# Status report on RHICf, FoCal, RANS test, etc.

RBRC exp group meeting May 17, 2022 Yuji Goto

#### RHICf results

- $\pi^0$  asymmetry
  - Phys.Rev.Lett. 124 (2020) 252501
  - Comparison with high  $p_T > 0.5$  GeV/c data of the past experiments
  - Nearly the same large asymmetry is reached at low  $p_T < 0.2 \text{ GeV}/c$
  - Contribution of other mechanisms, diffraction and resonance, may provide a hint to the mystery
- STAR comparison
  - Phys.Rev.D 103 (2021) 092009
  - √s = 200 GeV & 500 GeV
  - Forward  $\pi^{0},\,2.7<\eta<4.0$
- Significant part of the (isolated)  $\pi^0$  are from diffractive processes?



#### RHICf results

- Neutron asymmetry
  - Preliminary result
  - $A_N$  increases in magnitude with  $p_T$  up to 1 GeV/*c* at high  $x_F$
  - Clear  $x_F$  dependence at high  $p_T$
- Photon spectrum
  - arXiv:2203.1541 [hep-ex]
  - Comparison with LHCf photon results
  - First confirmation of collisionenergy scaling at zero degree photons





### RHICf-II proposal

- We proposed a second run for RHICf in 2024 (RHICf-II)
- RHICf-II Lol was discussed by the PAC in 2020.9
  - Parasitic beam-time
- We're collaborating with ALICE-FoCal group to use the FoCal-E technology
  - 8cm x 18cm detector
  - Kakenhi-Kiban-A (2021-2024) + RIKEN budget
  - The detector have enough radiation hardness to work for a small β\* and normal luminosity



## RHICf-II physics

- $K_{S}^{0}$  &  $\Lambda$  asymmetry measurement
  - Expected statistical uncertainty of asymmetry measurements for  $\pi^0$ ,  $K^0{}_S$ , and  $\Lambda$  compared to the RHICf (Run17)  $\pi^0$  assuming the similar luminosity
  - BRAHMS comparison
  - To understand the forward hadron production mechanism



Forward identified particles at BRAHMS

### RHICf-II physics

- $K_{S}^{0} \& \Lambda$  spectrum measurement
  - Impact on air-shower development
    - Muon excess issue
  - Cosmic neutrino background
    - For accelerator neutrino physics, too (FASER, SND)



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### STAR Spin/Cold-QCD PWG

- 2022.1.26
- ZDC performance issue
  - 9 o'clock blue-beam Snake failure
  - Luminosity measurement
  - Local polarimeter performance
- Peoplepower issue
  - BNL peoplepower necessary for installation and safety
- DAQ requirement
- Available space
- List of tasks
  - Simulation tasks

#### Conclusion

- 2022.4.8: Decision by the STAR management
  - They cannot accept the RHICf-II proposal.
  - The main reason is the lack of human resources.
- 2022.4.22: NPP meeting
  - Last discussion to avoid losing the RHICf-II physics for understanding the forward hadron production mechanism
    - Diffractive process
    - Air-shower development & cosmic neutrino background
  - No way to cover the lack of human resources
    - Minimum two FTE necessary, mechanical engineer and technician who understand the complex system of STAR
    - No way to obtain them from outside

#### Conclusion

- RHICf-I data analysis will continue
  - Combined analysis with STAR detectors
    - Event type categorization
    - Diffraction + resonance tagging with STAR + RHICf
  - Joint meeting to be held in the near future

#### FoCal activities

- RANS irradiation test
- Temperature dependence test @ NWU
- CERN-PS test beam
- FoCal-E pad trigger and readout scheme

## March 3 (Thu) - 4 (Fri)

- $\bullet$  RANS 7MeV proton beam up to 40  $\mu A$ 
  - Neutron irradiation from Be target







