

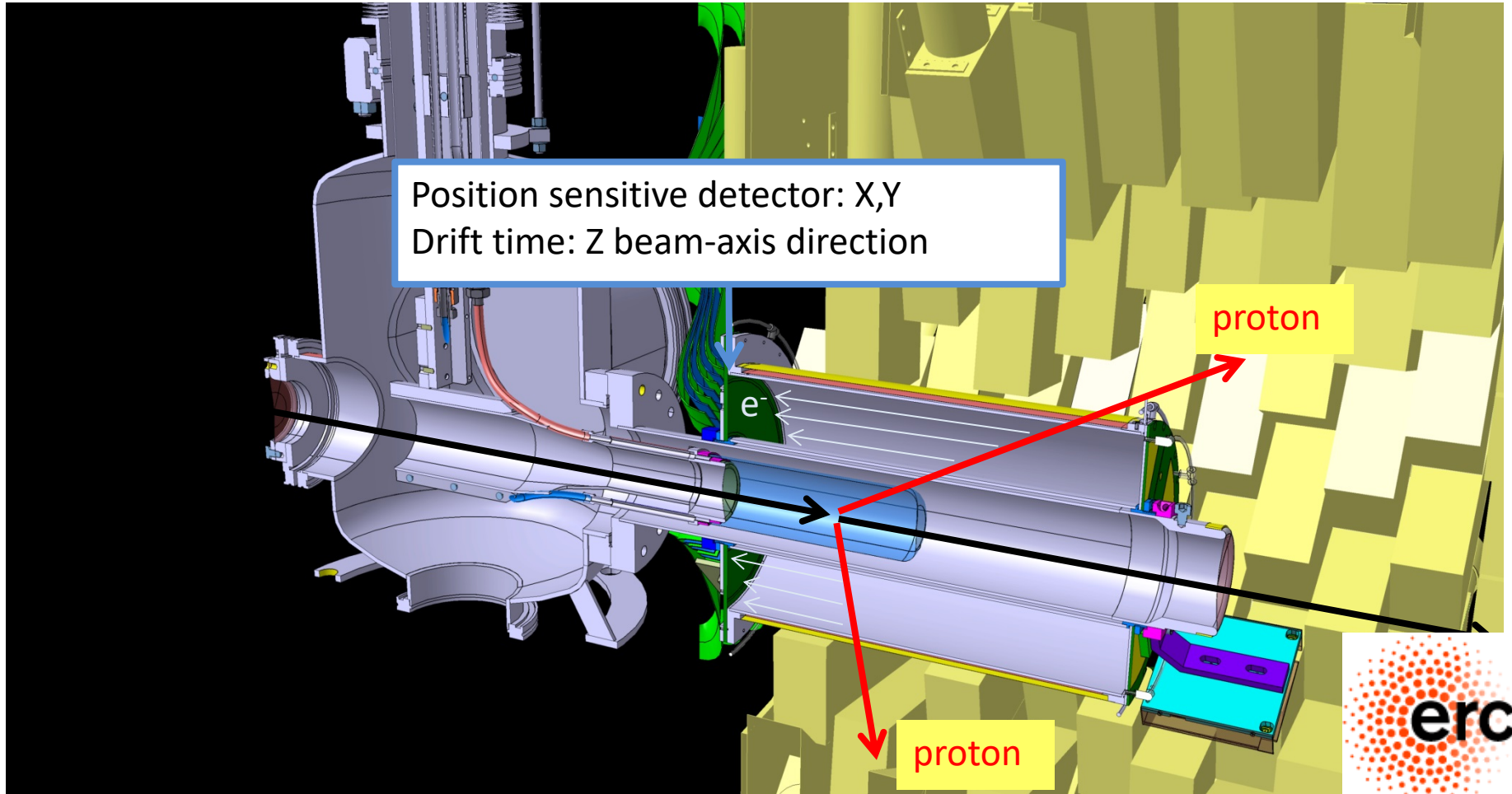
# Opportunities with MINOS+MINIBALL at RIBF

A.Corsi, CEA/IRFU/DPhN

High Resolution Gamma Spectroscopy at RIBF, Darmstadt, 10th-12th April 2019

Acknowledgments: D. Leboeuf, V.Panin (CEA), B.D.Linh (Vinatom)

