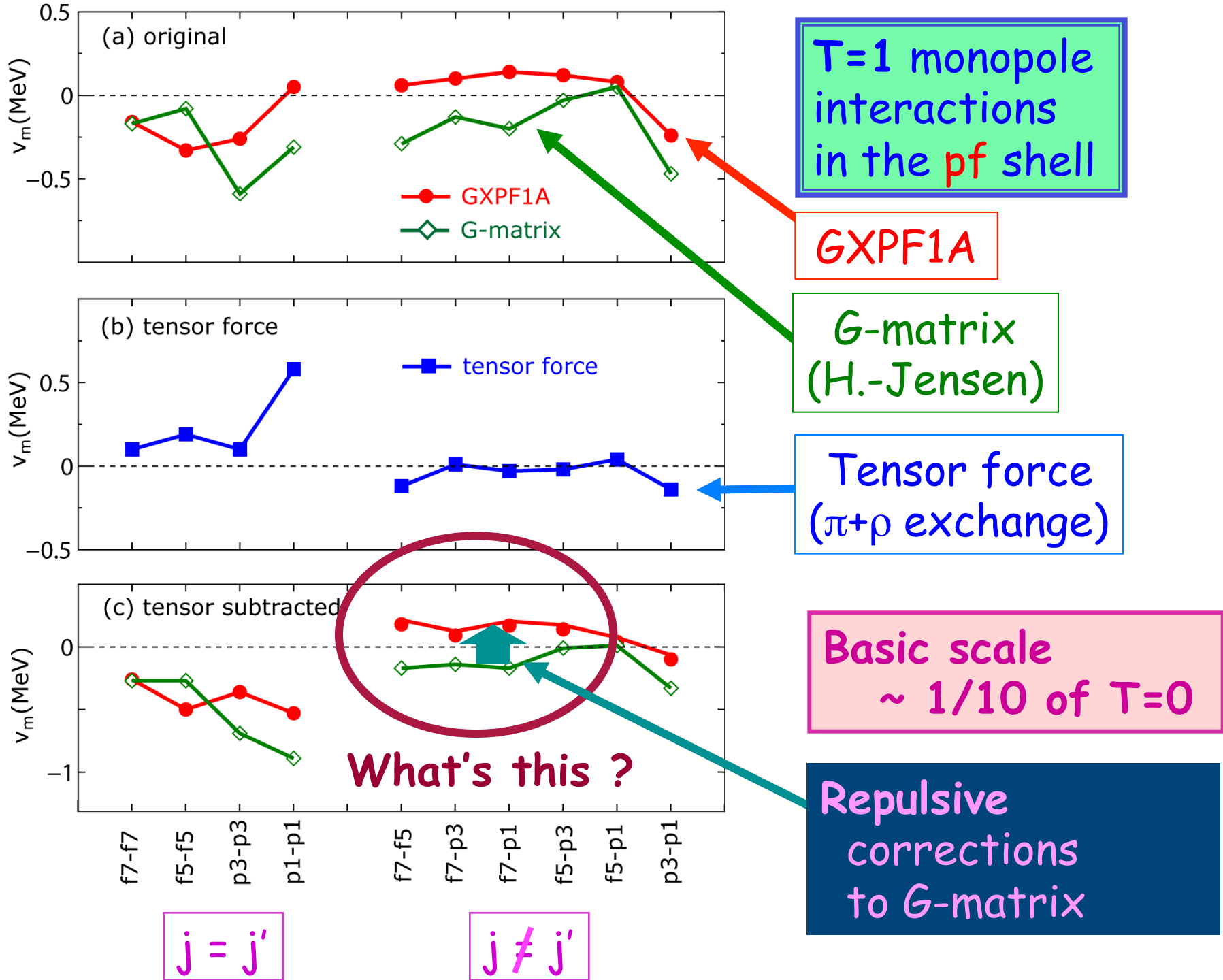
G-matrix (revised)
 3rd order Q-box
 24 hw int. states
 3NF : Fujita-Miyazawa

eff. single-particle energy of neutrons



3-5 MeV rising
 of SPEs
 => ~40 MeV shift
 of B.E. of ^{60}Ca
 from 2NF result

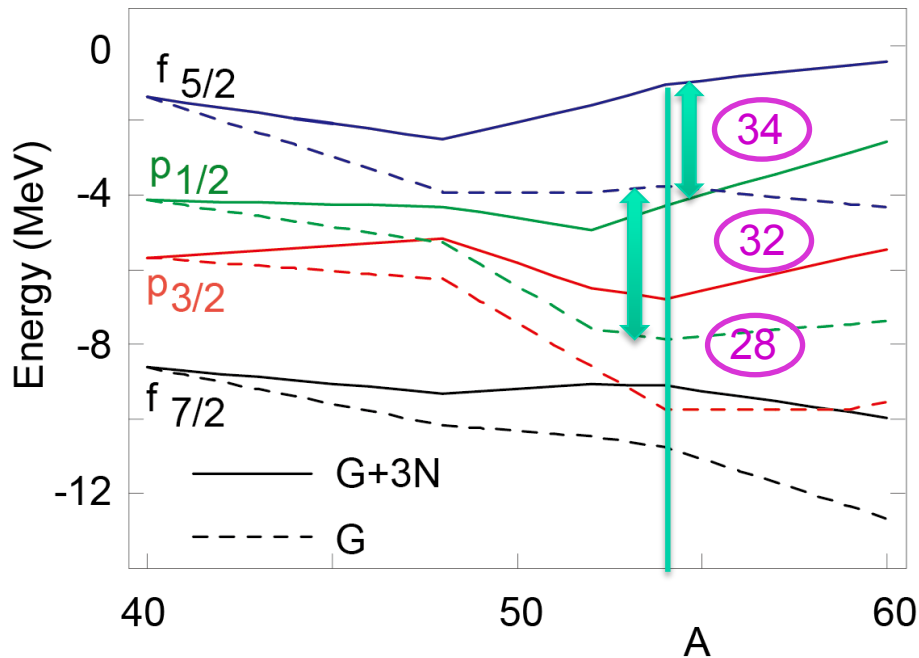
2 MeV enlargement of
 full shell height
 (anti-quenching)



3-body forces does produce another shell evolution

Ca isotopes

Neutron single-particle energy of Ca isotopes



Three-body force
reduces
 $f_{5/2} - p_{1/2}$ gap
by pushing up $p_{1/2}$

G-matrix

3rd order Q-box

24 hw int. states

3NF : Fujita-Miyazawa

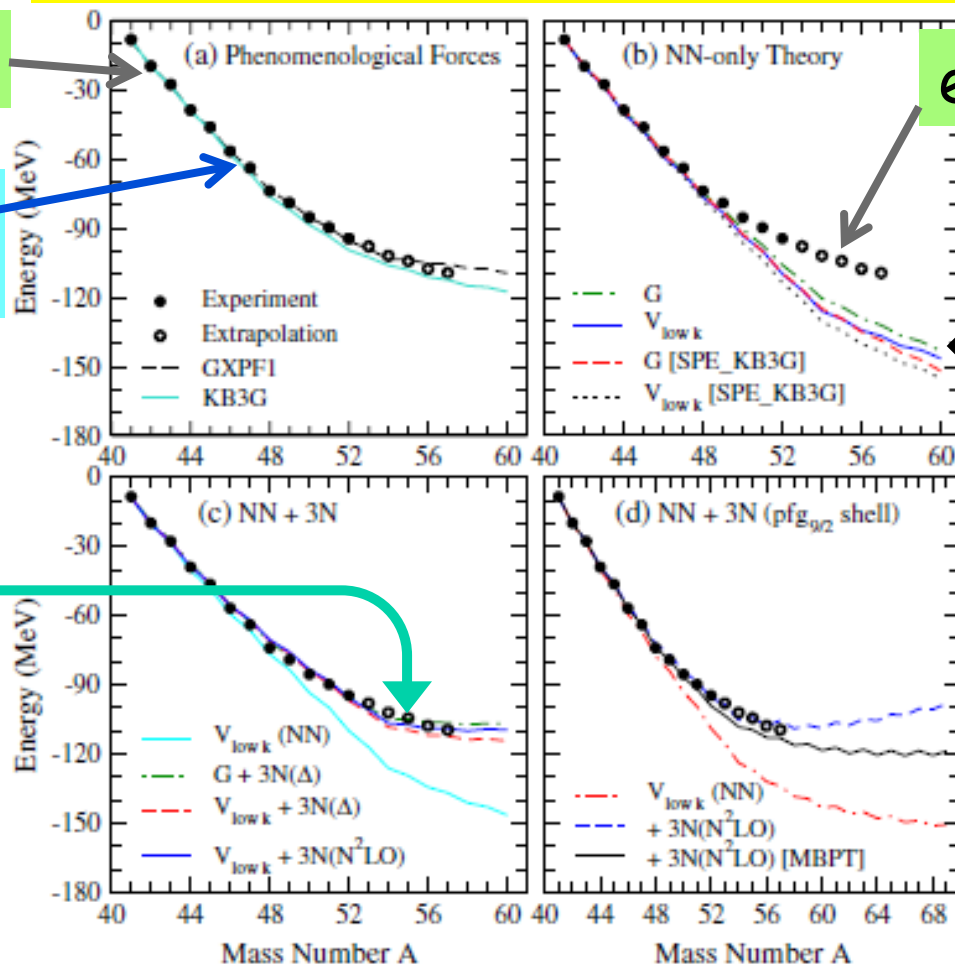
Three-body forces and shell structure in calcium isotopes

Jason D Holt^{1,2}, Takaharu Otsuka^{3,4}, Achim Schwenk^{5,6}
and Toshio Suzuki⁷

Ground-state energy of Ca isotopes

exp + extrap.

Phenomenological
NN



exp + extrap.

Microscopic
NN

Microscopic
NN + 3NF

+ $g_{9/2}$

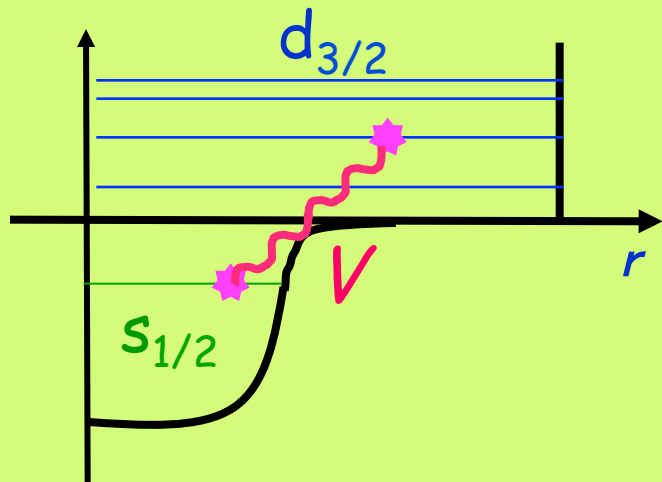
Continuum-coupled shell model (CCSM)

$$\text{Hamiltonian: } H = H_0 + \hat{V} = \sum_j \tilde{\epsilon}_j n_j + \hat{V}$$

$$H_0 = T + U_{WS} + V_{\text{wall}} = \sum_j \tilde{\epsilon}_j n_j$$

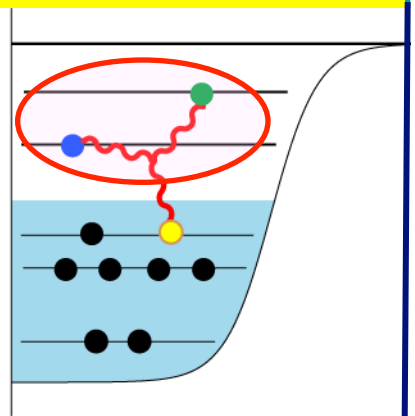
approximated
by Gaussian

basis state-vector (denoted by j):
bound states + discretized continuum states
wall very far (3000 fm, ~3000 basis states)



$V_{NN} +$

^{16}O core



included

$$\hat{V}(r) = \sum_{i=1,2} g_i (1 + a_i \sigma \cdot \sigma) e^{-r^2/d_i^2}$$

$$d_{1,2} = 1.4, 0.7 \text{ fm}$$

SDPF-M TBME = TBME of this $V(r)$
for HO wave functions

$$\langle 1s_{1/2} 0d_{3/2} | V | 1s_{1/2} 0d_{3/2} \rangle_{J=1,2}$$

$$\langle 0d_{3/2} 0d_{3/2} | V | 0d_{3/2} 0d_{3/2} \rangle_{J=0,2}$$

under the assumption that 3-body force effect
is included in SDPF-M interaction effectively

$V(r)$ is fixed only by interaction

$^{240}\text{O} = ^{220}\text{O} + 2n$ in the space

ground state : $2n$ in $1s_{1/2}$

excited states of 1^+ and 2^+ :

$$|iJ^+\rangle = |1s_{1/2} \otimes id_{3/2}; J^+\rangle$$



discretized continuum $id_{3/2}$ ($i = 1, 2, \dots$)

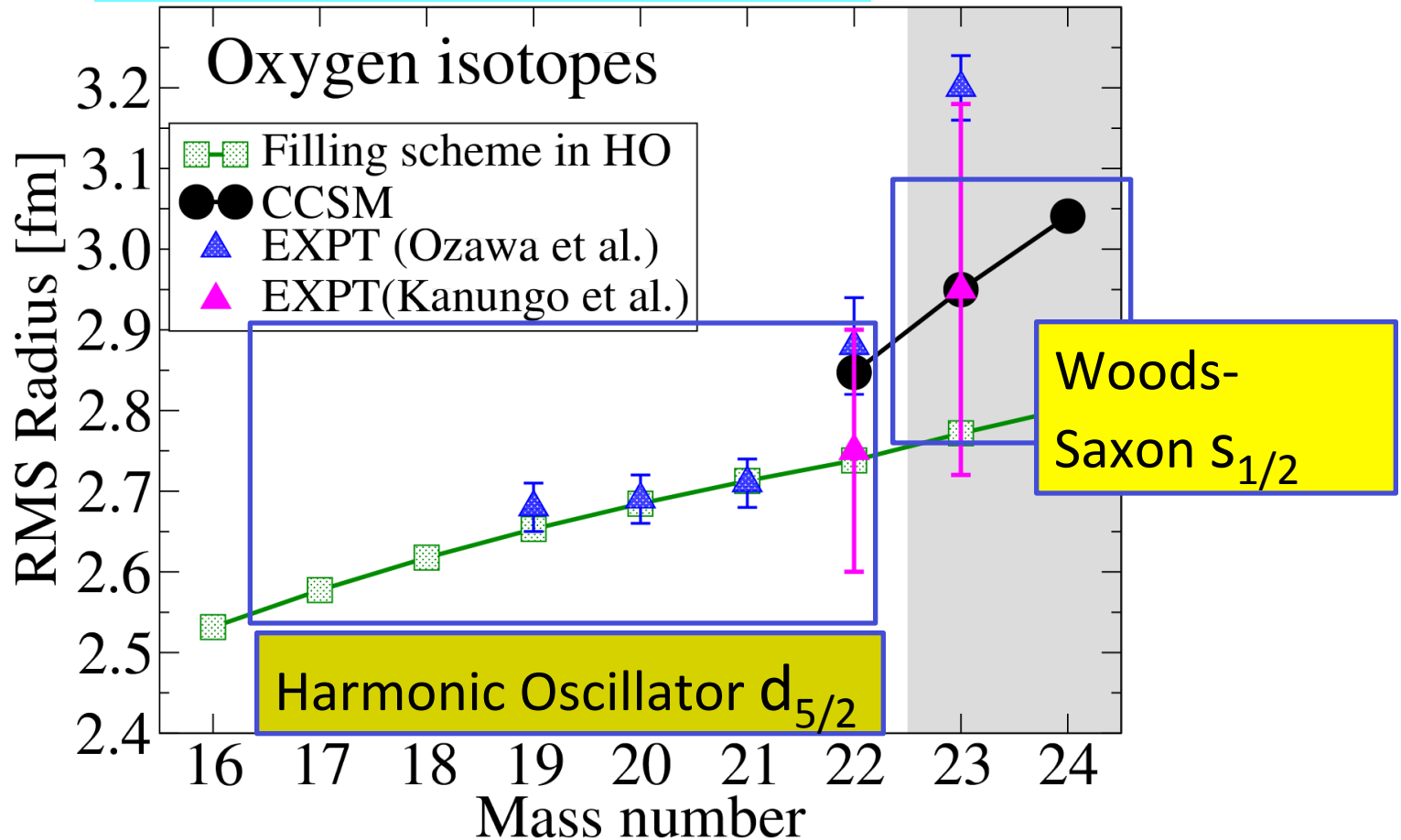
$1s_{1/2}$: solution of Woods-Saxon potential with observed S_n

diagonalize H

Eigenfunction :

$$|J_k^+\rangle = \sum_i c_i^{(J,k)} |iJ^+\rangle$$

RMS Radius: $^{16-24}\text{O}$



Exp: Ozawa et al., Nucl. Phys. A693, 32 (2001)
Kanungo et al., Phys. Rev. C84, 061304 (2011)

