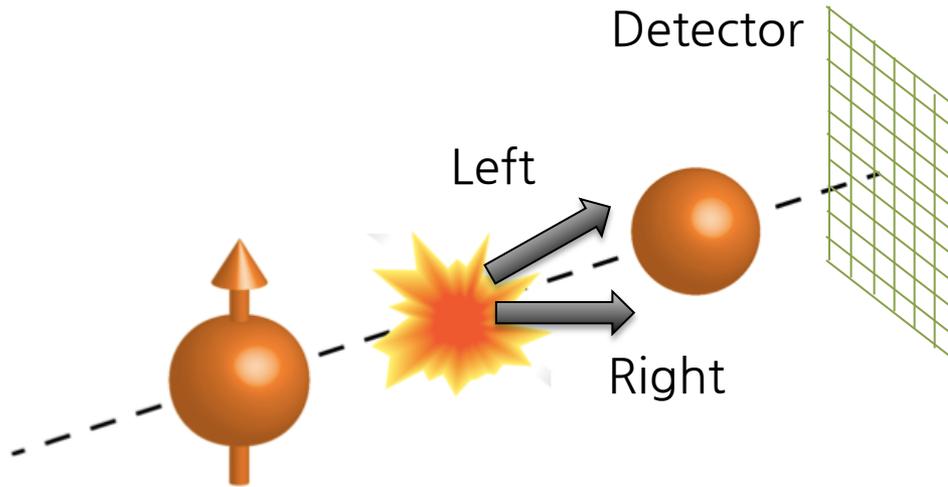


# Surprising transverse single spin asymmetry of $\pi^0$ at almost zero degree in 510 GeV polarized $p + p$ collisions

Minho Kim (Korea Univ./ RIKEN)  
on behalf of the RHICf collaboration



# Transverse single spin asymmetry ( $A_N$ )

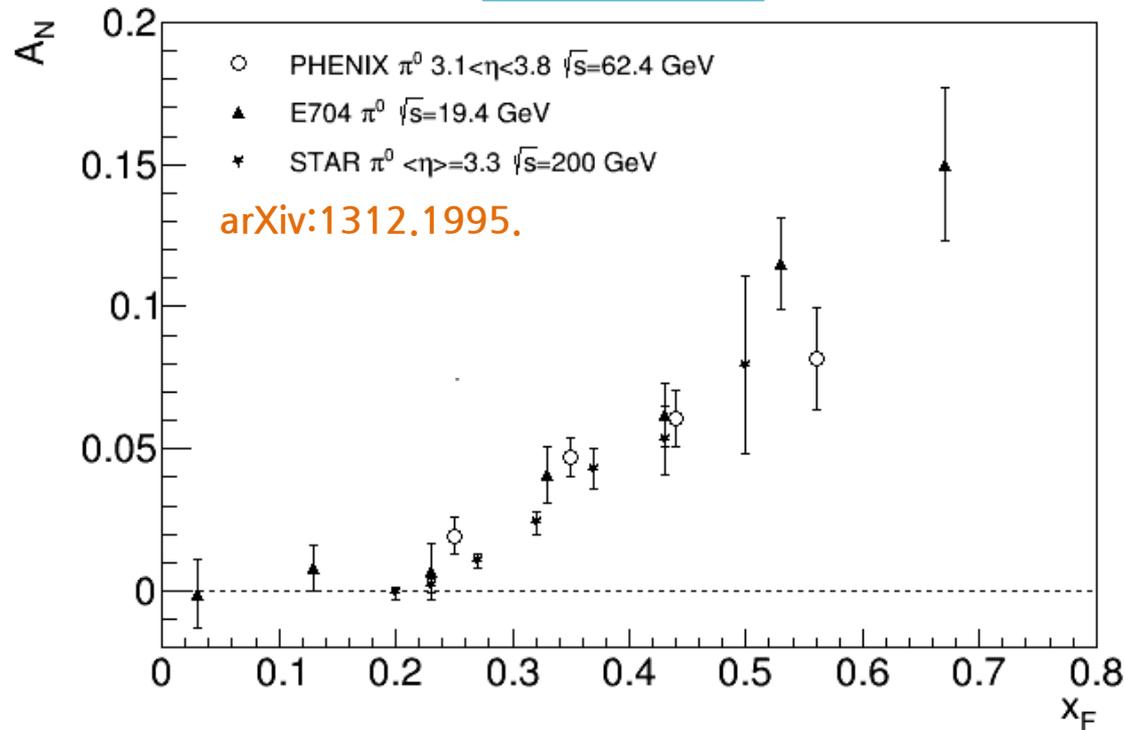


$$A_N = \frac{\sigma_L^\uparrow - \sigma_R^\uparrow}{\sigma_L^\uparrow + \sigma_R^\uparrow} = \frac{\sigma_L^\uparrow - \sigma_L^\downarrow}{\sigma_L^\uparrow + \sigma_L^\downarrow}$$

- In polarized p+p collision,  $A_N$  is defined as a left-right cross section asymmetry of a specific particle.
- Non-zero  $A_N$  of  $\pi^0$  has been a starting point of the study for the parton's spin-related transversity and more detailed interaction mechanism among quarks and gluons.

# $A_N$ in forward $\pi^0$ production

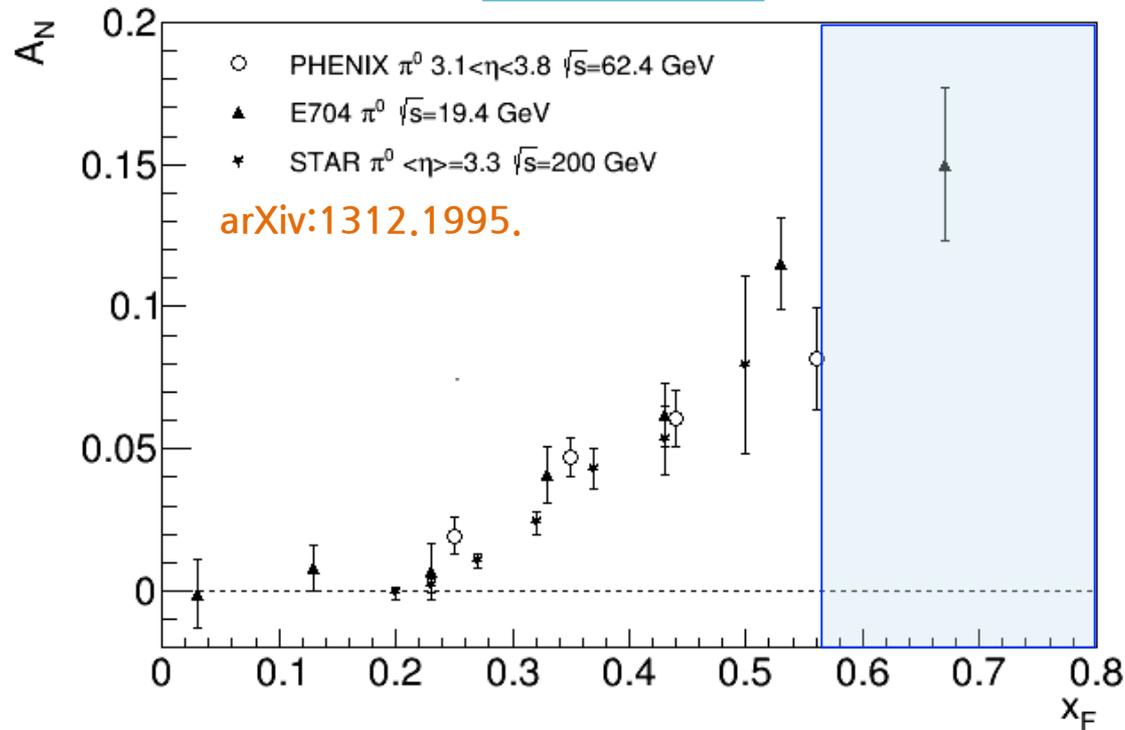
$$3 < \eta < 4$$



- Observed non-zero  $A_N$  of  $\pi^0$  ever has been interpreted based on only perturbative picture theoretically.

# $A_N$ in forward $\pi^0$ production

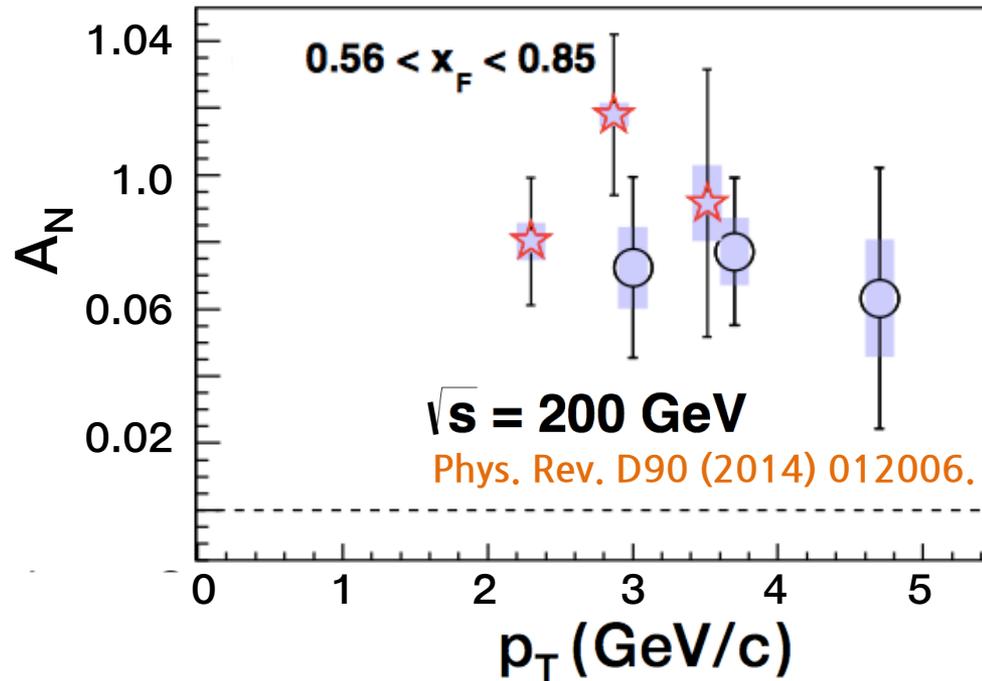
$$3 < \eta < 4$$



- Observed non-zero  $A_N$  of  $\pi^0$  ever has been interpreted based on only perturbative picture theoretically, **but NOT exactly understood yet.**