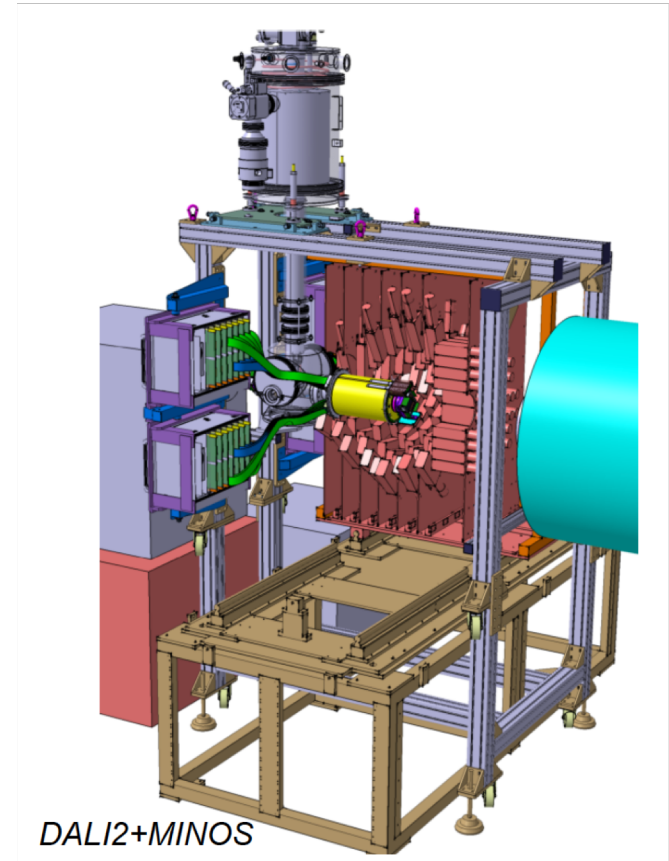
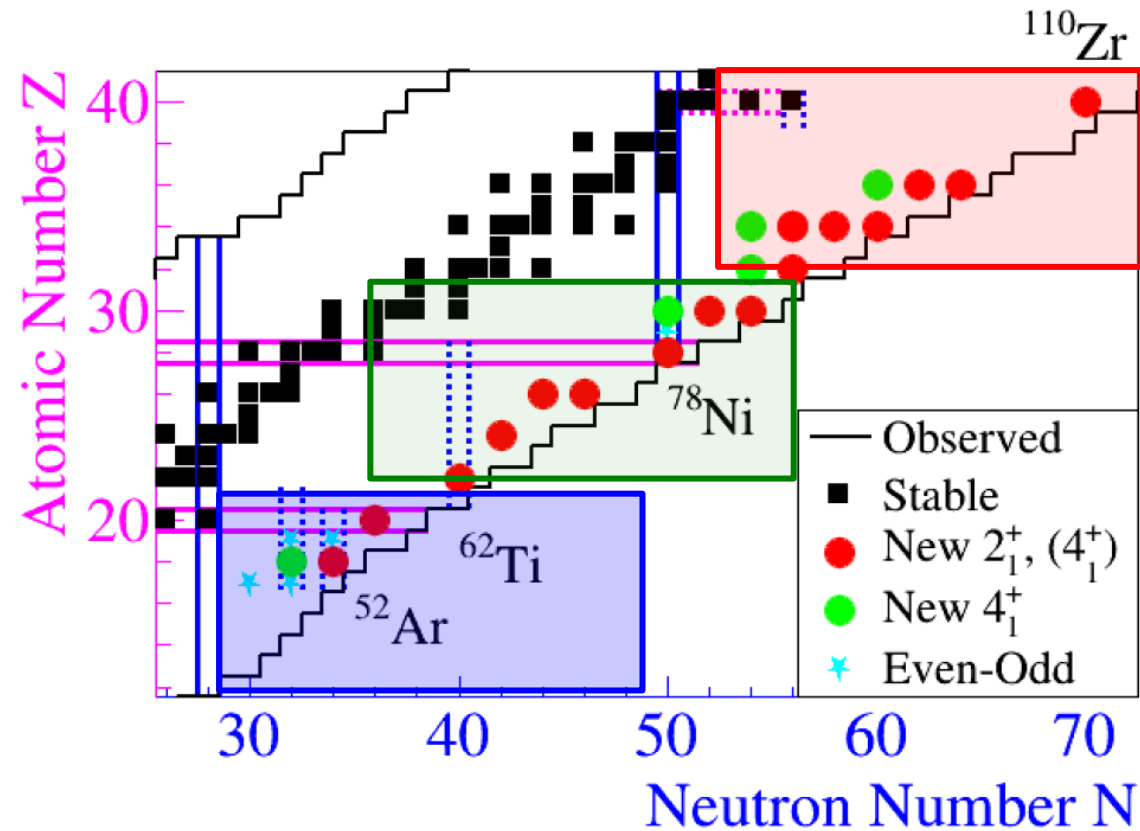
Thick target (5-15 cm) + vertex reconstruction for **Doppler and energy loss correction**