Removal of one proton and one neutron from ^{26}F

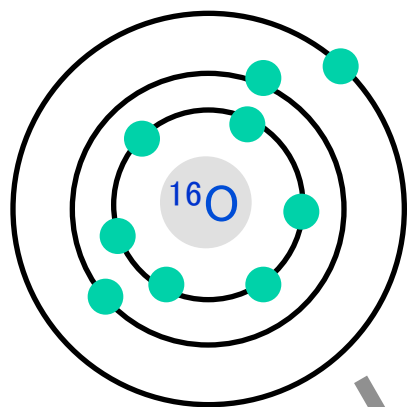
$^9\text{Be}(^{26}\text{F}, ^{24}\text{O})\text{X}$

C. Hoffman,
M. Thoennessen et al.

knockout reaction @MSU (2009)

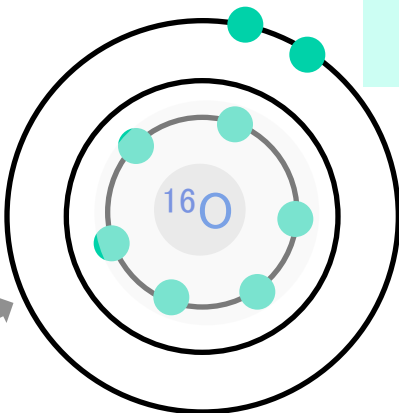
less probable

\Leftarrow large $s_{1/2}$ - $d_{3/2}$ neutron gap



bound nucleus
 ^{26}F

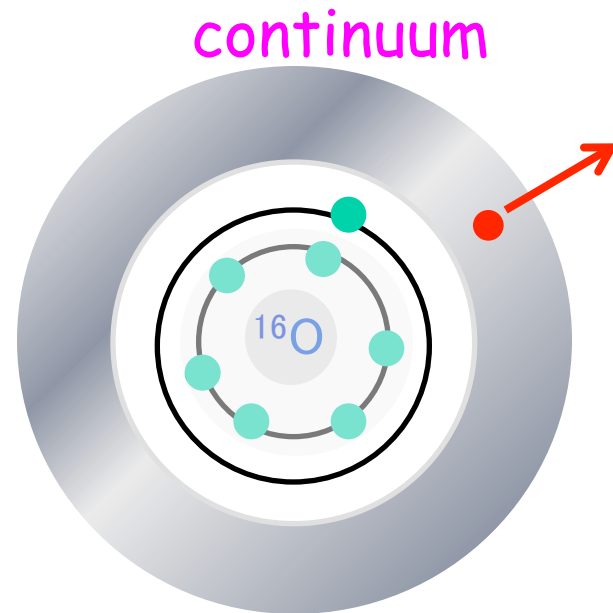
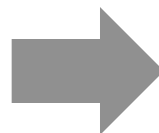
-p
-n



doorway state

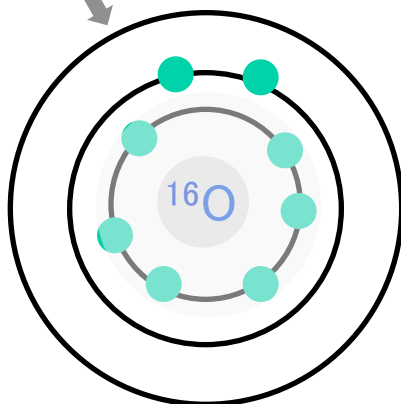
$$|1s_{1/2}0d_{3/2}; J_k^+\rangle$$

H.O.



excited states in ^{24}O

$$H^{\text{CCSM}}|J_k^+\rangle = E_k|J_k^+\rangle$$

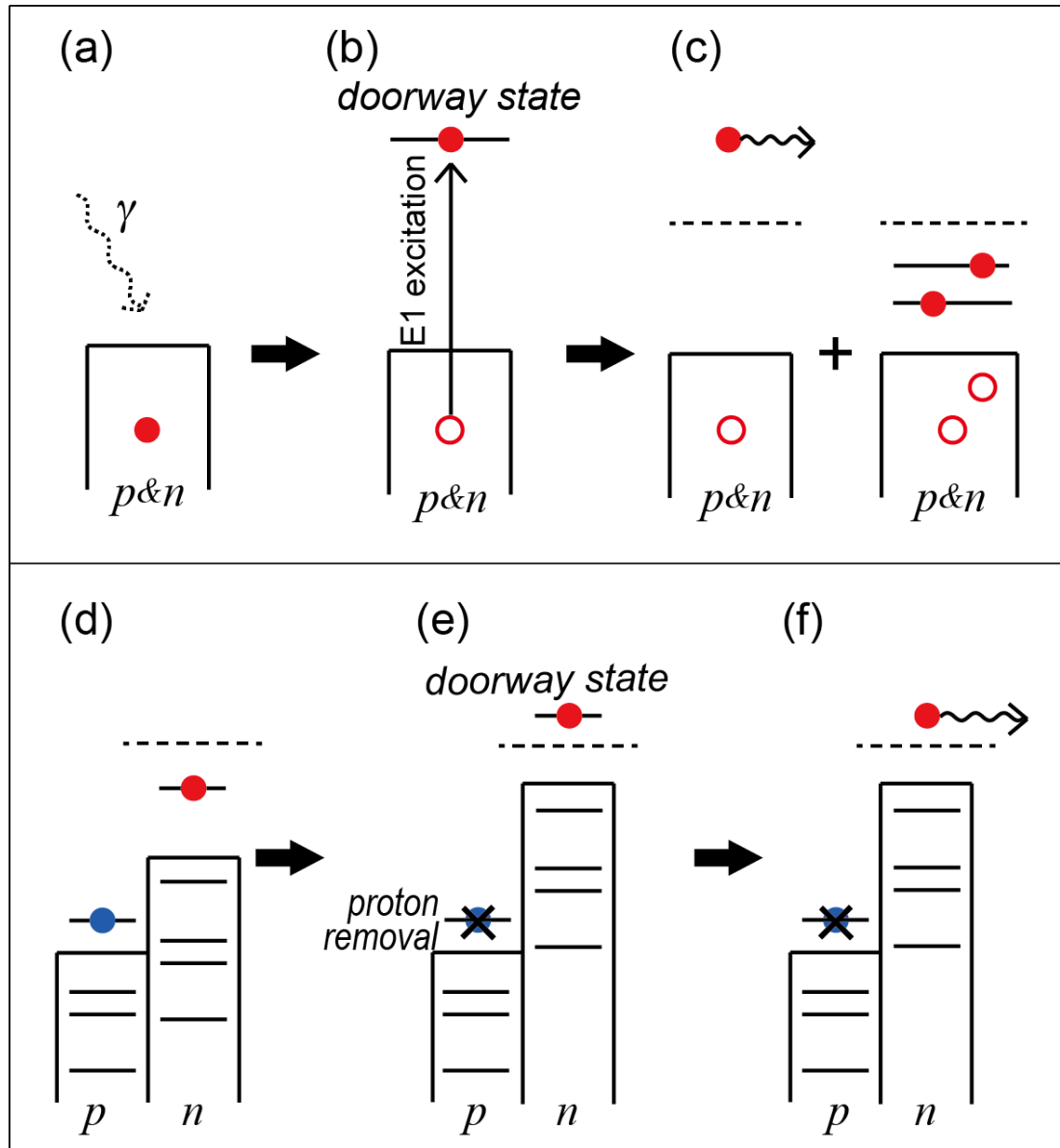


ground state

$1s_{1/2}$ is bound.

Kanungo et al. (2009)

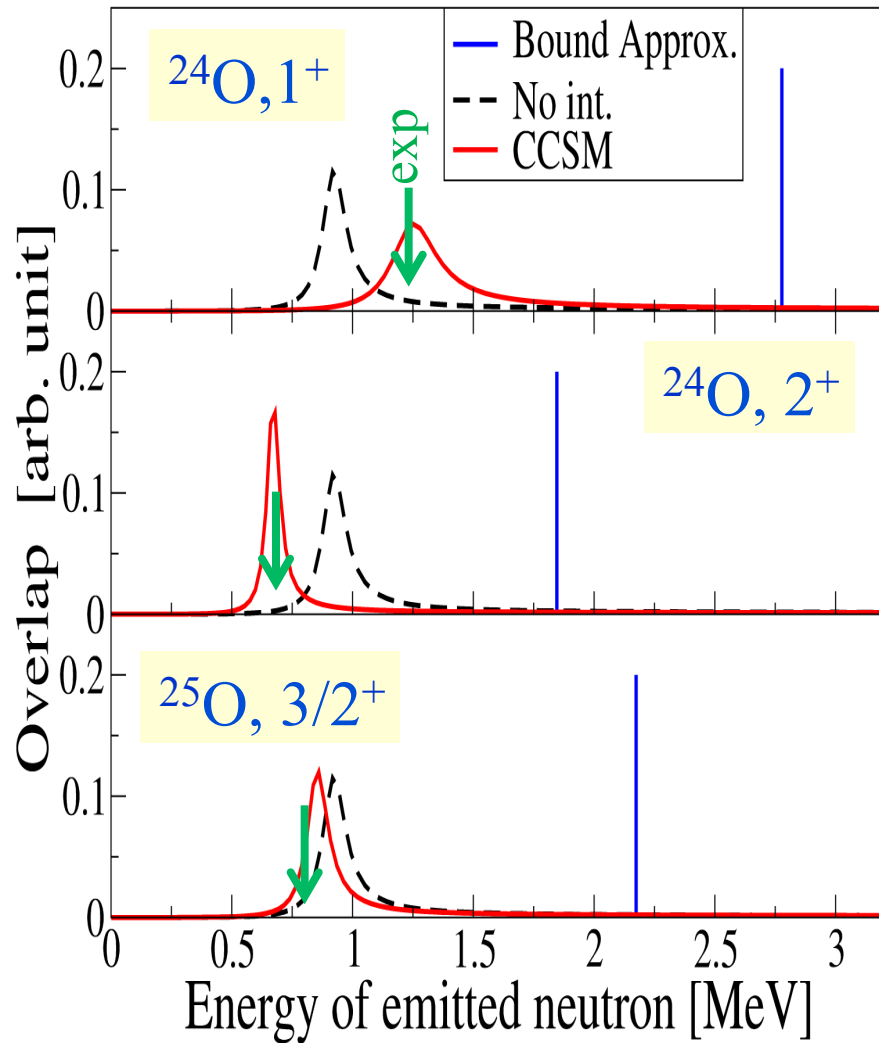
Doorway state \rightarrow doorway-state resonance



E1 excitation

nucleon transfer
(proton removal)

Low-lying Continuum Spectra in $^{24,25}\text{O}$



doorway-state resonance

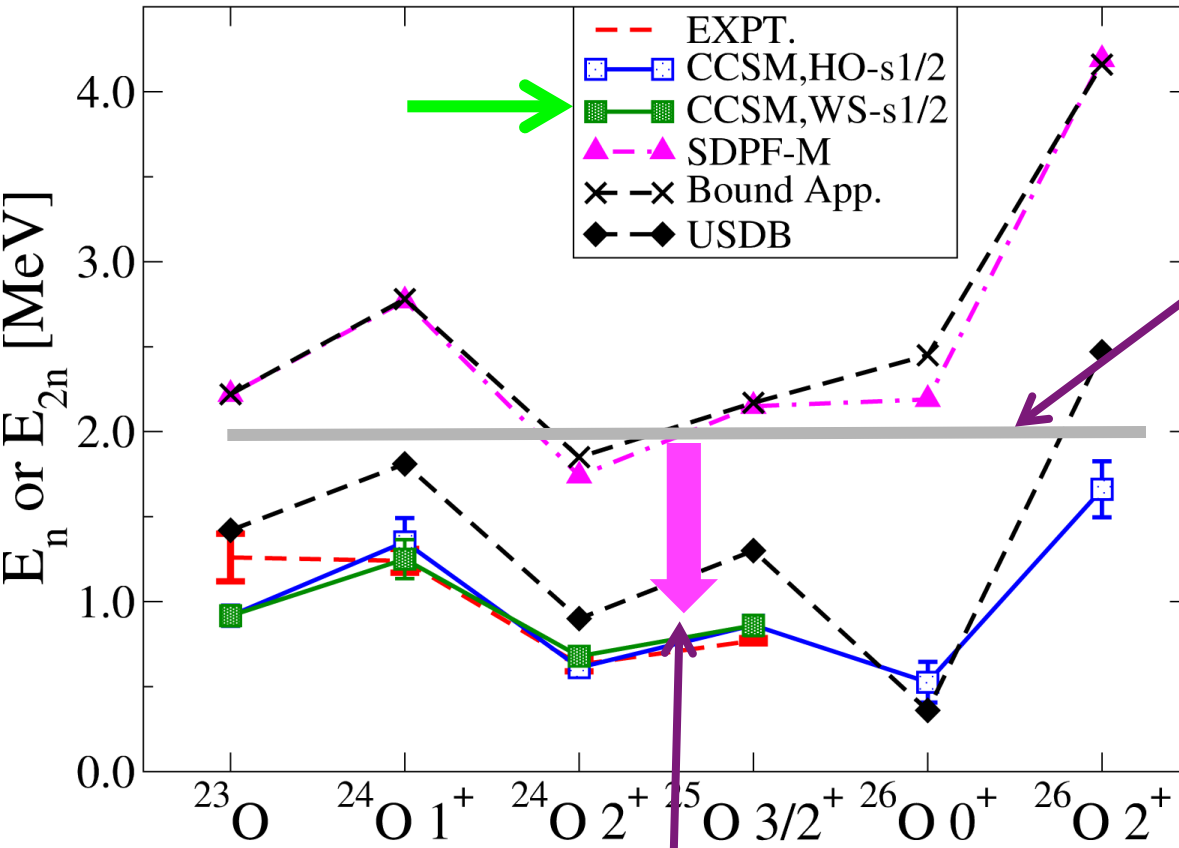
Doorway state \Rightarrow continuum states in ^{24}O

$$p_k^J = |\langle J_k^+ | \Phi_{\text{doorway}} \rangle|^2 = \left| \sum_i C_i^{(k)} \langle id_{3/2} | 0d_{3/2} \rangle \right|^2$$

- **bound approximation:**
Normal shell model with the same Hamiltonian : NO continuum effect
- **CCSM :** With continuum effect
incl. residual interaction
- **no int. :** With continuum effect but
no residual interaction.

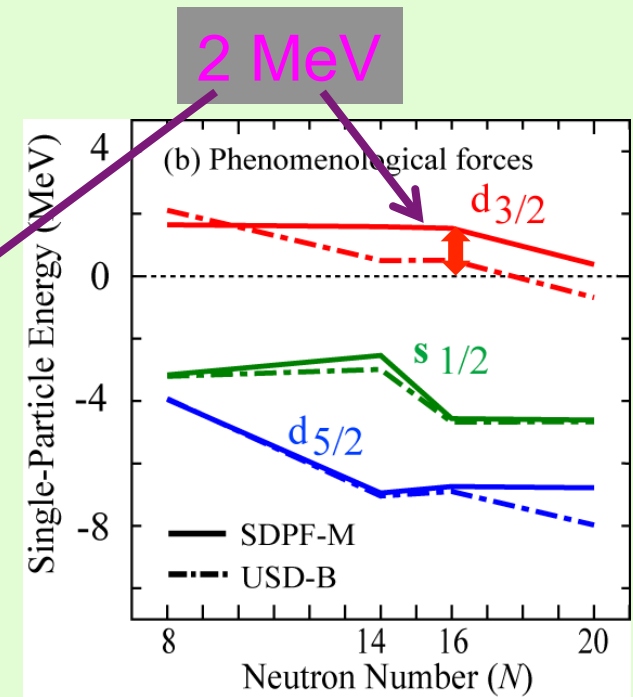
- Continuum effect is about 1 MeV
- No bound excited state.
- $1^+ - 2^+$ splitting by 2-body interaction
- $1^+ - 2^+$ splitting is in good agreement with experiments.

