

## Stellar Evolution and Nucleosynthesis in Massive Stars

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Presupernova models and nucleosynthesis in massive stars are reviewed in the context of supernovae. Evolutionary models of massive stars toward the onset of collapse for the stellar mass of  $M \geq 10M_{\odot}$  in the main-sequence stage are presented. It is stressed that silicon shell burning is important to determine the final size of the iron core. As the result, we obtain the iron core mass ranges from 1.2 to 1.9  $M_{\odot}$  for 13–70  $M_{\odot}$  stars. As for nucleosynthesis, the s-process during core helium burning and p-process in the subsequent burning stage are presented in detail. The r-process during supernova explosion is discussed.

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