Efficiency: 95% in (p,2p) & resolution on vertex: 5 mm (FWHM)

Gamma spectroscopy of very neutron rich nuclei

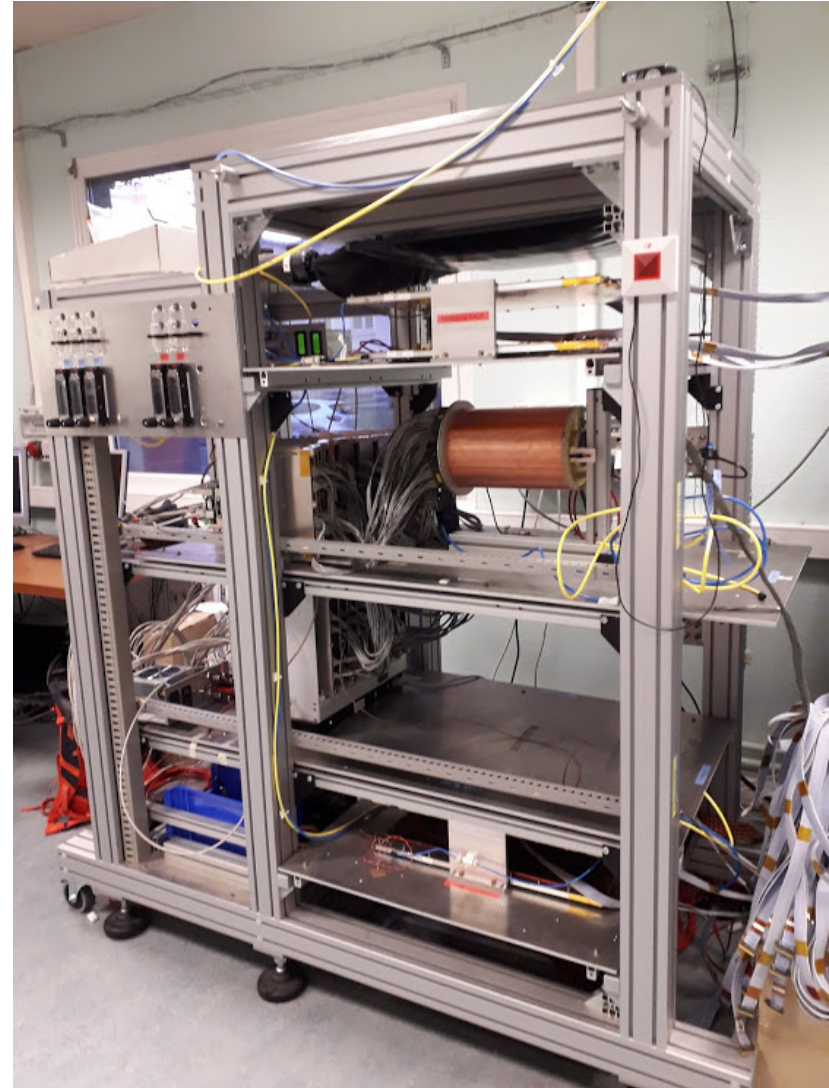
- May 2014 SEASTAR1
- May 2015 SEASTAR2
- May 2017 SEASTAR3