Peak Energies of neutron emission



Lowering due to continuum effect

SPE as bound state



Exp. : MSU (Hoffman et al),
RIKEN (Elekes et al)

Continuum spectra are consistent with the shell evolution

Oxygen isotopes



Fluorine isotopes

Neutron single-particle energies at $N=20$ for $Z=8\sim 20$

solid line : full (central + tensor)

dashed line : central only

A proton in $d_{5/2}$ moves
neutron orbits by

$d_{3/2}$ -2.0 MeV

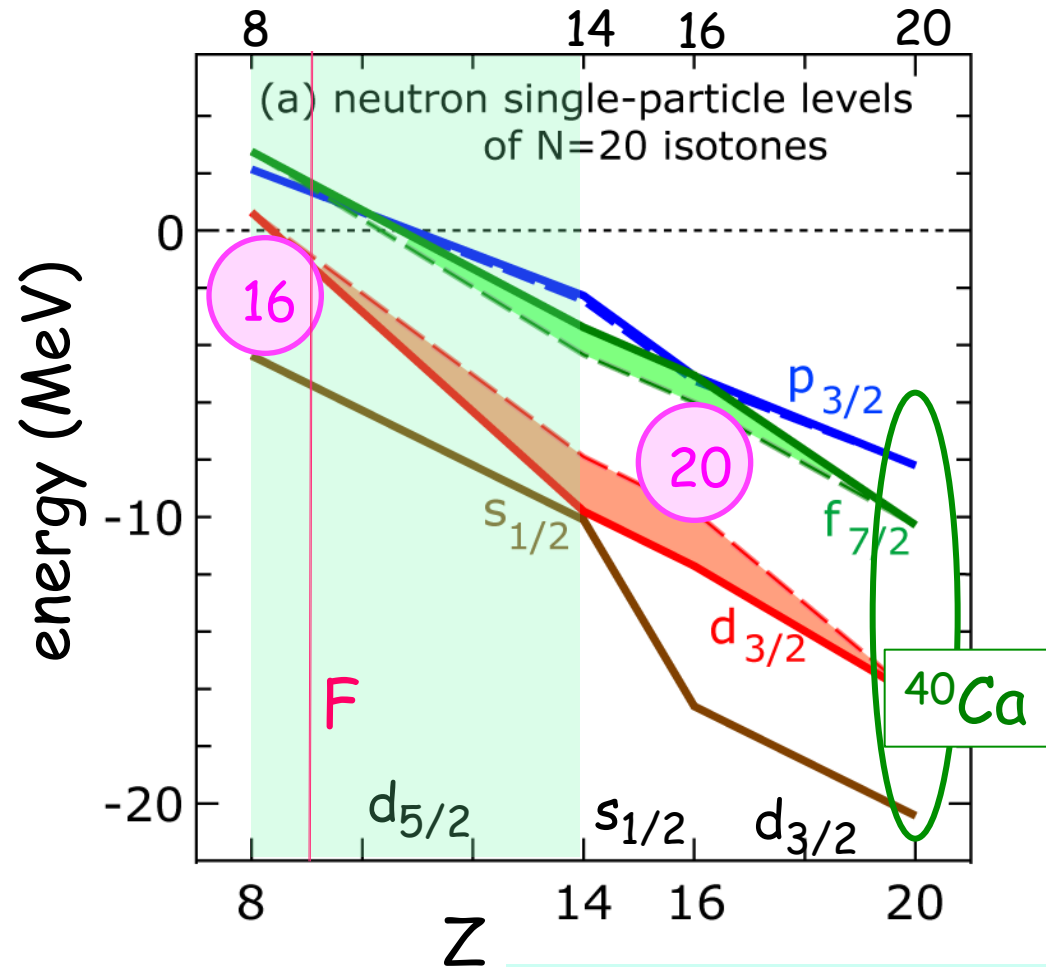
$s_{1/2}$ -1.1 MeV

$d_{5/2}$ -1.6 MeV



^{29}F well bound already
by s. p. e.

$^{31,\dots}\text{F}$ bound through
mixing with pf shell



TO, Suzuki, et al.
PRL 104, 012501 (2010)

Outline

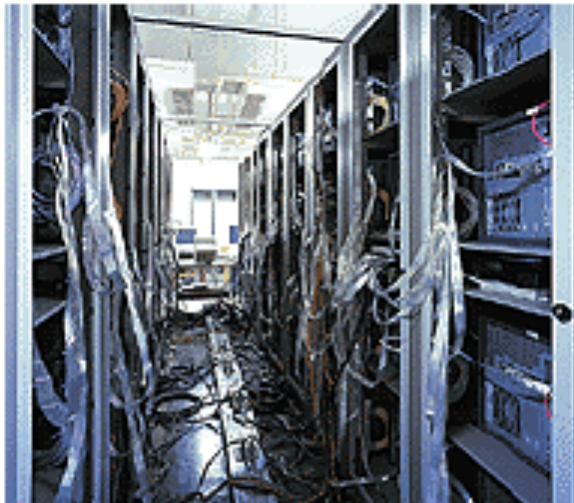
1. Introduction
2. Shell model and monopole interaction
3. Shell evolution and tensor force
4. Multiple quantum liquid in exotic nuclei
5. Shell evolution and three-nucleon force
6. Monte Carlo Shell Model
7. Summary



理化学研究所



クラスタ化した70台 (140CPU)のAlphaServerとCompaq Tru64 UNIXが高い並列処理パフォーマンスを実現し、ピクバン以降の原子核の構造解明を推進



70台のAlphaServer DS20をMyrinetで高速に並列結合し、最大のパラレル処理パフォーマンスを発揮するAlphleet

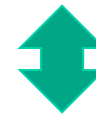
Major outcome

PHYSICAL REVIEW C, VOLUME 60, 054315

1999

Varying shell gap and deformation in $N \sim 20$ unstable nuclei studied by the Monte Carlo shell model

Yutaka Utsuno,^{1,2} Takaharu Otsuka,^{1,2} Takahiro Mizusaki,¹ and Michio Honma³
Computations have been carried out partly by the Alphleet computer system of RIKEN.



Physics Letters B 346 (1995) 9–14

Large deformation of the very neutron-rich nucleus ^{32}Mg from intermediate-energy Coulomb excitation

T. Motobayashi^{a,1}, Y. Ikeda^a, Y. Ando^a, K. Ieki^a, M. Inoue^a, N. Iwasa^a, T. Kikuchi^a, M. Kurokawa^a, S. Moriya^a, S. Ogawa^a, H. Murakami^a, S. Shimoura^a, Y. Yanagisawa^a, T. Nakamura^b, Y. Watanabe^b, M. Ishihara^{b,c}, T. Teranishi^c, H. Okuno^c, R.F. Casten^d

Present status

Advanced
MCSM

PTEP

Prog. Theor. Exp. Phys. 2012, 01A205 (27 pages)
DOI: 10.1093/ptep/pts012

New-generation Monte Carlo shell model for the K computer era

Noritaka Shimizu,^{1,*} Takashi Abe¹, Yusuke Tsunoda², Yutaka Utsuno³, Tooru Yoshida¹, Takahiro Mizusaki⁴, Michio Honma⁵, and Takaharu Otsuka^{1,2,6}

Project

HPCI Strategic Programs for Innovative Research (SPIRE)
Field 5 “The origin of matter and the universe”

Computer



Large-scale
calculations



Outline

- **Methodology**: advanced Monte Carlo shell model (MCSM)
 - intrinsic shape can be the objectives
 - ab initio (no core) MCSM and clustering in Be isotopes
- **Shape coexistence** and Quantum Liquid picture
 - exotic Ni isotopes (+ Co, Cu)
- **Shape evolution** (from seniority to rotor)
 - Xe and Ba isotopes
- **Extension**: Spectra to high-lying collective states
 - E1 excitation, GDR and PDR in Ca and Sr isotopes
 - Level density with shell model Hamiltonian

Advanced Monte Carlo shell model (MCSM)

Superposition of the projected Slater determinants
+ Extrapolation by energy variance

$$|\Psi\rangle = \sum_{k=1}^{N_{MCSM}} f_k P^{J,\pi} |\phi_k\rangle \quad |\phi_k\rangle = \prod_{\alpha=1}^N \left(\sum_{i=1}^{N_{sp}} c_i^\dagger D_{i\alpha}^{(k)} \right) |-\rangle$$

MCSM basis, deformed Slater det.

>10¹⁰ basis vectors

$$\mathbf{H} = \begin{pmatrix} * & * & * & * & * & \dots \\ * & * & * & * & * & \dots \\ * & * & * & * & * & \dots \\ * & * & \dots & \dots & \dots & \dots \\ \vdots & \dots & \dots & \dots & \dots & \dots \end{pmatrix} \xrightarrow{\text{diagonalization}} \begin{pmatrix} \varepsilon_1 & & & & & 0 \\ & \varepsilon_2 & & & & \\ & & \varepsilon_3 & & & \\ & & & \ddots & & \\ 0 & & & & \ddots & \end{pmatrix}$$

Conventional Shell Model
all Slater determinants

$$\mathbf{H} \approx \begin{pmatrix} * & * & * & \cdot \\ * & * & * & \cdot \\ * & * & \cdot & \\ \cdot & \cdot & & \end{pmatrix} \xrightarrow{\text{diagonalization}} \begin{pmatrix} \varepsilon'_1 & & 0 \\ & \varepsilon'_2 & \\ 0 & & \ddots \end{pmatrix}$$

Monte Carlo Shell Model
bases important for a specific eigenstate

~10² basis vectors

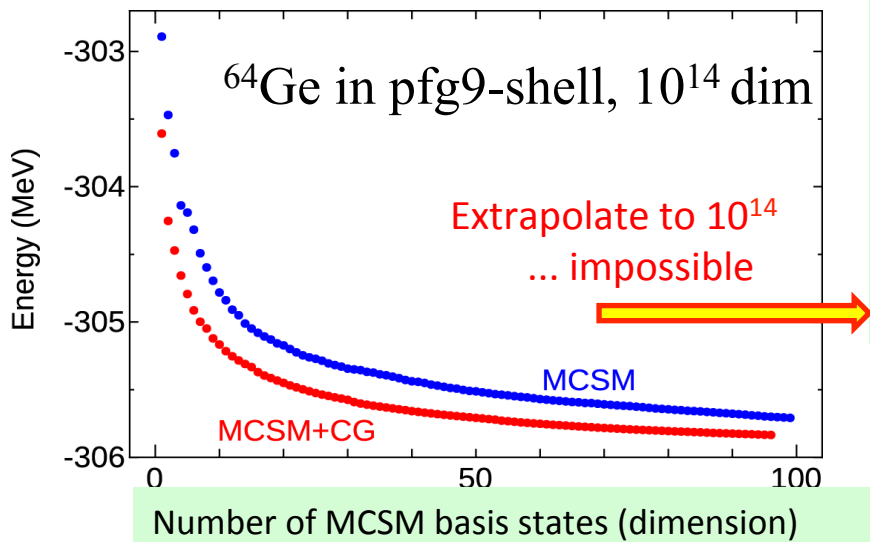
Step 1

- **Select** D stochastically from many candidates generated by the auxiliary-field Monte Carlo technique

Step 2

- **optimize** D variationally by the conjugate gradient method so as to minimize the energy eigenvalue of this small matrix

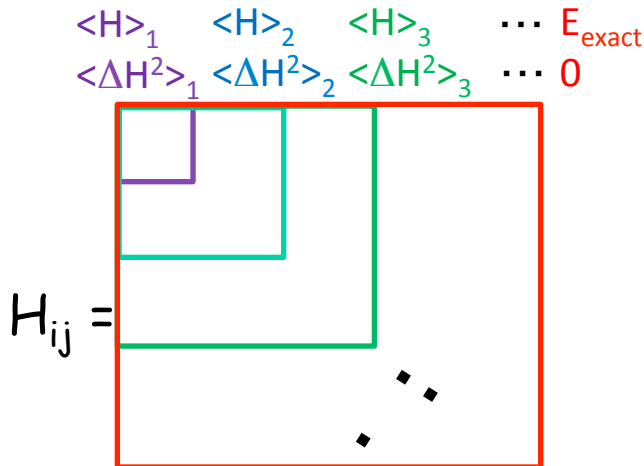
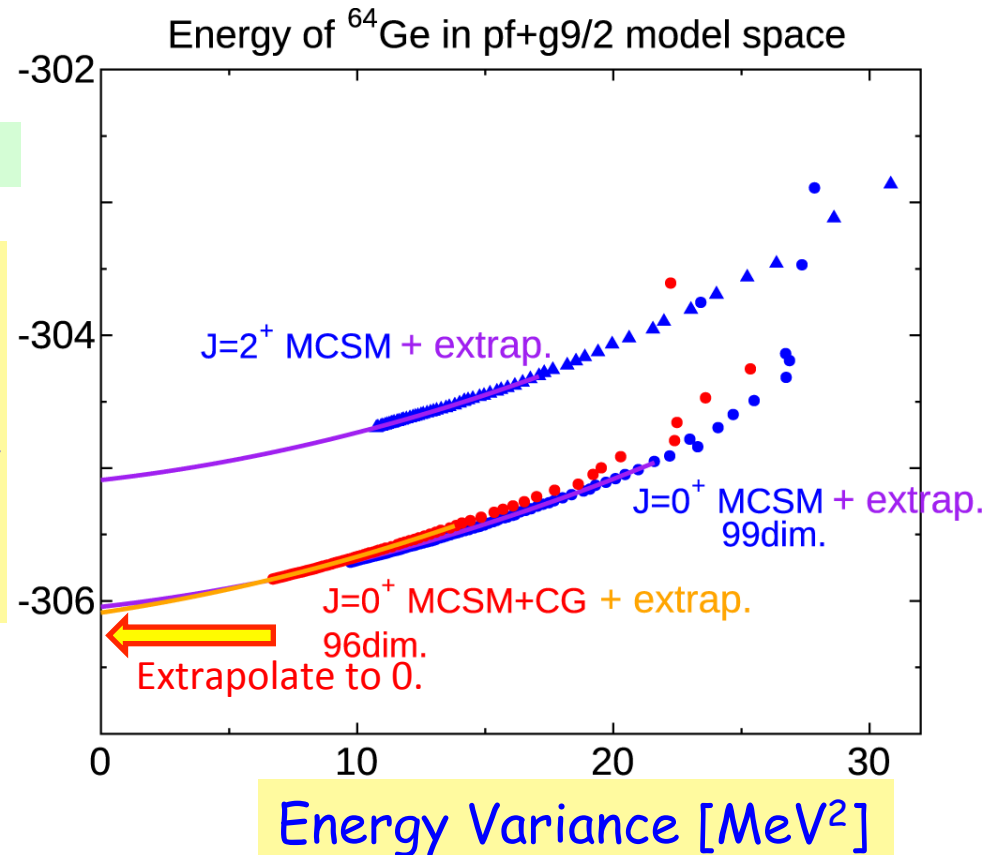
Step 3: Energy variance extrapolation



$$\text{Energy variance: } \langle \Delta H^2 \rangle = \langle H^2 \rangle - \langle H \rangle^2$$

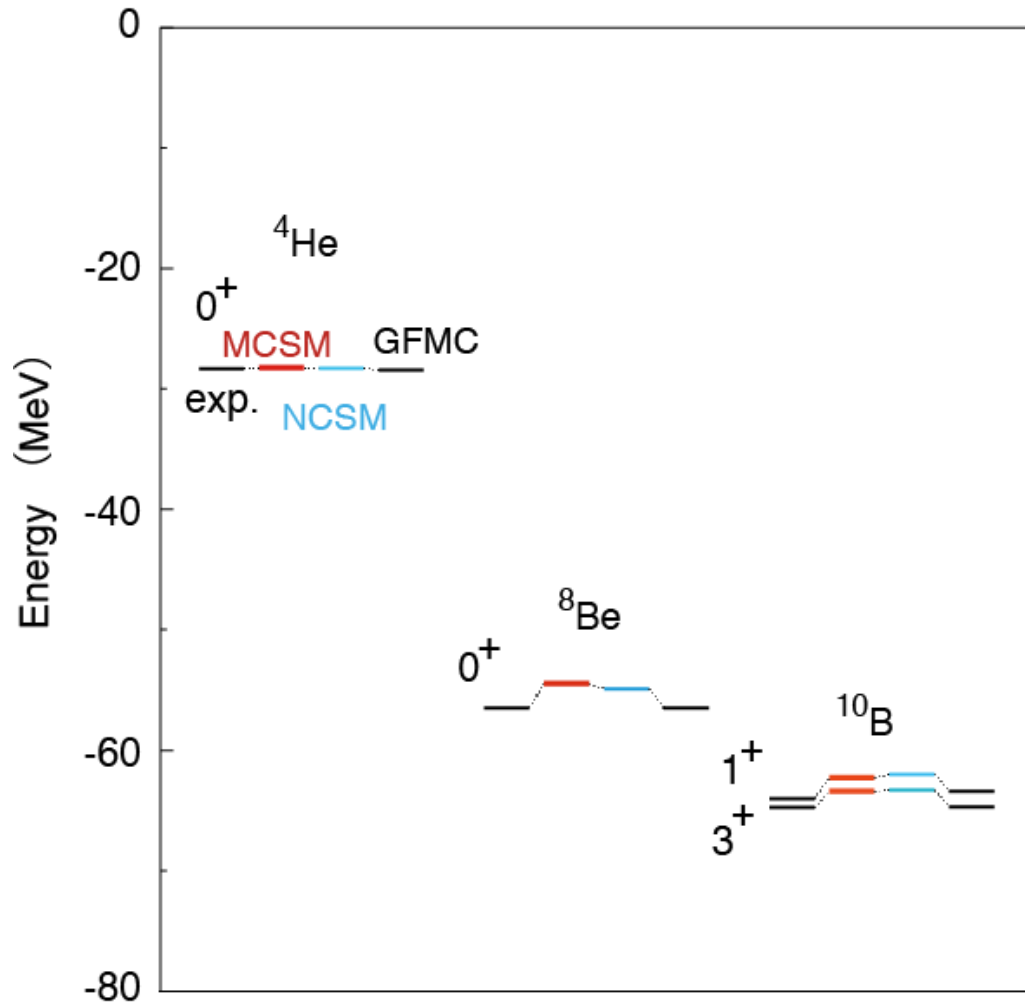
As the number of basis vectors increases, the approximated w.f. approaches the exact one and the energy variance approaches zero.