Mar. 4 5.99 hours in total



### March 3 (Thu) - 4 (Fri)

- Tested
  - FoCal-E pad
    - p-type 1x1 baby chip
    - p-type monitor PD
    - n-type monitor PD
    - P-type 2x2 baby chip
  - APD
    - by Yamazaki & Shimizu
  - sPHENIX-INTT cable
    - by Nakagawa
- Monitored by
  - Monitor PD
    - from Kyushu Univ.
  - Indium foil
  - Thermistor





#### Neutron fluence analysis

- Estimation with Indium foil
  - by Shimizu-san
  - In Run1, 1092 s, 32.85 $\mu$ A average
  - 336 keV  $\gamma$ -ray measured by Ge detector



#### Status

- Residual radiation levels are still too high to take all out of the area.
- We have disassembled the setups and identify items with high radiation levels.
  - 4/8: board #4 60 cpm  $\rightarrow$  took outside
    - Sent to NWU for the I-V measurement
  - 4/13: board #2 & #3, rod+box 80cpm  $\rightarrow$  took outside
  - 4/28: board #5 & #6 70cpm  $\rightarrow$  took outside
  - INTT micro-coax cable, APD, board #1 still in the area
    - To be checked again in this week
- Inaba-san is working on the analysis of the online measurement data and continuing I-V and C-V measurement

# **RANS** irradiation test

13/20



#### - Monitor PD, baby chips were used

• 16:42:29

\* 12:17:41

16:01:21

0

- irradiated ~1014 neutron /cm2 in two days
- Future: IV, CV test, components irradiation test



p-type



PCBs with sensors

#### (M. Inaba)

Slide by T. Chujo

#### What we have learned

- Neutron fluence was 10 times higher than expected by some miscalculations. The distance dependence from the target was smaller than expected.
  - Good to be able to confirm by simulation calculation
- It would be good to create a system using MPD as an online monitor of irradiation doses.
  - It would be good to collaborate with RANS to develop such a system, which can be offered to other users.
- Since the indium foil irradiation measurement is performed under high radiation dose, it is necessary to devise a way to mount and dismount the foil in a short period of time.
  - The amount of irradiation (time) and the number of sheets at a time should be considered, taking into account the measurement time with the Ge detector.
- Online measurements provided very useful data. We hope to succeed next time with the measurements we were not able to make this time.

#### Temp. dep. of I/V for p-type sensor



- I/V curves for 2x2 and 1x1 babies have been measured at Nara Women's Univ. before the irradiation.
- Initial measurements after irradiation was done (April, 2022)

(M. Hata, T. Hachiya, M. Shimomura)

Slide by T. Chujo

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#### FoCal PS/ SPS test beam in 2022

- · June @ PS (dedicated for PAD (HCal))
- September @ SPS (all subsystems
- November @ SPS (all subsystems)

#### FoCal-E

- ·18 single pad, and 2 pixel layers
- $\bullet$  PAD: HGCROC for PAD w/ aggregator board and O² (FLP and CRU) system
- PIXEL: under repair of damaged half-layers. Another option is old prototype (mTower with 8 ALPIDEs)

#### FoCal-H

- 9 modules, 3x3 construction underway
- •Each module: 6.5 x 6.5 x 110 cm<sup>3</sup>







Slide by T. Chujo

#### FoCal-E pad trigger and readout scheme discussion



• Started Bi-weekly meeting on trigger using FoCal-E pad data.

(S. Shimizu)

Layer ID

• Skimming simulations for trigger simulation, and trigger studies are ongoing.

Slide by T. Chujo

# Backup Slides

#### The p-type monitor PD (Tsukuba)



#### The C-V characteristics of n-type monitor PDs



#### Discussion

- 2022.4.8: Decision by the STAR management
  - They cannot accept the RHICf-II proposal.
  - The main reason is the lack or human resources.
- 2022.4.22: NPP meeting
  - Any way to change this situation?
  - To avoid losing the RHICf-II physics for understanding the forward hadron production mechanism
    - Diffractive process
    - Air-shower development & cosmic neutrino background
- RHICf-I data analysis will continue
  - Combined analysis with STAR detectors
    - Event type categorization
    - Diffraction + resonance tagging with STAR + RHICf

#### Status

- Residual radiation levels are still too high to take anything out of the area.
- We have been able to disassemble some of the setups and identify areas with high radiation levels.
  - Substrate
- Expect to be able to take out the ones that are low enough next week.
- Setups are being stored together in the area and online measurements are continuing.
- Inaba-san is working on the analysis of the online measurement data.