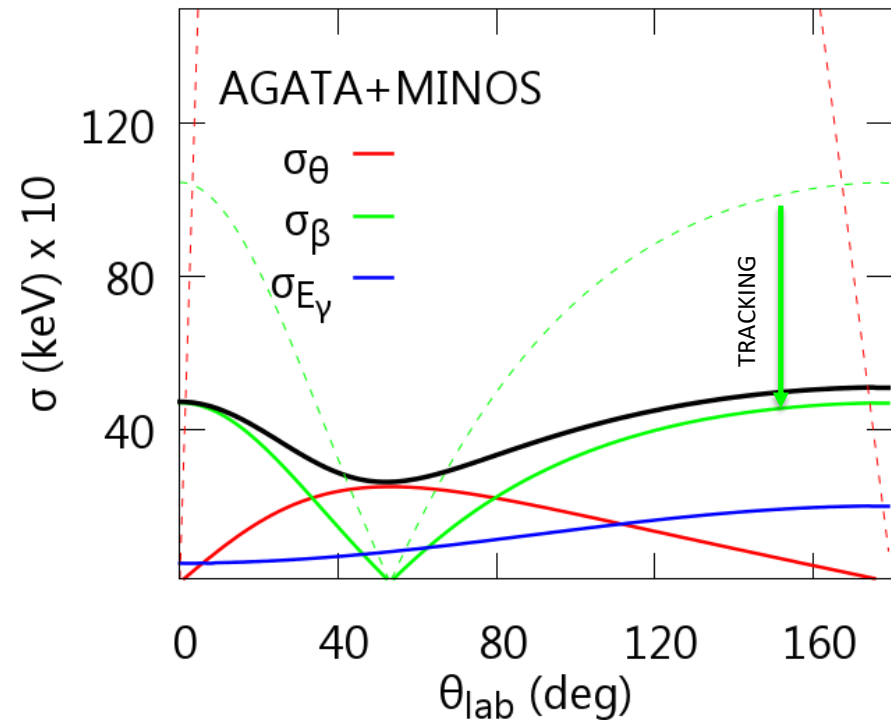
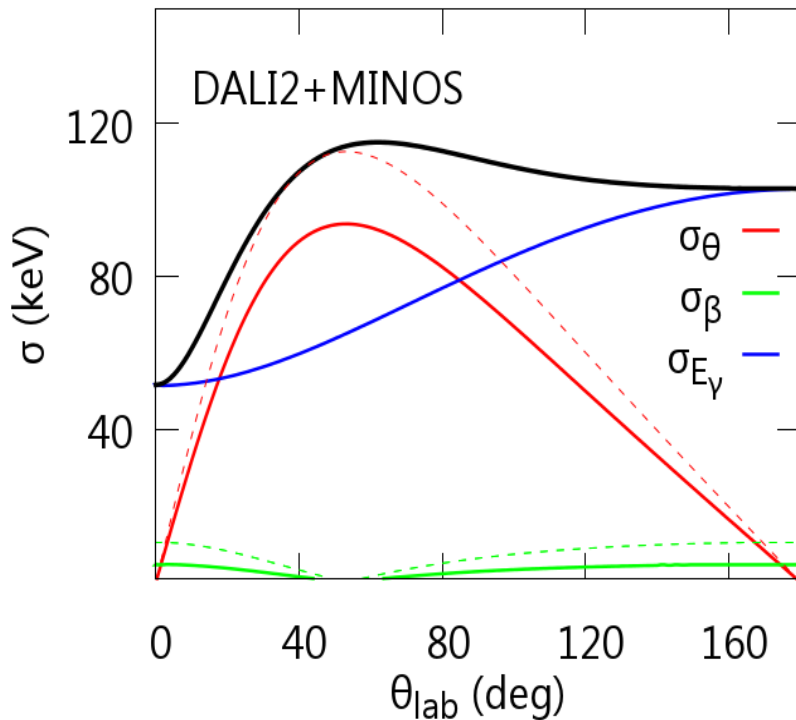
Spokespersons: P. Doornenbal (RIKEN), A. Obertelli (CEA, TU Darmstadt)

## Plans / ongoing

- Refurbishment of slow control system (hardware and software) ongoing at Saclay
- Test of cryogenic system in RIKEN this summer
- Reparation of TPC damaged before SEASTAR3
- Test of TPC in the cosmic bench at Saclay
- Support from EXPAND grant (N.Orr)
- MoU CEA-RIKEN under renewal until 2022

