

# Nuclear Physics

Sakura-i    Hiro-yoshi  
櫻        井    博    儀

Fundamental interactions

Elementary particles

Composite particles

Nuclear Physics is not Particle Physics,  
not Condensed Matter Physics

Interaction?

Effective interaction ?

Correlations ?

Isospin, Density, temperature dependences ?

Surface boundary, non-linear, finite system

Collective motions

Q. 1 Life time of neutron ?

Q. 2 Age of universe is 13.7B Years after BigBang.  
At present, there are neutrons in materials. Why?

Q. 3 Spin-parity for ground state of deuteron ?

Q. 4 Limits of existence of nuclei ?

Q. 5 Magic numbers of nuclei ?

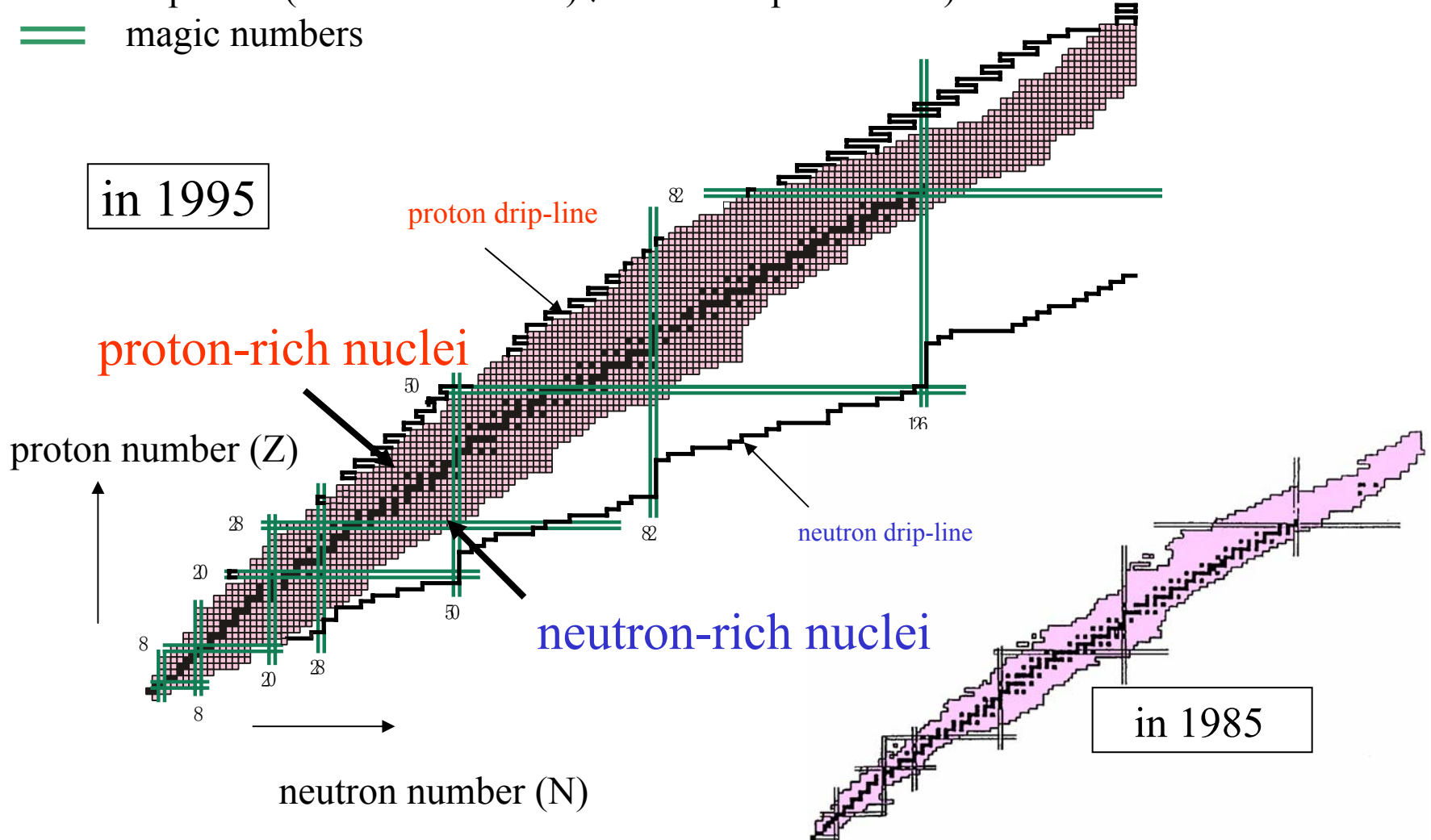
Q. 6 Size of nuclei ?

Q. 7 Collective motions of nuclei ?

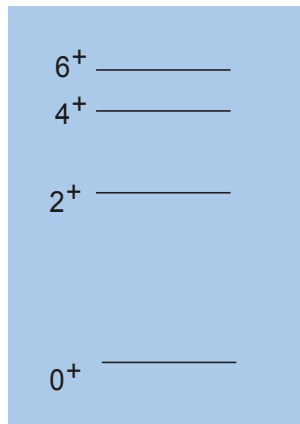
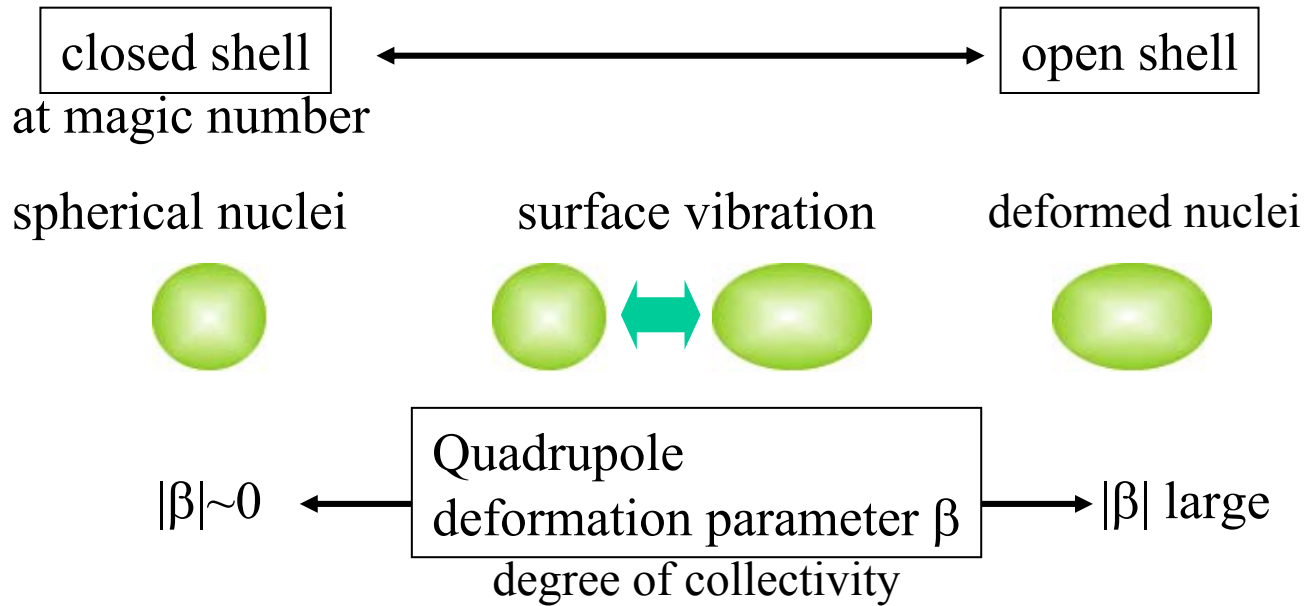
Q. 8 How and where elements around us have been created ?

# Exploration of the Limit of Existence

- stable nuclei ~300 nuclei
- unstable nuclei observed so far ~2700 nuclei
- drip-lines (limit of existence) (theoretical predictions) ~6000 nuclei
- == magic numbers