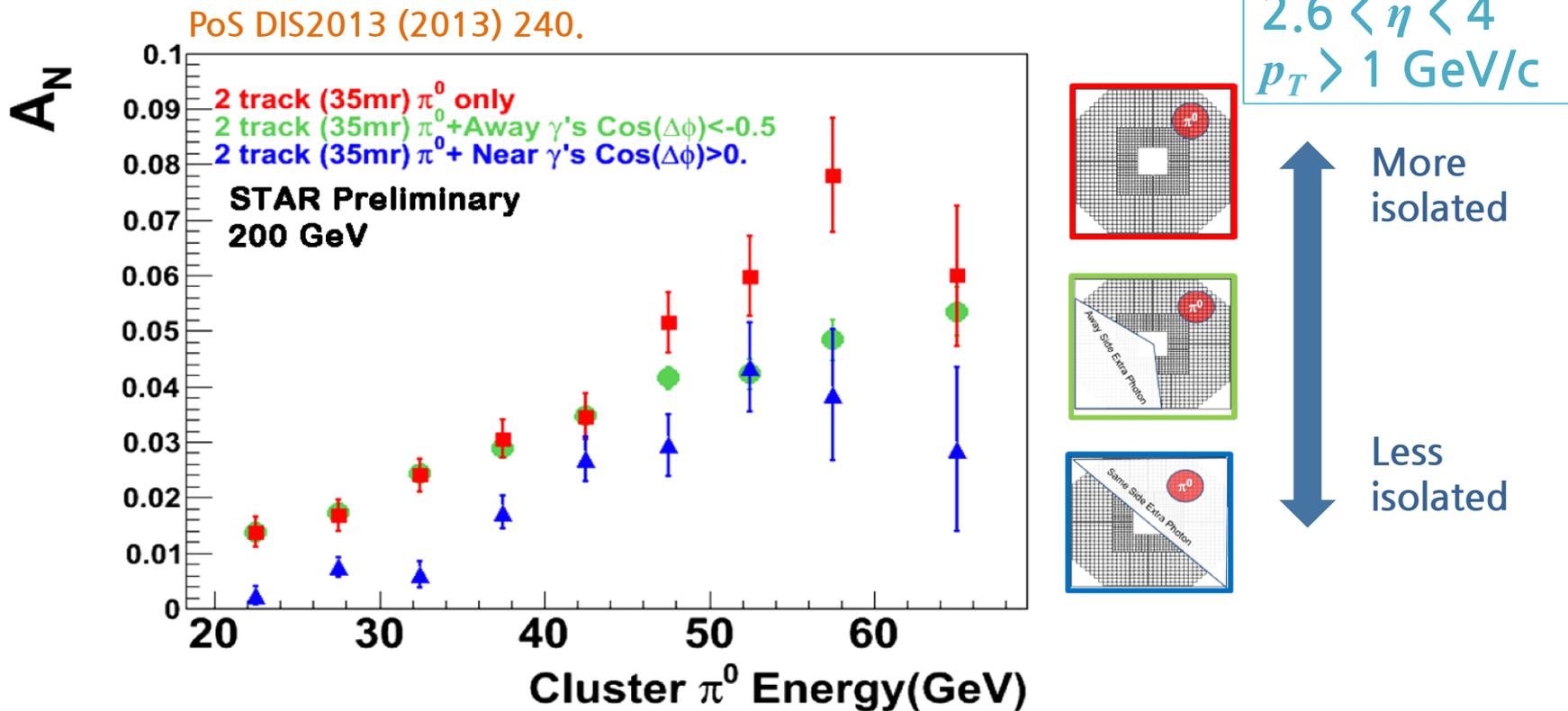
# $A_N$ in forward $\pi^0$ production

- $p^\uparrow + p \rightarrow \text{Cluster} + X$  (PHENIX)
- ☆  $p^\uparrow + p \rightarrow \pi^0 + X$  (STAR)



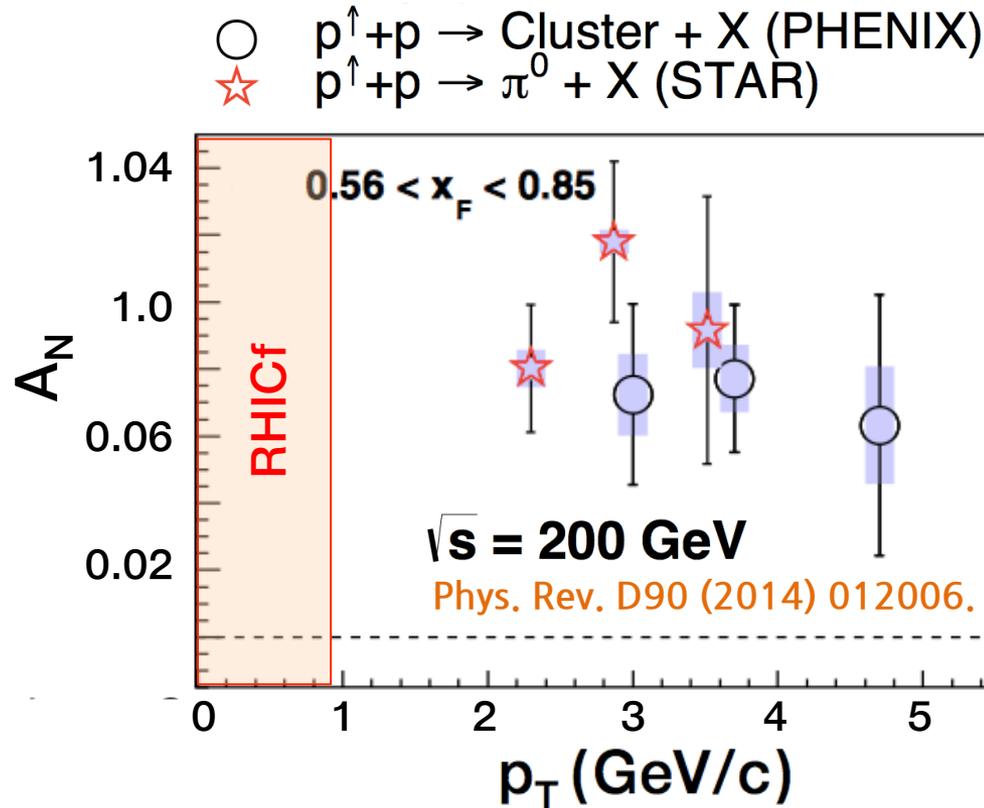
- Observed non-zero  $A_N$  of  $\pi^0$  ever has been interpreted based on only perturbative picture theoretically.
- No experimental necessity and effort to measure the very forward  $\pi^0$  production below 1 GeV/c.

# New question to the $A_N$ of forward $\pi^0$



- Larger  $A_N$  was observed by more isolated  $\pi^0$  than less isolated one.
- Non-perturbative process may have a finite contribution to the  $\pi^0$   $A_N$  as well as perturbative one.