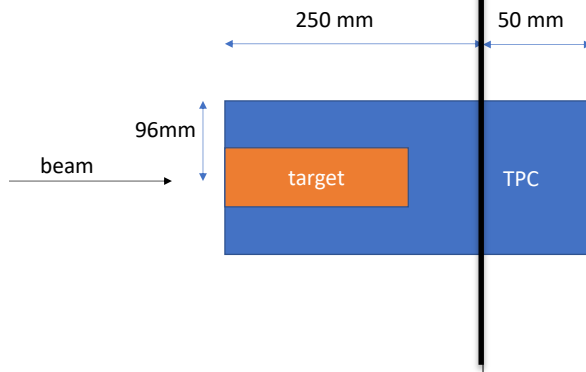
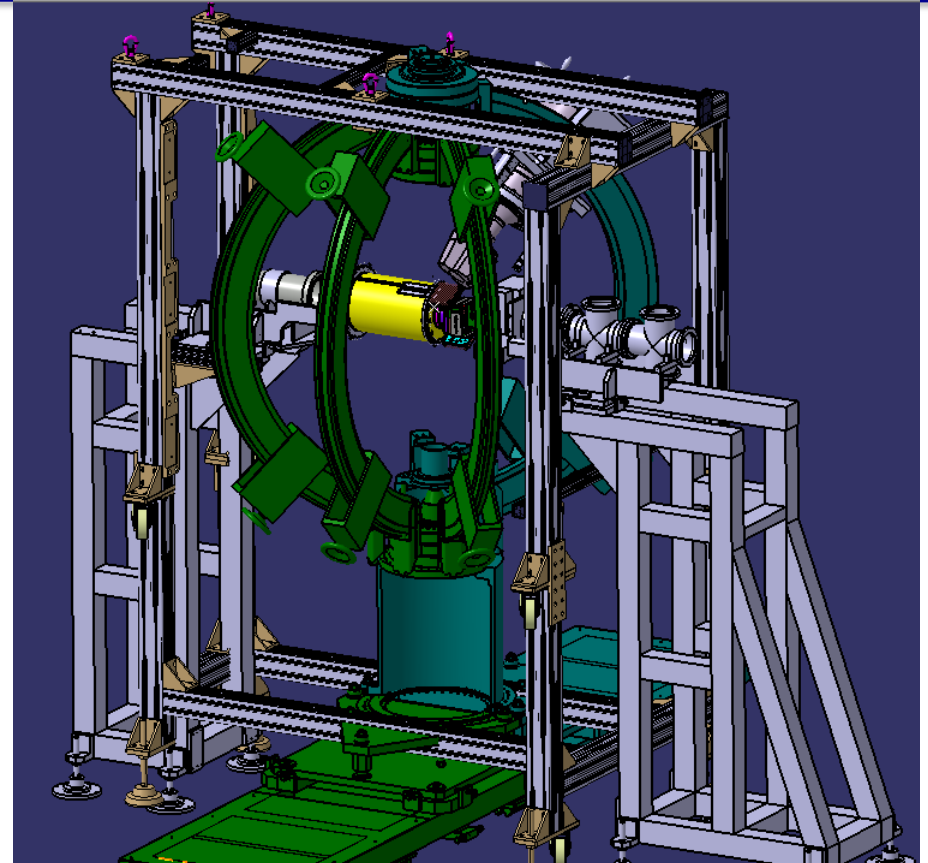
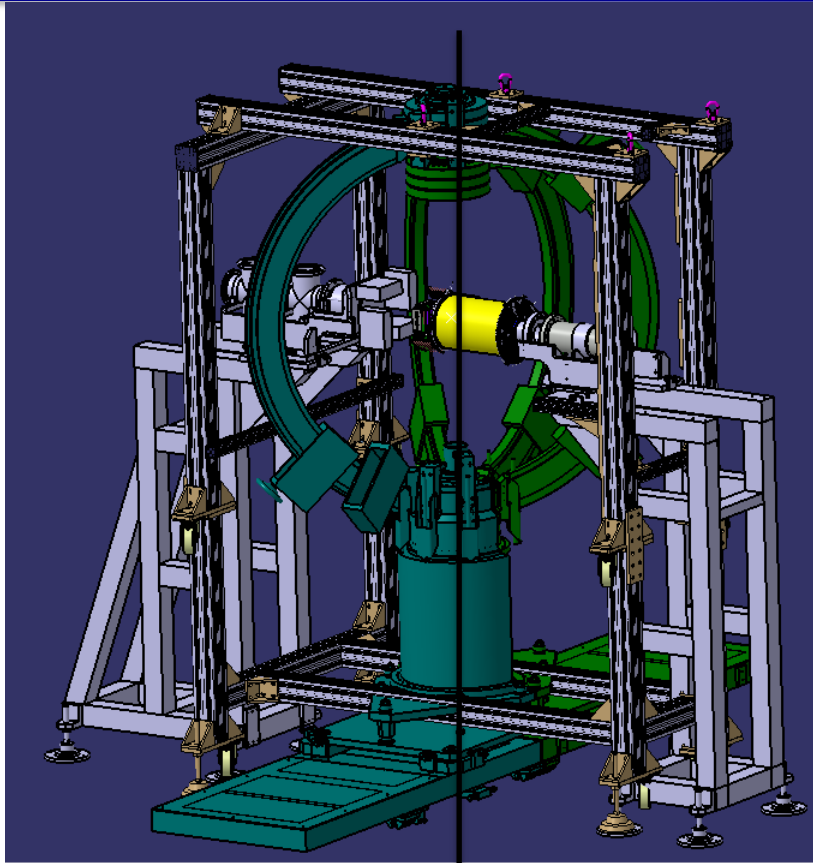


$E_\gamma=1$  MeV,  $\beta=0.6$ , 15 cm LH2 target



- Gamma energy and interaction point, and vertex resolution needed for good resolution!
- Efficiency?
- See Kathrin's talk

# Coupling MINOS and MINIBALL



- Few (minor) conflicts
- Simulations needed...