Collision system & Polarization	Science goals & objects	Measurement time, luminosity or number of events	Trigger rate / DAQ requirement
p+p Radial polarization	High-p <sub>T</sub> π <sup>0</sup> , K <sup>0</sup> <sub>S</sub> , Λ SSA	1 pb <sup>-1</sup> , a few hours with 200 Hz rare trigger	200 Hz rare trigger for high- p <sub>T</sub> $\pi^0$ , K <sup>0</sup> <sub>S</sub> , $\Lambda$ with no-prescale & high efficiency
p+p Vertical polarization	K <sup>0</sup> <sub>S</sub> , Λ Spectrum	10 <sup>8</sup> events, about a week with 200 Hz shower trigger (with prescale)	200 Hz shower trigger (with prescale)
p+A Radial polarization	High- $p_T \pi^0$ SSA nuclear dependence	Similar to p+p Radial polarization	200 Hz rare trigger for high- $p_{\rm T}  \pi^0$ with no-prescale
p+A Vertical polarization	Photon, π <sup>0</sup> , neutron Spectrum	< 10 <sup>8</sup> events, < 1 week with 200 Hz shower trigger (with prescale)	200 Hz shower trigger (with prescale)

#### ZDC performance issue

- Luminosity measurement
  - No effect found in 2017 Vernier scan data
    - Can we understand this?
  - Calibration by Vernier scans if necessary
- Polarization measurement
  - Especially, problematic blue-beam snake failure requires a stable measurement
  - How stable we can monitor & evaluate polarization of the blue beam?
    - With shifted threshold energy of ZDC by our detector
  - We're studying the effect of additional material in front of the ZDC, or W+ZDC by simulation.
  - We'll consider to study it with existing data in 2022.

#### Peoplepower issue

- BNL's peoplepower necessary for installation and safety
  - We will do everything we can to support this.
- Hardware design and fabrication to be done by the RHICf-II collaboration
  - Remote manipulator in front of the ZDC
  - Other materials and supplies
- Participation in the STAR shift from 2023
- New collaborators in the US
  - Stony Brook Univ, Kansas Univ
- Other new collaborators
  - Sejong Univ

#### DAQ requirement

- STAR data recording with 200 Hz RHICf trigger
  - 10% TPC data recording if possible
  - Remaining 90% without TPC but all other STAR data recording for combined analysis of RHICf + STAR
- Standalone RHICf-II DAQ with independent data stream
  - Event correspondence between STAR DAQ & RHICf-II DAQ with event number sharing
  - Established in 2017 run

#### Available space

- We installed RHICf calorimeter (LHCf Arm-1 calorimeter) in 2017.
  - W:9cm x H:62cm x D: 29cm
    - by removing the top structure as shown in the right picture





#### Available space

- There is a ZDC support frame under the detector.
- Due to this limitation, there is only about 5cm space below the beam pipe.
- The space between the beam pipes is about 12.5cm in front of the ZDC.
  - (9.5 cm at the exit of the vacuum section)







May 17, 2022

#### Configuration 1





#### Aggregator and interface boards for 2022

- For SPS test beam in 2022, PCB v2, aggregator and interface board have been produced and largely programmed
- Logic tests are ongoing
- Built-up of cosmic test bench in progress
  - Grenoble group is preparing the firmware and online monitoring software





(Olivier Bourrion, Damien Tourres, Fatah Rarbi, Rachid <u>Guernane</u> and Grenoble LPSC CAD team)





(Grenoble)