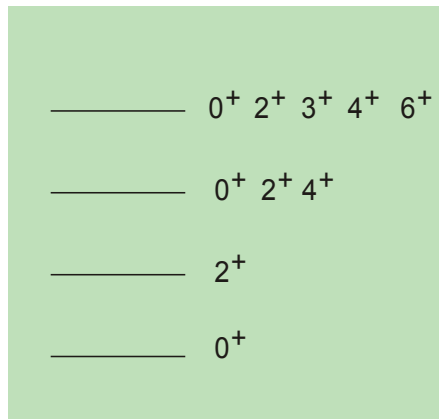


# Nuclear Collective Motion

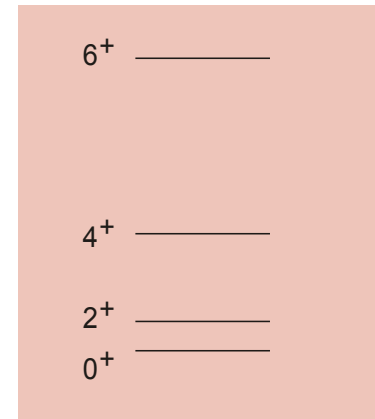


$$E(4^+)/E(2^+)$$

$$\sim 1.8$$

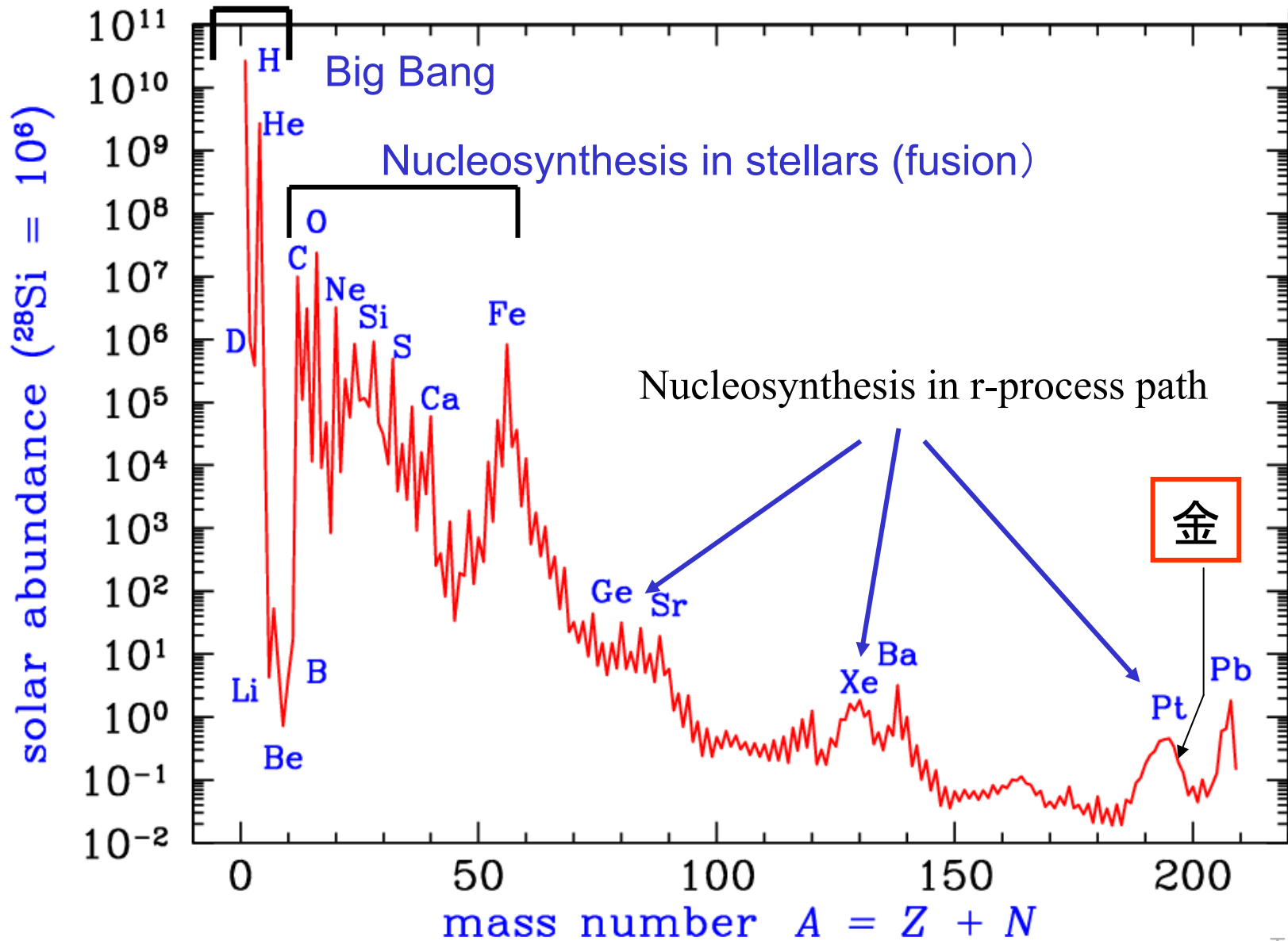


$$\sim 2.2$$



$$\sim 3.3$$

# Solar Abundance of Elements



# Gold . . .



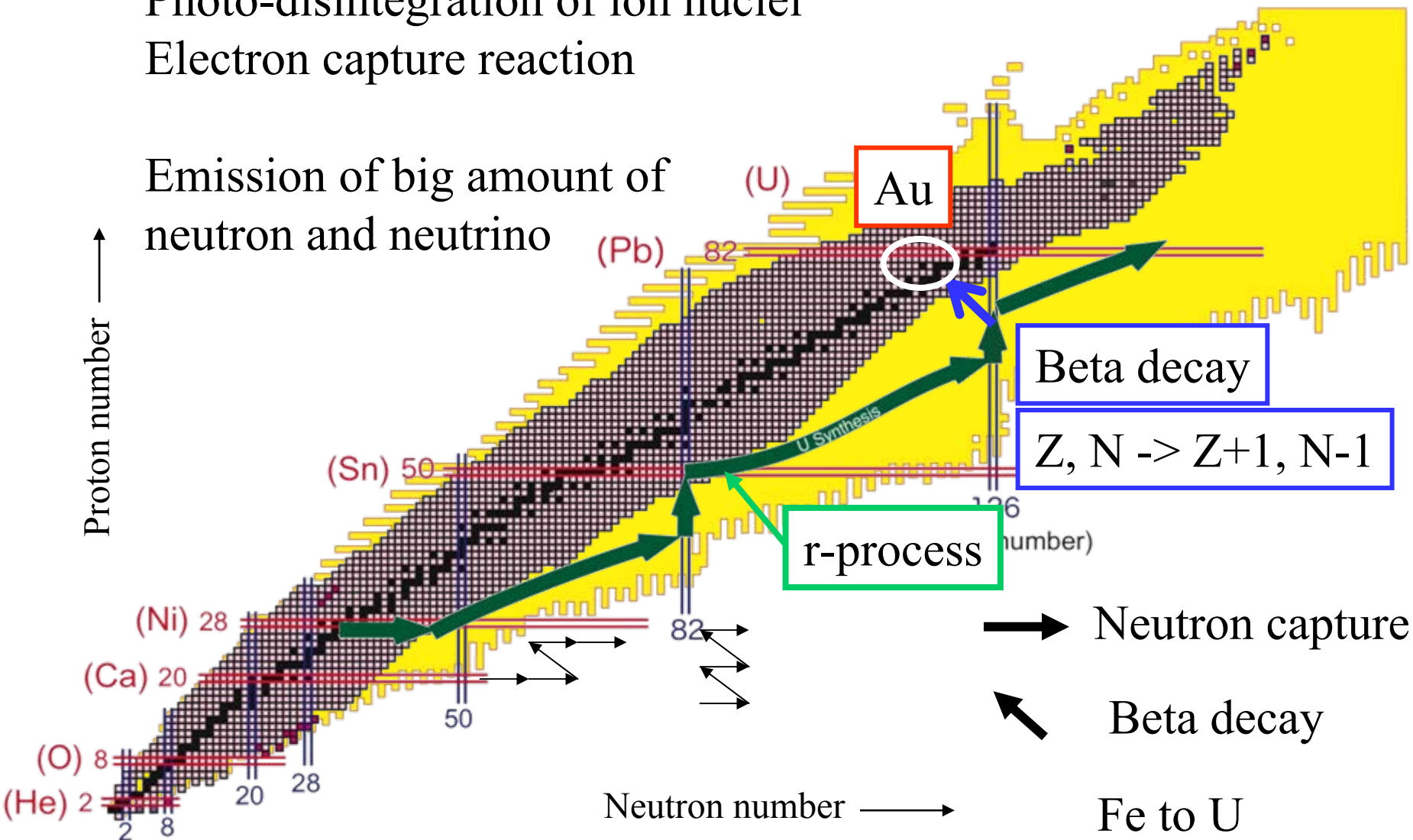


# r-process path in supernova explosion

Photo-disintegration of ion nuclei

Electron capture reaction

Emission of big amount of neutron and neutrino



Beta decay