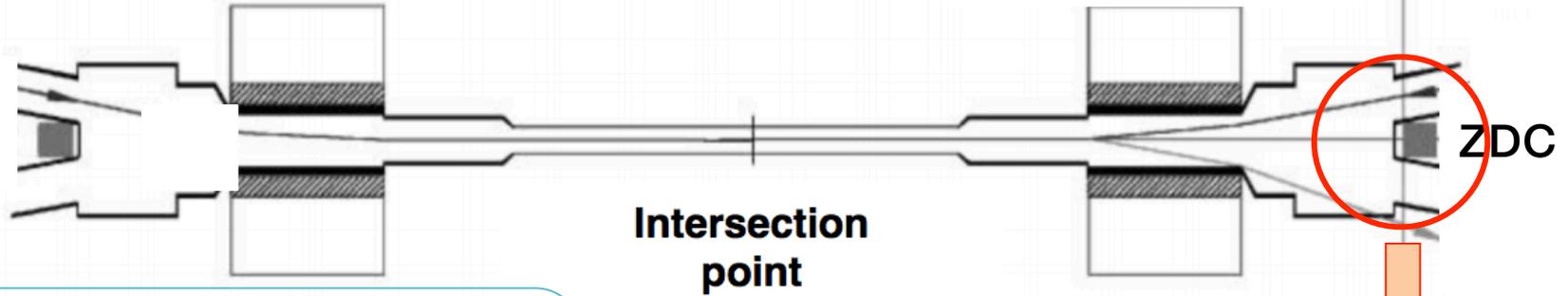
# $A_N$ in forward $\pi^0$ production



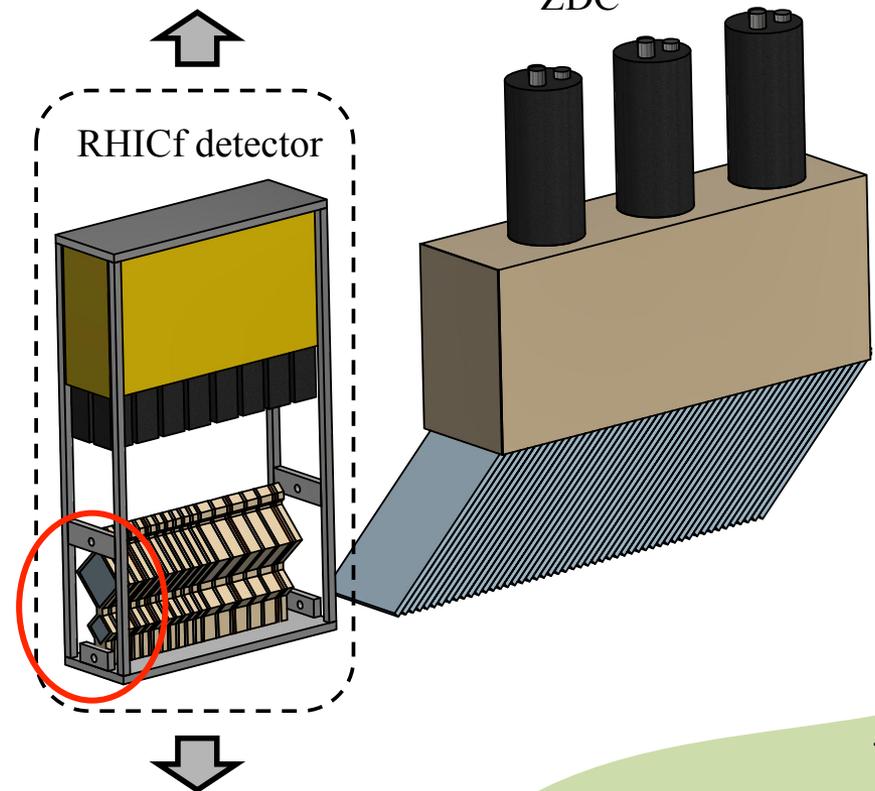
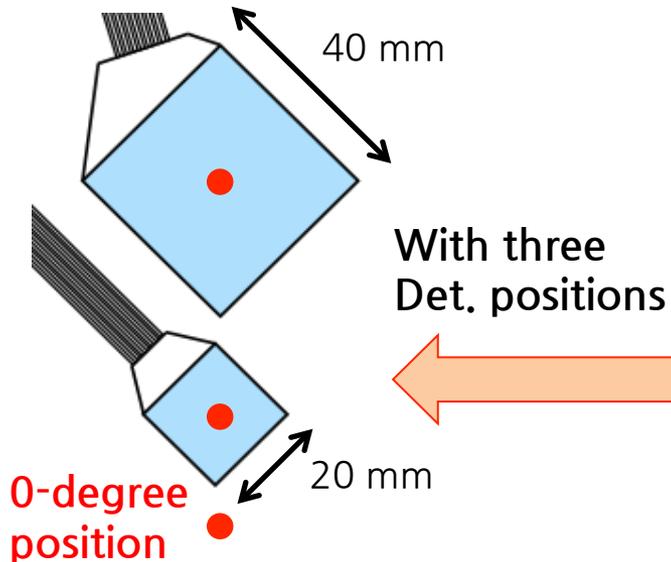
- Now, we have a strong motivation to measure the  $A_N$  of very forward  $\pi^0$  production.
- RHICf experiment measured very forward ( $6 < \eta$ )  $\pi^0$   $A_N$  to study the non-perturbative contribution to it in detail.

# RHIC forward (RHICf) experiment

STAR experiment



- 18 m away from IP.
- $0.2 < x_F < 1.0$ .
- $0.0 < p_T < 1.0$  GeV/c.

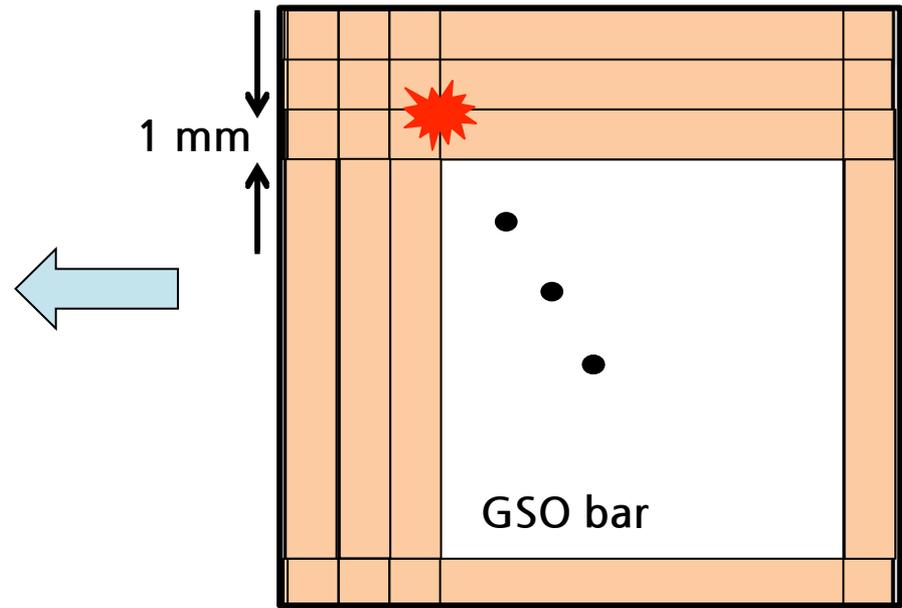
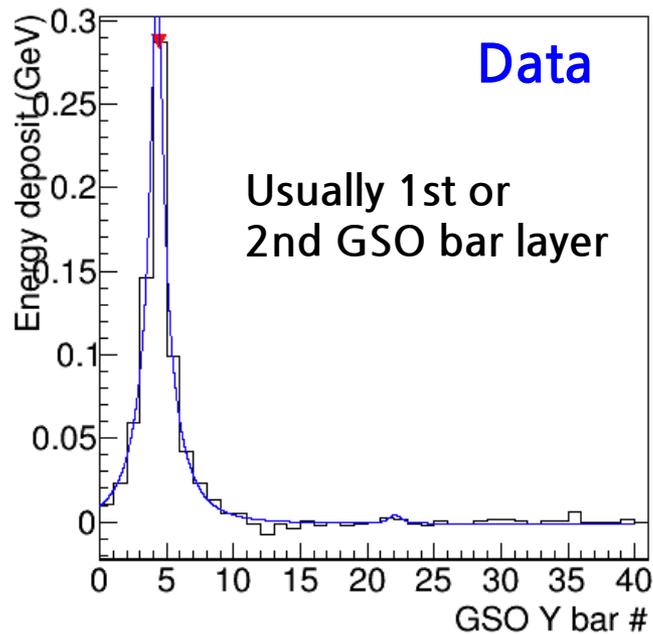
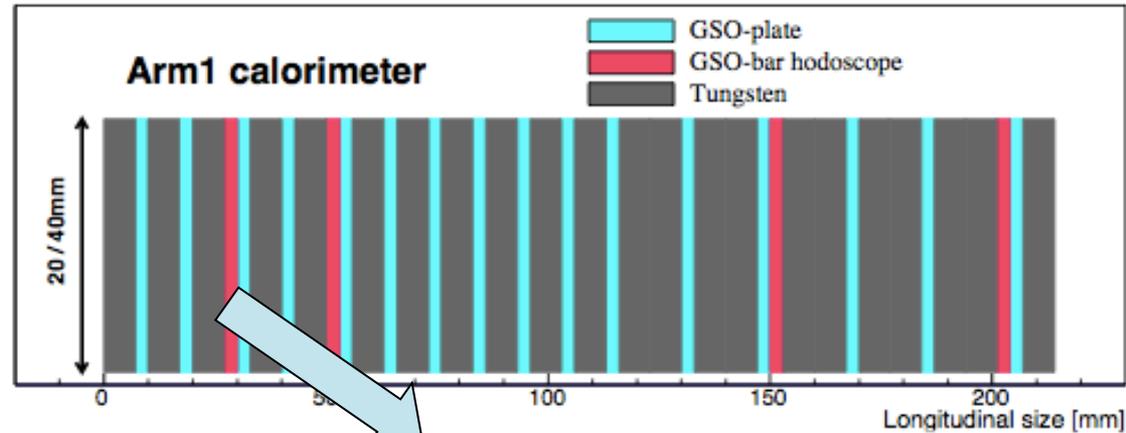