$$\text{Extrapolate towards } \langle \Delta H^2 \rangle \rightarrow 0$$



Systematic calculations in terms of
ab initio Monte Carlo Shell Model
with JISP-16 interaction

*Extrapolation to infinite
model space is included.*



MCSM: same as present

NCSM:

No-core shell model
with JISP-16

T. Abe, P. Maris, et al.
PRC 86, 054301 (2012)

GFMC

AV18 + IL7

J. Carlson, et al.,
arXiv:1412.3081 (2014).

Ab initio (no-core) MCSM : Be isotopes

T. Yoshida, T. Abe *et al.*

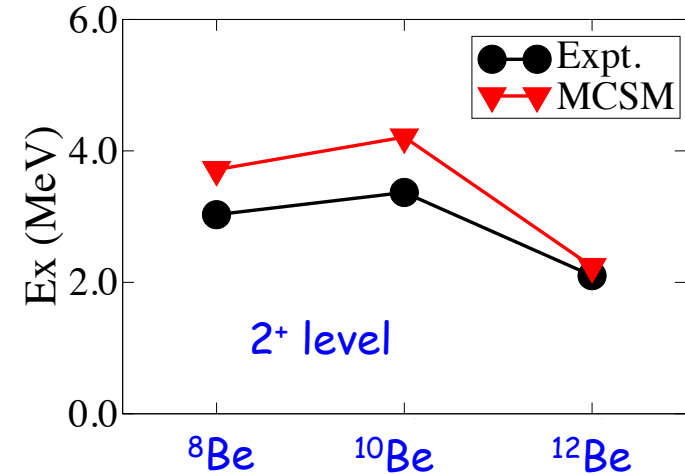
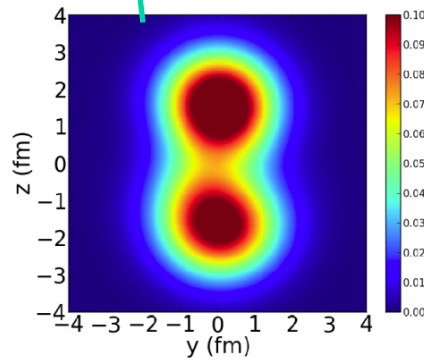
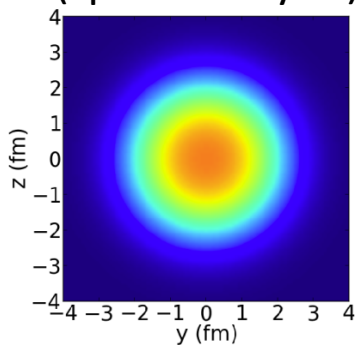
JISP-16 interaction used

Matter density of the ground state $J^\pi = 0^+$ of ${}^8\text{Be}$

$$|\Psi\rangle = \sum_{k=1}^{N_{\text{MCSM}}} f_k P^{J,\pi} |\phi_k\rangle = P^{J,\pi} \left(\sum_{k=1}^{N_{\text{MCSM}}} f_k |\phi_k\rangle \right)$$

laboratory frame
(spherical sym.)

Intrinsic frame



Matter density of Be isotopes

Cluster structure emerges spontaneously in ${}^8\text{Be}$, and fades away as N increases

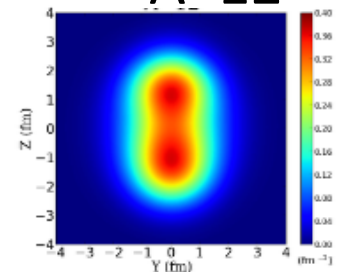
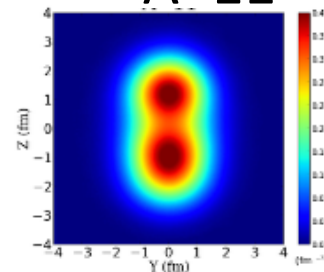
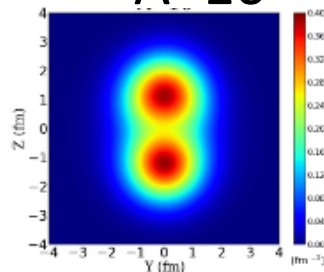
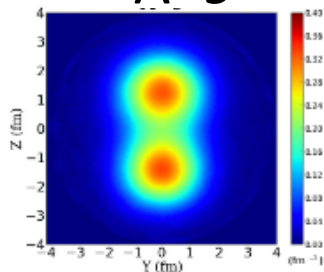
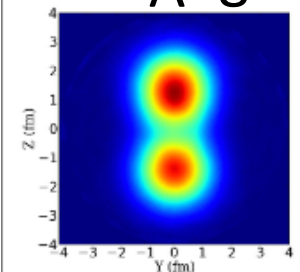
A=8

A=9

A=10

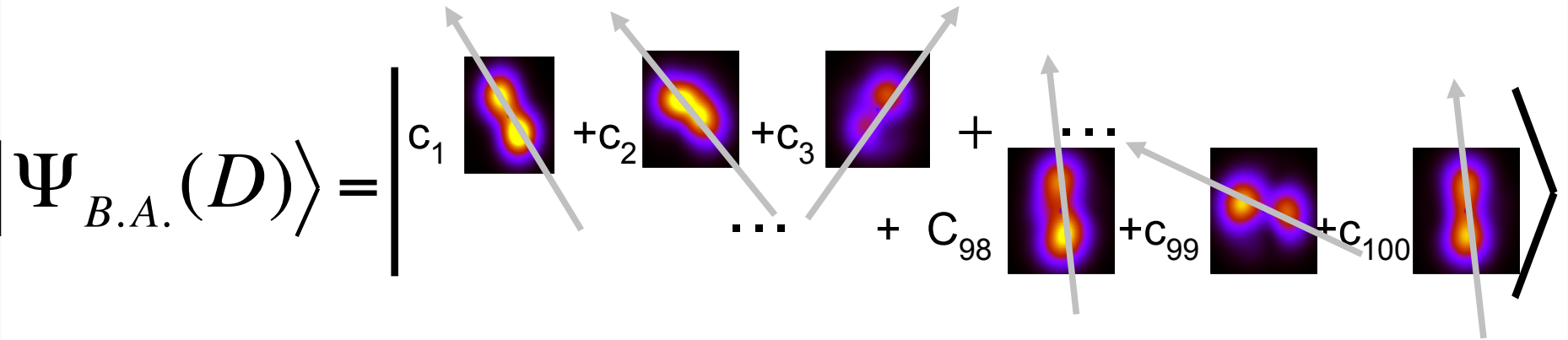
A=11

A=12

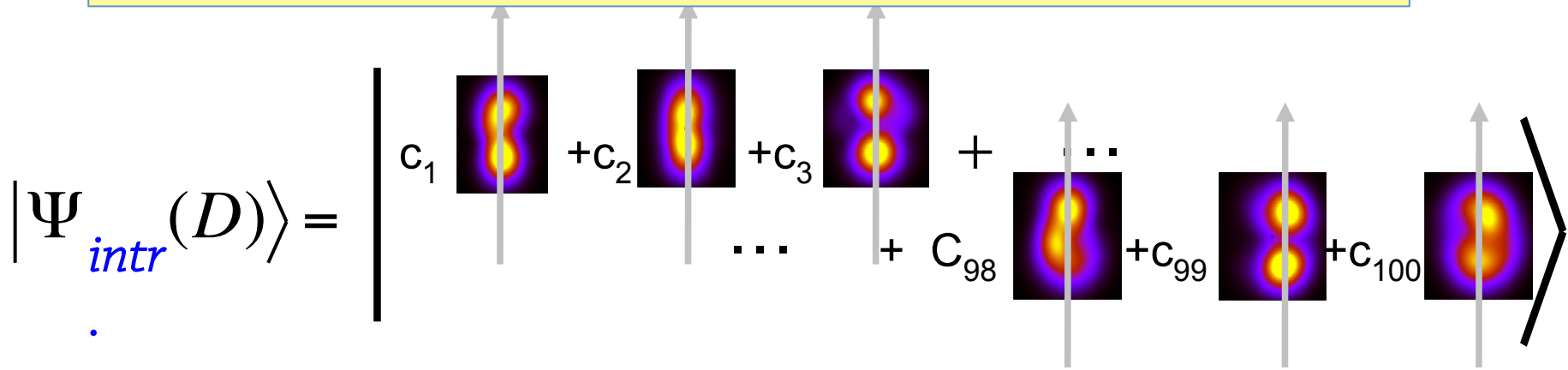


Alignment of each basis state (Slater determinant)

Before the alignment; orientations are random



All basis states are aligned → “intrinsic state”

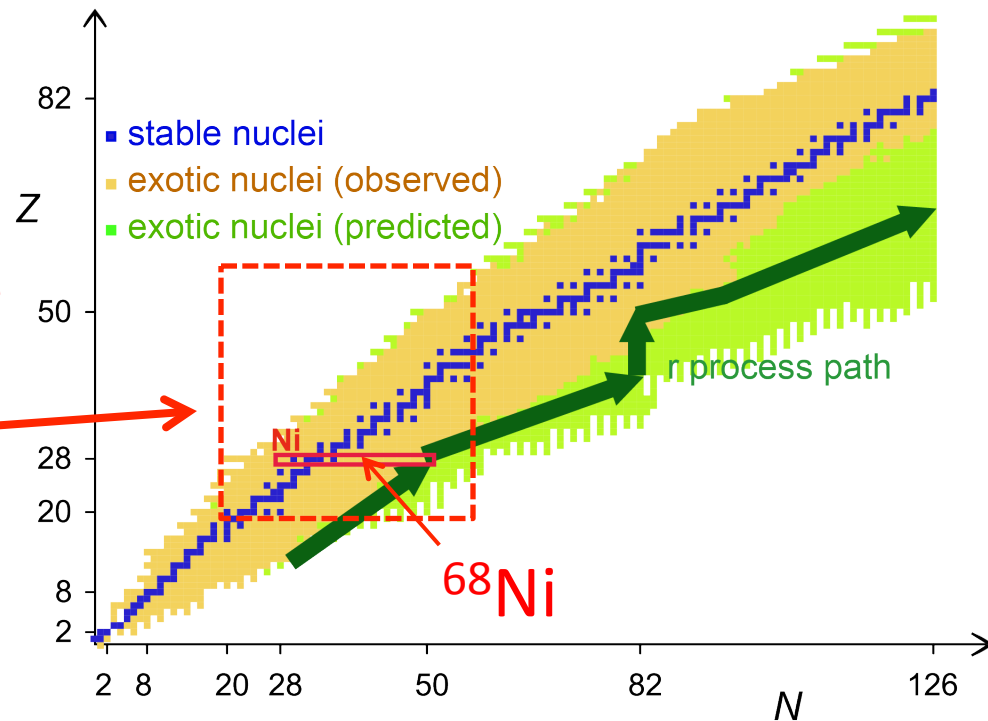
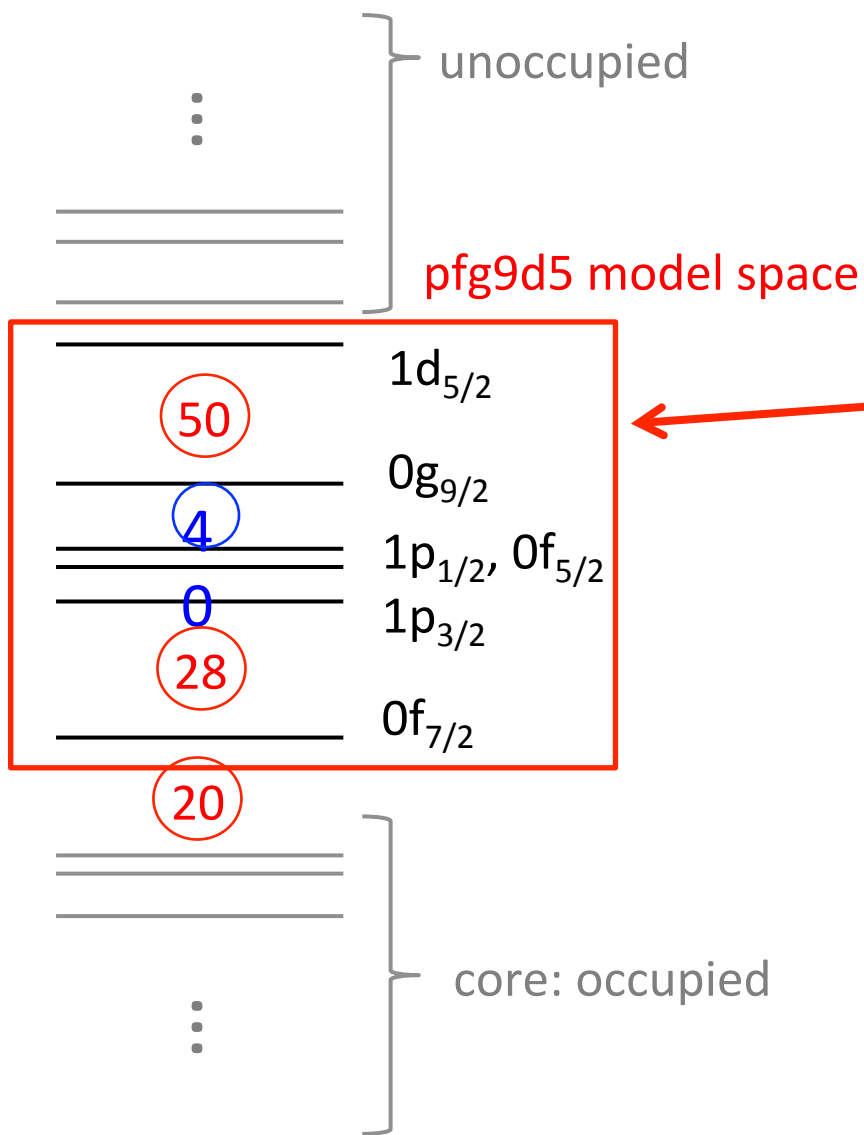


Outline

- **Methodology**: advanced Monte Carlo shell model (MCSM)
 - intrinsic shape can be the objectives
 - ab initio (no core) MCSM and clustering in Be isotopes
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 - exotic Ni isotopes (+ Co, Cu)
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 - Level density with shell model Hamiltonian

MCSM calculation on Ni isotopes

Y. Tsunoda *et al*



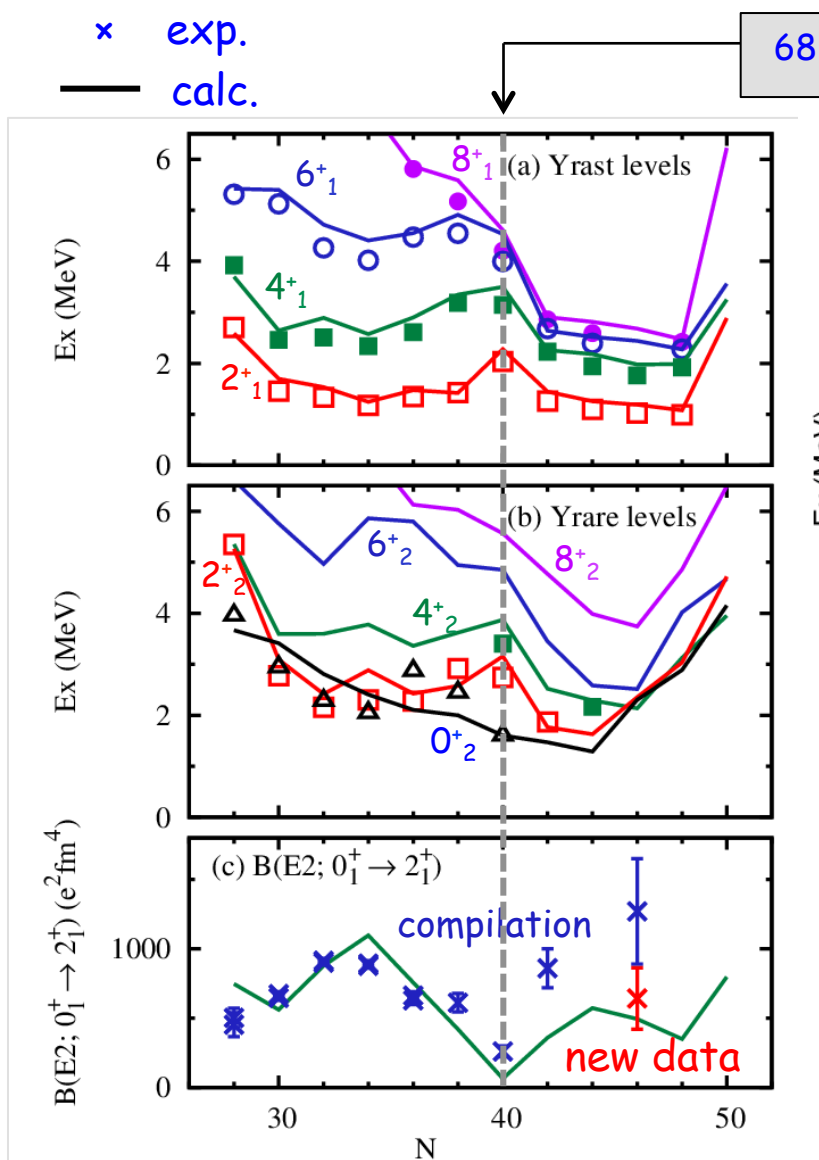
This model space is wide enough to discuss how **magic numbers 28, 50** and **semi-magic number 40** are visible or smeared out.