$Z, N \rightarrow Z+1, N-1$

r-process

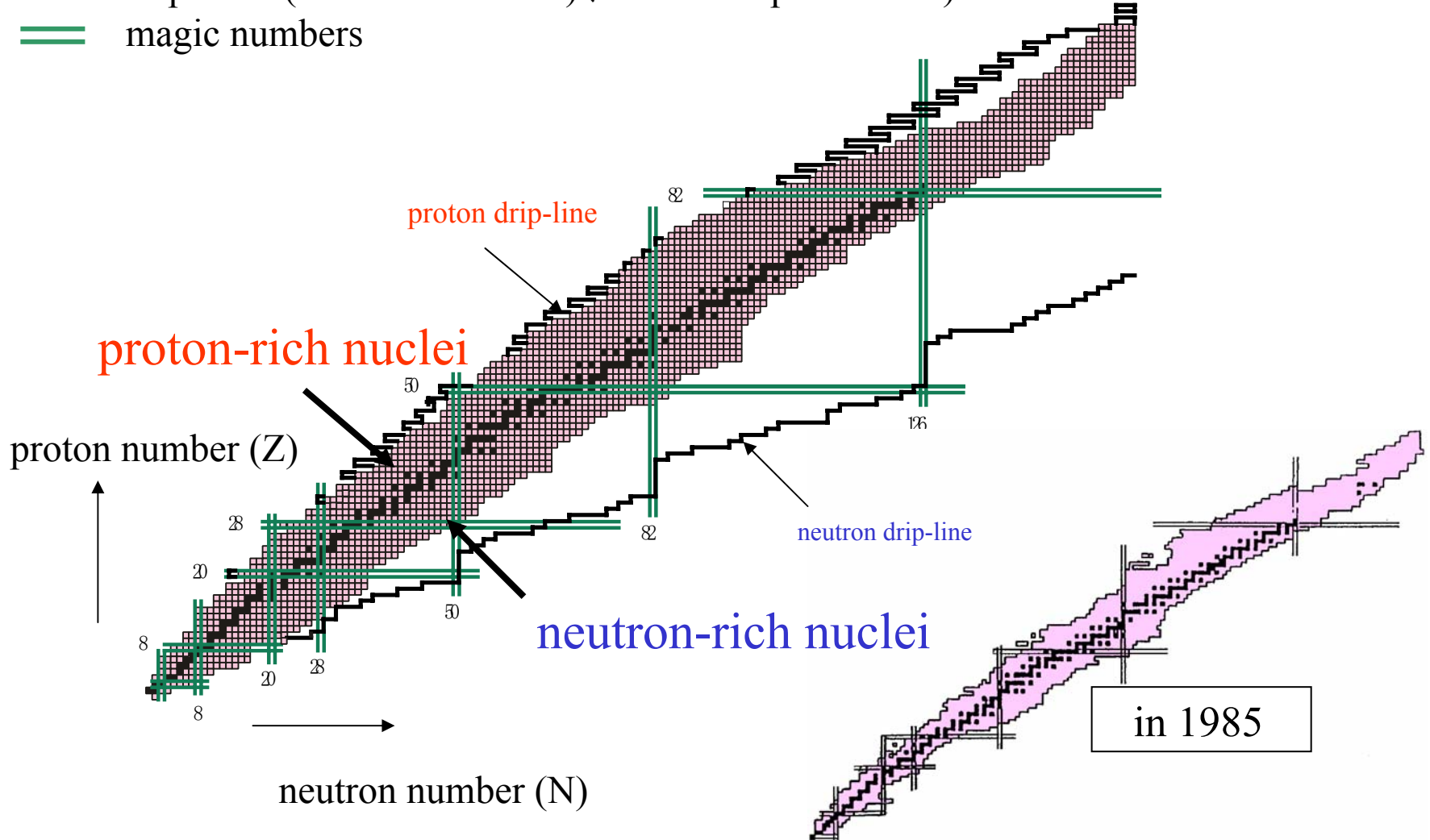
Neutron capture

Beta decay

Fe to U

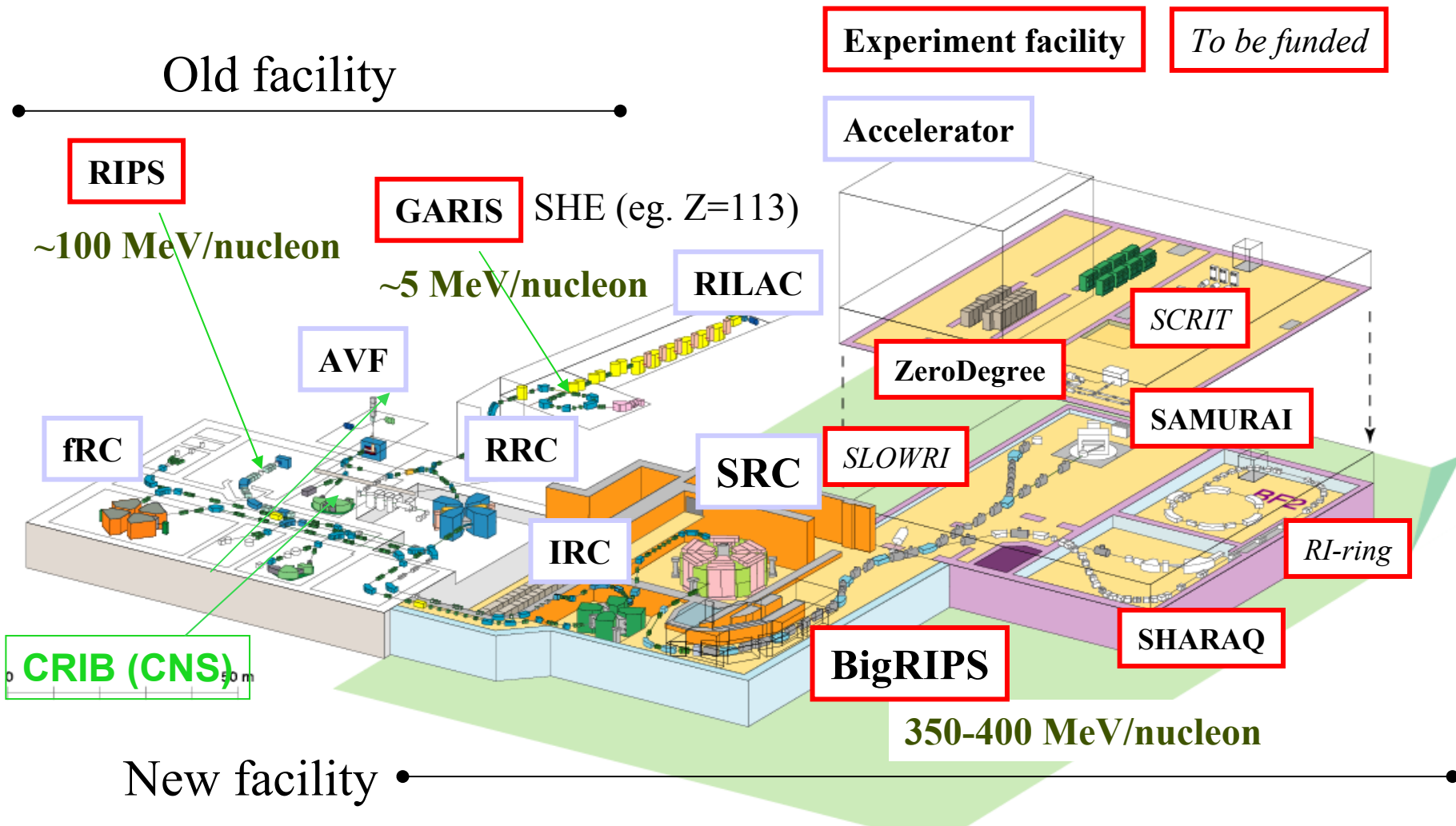
# Exploration of the Limit of Existence

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# RIKEN RI Beam Factory (RIBF)

Old facility



New facility

**Intense (80 kW max.) H.I. beams (up to U) of 345A MeV at SRC**  
**Fast RI beams by projectile fragmentation and U-fission at BigRIPS**  
**Operation since 2007**