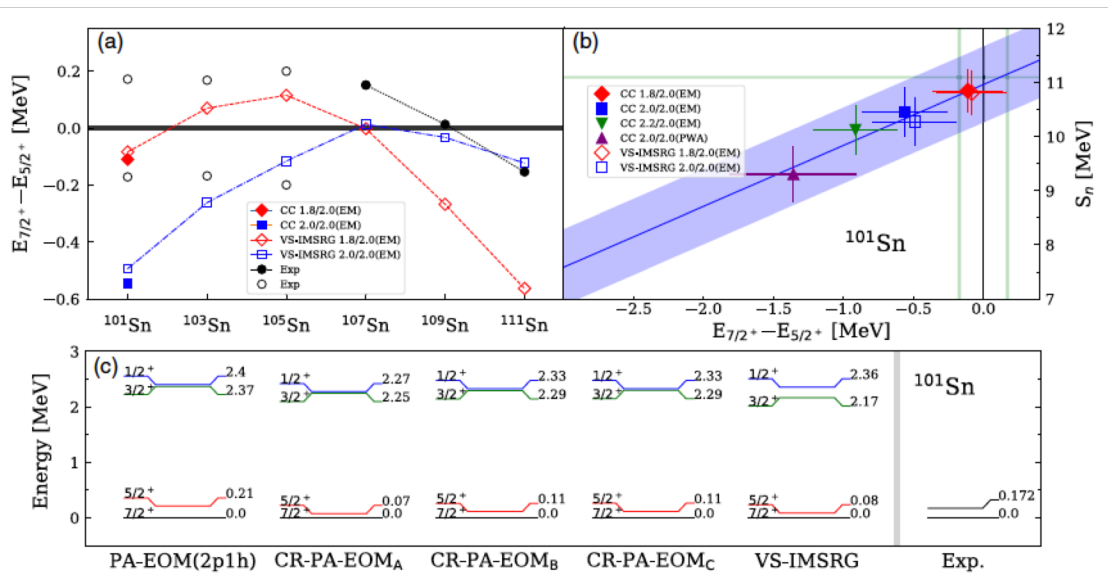
# 101Sn: ordering of d5/2 and g7/2

## STATE OF THE ART

- First excited state at 172 keV, no spin assignment
- Theory not (yet) enough precise

## PROPOSAL

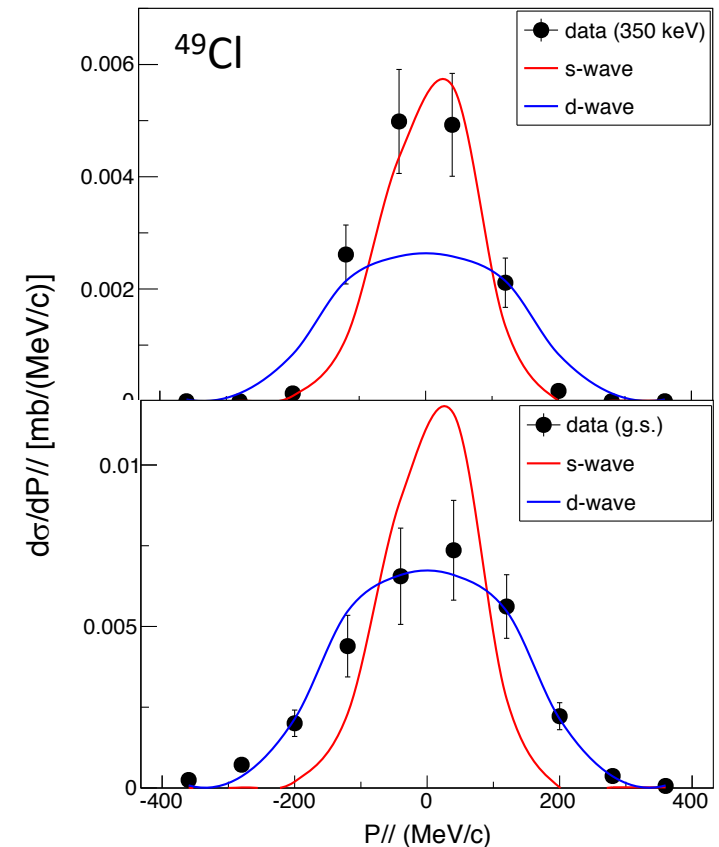
- $l=2$  and  $l=4$  from parallel momentum after  $^{102}\text{Sn}(p,pn)^{101}\text{Sn}$
- **Need to identify final state (MINIBALL) and measure  $p_{\parallel}$  (MINOS)**
- To be simulated:  $p_{\parallel}$  resolution for  $Z=50$  and 5 cm LH2,  $\gamma$  efficiency