Interaction:
A3DA interaction is used with minor corrections

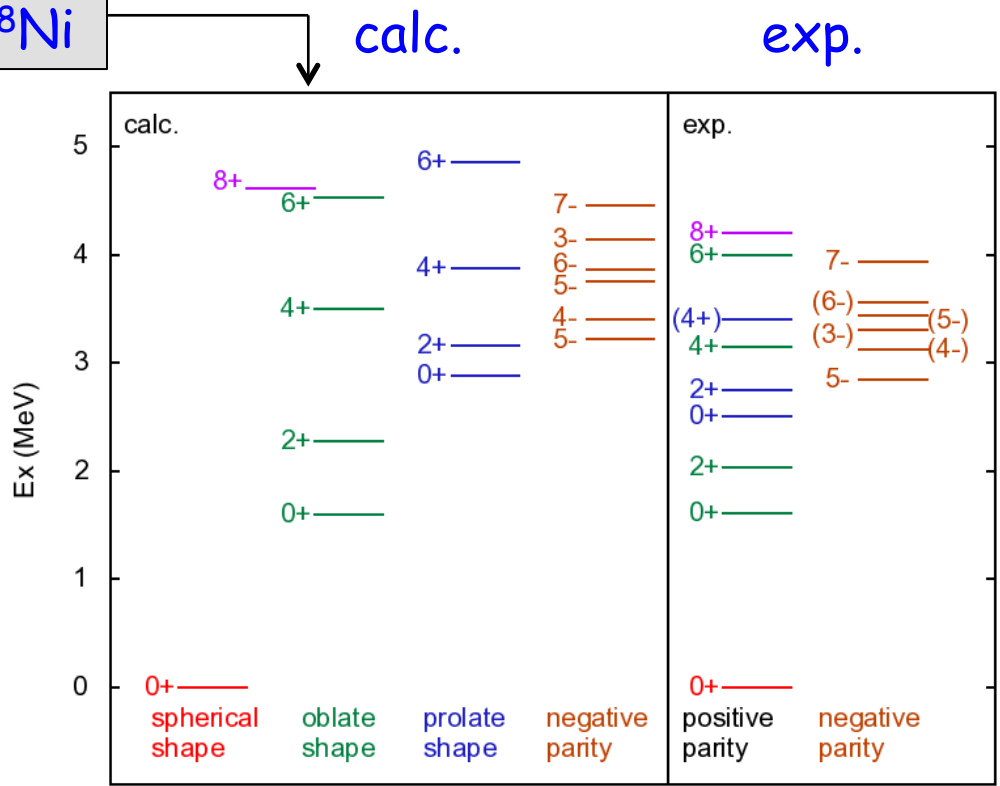
Energy levels and B(E2) values of Ni isotopes

Description by the same Hamiltonian

Shape coexistence in ^{68}Ni



^{68}Ni

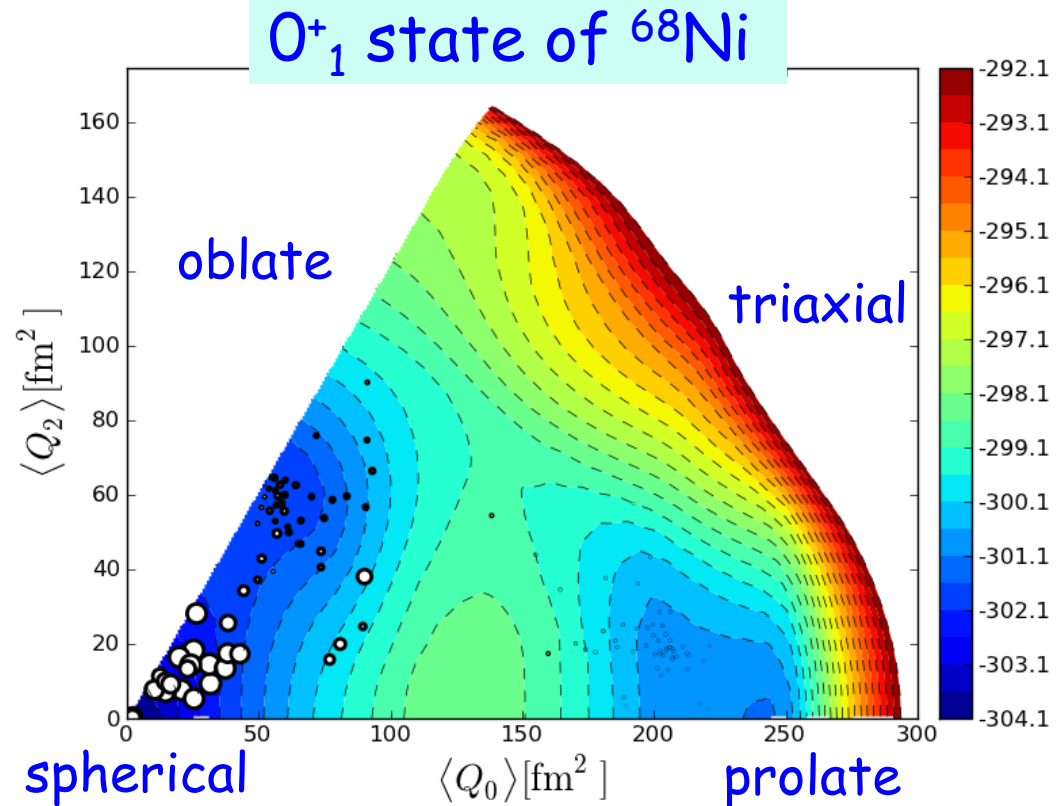


Y. Tsunoda, TO, Shimizu, Honma and Utsuno, PRC 89, 031301 (R) (2014)

MCSM basis vectors on Potential Energy Surface

eigenstate $\Psi = \sum_i c_i P[J^\pi] \Phi_i$ ← Slater determinant → intrinsic shape

- **PES** is calculated by CHF for the shell-model Hamiltonian
- **Location of circle** : quadrupole deformation of unprojected MCSM basis vectors
- **Area of circle** : overlap probability between each projected basis and eigen wave function



Called ***T-plot*** in reference to

Y. Tsunoda, TO, Shimizu, Honma and Utsuno,
PRC 89, 031301 (R) (2014)

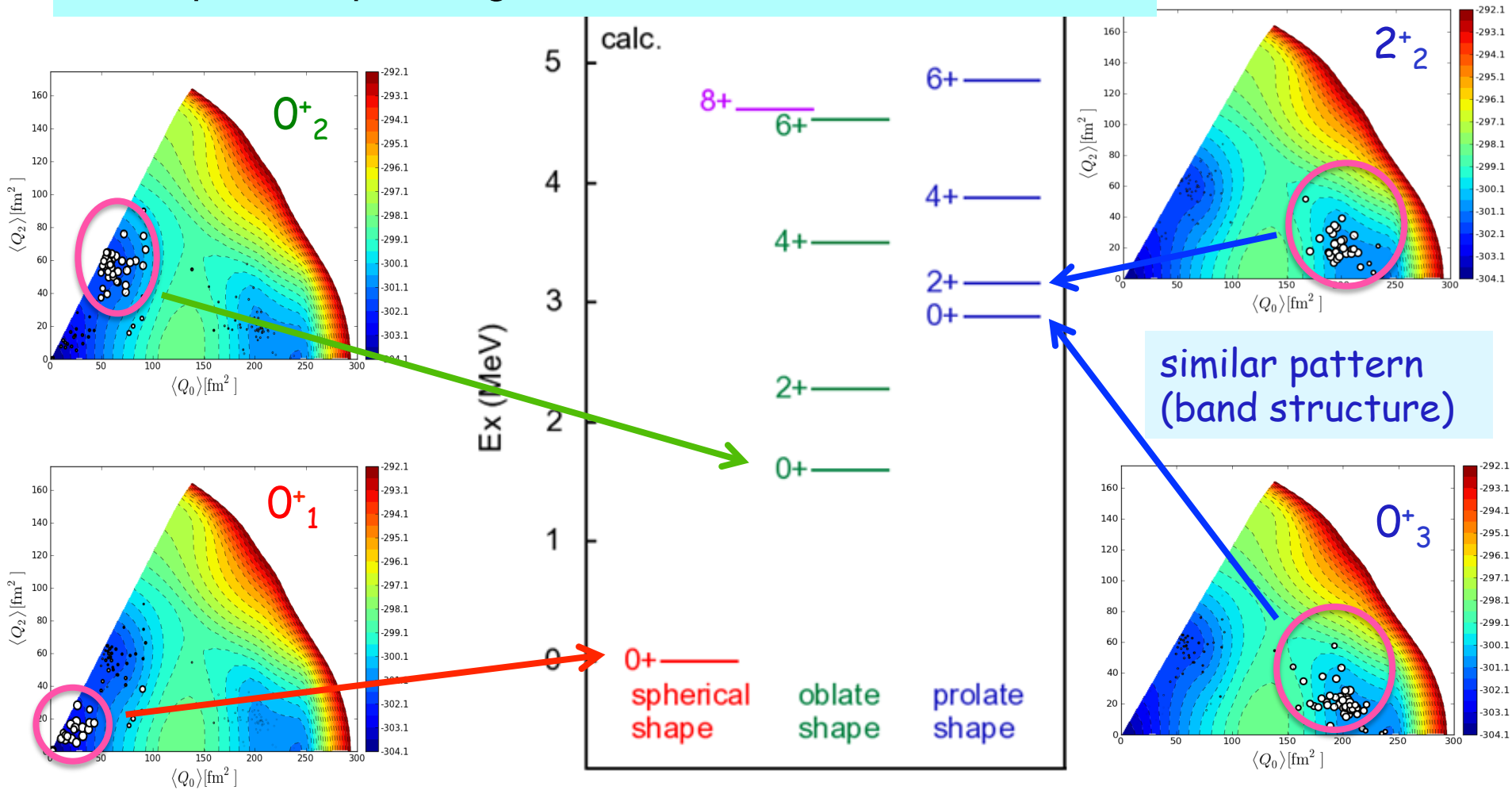
General properties of T-plot :

Certain number of large circles in a small region of PES

⇔ pairing correlations

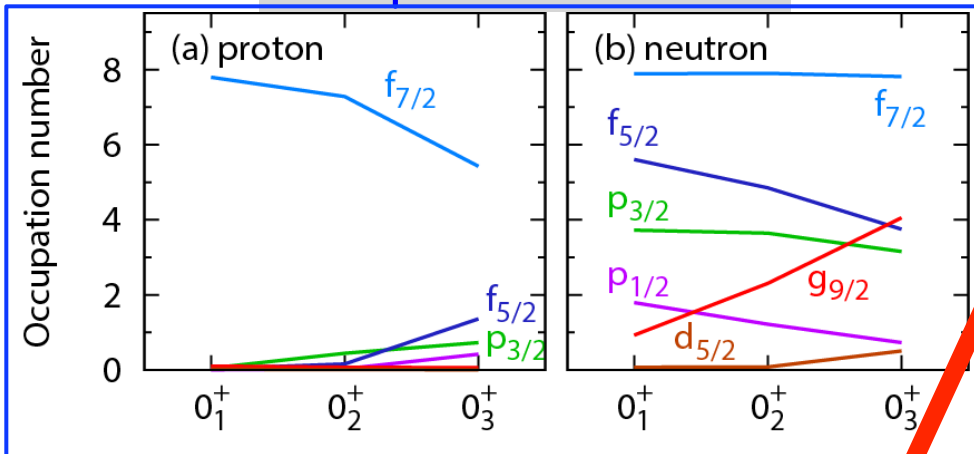
Spreading beyond this can be due to shape fluctuation

Example : shape assignment to various 0^+ states of ^{68}Ni



^{68}Ni 0^+ states

occupation numbers



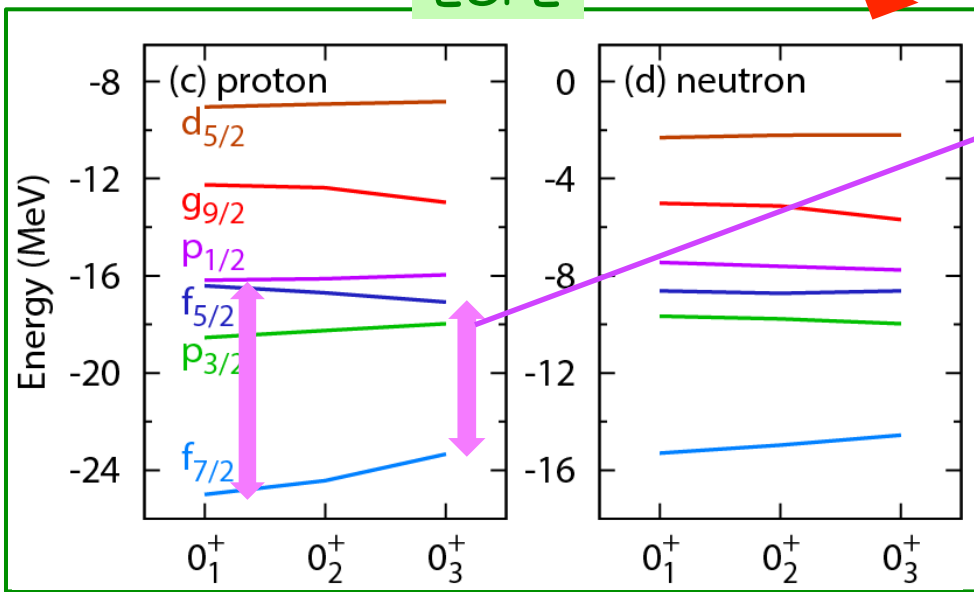
effective single-particle energies (ESPE) for correlated eigenstate