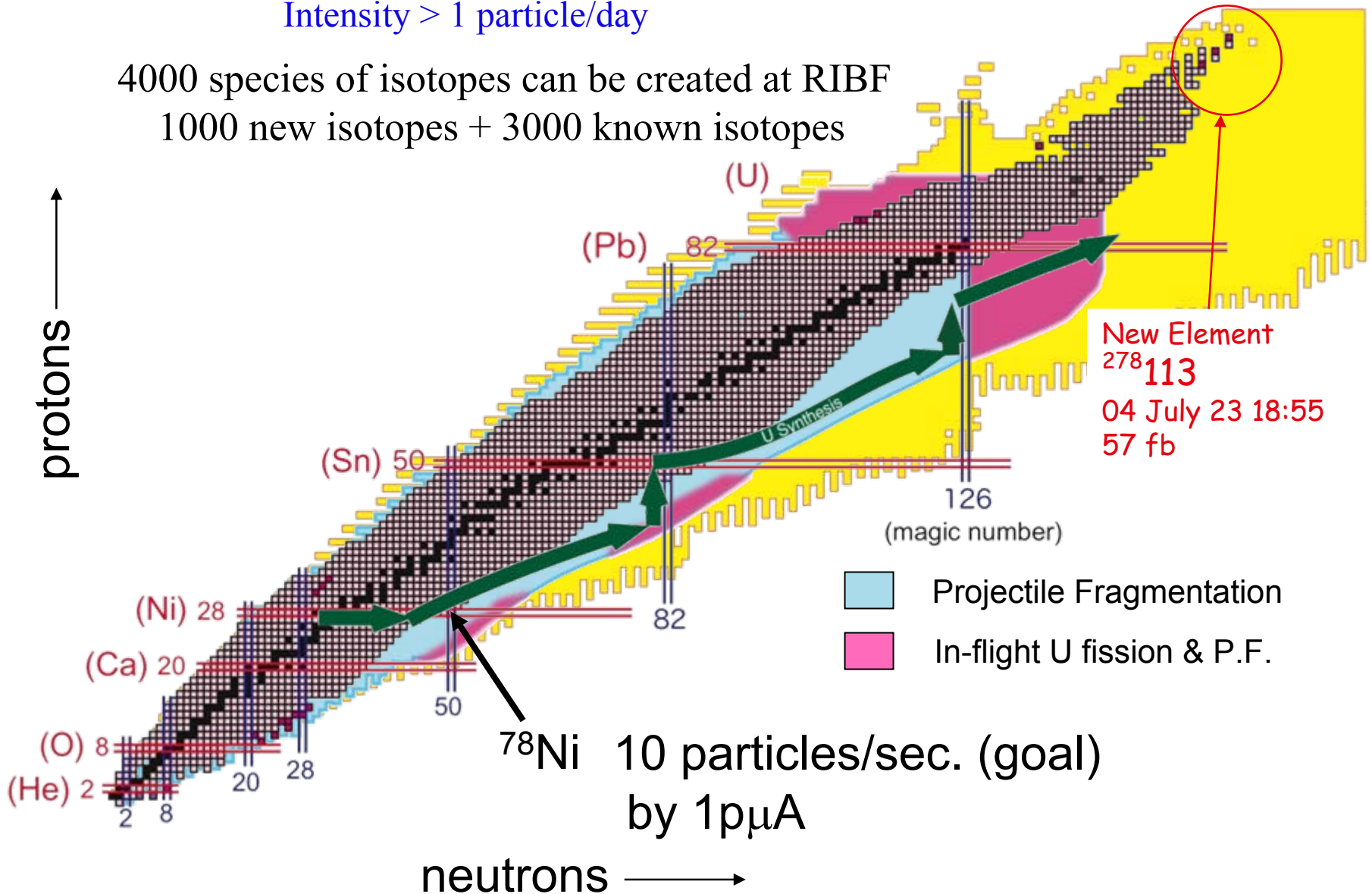
# Exploration of the Limit of Existence

Great expansion of nuclear world by RIBF

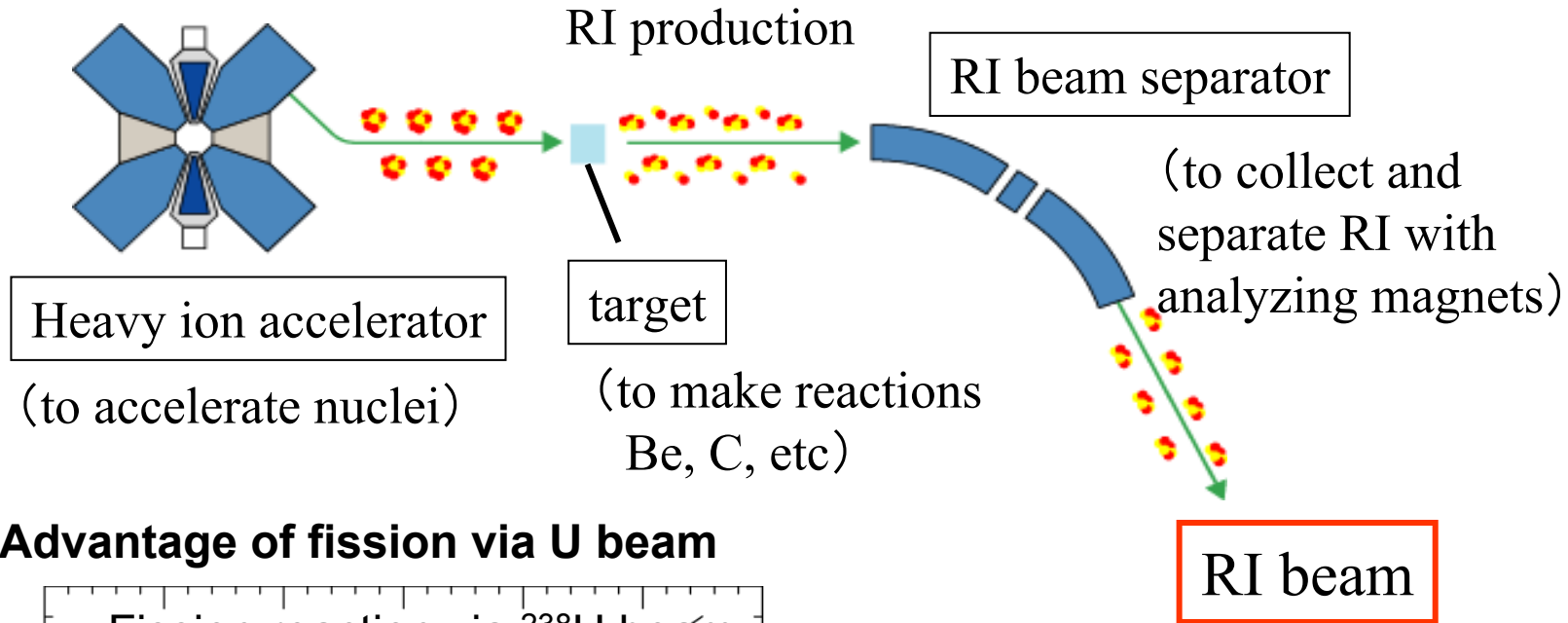
Intensity  $> 1$  particle/day

4000 species of isotopes can be created at RIBF

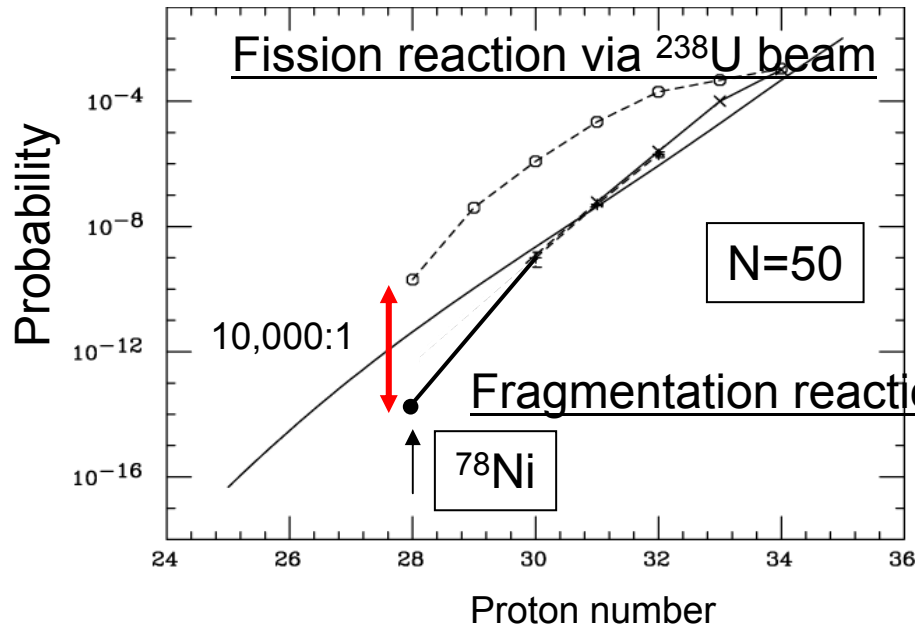
1000 new isotopes + 3000 known isotopes



# RI beam production via in-flight method

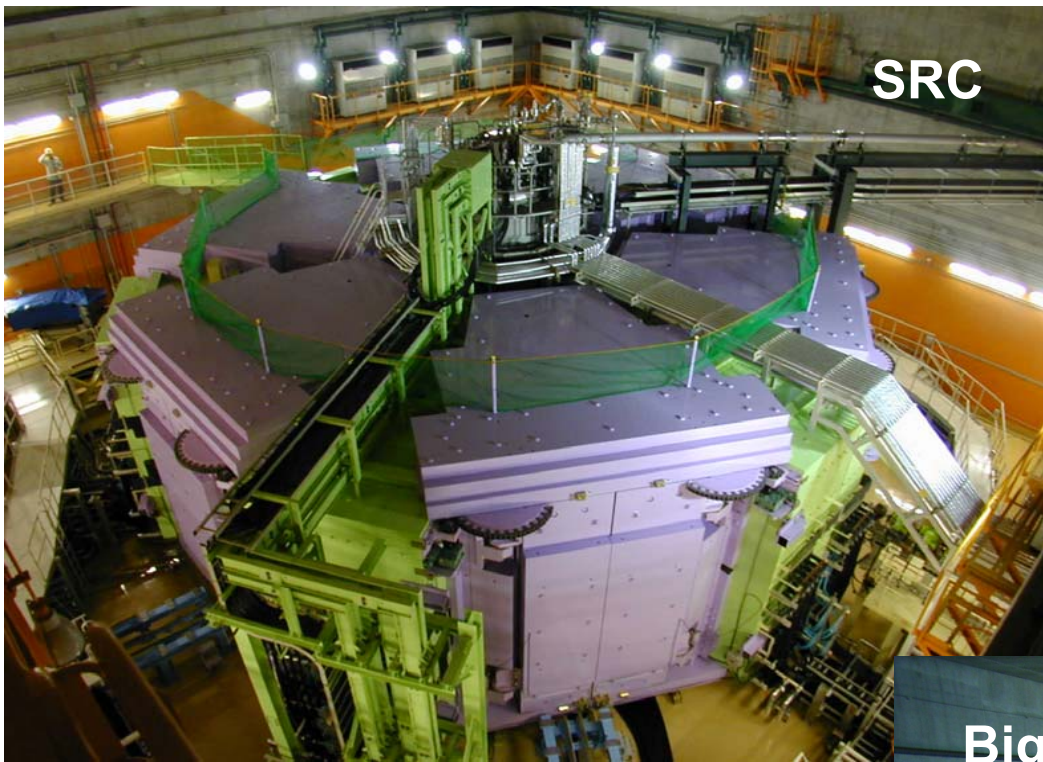


## Advantage of fission via U beam



Yield rate of  $^{78}\text{Ni}$  via fission is **about 1,000** times higher than via fragmentation.