# Position reconstruction of photon

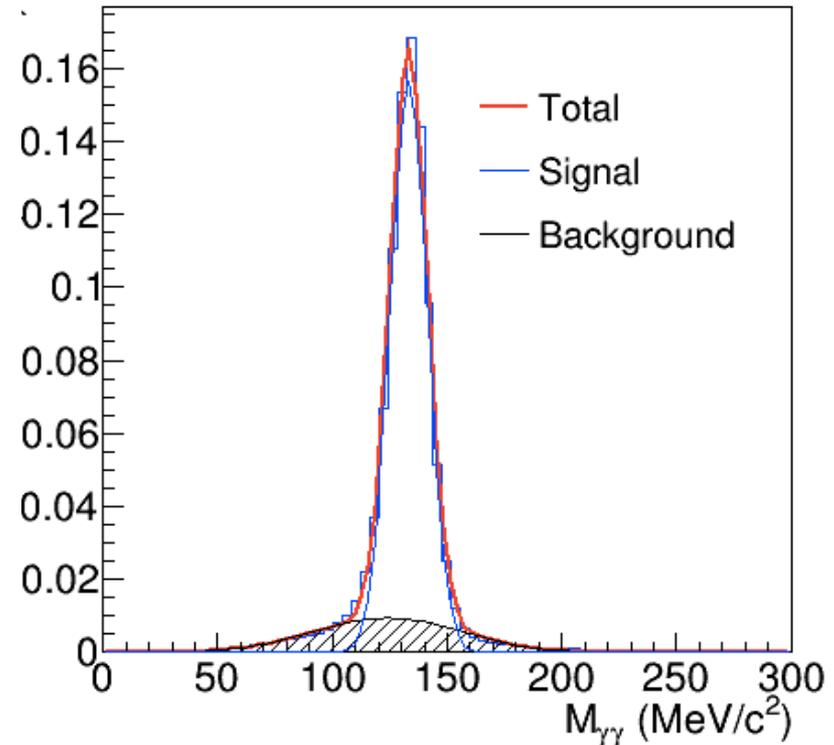
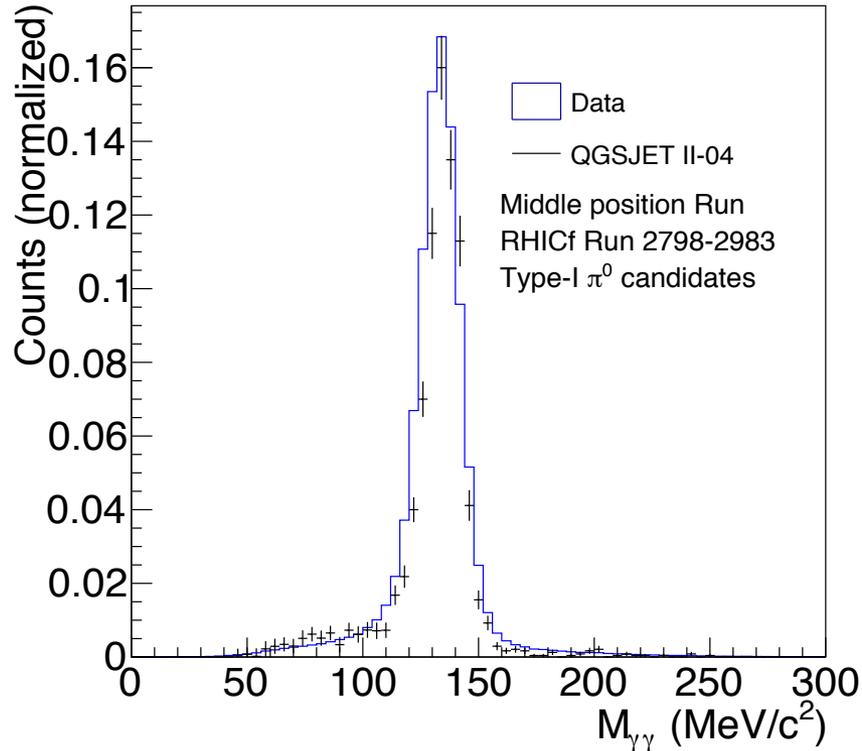
- If a photon hit a tower,

$\gamma$  →



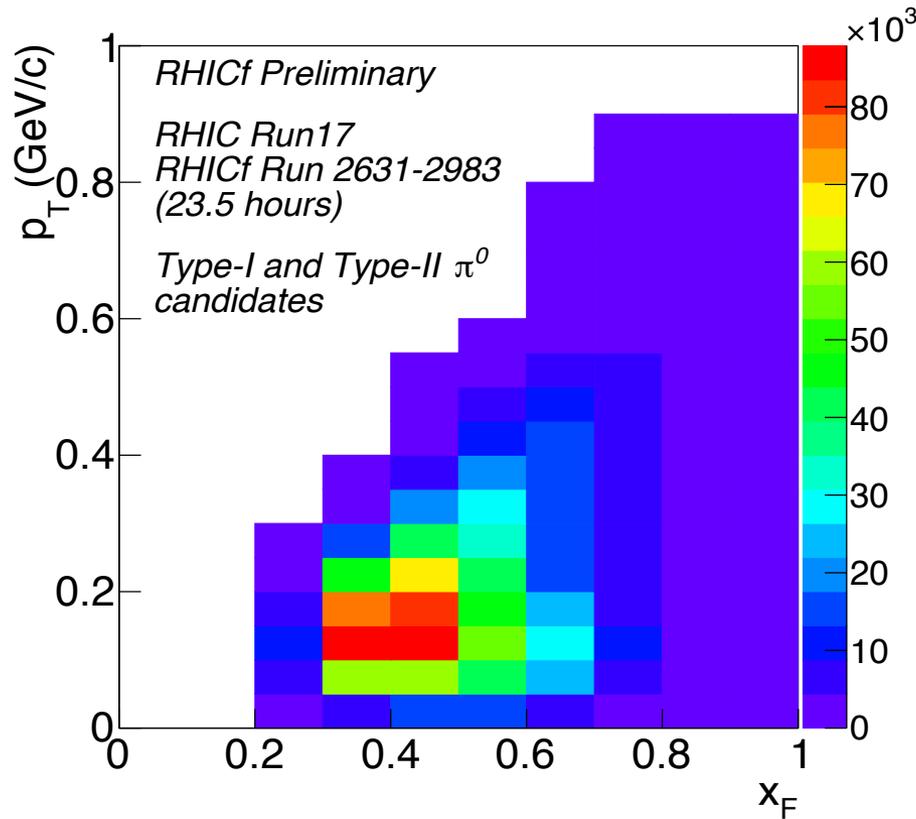


# Invariant mass of two photons



- Data is well matched with simulation showing clear  $\pi^0$  peak around 135 MeV/c<sup>2</sup> with  $\sim 8$  MeV/c<sup>2</sup> peak width.
- Invariant mass was fitted by polynomial for background and Gaussian for  $\pi^0$ .
- Background part usually comes from coincidence of two photons from different  $\pi^0$ , not wrong reconstruction.

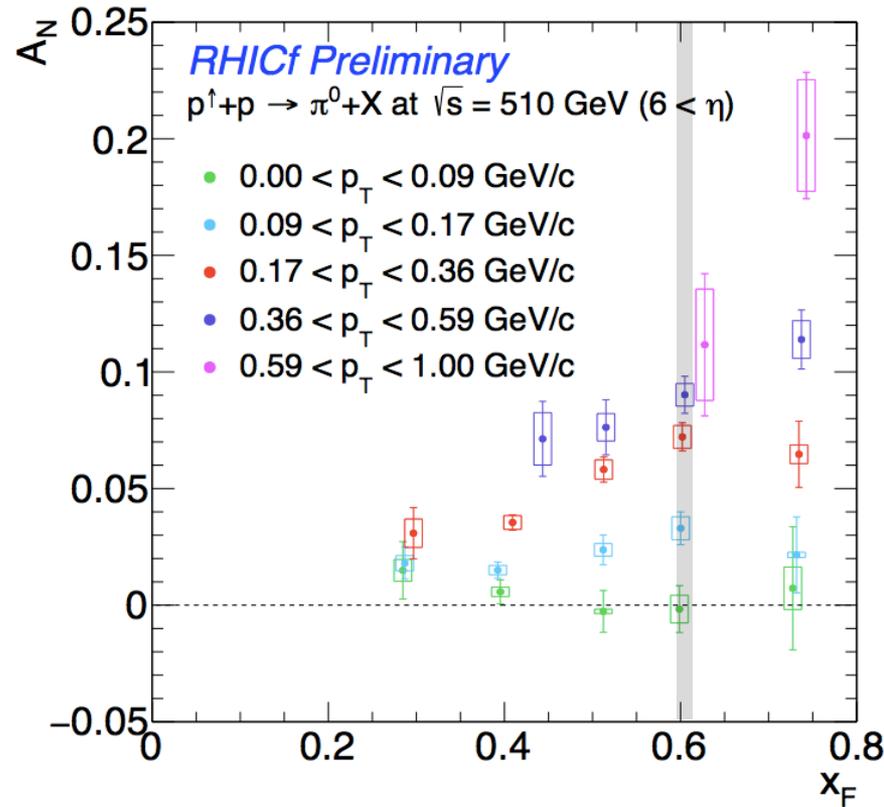
# $\pi^0$ kinematics & $A_N$ calculation



$$A_N = \frac{1}{P} \frac{1}{D_\phi} \left( \frac{N^\uparrow - RN^\downarrow}{N^\uparrow + RN^\downarrow} \right)$$