$$\epsilon_j = \left\langle \frac{\partial H_m}{\partial n_j} \right\rangle$$

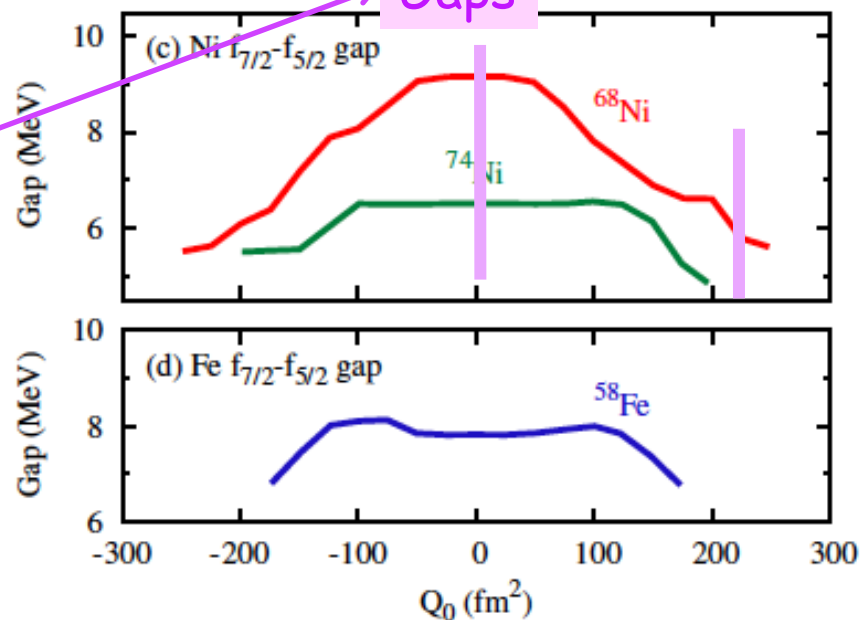
H_m monopole part of H

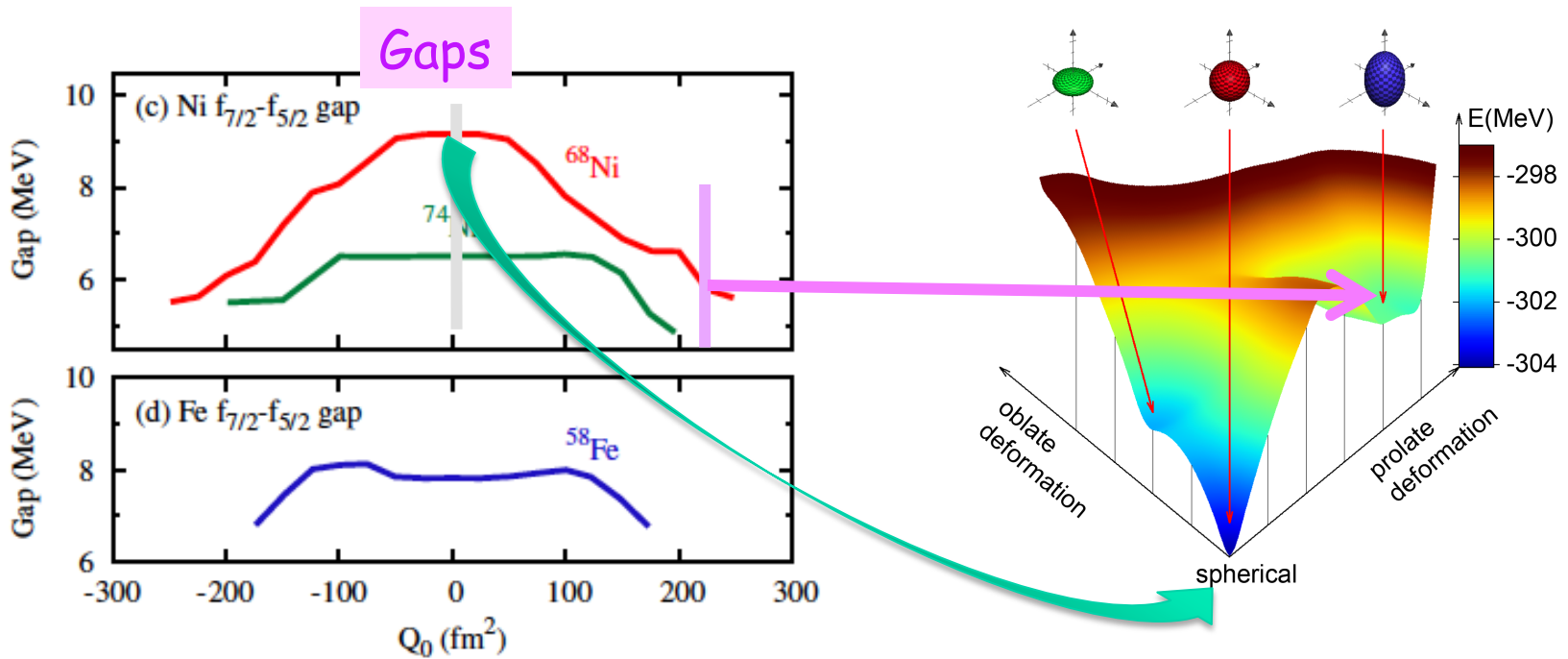
$\langle \rangle$: by actual occup. numbers

ESPE



Gaps



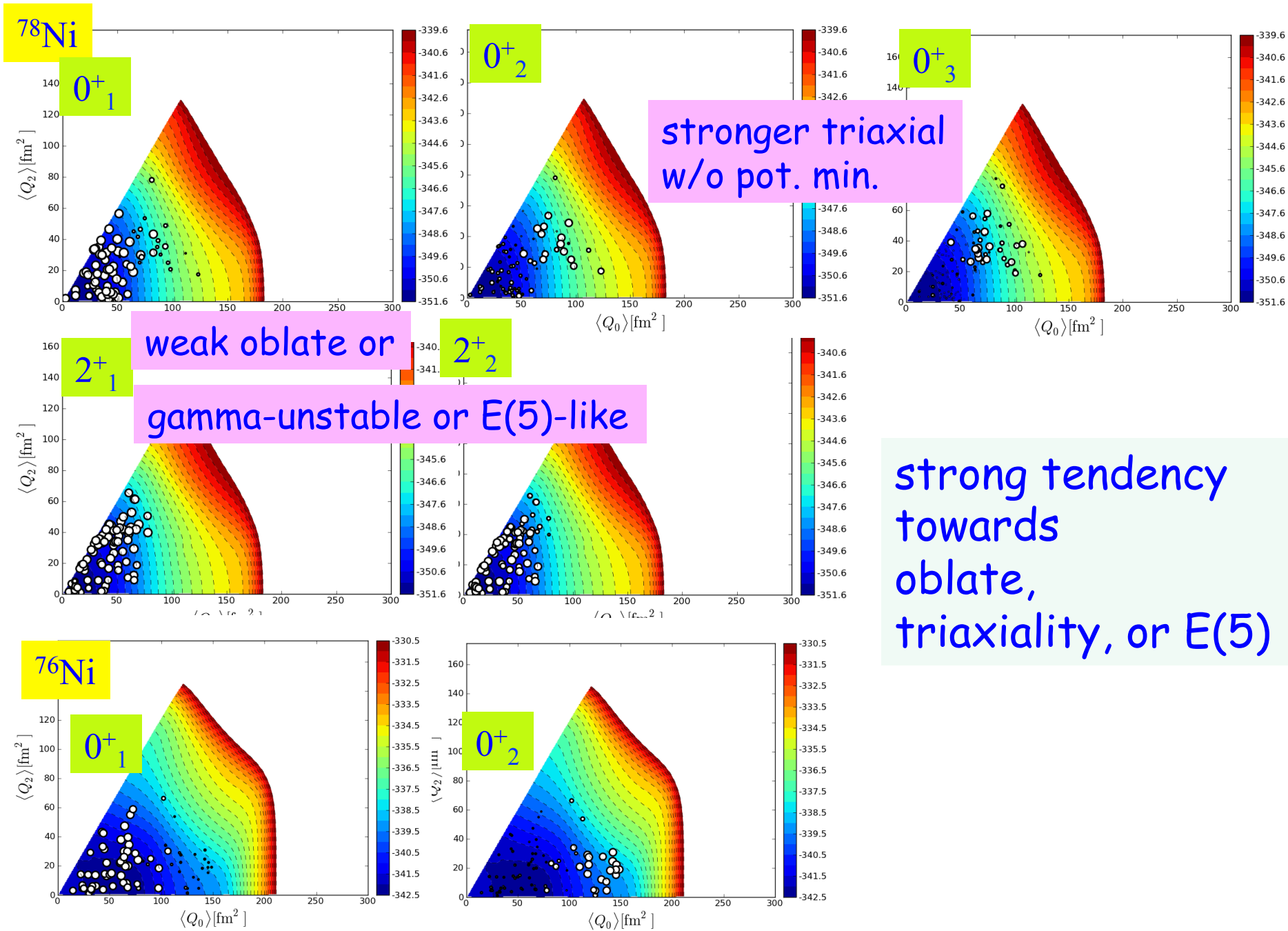


shell gap (spherical)
configurations = shapes

determined self-consistently
and non-linearly

Type II Shell Evolution
shell evolution within the same nucleus
driven by the tensor force

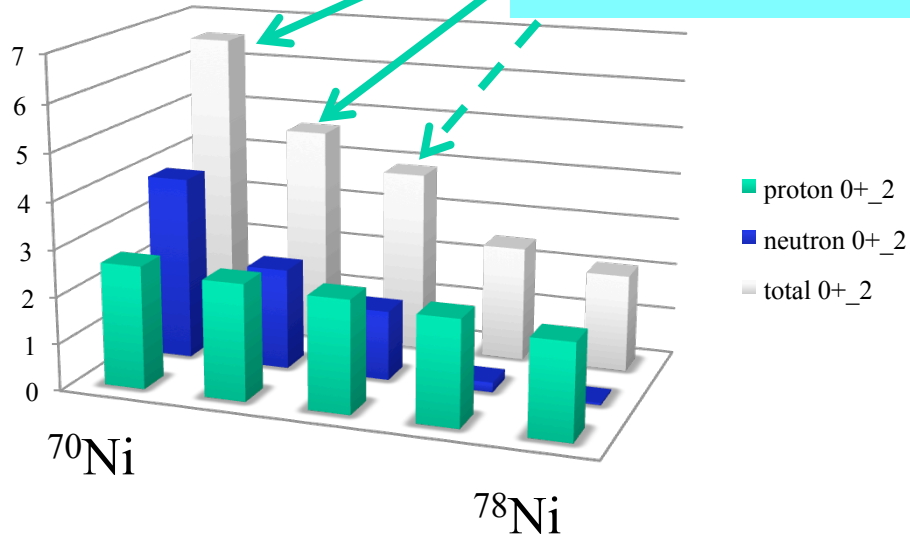
Spherical shell gap is not changed in Nilsson model



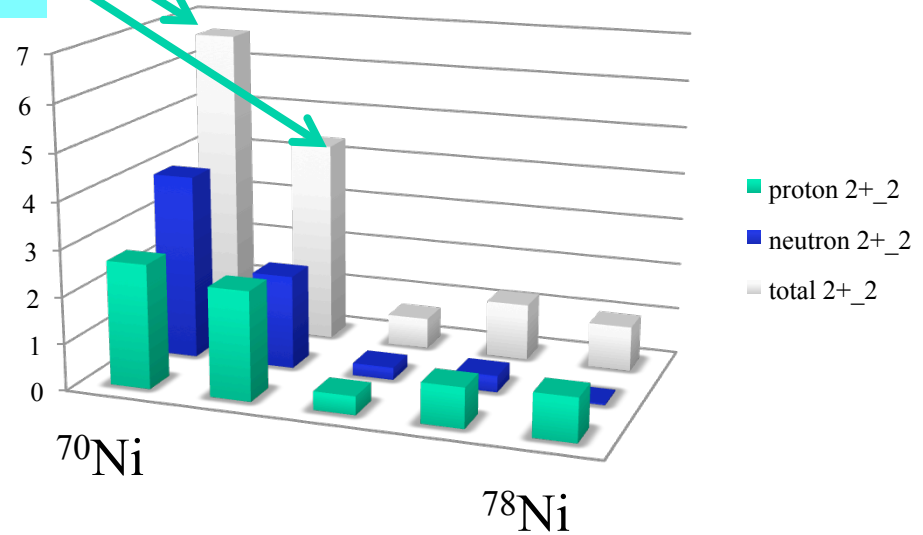
Number of protons and neutrons excited across Z=28 or N=40 magic numbers

0^+_2 state

strong prolate deformation



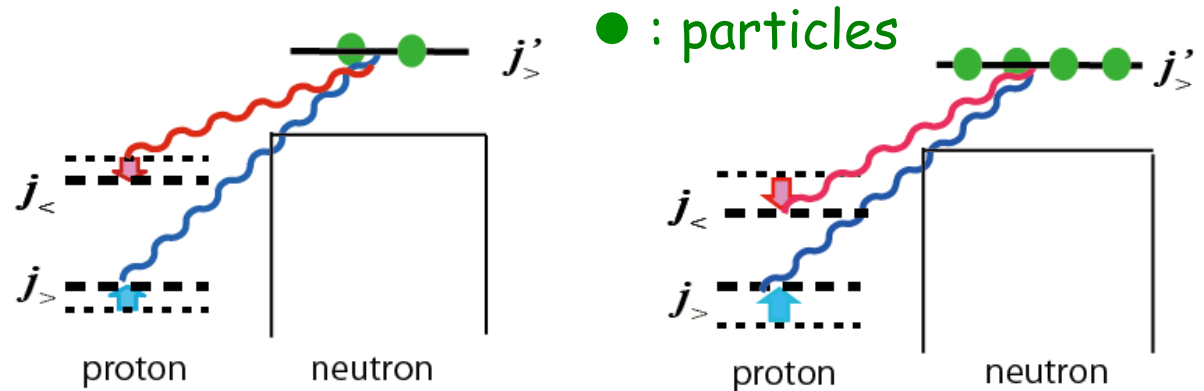
2^+_2 state



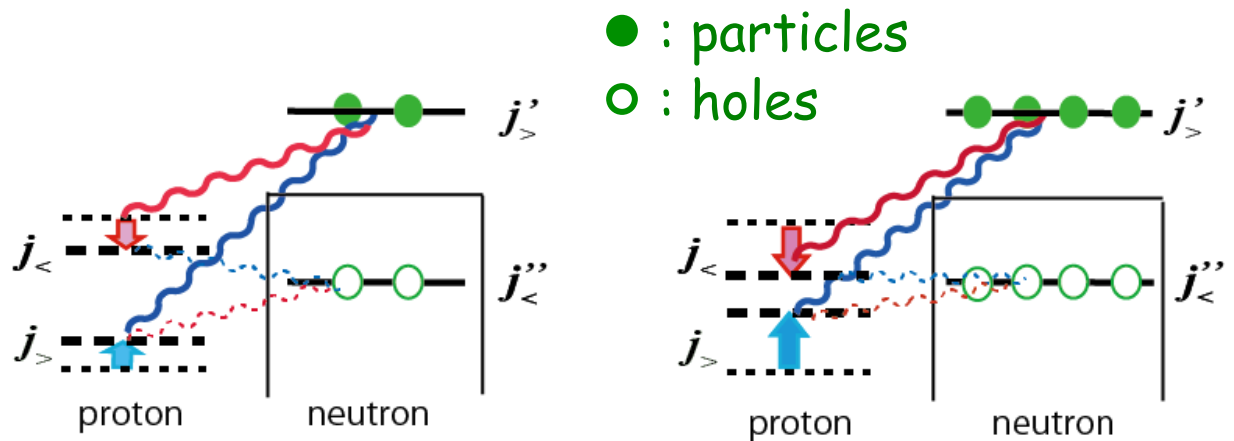
Underlying mechanism of the appearance of low-lying deformed states : Type II Shell Evolution

Monopole effects on the shell structure from the tensor interaction

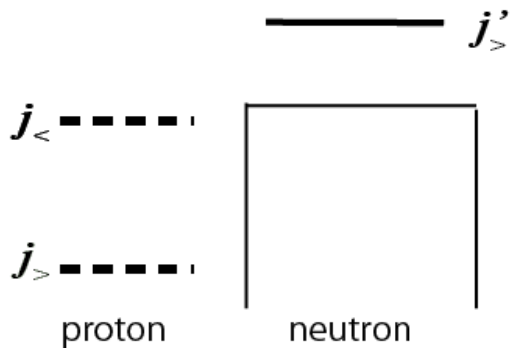
Type I Shell Evolution : different isotopes