Morris et al. PRL 120, 152503 (2018)

## SEASTAR3

$\sigma=44$  MeV/c with Ar and 15 cm LH2 target

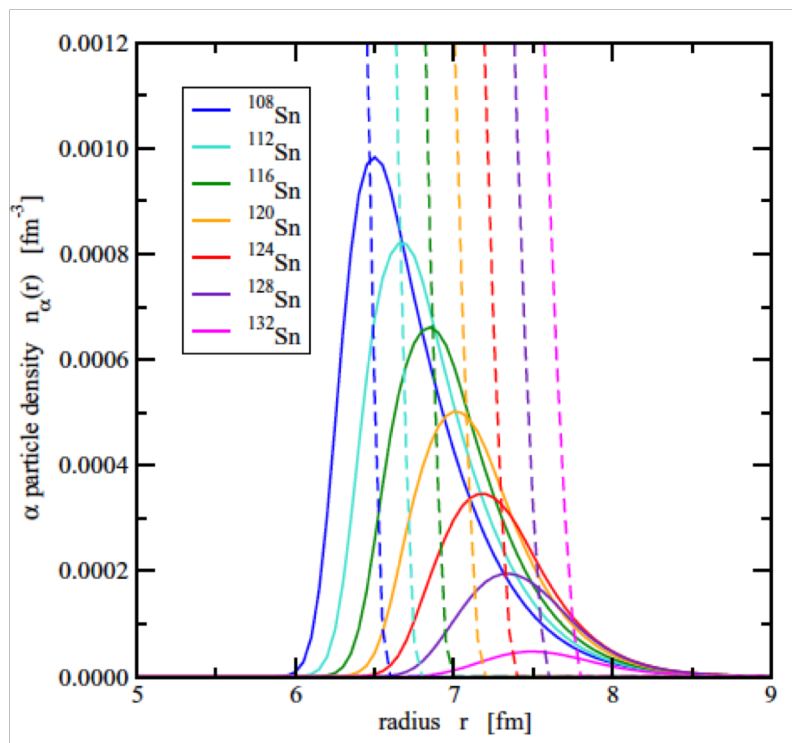


Courtesy B.D.Linh, Calculations: A.Moro

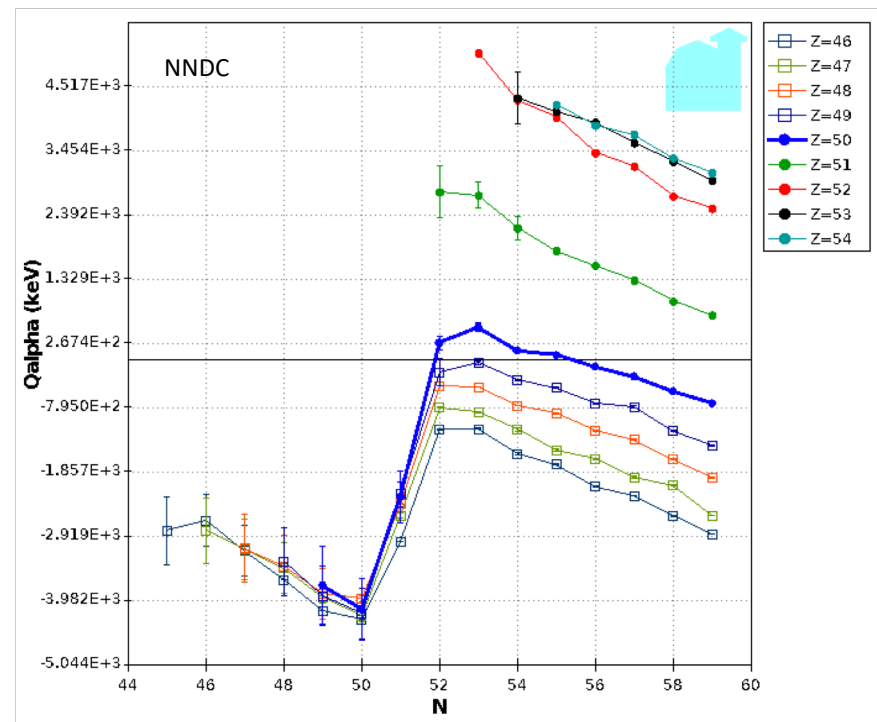
Spectroscopy of  $^{100}\text{Sn}$ , J.Lee, K.Wimmer, A.Corsi  
3 days approved for cross section measurement, scheduled in May

## STATE OF THE ART

- Clustering at the surface predicted to decrease when neutron skin develops
- Recent results from  $(p,p\alpha)$  at RCNP (112-124Sn) submitted to Nature
- Resolution on the final state critical, especially in inverse kinematics



S.Typel et al., PRC 89, 064321 (2014)



$\alpha$  separation energy sensitive to shell closure!