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2022 SPS test beam setup

#### List of tasks

- Simulation tasks
  - ZDC + W simulation for luminosity measurement and polarimetry performance with shifted threshold energy of ZDC
  - $\Lambda \rightarrow$  n +  $2\gamma$  background simulation for reconstruction and resolution
  - Detector configuration and trigger scheme
  - Minho Kim is working on the simulation studies
- Blue beam snake failure
  - 2022 data analysis
  - Hope someone can participate in 2022 data analysis
- Timeline for the RHICf-II calorimeter construction
  - ALICE-FoCal-E prototype beam test at CERN-SPS in 2022 (September & November)
  - ALICE-FoCal-E prototype will be used as the first module of the RHICf-II calorimeter and commissioned at RHIC in 2023
  - The second module will be constructed in 2022-2023

### Physics at RHICf & RHICf-II

- Measurements of neutral particle production at zero degree with RHIC polarized proton collisions
- Cosmic-ray study
  - Cross section measurement to understand ultra-high energy cosmic rays
- Asymmetry measurement
  - To understand the hadronic collision mechanism based on QCD





#### RHICf at STAR in 2017

• EM calorimeter (RHICf detector) installed in front of the Zero-Degree Calorimeter (ZDC) of the STAR experiment



- Two position-sensitive sampling calorimeters
  - TS (small tower): 20mm x 20mm
  - TL (large tower): 40mm x 40mm
  - Tungsten absorber (44  $X_0$ , 1.6  $\lambda_{int}$ )
  - 16 GSO sampling layers
  - 4 XY pairs of GSO-bar position layers



May 17, 2022

#### RHICf at STAR in 2017

- $\pi^0$  asymmetry
  - Phys. Rev. Lett. 124, 252501 (2020)
  - Comparison with high  $p_T > 0.5 \text{ GeV}/c$  data of the past experiments
  - Nearly the same large asymmetry is reached at low  $p_T < 0.2~{\rm GeV}/c$
  - Contribution of other mechanisms, diffraction and resonance, may provide a hint to the mystery



#### RHICf at STAR in 2017

- Other analyses ongoing
  - $\pi^0$  & neutron cross section analysis
  - Neutron asymmetry (RHICf + ZDC)
  - Combined analysis with STAR detectors
    - Event type categorization
    - Diffraction + resonance tagging with STAR + RHICf combined data analysis
    - Event type, multiplicity (FMS) dependence of cross section & asymmetry to be obtained



#### New STAR results

- Phys.Rev.D 103 (2021) 092009
  - √s = 200 GeV & 500 GeV
  - Forward  $\pi^0$ , 2.7 <  $\eta$  < 4.0
  - Asymmetries for the isolated  $\pi^0$  are larger than these for the non-isolated  $\pi^0$
  - Possible explanation is that a significant part of the isolated  $\pi^0$  are from diffractive processes



#### New STAR results

- Phys.Rev.D 103 (2021) 092009
  - Small EM-jet asymmetry, consistent with AnDY result
  - $z_{em} = E_{\pi 0} / E_{jet}$
  - Hadron in jet Collins asymmetries small
    - Cancellation of the Collins effect of the u/d quark?



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  - Parasitic beam-time
- We're collaborating with ALICE-FoCal group to use the FoCal-E technology
  - 8cm x 18cm detector
  - Kakenhi-Kiban-A (2021-2024) + RIKEN budget
  - The detector have enough radiation hardness to work for a small β\* and normal luminosity



#### RHICf-II Collaboration

- Y. Goto, I. Nakagawa, R. Seidl (RIKEN)
- B. Hong, M.H. Kim (Korea Univ.)
- K. Tanida (JAEA)
- T. Chujo (Tsukuba Univ.) ← New
- Y. Itow, H. Menjo (Nagoya Univ.)
- T. Sako (ICRR, Univ. of Tokyo)
- K. Kasahara (Shibaura Tech.)
- O. Adriani, L. Bonechi, R. D'Alessandro (INFN Firenze)
- A. Tricomi (INFN Catania)
- New collaborators expected or under discussion from:
  - Sejong Univ.
  - Univ. of Kansas
  - Nara Women's Univ. and EIC Japan group
- Cooperation from FoCal collaboration
  - ORNL