Type II Shell Evolution : within the same nucleus



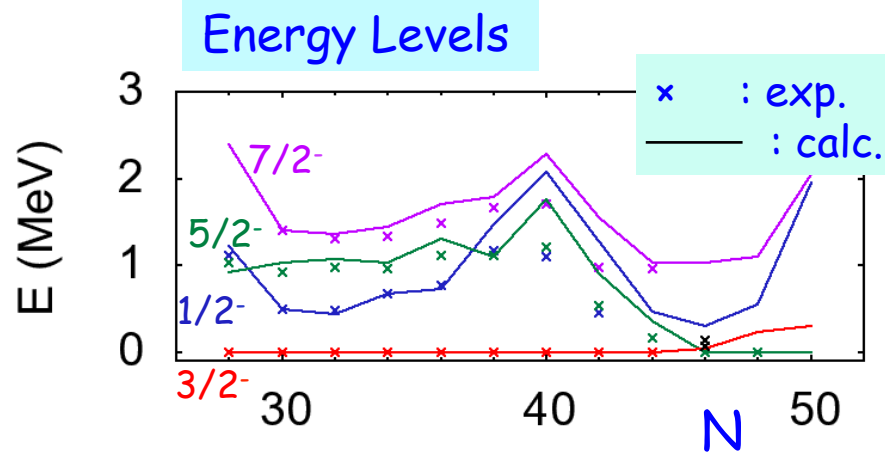
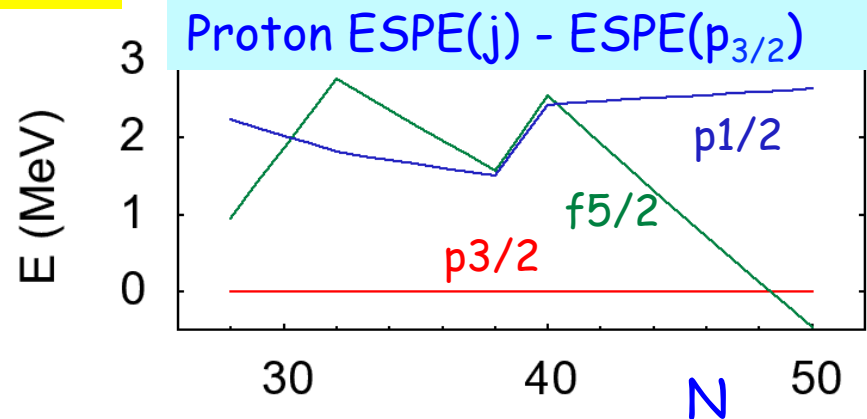
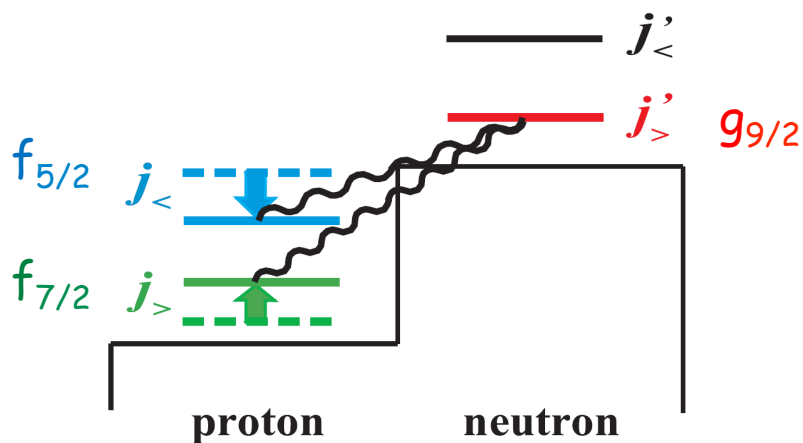
(a)



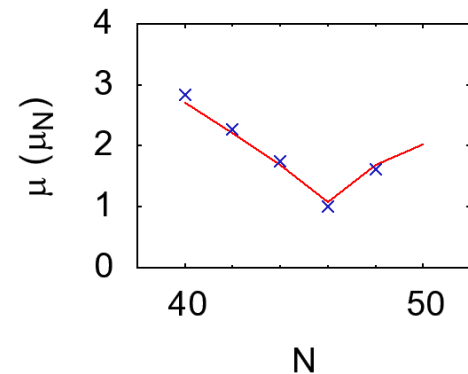
Cu isotopes

- proton $p_{3/2}$ - $f_{5/2}$ level crossing from $N = 40$ to $N = 50$ (type I shell evolution)
- Calculated states show agreement with experiments, although they are not pure single-particle states.

Monopole effect of tensor force



magnetic moment of ground state



Shape coexistence of ^{70}Co (Z=27, N=43)

g.s. and an isomer in ^{70}Co are known experimentally (PRC **61**, 054308 (2000))

High-spin state (6⁻, 7⁻)

$$\pi f_{7/2}^{-1} \nu g_{9/2}^{+3}$$

and **Low-spin state (3⁺)**

$$\pi f_{7/2}^{-1} \nu p_{1/2}^{-1} \nu g_{9/2}^{+4}$$

were suggested

From our calculations,

High-spin state (7⁻) is near-spherical

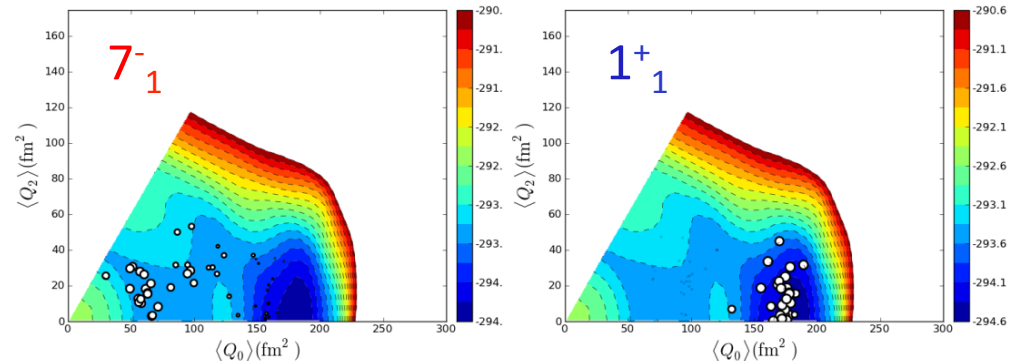
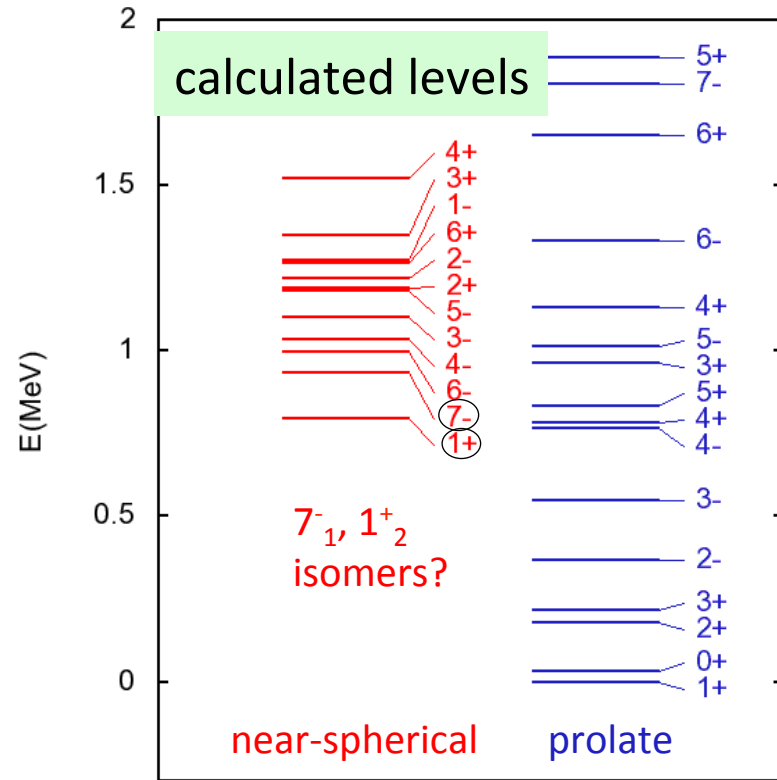
Low-spin state (1⁺) is prolate deformed

In the **prolate state 1⁺₁**,

many nucleons are excited

~2.7 protons above Z=28 gap

~3.1 neutron holes below N=40 gap



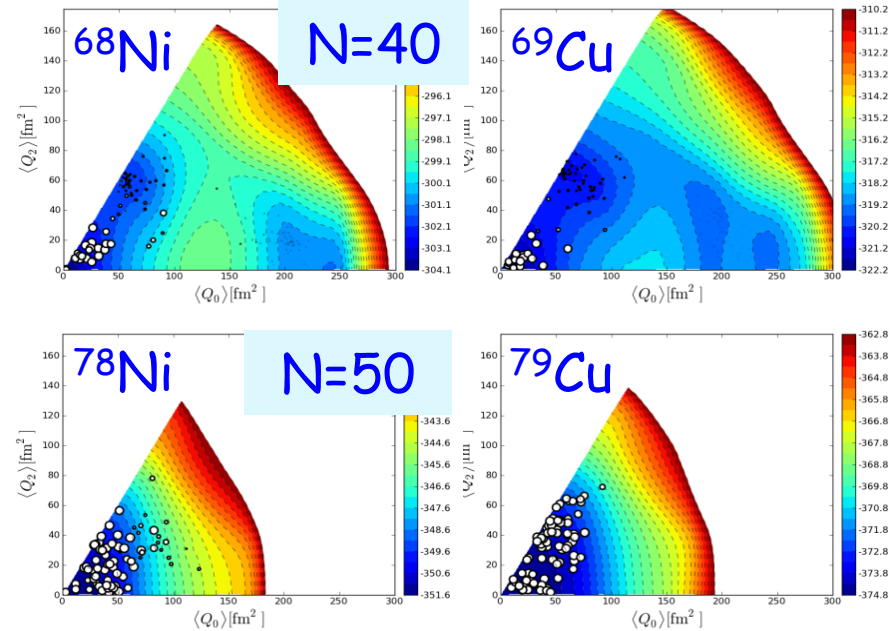
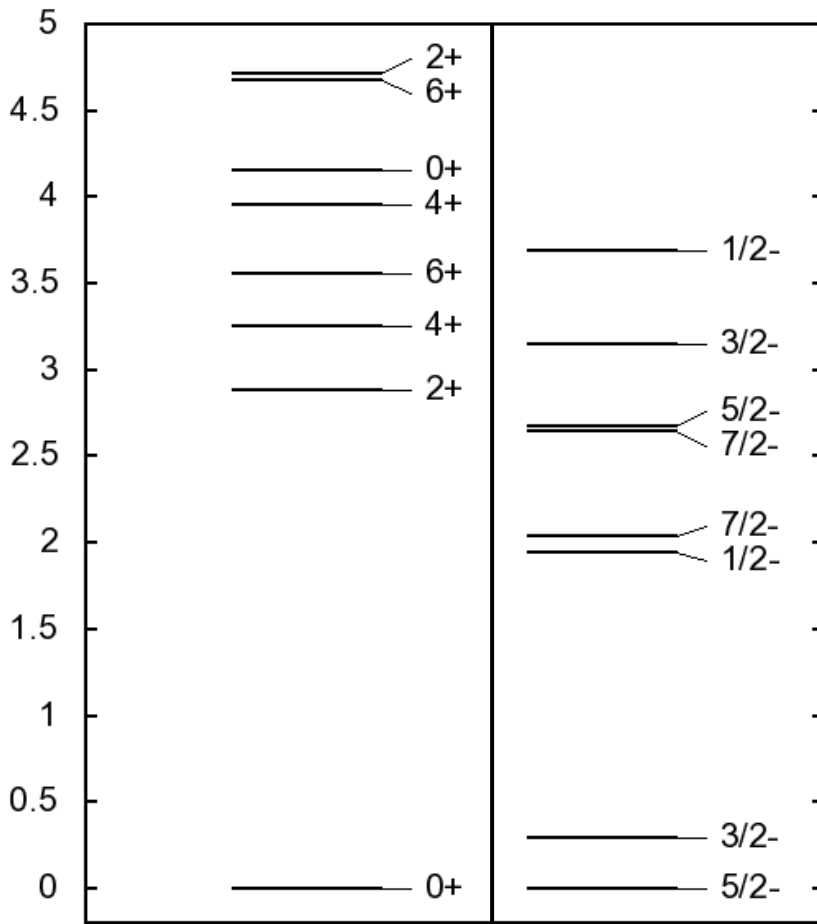
$^{68}\text{Ni} - ^{69}\text{Cu}$ v.s. $^{78}\text{Ni} - ^{79}\text{Cu}$

T-plot of ground state

^{78}Ni

calc.

^{79}Cu

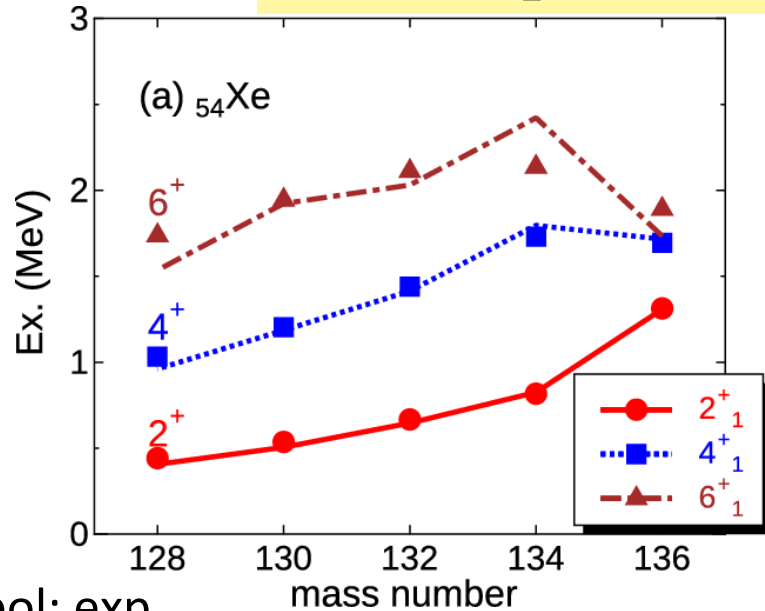


- **Similar distribution patterns** between Ni and Cu, while Cu is somewhat more deformed
- **Shape fluctuations** are larger in N=50 isotones

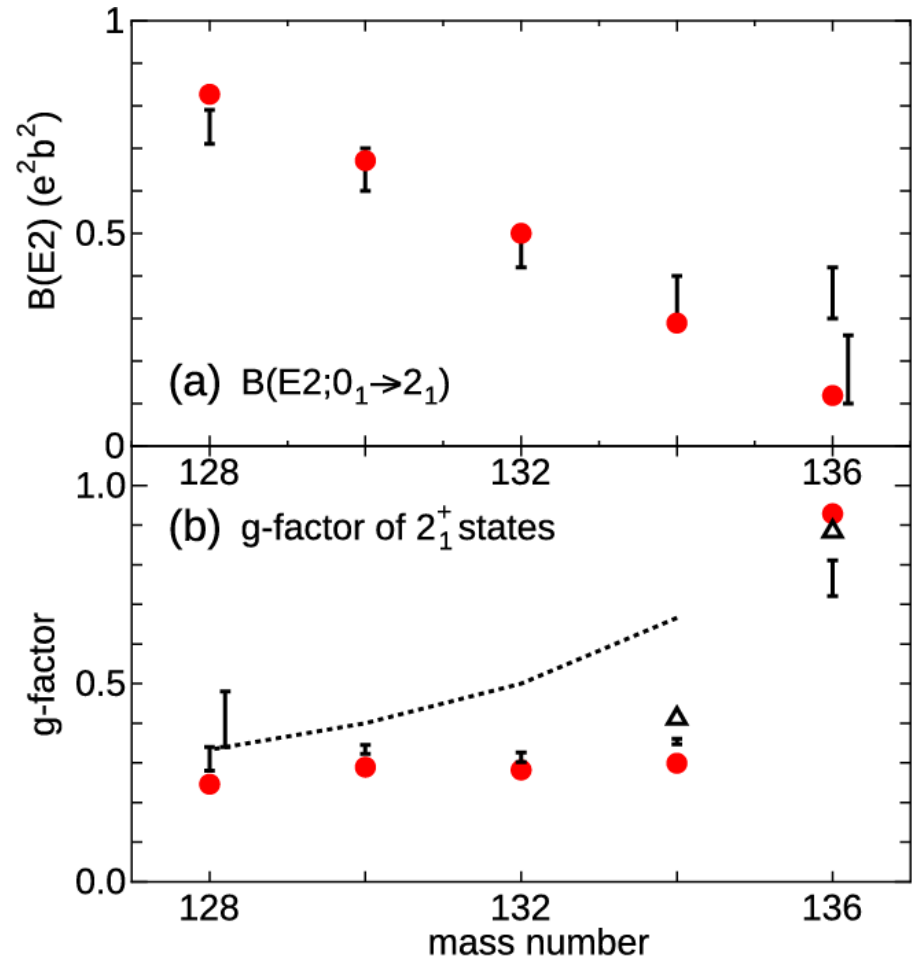
Outline

- **Methodology**: advanced Monte Carlo shell model (MCSM)
 - intrinsic shape can be the objectives
 - ab initio (no core) MCSM and clustering in Be isotopes
- **Shape coexistence** and Quantum Liquid picture
 - exotic Ni isotopes (+ Co, Cu)
- **Shape evolution** (from seniority to rotor)
 - Xe and Ba isotopes
- **Extension**: Spectra to high-lying collective states
 - E1 excitation, GDR and PDR in Ca and Sr isotopes
 - Level density with shell model Hamiltonian