Spectroscopy of  $^{100}\text{Sn}$ , J.Lee, K.Wimmer, A.Corsi  
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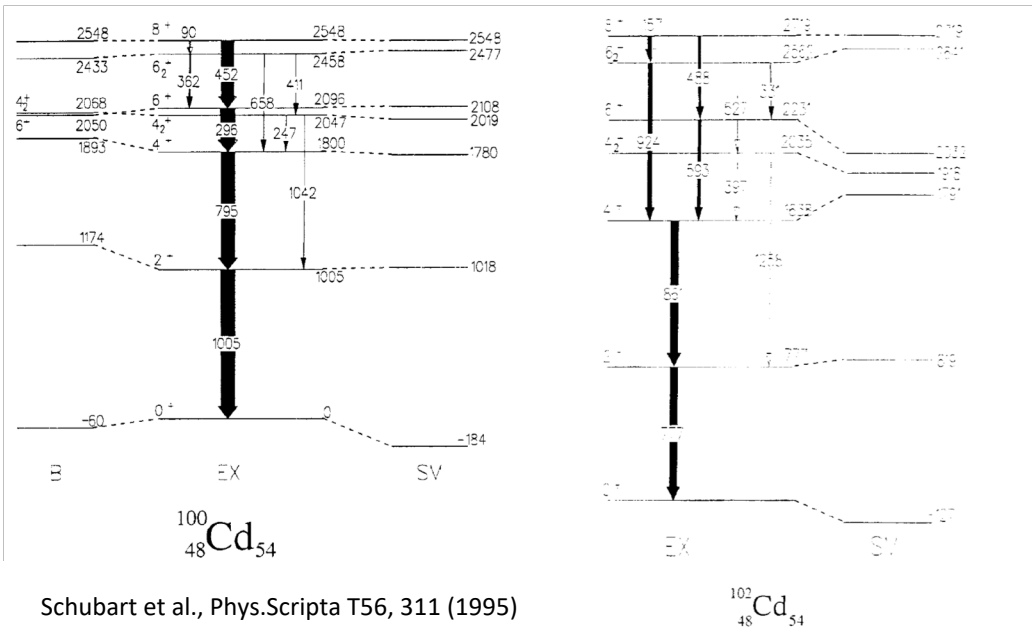


## PROPOSAL:

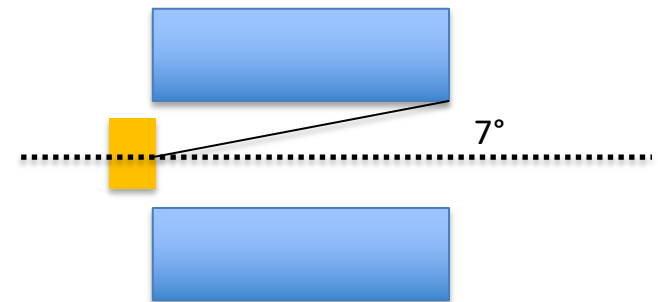
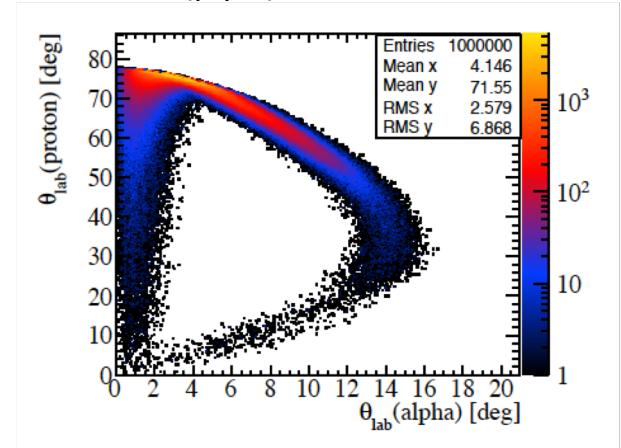
- Explore clustering for neutron deficient/rich Tin
- $^A\text{Sn}(p,p\alpha)^{A-4}\text{Cd}$
- projectile and A-4 have almost same magnetic rigidity (ZDS ok)

## REQUIREMENTS

- energy resolution for the final state
- measure scattering angles down to few degrees ( $\alpha$ )



$^{104}\text{Sn}(p,p\alpha)^{100}\text{Cd}$  at  $\sim 150$  MeV/u



Spectroscopy of  $^{100}\text{Sn}$ , J.Lee, K.Wimmer, A.Corsi  
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