**SRC**

**World's First and Strongest  
K2600MeV  
Superconducting Ring Cyclotron**

400 MeV/u Light-ion beam  
345 MeV/u Uranium beam

**World's Largest Acceptance  
9 Tm  
Superconducting RI beam Separator**

~250-300 MeV/nucleon RIB



**BigRIPS**



Press-Conference on June 8<sup>th</sup>, 2010

June 8, 2010  
RIKEN

### Scientists discover 45 new radioisotopes in 4 days

毎 日 新 聞 2010年(平成22年)6月22日(火)

23

## 放射性同位元素 効率的発見 4日で新たに45種

理 研

理化学研究所は、超短波長放射線を用いた実験で、4日間で世界に初めて発見した。世界の年間発見数(平均約40種)を大幅に上回る効果で「世界的に類例がほとんどない」といって、成果は日本物理学会論文誌「Journal of Physics」に掲載された。

放射性同位元素は、原子核の不安定な状態にある原子核から放射線を出して安定な状態になる。この性質を利用して、がんの治療や放射線治療、長い年月を経た放射性同位元素を集中的に探し出すことが可能になる。宇宙線の性質を詳しく調べれば、宇宙線の理論や加速器の性能を知ることもできる。同位元素の発見は、加速器の性能を向上させる重要な指標である。【田中 健】

2010年(平成22年)6月9日(水曜日) 3

## 同位元素45種 一度に発見

### 加速器施設 世界最高の性能

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日 刊 工 業 新 聞 2010年(平成22年)6月9日 水曜日 21

## 中性子多い放射性同位元素 新たに45種類発見

### 理研など

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IOP physicsworld.com  
 “Radioisotopes galore at RIKEN”  
<http://physicsworld.com/cws/article/news/42915>

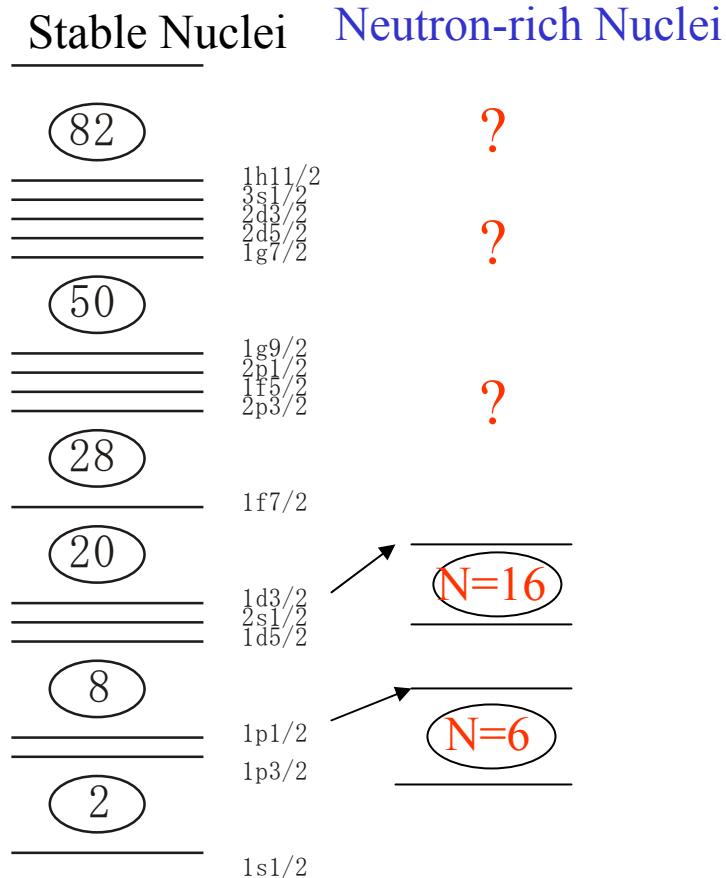
# Three Challenges of RIBF

- 1 To Establish New Framework of Nuclear Physics
- 2 To Elucidate the Origin of Elements
- 3 To Explore New Applications with fast RI Beams



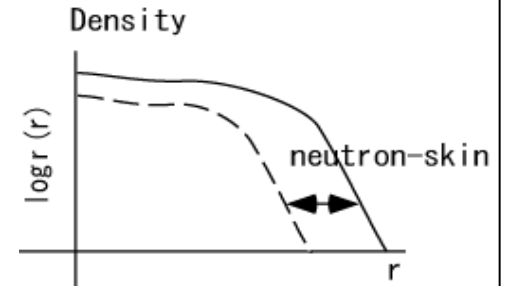
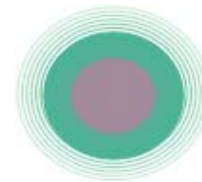
# New frameworks for the new region of nuclear chart

## Nuclear Structure: Shell evolution

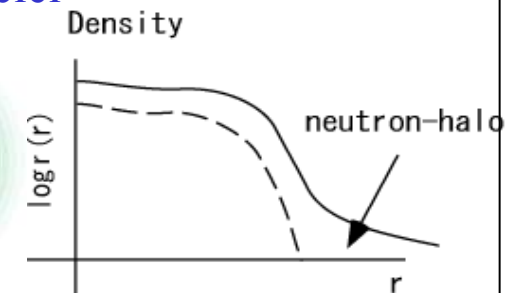
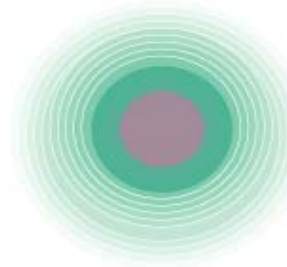


## Nuclear Matter: New forms

### neutron-skin nuclei



### neutron-halo nuclei



$^{11}\text{Be}$ ,  $^{11}\text{Li}$ ,  $^{19}\text{C}$ ...

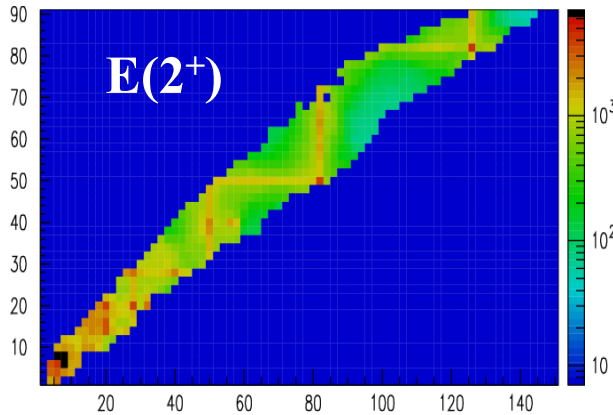
To write up new text book: Exotic phenomena, Systematics, etc.

Isospin-, density-dependences of effective interactions, nucleon-corrections

Microscopic system (nuclei) to Macroscopic system (neutron stars)

# Liberation from Stable Region and Exotic Nuclei

## Shell Evolution : magicity loss and new magicity



Shape ?  
 Shell gap ?  
 Single particle level ?  
 Cluster formation?  
 Role of 3NF ?  
 Magicity loss ?

50, 82, 126, 184

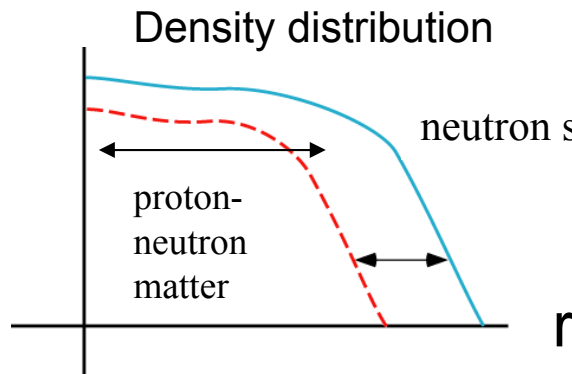
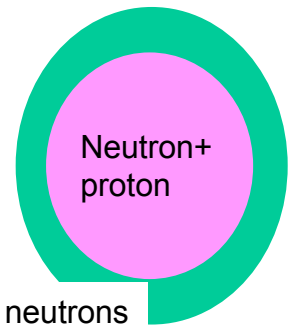
Spherical



Deformed



## Dynamics of new "material" : Neutron-skin (halo)



New quantum objects with two surfaces  
 Skin thickness ? Density distribution ?  
 Role of skin in reactions ?  
 Pairing in skin ? di-neutrons?  
 Exotic modes of skin ?