- ${\rm K0}_{\rm S}$  and  $\Lambda$  measurement
  - Spectrum and cross section
  - Asymmetry
- A-dependence of the  $\pi^0$  asymmetry
  - Correlation between asymmetries of forward neutron and  $\pi^0$
  - Strong A-dependence of the neutron asymmetry measured at PHENIX in Run 15
    - Phys. Rev. Lett. 120, 022001 (2018) <sup>0.4</sup>
    - UPC vs hadronic component



### ALICE FoCal-E

- Led by Tsukuba Univ. group
- Tungsten absorber
- Low granularity (LG) silicon pad for energy measurement
  - $\sigma_{\rm E}$  / E = 25% /  $\sqrt{\rm E}$  (GeV)  $\oplus$  2% for photon energy resolution (simulation)
- High granularity (HG) silicon pixel (CMOS-MAPS) for accurate position measurement



#### ALICE FoCal-E for RHICf-II

- Space restriction at RHICf
- Pad layer
  - Lead by Tsukuba Univ. group
  - Readout electronics based on HGCROC ASIC (CMS) working with Grenoble group leading the development
- Pixel layer
  - Lead by European group
- Trigger
  - Rare trigger for asymmetry measurement
  - Shower trigger for cross section measurement
- DAQ
  - Standalone ALICE DAQ
  - Event correspondence with STAR DAQ



- p + A collisions
  - Measurement of nuclear effect (p+A / p+p)
- Strong A-dependence of the neutron asymmetry
  - Measured at PHENIX in Run 15
  - Phys. Rev. Lett. 120, 022001 (2018)
  - UPC vs hadronic component
- A-dependence of the  $\pi^0$  asymmetry
  - Correlation between asymmetries of forward neutron and  $\pi^0$
- p + Oxygen collision
  - Ideal condition for cosmic-ray interaction studies measuring  $\pi^0,$  neutron, photon,  ${\rm K^0}_{\rm S}$



RHIC

100

-100

18.0m from IP

Crossing angle (half): 0.0 urad Detector posit on: 24.0 mm

100

- Large acceptance detector
  - 8cm x 18cm
  - For more particles:  ${\rm K0}_{\rm S}$  and  $\Lambda$



- 0.2  $\text{K}^{0}_{\text{S}}$  /sec = 10<sup>4</sup>  $\text{K}^{0}_{\text{S}}$ s in 14 hours operation
- $\Lambda \rightarrow n + \pi^0 \rightarrow n + 2\gamma$  (B.R. 35.9%)
  - 12  $\Lambda$  /sec = 10<sup>5</sup>  $\Lambda s$  in 2.5 hours operation
- Geometric acceptance of  $\pi^0,\,\mathsf{K}^0{}_{\mathsf{S}}$  and  $\Lambda$



- K<sup>0</sup><sub>S</sub> for studying impact on the high-energy atmospheric neutrino flux
  - Differences in p+p collisions at 200 GeV between models: EPOSLHC (magenta), QGSJET II-4 (blue), SIBYLL 2.3 (green)



# Kaons in atm. v productions

6 years (ICRC 2017)

La Contraction of the second s

10<sup>2</sup>

 $10^{1}$ 

 $10^{0}$ 

10<sup>-1</sup>

Events per 2078 Days

IceCube detected astronomical neutrinos. Better understanding of background (Atmospheric neutrinos) is required.

Slide by H. Menjo



- Asymmetry measurement of  ${\rm K^0}_{\rm S}$  and  $\Lambda$ 
  - Expected statistical uncertainty of asymmetry measurements for  $\pi^0,~{\rm K^0}_{\rm S},$  and  $\Lambda$  compared to the RHICf (Run 17)  $\pi^0$
  - Assuming the similar luminosity