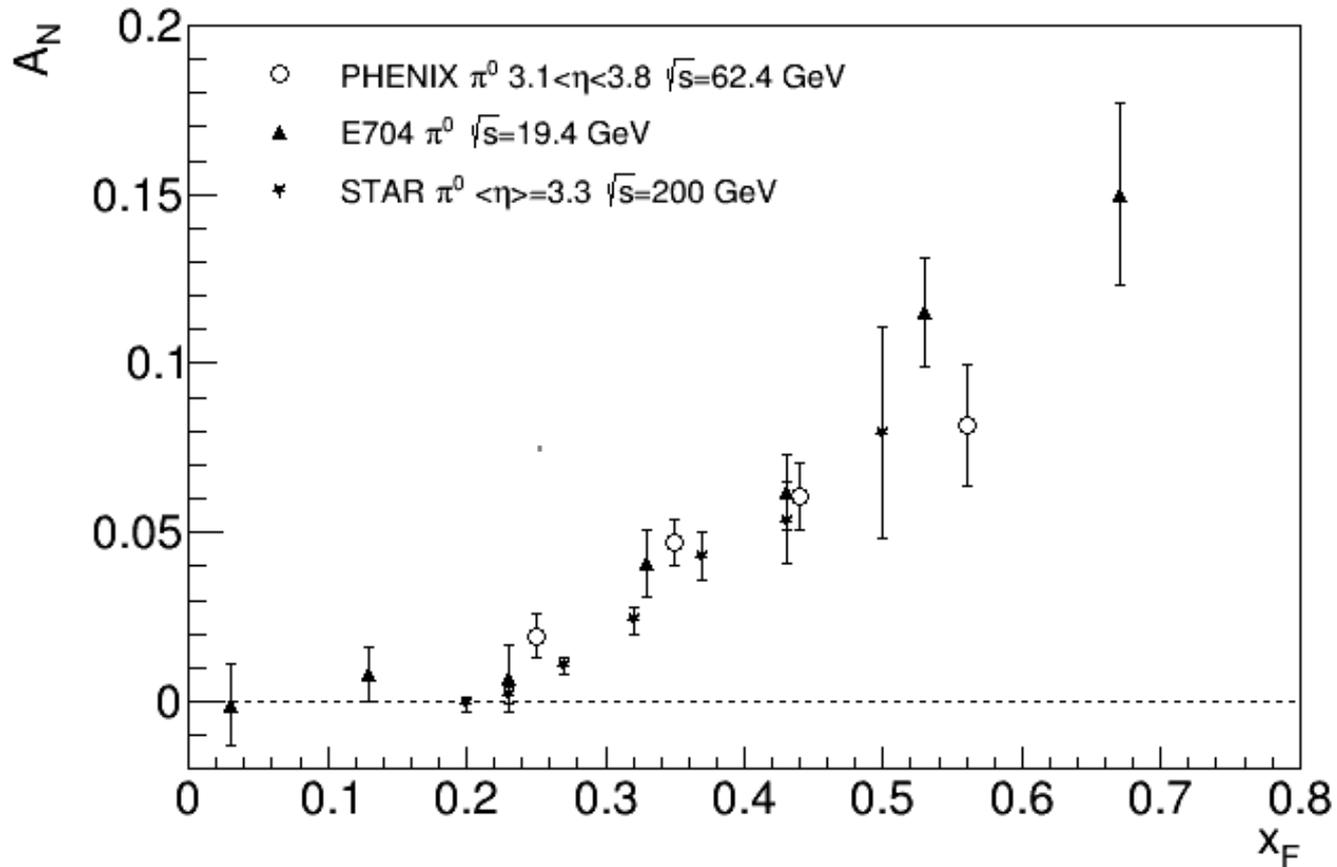
- Very forward  $\pi^0$  over the  $x_F$  range of  $0.2 < x_F < 1.0$  and  $p_T$  range of  $0.0 < p_T < 1.0$  GeV/c was measured.
- Systematic uncertainties by polarization,  $\pi^0$  azimuthal angle, background  $A_N$  subtraction, and beam center was included.

# Very forward $\pi^0 A_N$ as a function of $x_F$

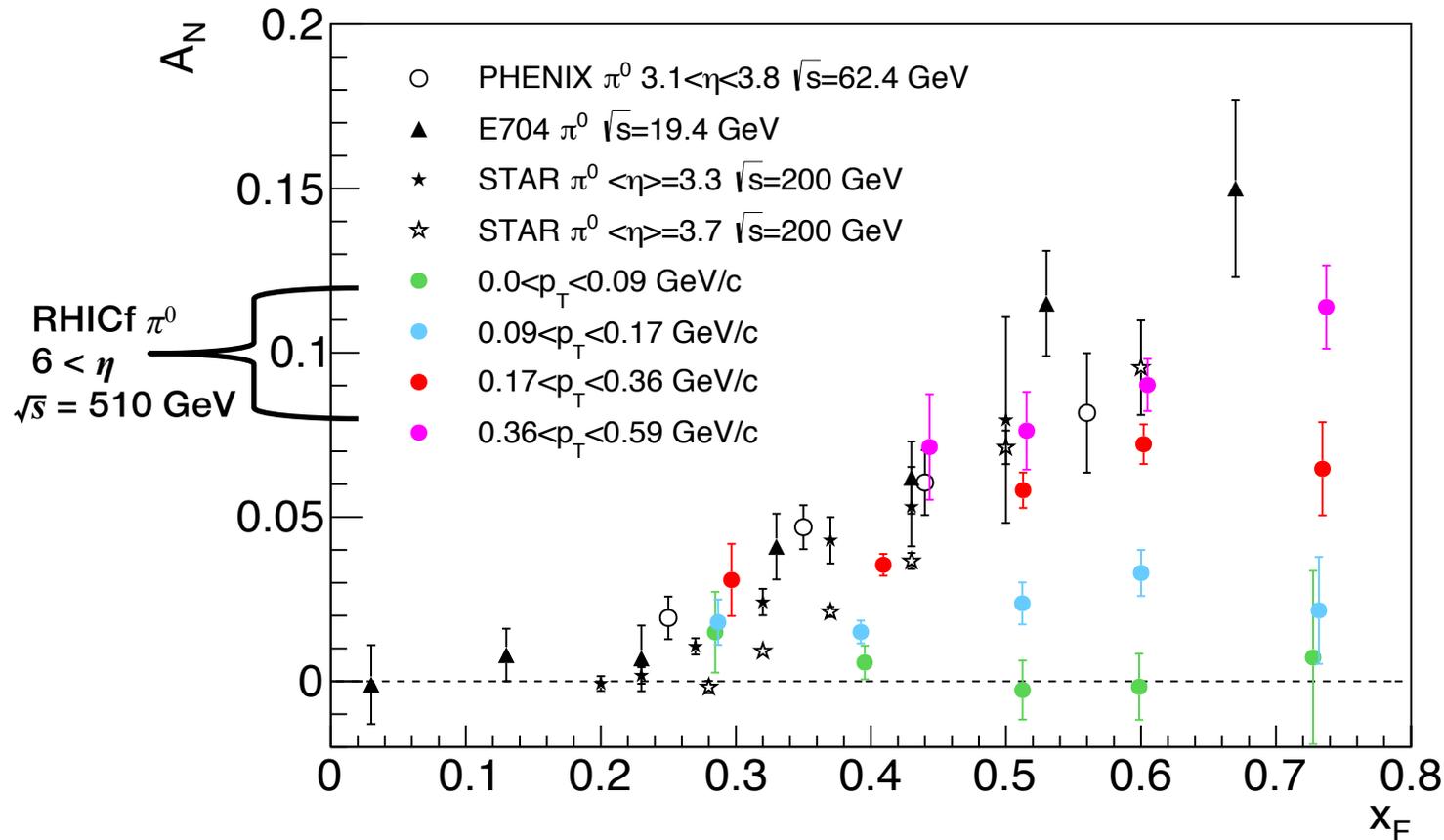


- The higher  $p_T$  range the  $A_N$  is measured in, the more clearly it increases as a function of  $x_F$ .
- Note that  $x_F$  resolutions of the RHICf detector are also much finer than  $x_F$  binning in the preliminary plot.

# Comparison with previous measurements

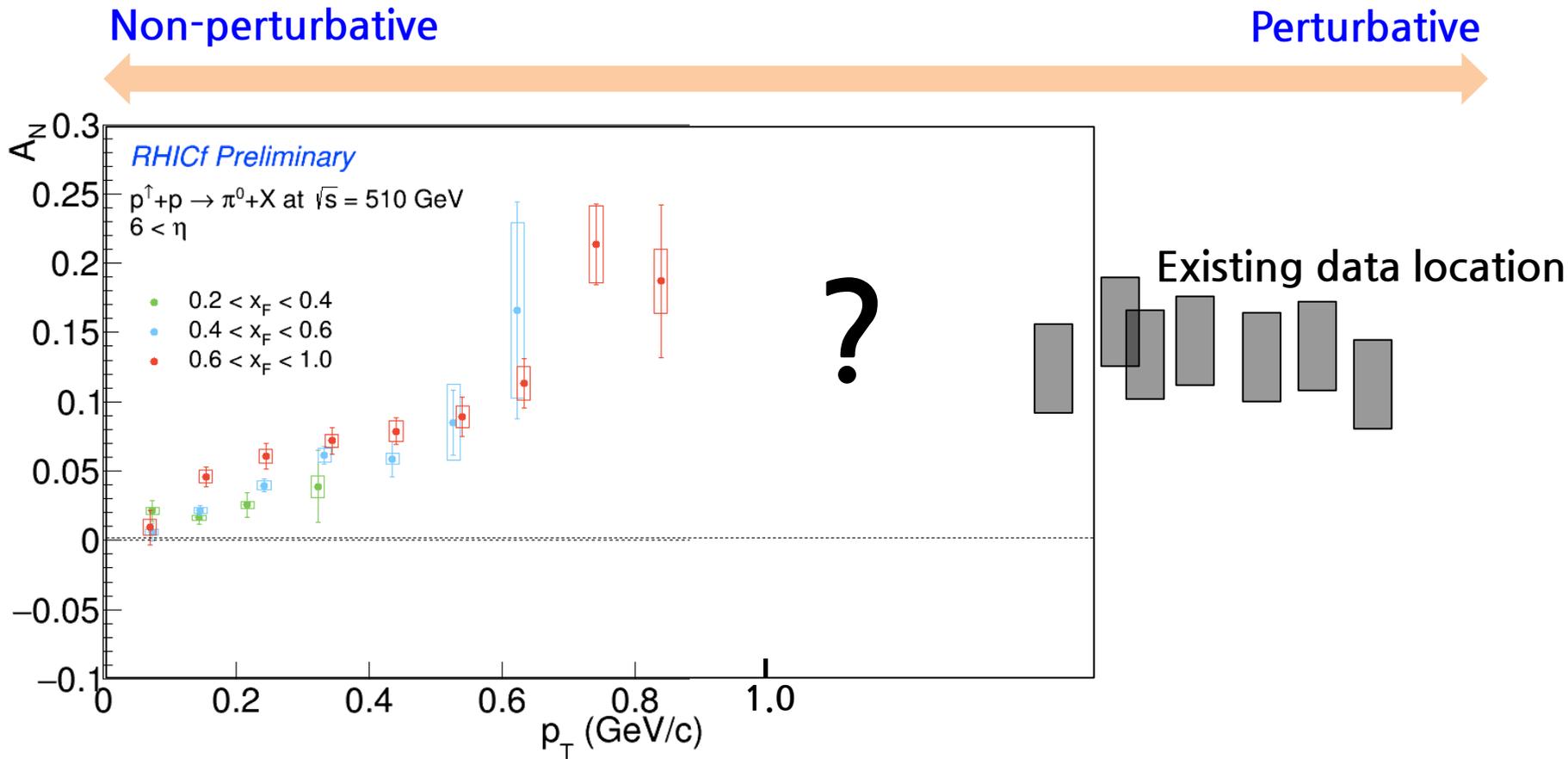


# Comparison with previous measurements



- Is there perturbative contribution even in lower  $p_T$  area?
- The origin of  $x_F$  scaling may be non-perturbative process.
- Now, understanding the low momentum  $\pi^0 A_N$  ( $< 1$  GeV/c) should be a necessary tool to understand the origin of the  $\pi^0 A_N$  itself.