**RIBF provides data for nuclei far from the stability line**

Challenges in establishing new frame work of nuclear physics

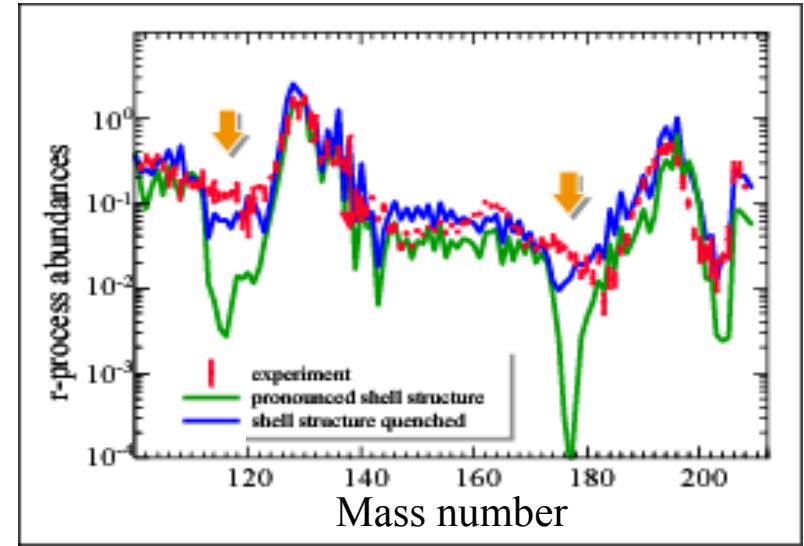
# Challenge for r-process path and explosion in supernovae

## Synthesis up to U (r-process)

unknown neutron-rich nuclei  
theoretical predictions only

Necessary of experimental investigation  
for nuclear properties of heavy and  
neutron-rich nuclei

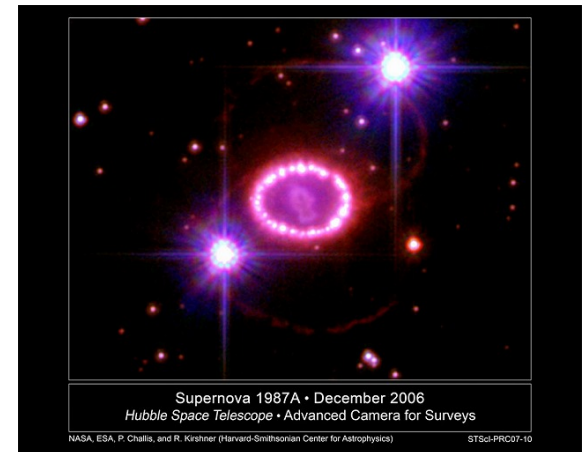
Mass, life-time, decay mode



## Explosion mechanism of supernova

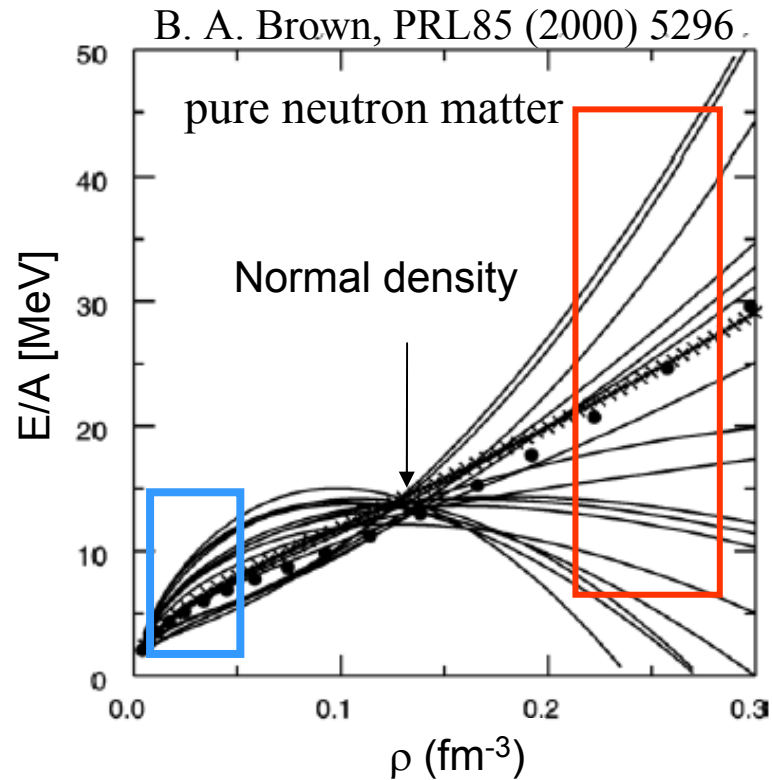
No explosion in theoretical works  
Outer clast of neutron star

Necessary of experimental study for  
Equation-of-State for nuclear matter



1987A

# Challenge to investigate EOS of neutron matter from nuclei to neutron stars



## **$^1S$ correlation**

BCS-BEC crossover  
in dilute system ( $\rho \sim 0.1\rho_0$ ) ?

## **$^3N_F$**

$T=3/2$  channels?  
density dependence?

Elastic  $d+p$  for  $T=1/2$

Nuclear structure in

very neutron-rich nuclei for  $T=3/2$ ?

Heavy-ion Collisions to achieve  $\rho \sim 2-3\rho_0$  ?

## **$^3P_2$ correlation**

pairing gap?

Density dependence?

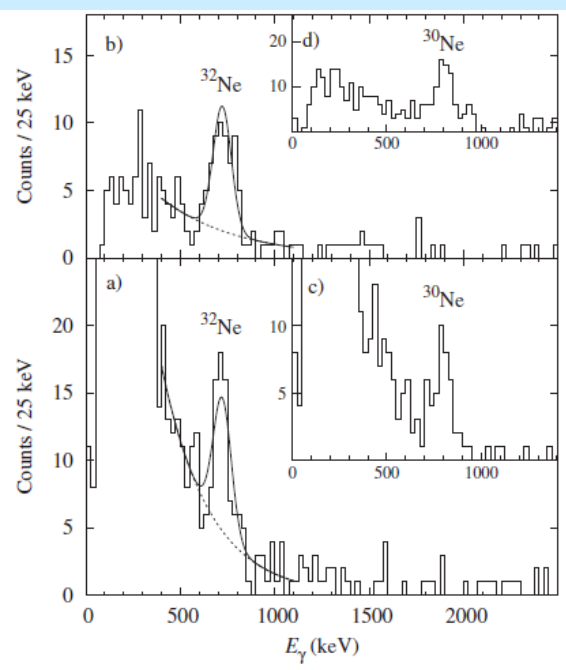
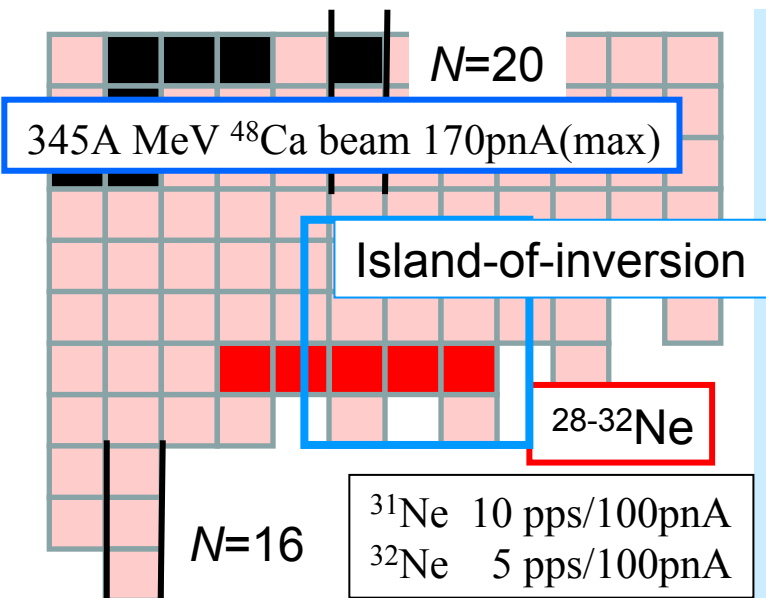
????

Role of di-neutron in skin? : collectivity, transfer reactions

# DayOne Experiments in Dec., 2008

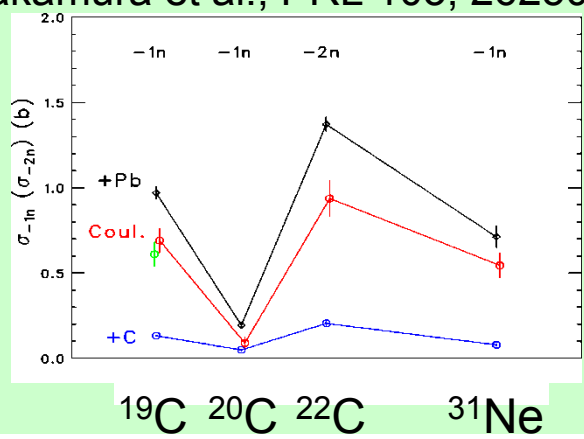
-The first data in the “island-of-inversion” -