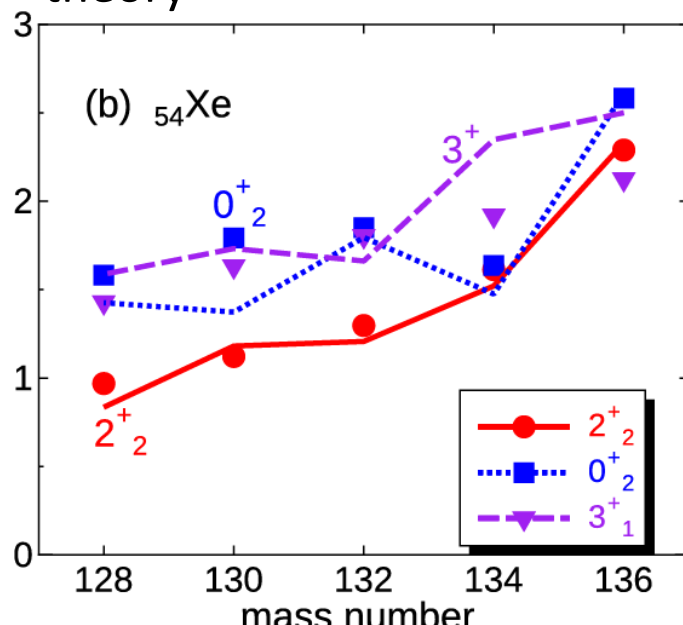
Shape evolution of Xe isotopes



P+QQ int.
model space: $50 < N, Z < 82$

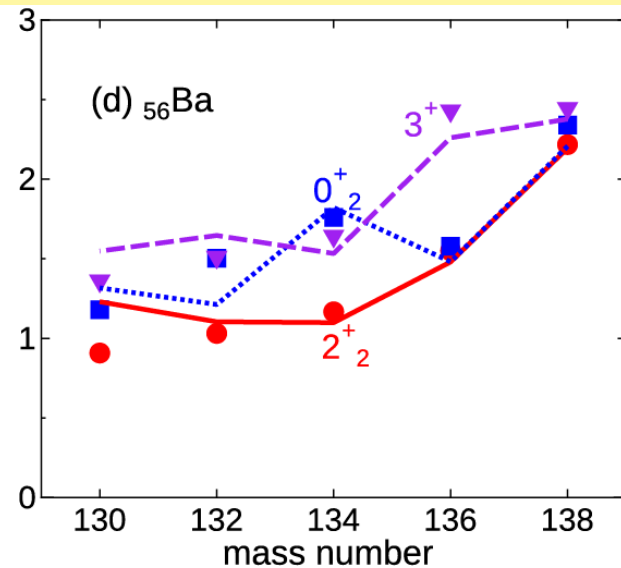
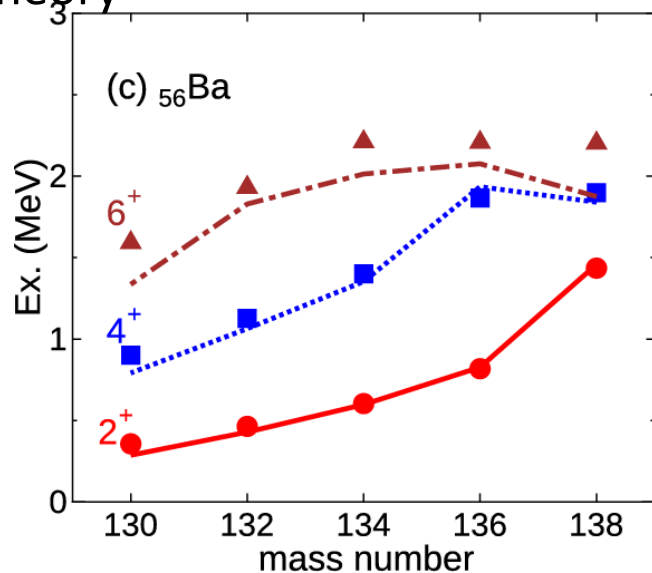


symbol: exp.
line: theory

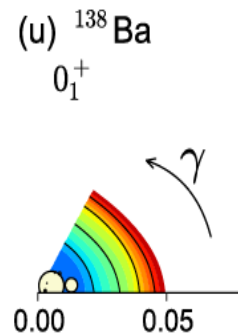
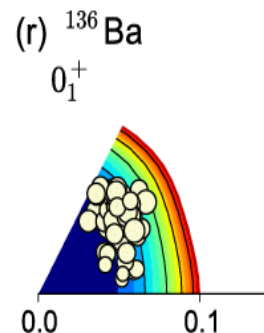
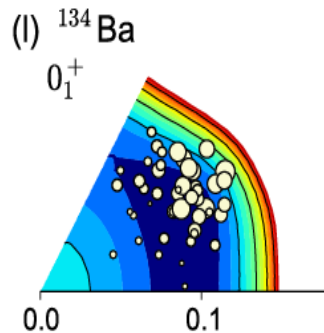
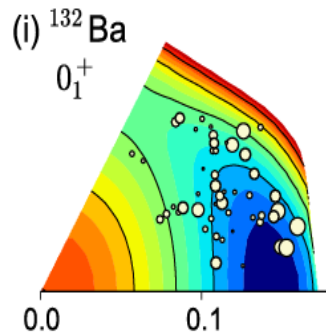
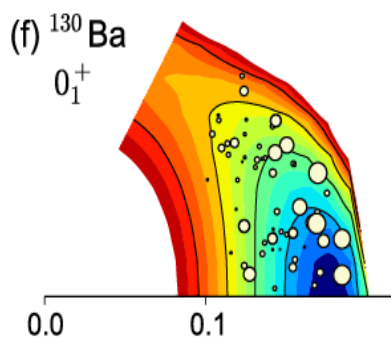


symbol: exp.
line: theory

Shape evolution of Ba isotopes



P+QQ int.
model
space:
 $50 < N, Z < 82$



MCSM w.f.

$$|\Psi\rangle = \sum_{k=1}^{N_{MCSM} \sim 100} f_k P^{J,\pi} |\phi_k\rangle$$

PES by Q-constraint HFB calc. for the SM Hamiltonian

Location of circle : quadrupole deformation of $|\phi_k\rangle$

Area of circle : overlap probability with the eigenstate

Outline

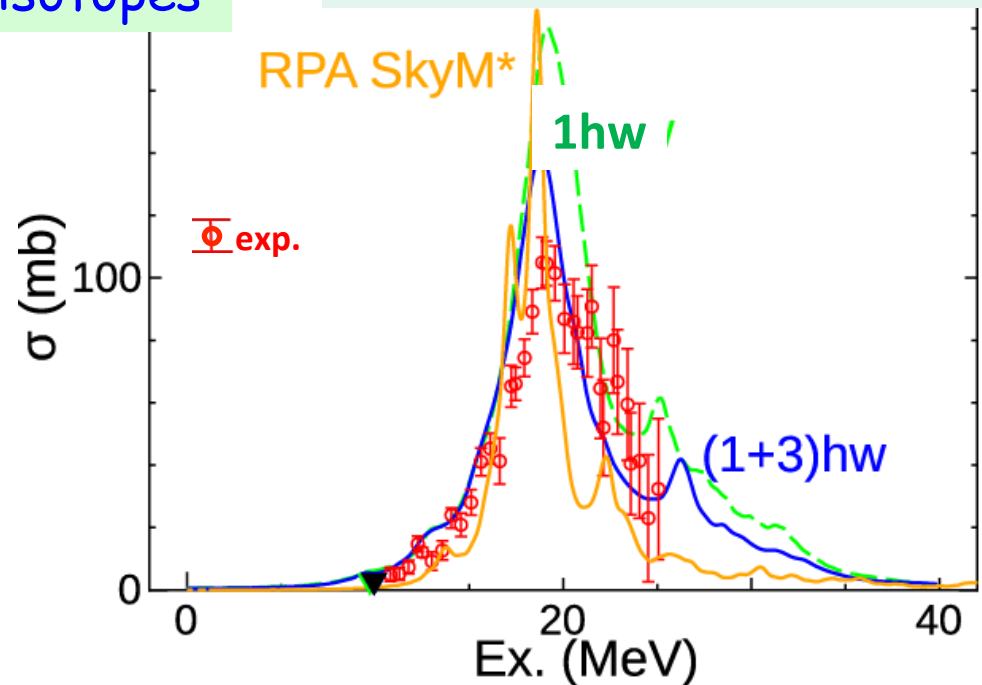
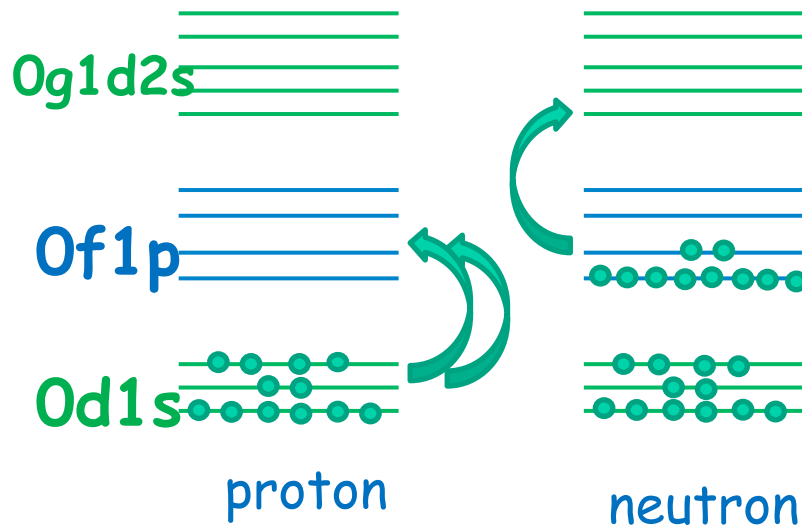
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E1 excitation of Ca isotopes in (conventional) LSSM

NS, Y. Utsuno, S. Ebata, T. Otsuka, M. Honma and T. Mizusaki, in preparation

1hw/3hw sd-pf-sdg shell calculations
for negative-parity states of Ca isotopes

Photo-absorption cross section
of ^{48}Ca



M-scheme shell-model code
"KSHELL"
by Lanczos method
on massive parallel computer



1hw : upto 1hw excitation in sd-pf-sdg shell
 4.1×10^6 M-scheme dim. at PC

(1+3)hw: up to 3hw excitation in sd-pf-sdg shell
 1.2×10^{10} M-scheme dim. at supercomputer


B(E1) sum rule
by Monte Carlo shell model
 $N_{\max} \hbar\omega$ configuration

E1 sum rule from fully correlated ground state wave function

- ^{48}Ca B(E1) sum rule ($e^2\text{fm}^2$)

– $1\hbar\omega$... 16.5			-18%
– $(1+3)\hbar\omega$... 13.6			
– MCSM 50dim.	... 10.1			

- ^{51}V B(E1) sum rule ($e^2\text{fm}^2$)

– $1\hbar\omega$... 18.1		-32%
– $(1+3)\hbar\omega$... NA		
– MCSM 50dim	... 12.4		

Many-body correlations
reduces B(E1) sum rule

E1 excitation spectrum can be calculated by MCSM

Ground state:

$$|\Psi\rangle = \sum_{k=1}^{N_{MCSM}} f_k P^{J,\pi} |\varphi_k\rangle$$

Basis vector of the ground state

Basis vectors for $E1$ spectrum (a,b,c,d, \dots : orbits)

$$\exp(i\varepsilon \cdot E1(a \rightarrow b)) |\varphi_k\rangle, \exp(i\varepsilon \cdot E1(c \rightarrow d)) |\varphi_k\rangle, \dots (k=1,2,\dots)$$

These are still Slater determinants

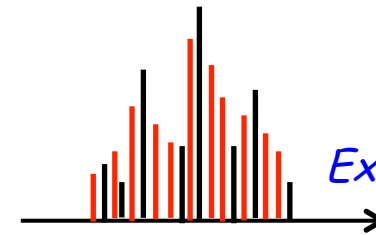
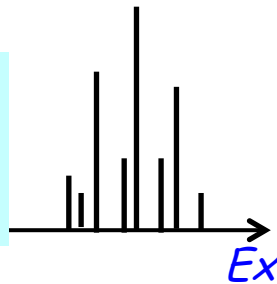
+

Additional bases for fine tuning :
variation for energy average by the conjugate gradient

$$|\varphi_k(E1(a \rightarrow b))^{Var}\rangle, |\varphi_k(E1(c \rightarrow d))^{Var}\rangle, \dots$$

The global feature of excitation spectrum can be calculated.

Diagonalization with these basis vectors after projection to 1-

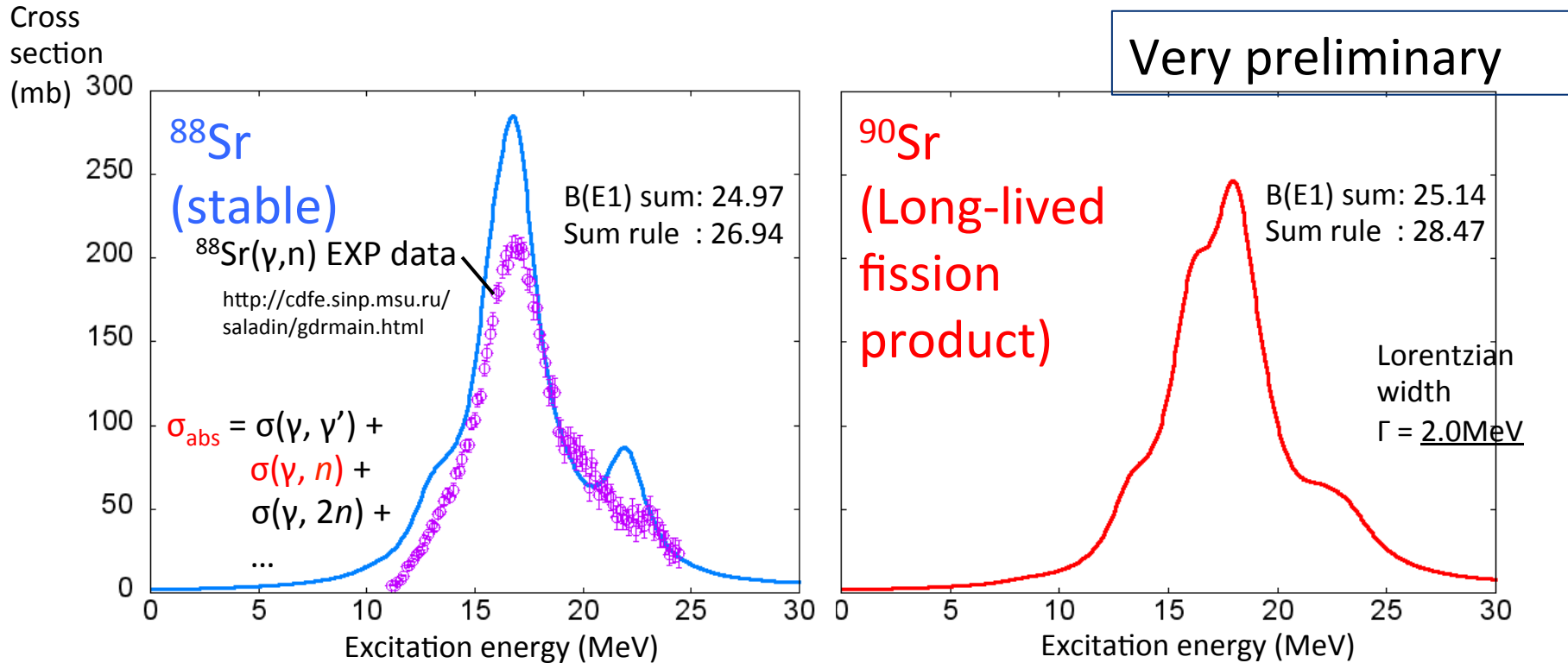


E1 spectrum with ~ 3000 levels connected somehow to the g.s. (confirmed by E1 sum rule)

E1 excitation spectrum by MCSM

T. Togashi, T. Otsuka *et al.*

Photoabsorption cross section of ^{88}Sr , ^{90}Sr



Application to the nuclear transmutation
for the purpose of nuclear waste processing (ex. ^{90}Sr LLFP)
as basic data for nuclear technology

Perspectives

1. Ab initio at the level of shell-model Hamiltonian based on EKK method (for multi-shells)
tensor + 3NF inclusive
2. Variety of magic nuclei (magic index)
3. Quantum liquid - single or dual -
implications ?
4. Shell evolution in continuum and doorway-state resonance - reminiscence of bound states -
5. Shapes and clusters by MCSM
6. Excitation spectrum by MCSM and their application to astrophysics, nuclear energy, etc.

Collaborators

HPCI project, Field 5

- Takashi Abe (Tokyo)
- Shuichiro Ebata (Hokkaido)
- Michio Honma (Aizu)
- Yoritaka Iwata (CNS, Tokyo)
- Takahiro Mizusaki (Senshu Univ.)
- Noritaka Shimuzu (CNS, Tokyo)
- Tomoaki Togashi (CNS, Tokyo)
- Naofumi Tsunoda (CNS, Tokyo)
- Yusuke Tsunoda (CNS, Tokyo)
- Yutaka Utsuno (JAEA)
- Tooru Yoshida (CNS, Tokyo)

- James Vary (Iowa)
- Pieter Maris (Iowa)

END

ご清聴ありがとうございました