# What's the next step?



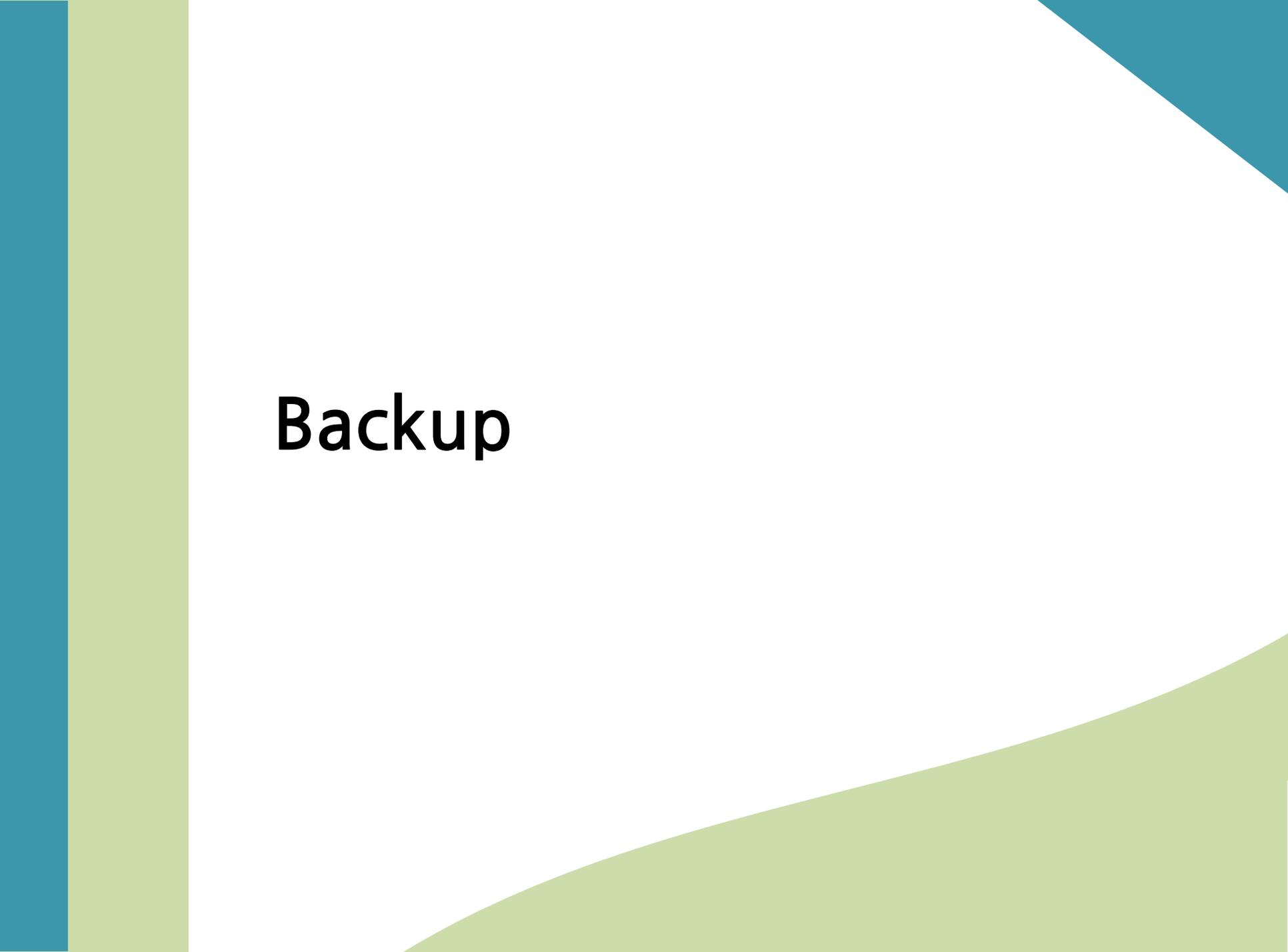
- What made the non-zero  $\pi^0 A_N$ ?
- How competitively each perturbative and non-perturbative process contribute to the  $\pi^0 A_N$  will certainly provide a powerful input for it.

# Summary

- RHICf experiment measured the  $A_N$  of very forward ( $6 < \eta$ ) neutral particles (neutron,  $\pi^0$ , single  $\gamma$ ).
- Preliminary results of RHICf  $\pi^0 A_N$  are meaning possible non-negligible contribution from the non-perturbative process.
- Further analysis with other STAR detectors and data taking will provide a powerful input to understand the origin of the  $\pi^0 A_N$ .
- **We hope many interest from many professors and students!**

RIKEN [ Itaru Nakagawa: [itaru@riken.jp](mailto:itaru@riken.jp)  
Yuji Goto: [goto@bnl.gov](mailto:goto@bnl.gov)  
Ralf Seidl: [rseidl@ribf.riken.jp](mailto:rseidl@ribf.riken.jp)

Korea Univ. [ Minho Kim: [jipangie@korea.ac.kr](mailto:jipangie@korea.ac.kr) (Speaker)  
Prof. Hong: [bhong@korea.ac.kr](mailto:bhong@korea.ac.kr)



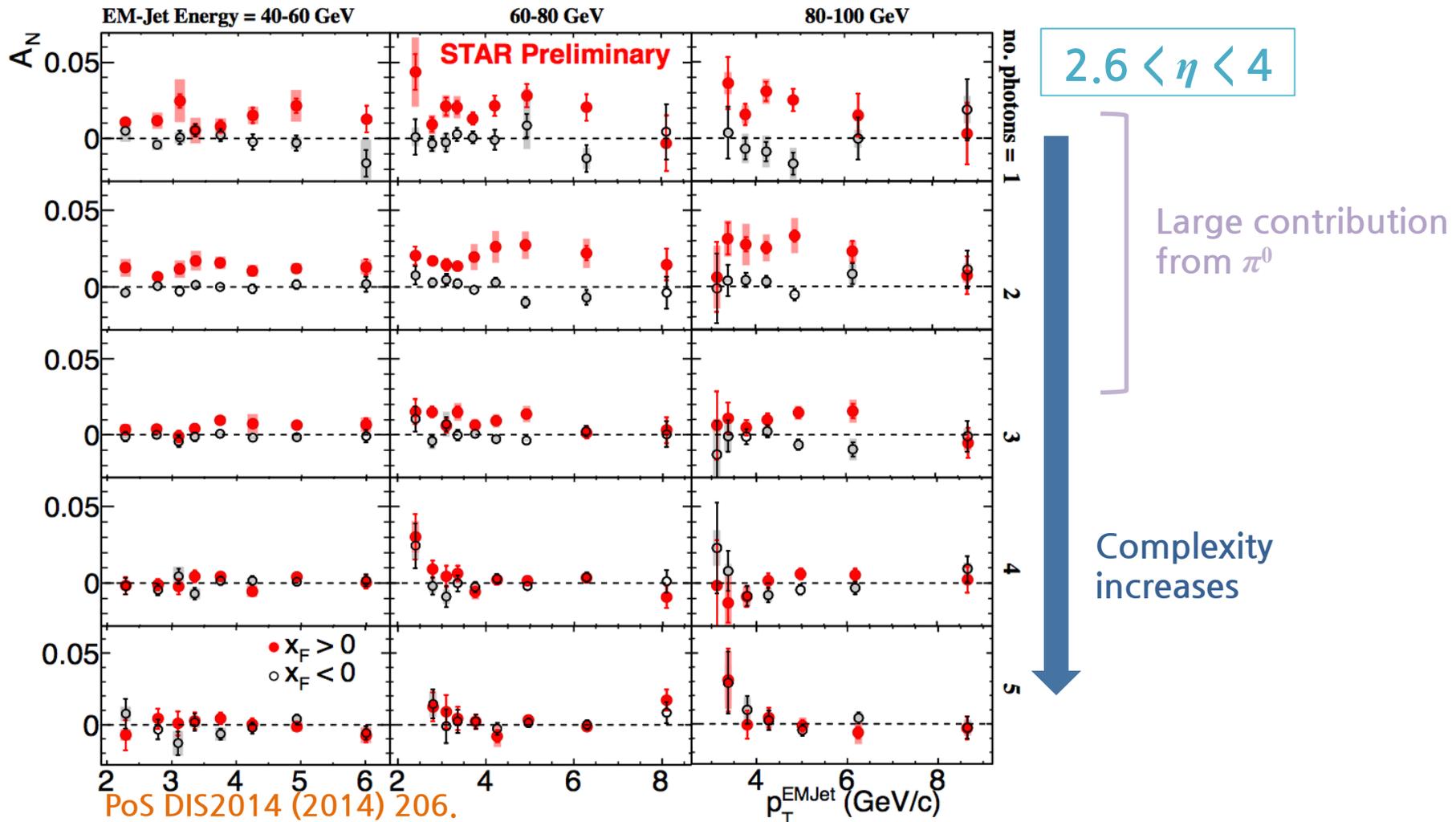
**Backup**

# Subject..

- Contents..