Coordinated by Aoi

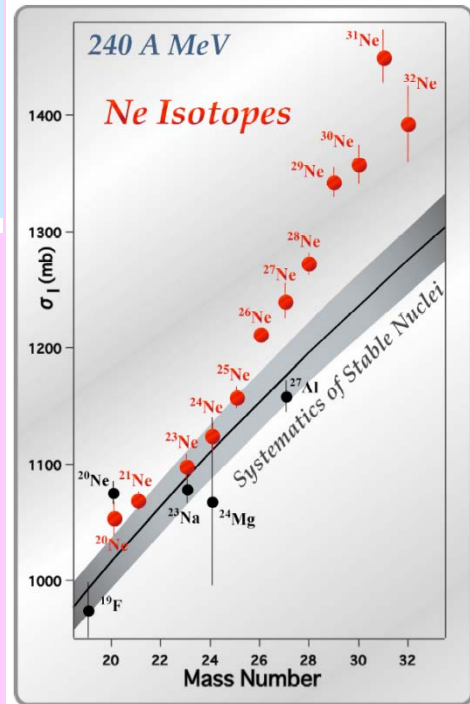


Spectroscopy of  $^{32}\text{Ne}$  and the “island-of-inversion”  
 $E(2+) = 722 \text{ keV}$   
 Doornenbal, Scheit et al. PRL 103, 032501 (2009)  
 New states in  $^{31,32,33}\text{Ne}$  PRC 81, 041305R (2010)

A new candidate of halo nuclei  $^{31}\text{Ne}$  via Coulomb breakup  
 Nakamura et al., PRL 103, 262501(2009)



Total interaction cross sections for the neutron-rich Ne isotopes  
 Takechi, Otsubo et al., Niigata 2010 symposium



# Further investigation for nuclei in the island of inversion region and beyond

Next campaign  
Nov.-Dec 2010

## 1. Halo states with deformed core?

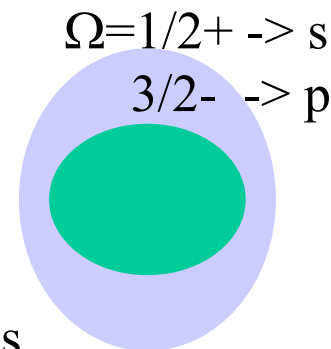
Most of halo nuclei so far are located in deformed regions

Be-11, Li-11, C-19, Ne-31 ...

Deformed core + valence neutron

low-L dominance in Nilsson orbits for weakly-bound neutrons

M. Misu et al., Nucl. Phys. A 614 (97) 44; I. Hamamoto, Phys. Rev. C69 (04) 041306R



Many halo nuclei could be found on the nuclear chart?? Next candidates??

## 2. Mechanism of shell evolution ?

Tensor interactions

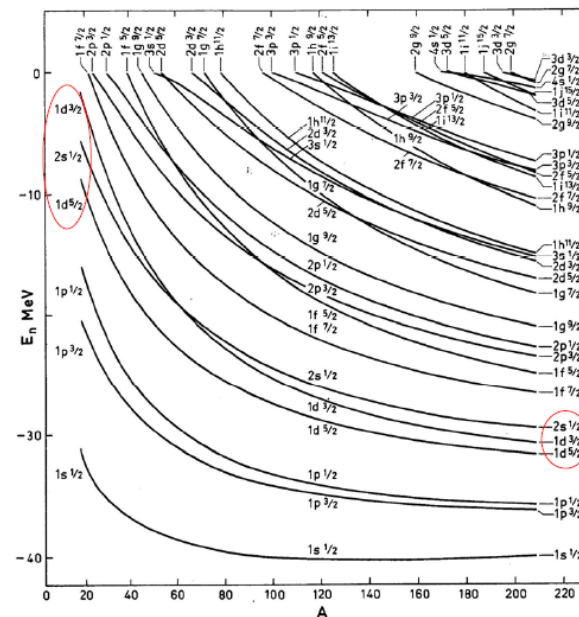
Otsuka et al.

Weakly bound natures

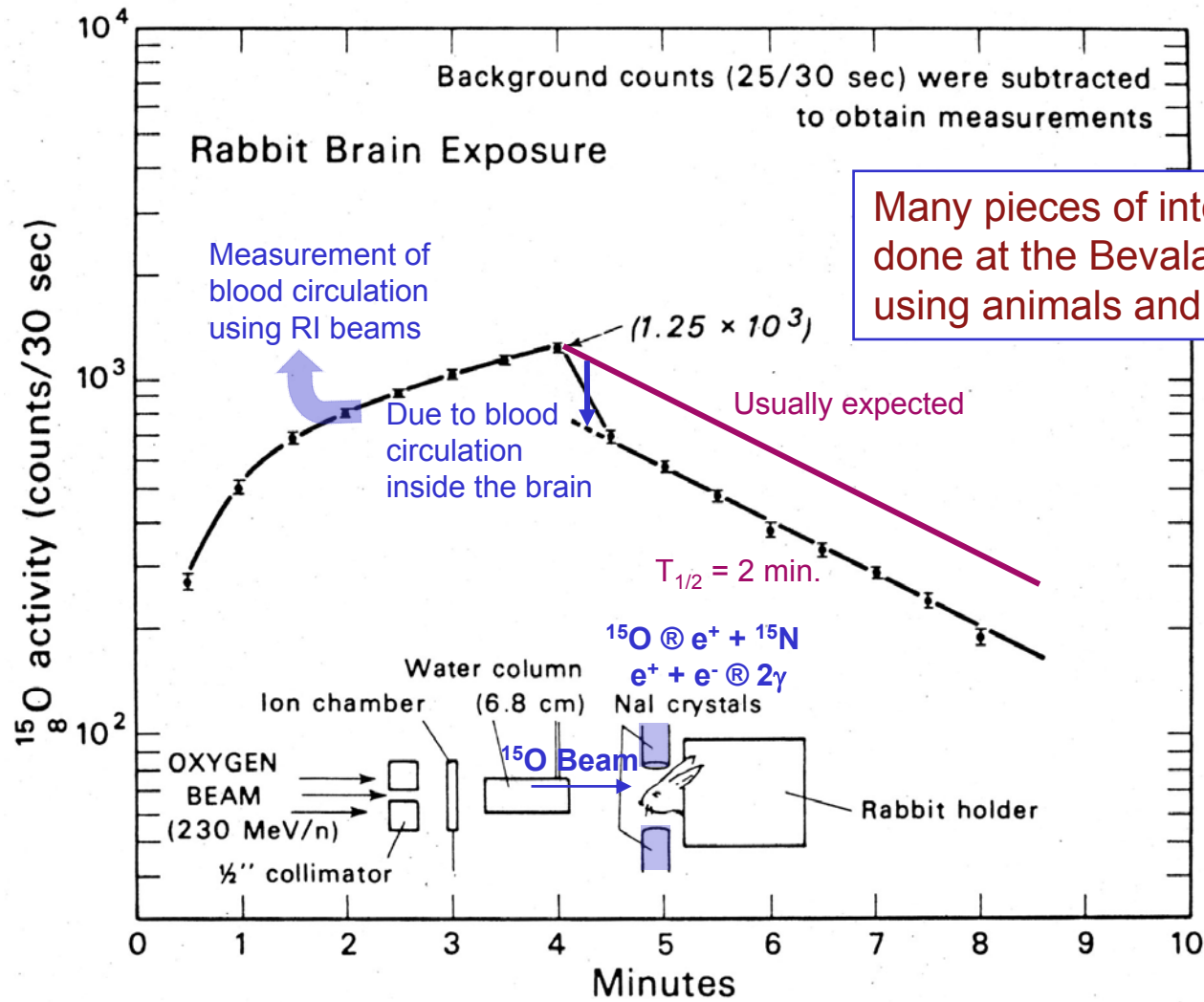
Low-L orbits behaviors  
(Bohr-Mottelson Vol.1)

Cluster formation, What else....