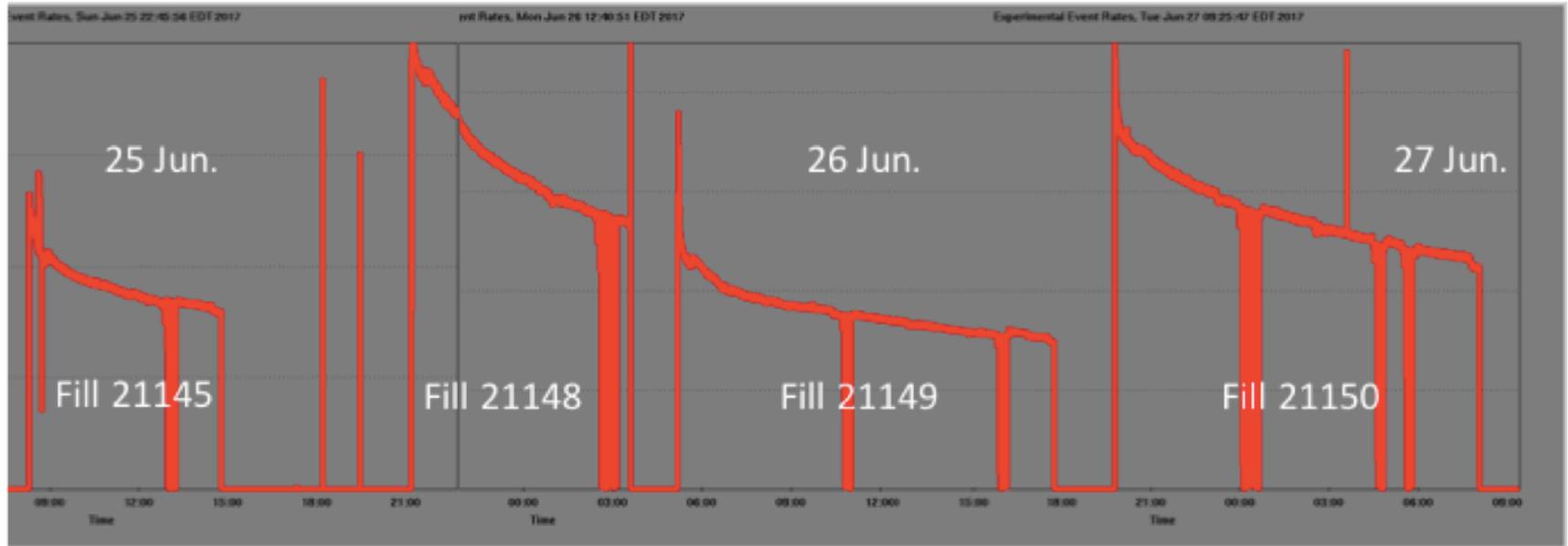
-

# New question to the $A_N$ of forward $\pi^0$



- Smaller  $A_N$  was observed with increasing multiplicity of photons (closer to hard scattering event topology).

# Operation summary



- RHICf experiment was successfully operated in June 2017.
- Total 110 M events were accumulated for neutral particles (neutron,  $\pi^0$ , and single photon) during 28 hours.
- Radial polarization.
- Higher  $\beta^*$ : 8 m and lower luminosity:  $10^{31} \text{ cm}^{-2}\text{s}^{-1}$  than usual.

# Triggers of RHICf detector



OR

**Shower trigger:** Energy deposits of three successive layers at large or small tower are larger than 45 MeV.

(for neutron and single photon)



**High EM trigger:** Energy deposit of 4th layer at large or small tower is larger than 500 MeV.

(for high energy photon and Type-II  $\pi^0$ )



OR



**Type-I  $\pi^0$  trigger:** Energy deposits of three forward (up to 7th) successive layers at large and small tower are larger than 45 MeV.

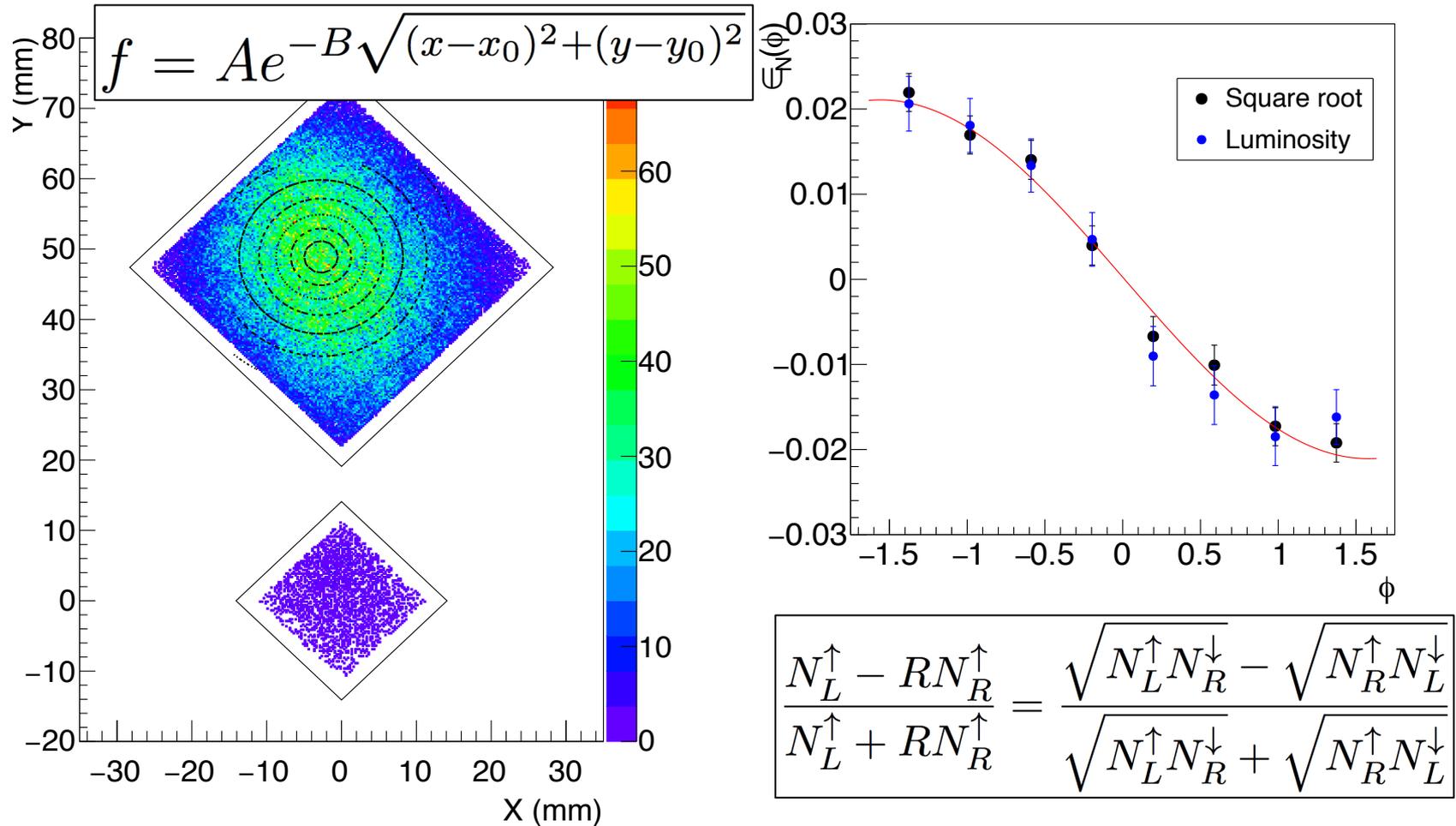
(for Type-I  $\pi^0$ )



AND

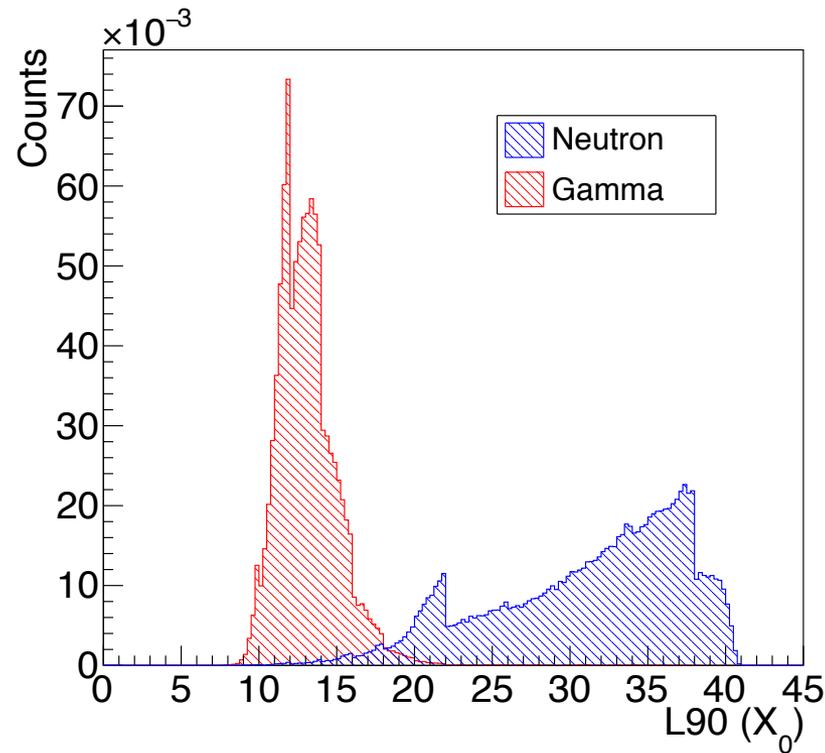


# Beam center calculation (by neutron)



- Neutrons were used for beam center calculation.
- Square root formula shows good agreement with luminosity one.

# Neutron and gamma PID



- L90 represents the longitudinal depth where the energy deposit reaches 90 % of total energy deposit.
- Gamma events can be distinguished from neutron ones using that EM shower develops more rapidly than hadronic one.

# $A_N$ calculation

Luminosity ratio between  
spin up and down

Number of  $\pi^0$  in  
specific  $x_F$  and  
 $p_T$  range

$$A_N = \frac{1}{P\epsilon} \frac{N^\uparrow - RN^\downarrow}{N^\uparrow + RN^\downarrow}$$

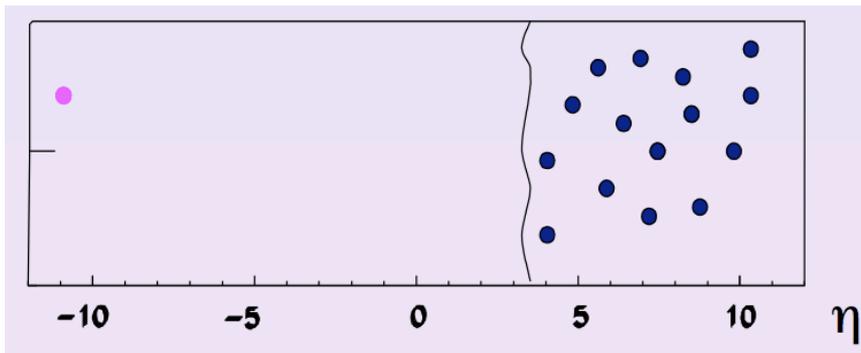
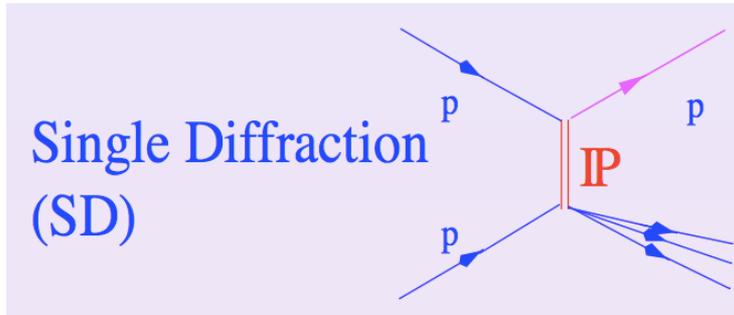
Beam polarization

Smearing by beam emittance,  
azimuthal angle distribution of  $\pi^0$ , and  
detector position resolution

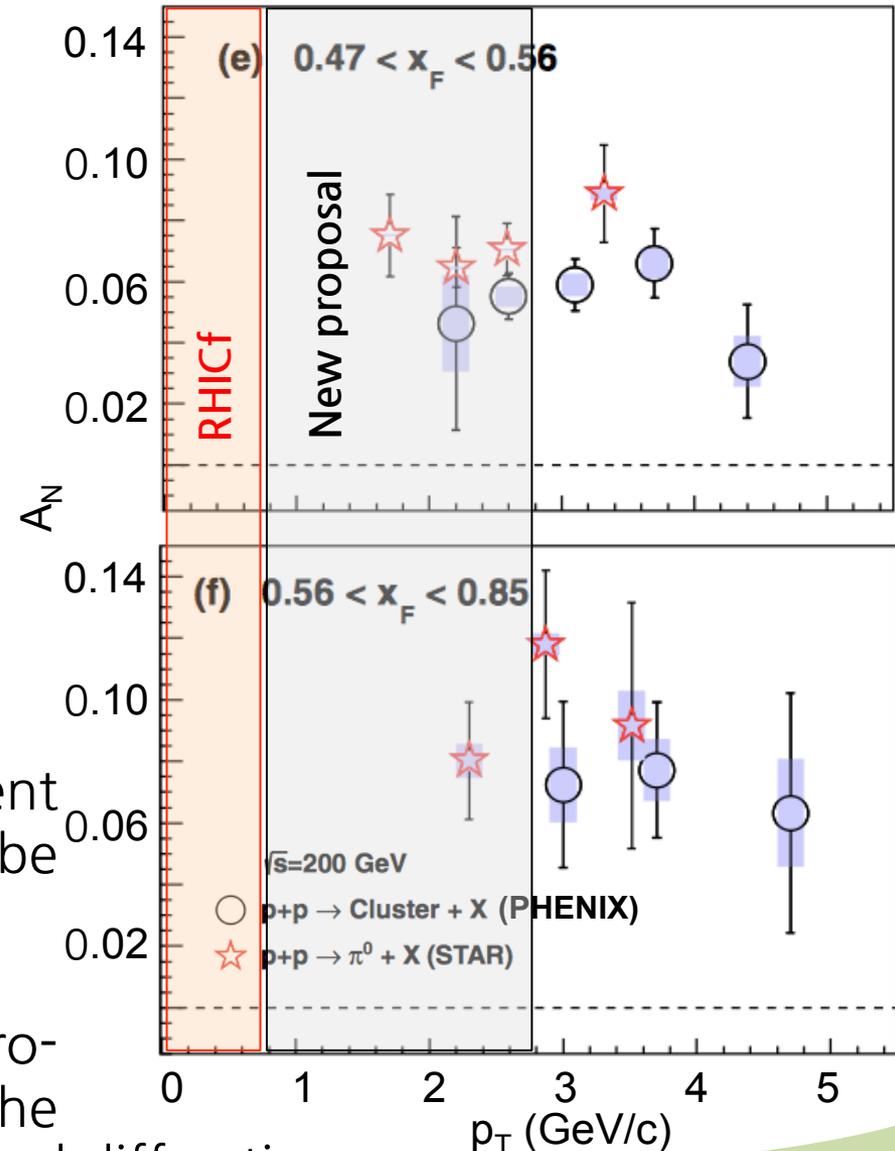
- $P$  ( $\sim 0.55 \pm 0.05$ ) can be calculated by polarization monitor.
- $R$  ( $\sim 0.970 \pm 0.02$ ) is estimated by luminosity ratio of charged particles near IP.
- $\epsilon$  ( $\sim 0.95 \pm 0.05$ ) can be studied by comparing actual and diluted  $A_N$  in simulation.

# What's the next?

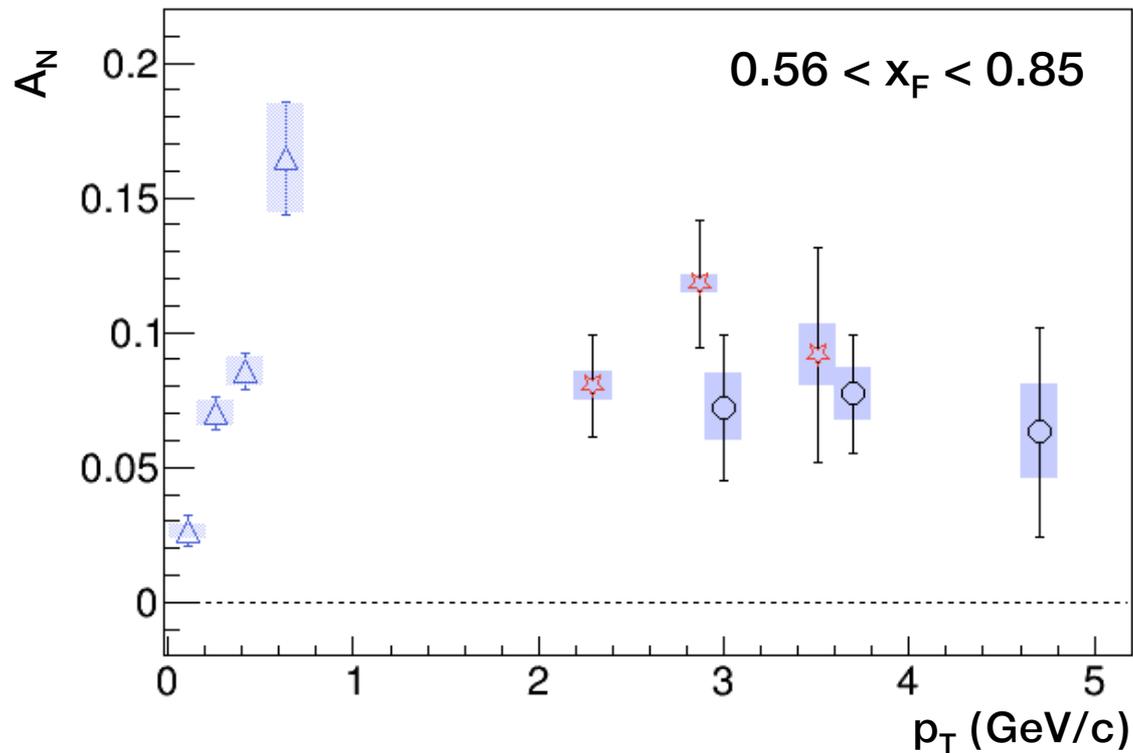
Phys. Rev. D90 (2014) 012006.



- Using other STAR detectors, event type dependence for the  $A_N$  can be studied.
- A follow-up experiment will be proposed to practically compare the each contribution from partonic and diffractive process.



# What's the next?



$\sqrt{s} = 200$  GeV

○  $p^\uparrow + p \rightarrow \text{Cluster} + X$  (PHENIX)  
☆  $p^\uparrow + p \rightarrow \pi^0 + X$  (STAR)

$\sqrt{s} = 510$  GeV

△  $p^\uparrow + p \rightarrow \pi^0 + X$  (RHICf)

- How much both perturbative and non-perturbative process contribute to the  $\pi^0 A_N$  should be exactly studied.
- Combined analysis with STAR will make the comparison of two processes possible.