E(2+), E(4+), B(E2), ..as many as possible for Ne and Mg isotopes

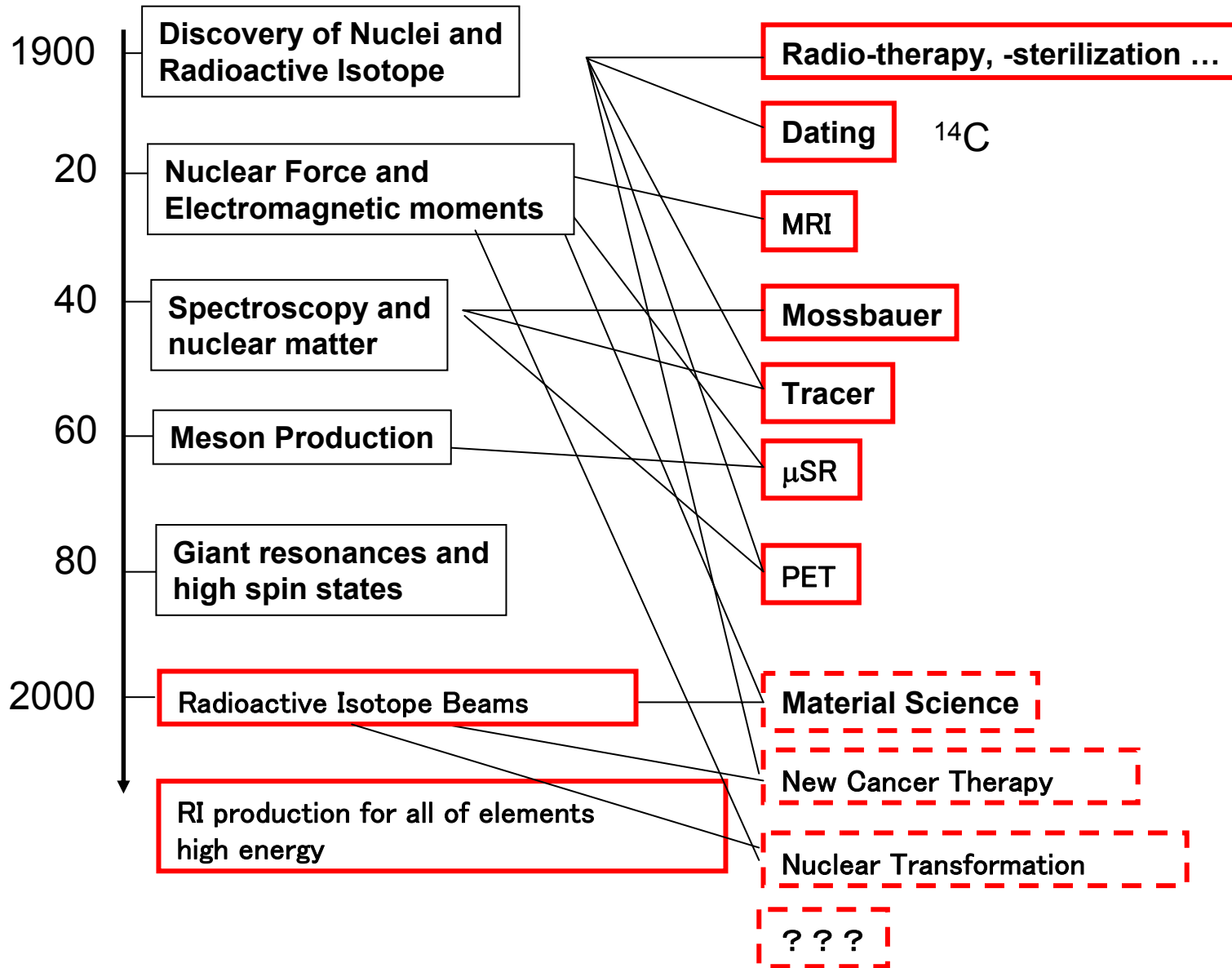


# The First RI Beam Experiment (1974)



Many pieces of interesting work done at the Bevalac in biology using animals and plants

# Development of Nuclear Physics and Applications



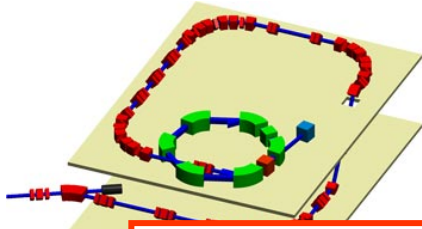


# New Devices of RIBF

To maximize the potentials of intense RI beams available at RIBF

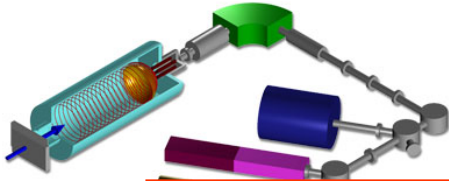
for several 100 – 1000 species

Rare RI ring



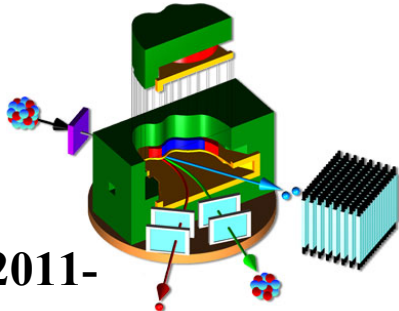
to be funded

SLOWRI



to be funded

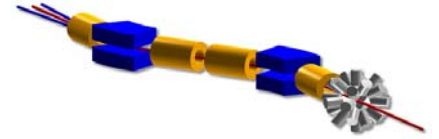
SAMURAI



2011-

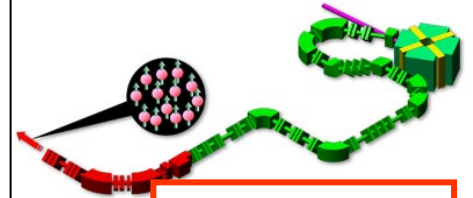
mass  
half-life  
excited states  
deformation  
charge radii  
matter radii  
charge distribution  
matter distribution  
EM moments  
single particle states  
astrophysical reactions  
giant resonances  
exotic modes  
HI collisions (EOS)

ZeroDegree



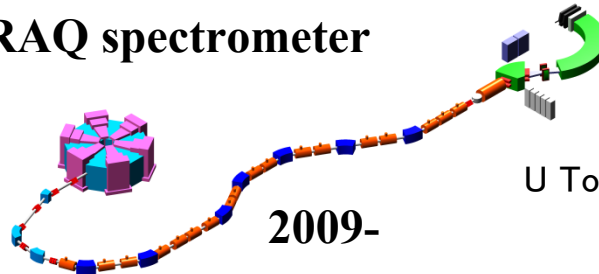
2008-

IRC-to-RIPS BT



to be funded

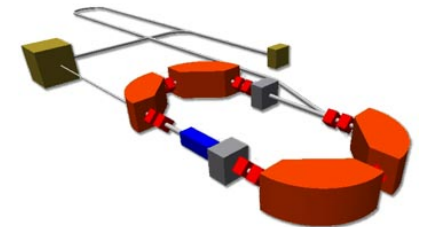
SHARAQ spectrometer



2009-

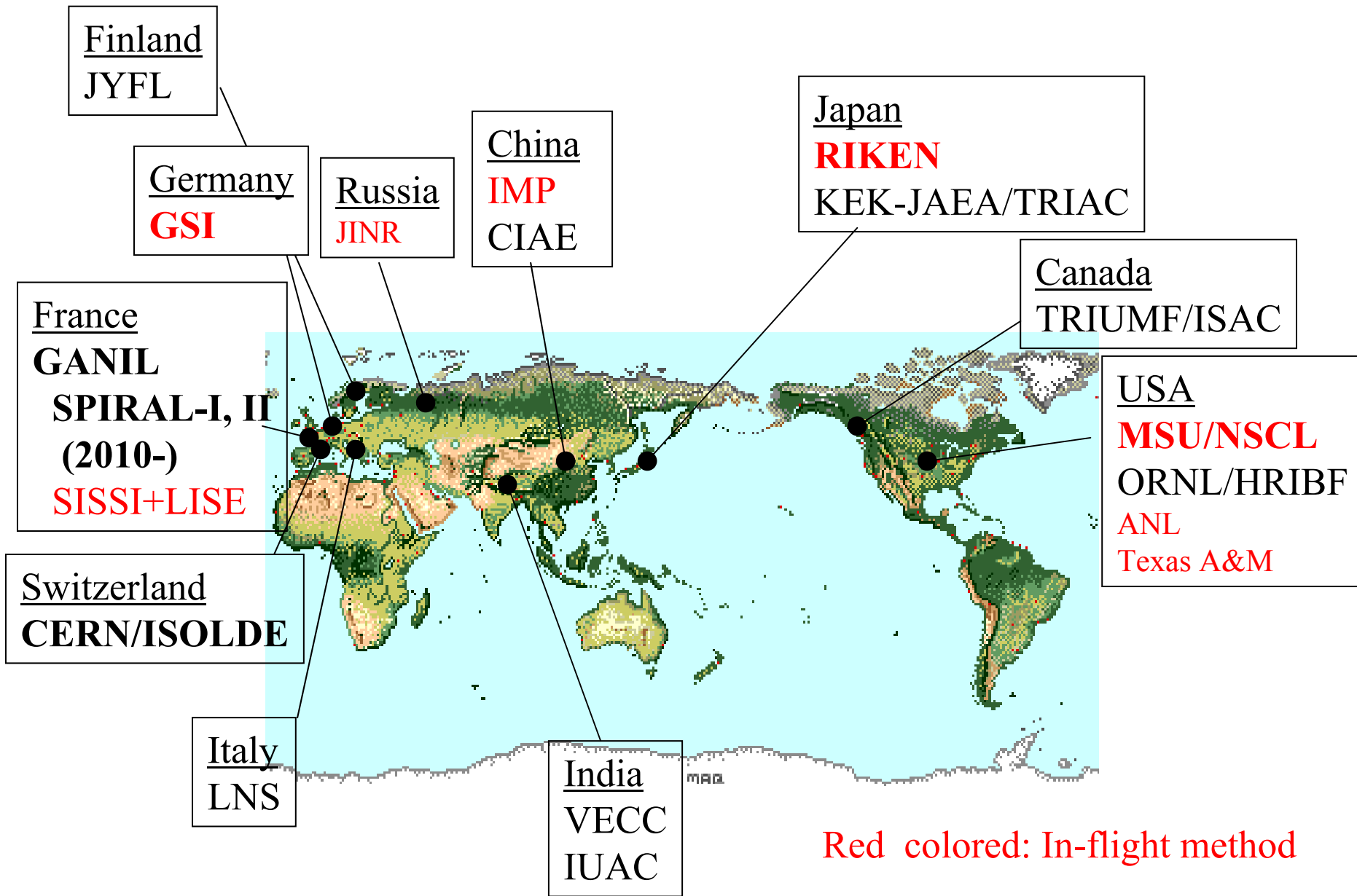
U Tokyo

SCRIT



2010-

# RI Beam Facilities in the world



Challenge

Action

Discussion

Enjoy