

Pion production in heavy-ion collisions near threshold energies

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OUTLINE

•Neutron star matter and pion probes in heavy-ion collisions

• Transport approach for pion production (LQMD)

• π^{-}/π^{+} ratio for constraining the high-density symmetry energy

Summary and perspective

I. Neutron star matter and pion probes in heavy-ion collisions



EOS-Equation of State

Van der Waals equation [p+a(

$$[p+a(\frac{n}{v})^2](v-nb) = nRT$$



Johannes Diderik van der Waals 1837-1923 Nobel Laureate in physics 1910 for study of gas and liquid equation of state



QCD phase structure, WA Zajc / Nuclear Physics A 967 (2017) 265-272



May 23-26, 2023

Nuclear symmetry energy

$$E_{\text{sym}}(\rho) = E_{\text{sym}}(\rho_0) + \frac{L}{3} \left(\frac{\rho - \rho_0}{\rho_0} \right) + \frac{K_{\text{sym}}}{18} \left(\frac{\rho - \rho_0}{\rho_0} \right)^2,$$

(\rho \sigma \rho_0)
Liquid drop model: $E_{\text{sym}}(\rho_0) \approx 32 \text{ MeV}$

Slope parameter and curvature parameter

$$L \equiv 3\rho_0 \frac{\partial E_{\rm sym}(\rho)}{\partial \rho} \bigg|_{\rho = \rho_0} \qquad \qquad K_{\rm sym} \equiv 9\rho_0^2 \frac{\partial^2 E_{\rm sym}(\rho)}{\partial \rho^2} \bigg|_{\rho = \rho_0}$$

Hajime Sotani, Nobuya Nishimura, and Tomoya Naito, Prog. Theor. Exp. Phys. 041D01 (2022)





Neutron star matter and heavy-ion collisions

Article

Nature 606 (2022) 276

Chiral effective field theory

Prior

a

Mass M (Mo)

C

Mass M (M_o)

3.0

2.5

2.0

1.0

3.0

2.5

2.0

1.5

1.0

10

HIC experiments

10

HIC

Prior

12

Radius R (km)

12

Radius R (km)

14

Constraining neutron-star matter with microscopic and macroscopic collisions



Hyperons in neutron stars

S. Weissenborn, D. Chatterjee, J. Schaffner-Bielich, Nucl. Phys. A 881, 62 (2012) W. Z. Jiang, R. Y. Yang, and D. R. Zhang, Phys. Rev. C 87, 064314 (2013) D. Logoteta, I. Vidaña, I. Bombaci, Eur. Phys. J. A 55, 207 (2019)



Transport models for heavy-ion collisions

Progress in Particle and Nuclear Physics 125 (2022) 103962



Con	tents lists availa	able at Scien	iceDirect	
Progress in	n Particle	and Nu	clear Phy	vsics

journal homepage: www.elsevier.com/locate/ppnp

Review

Transport model comparison studies of intermediate-energy heavy-ion collisions

Hermann Wolter ^{1,*}, Maria Colonna ², Dan Cozma ³, Pawel Danielewicz ^{4,5}, Che Ming Ko ⁶, Rohit Kumar ⁴, Akira Ono ⁷, ManYee Betty Tsang ^{4,5}, Jun Xu ^{8,9}, Ying-Xun Zhang ^{10,11}, Elena Bratkovskaya ^{12,13}, Zhao-Qing Feng ¹⁴, Theodoros Gaitanos ¹⁵, Arnaud Le Fèvre ¹², Natsumi Ikeno ¹⁶, Youngman Kim ¹⁷, Swagata Mallik ¹⁸, Paolo Napolitani ¹⁹, Dmytro Oliinychenko ²⁰, Tatsuhiko Ogawa ²¹, Massimo Papa ², Jun Su ²², Rui Wang ^{9,23}, Yong-Jia Wang ²⁴, Janus Weil ²⁵, Feng-Shou Zhang ^{26,27}, Guo-Qiang Zhang ⁹, Zhen Zhang ²², Joerg Aichelin ²⁸, Wolfgang Cassing ²⁵, Lie-Wen Chen ²⁹, Hui-Gan Cheng ¹⁴, Hannah Elfner ^{12,13,20}, K. Gallmeister ²⁵, Christoph Hartnack ²⁸, Shintaro Hashimoto ²¹, Sangyong Jeon ³⁰, Kyungil Kim ¹⁷, Myungkuk Kim ³¹, Bao-An Li ³², Chang-Hwan Lee ³³, Qing-Feng Li ^{24,34}, Zhu-Xia Li ¹⁰, Ulrich Mosel ²⁵, Yasushi Nara ³⁵, Koji Niita ³⁶, Akira Ohnishi ³⁷, Tatsuhiko Sato ²¹, Taesoo Song ¹², Agnieszka Sorensen ^{38,39}, Ning Wang ^{11,40}, Wen-Jie Xie ⁴¹, (TMEP collaboration)

Table 1

List of transport models that participated in the TMEP code comparisons discussed in this paper. The columns give the information on the name of the code, the main correspondents of the code, the energy range intended for the code, the treatment of effects of relativity (see Section 2.1), and the comparisons in which the code participated. The different comparisons are listed in the last column in the table by a numbers n, which refer to the subsections 3.n, where they are described in detail: n = 1 for Au+Au collisions around 1 AGeV, n = 2 for Au+Au collision at 100 and 400 AMeV, n = 3 for box-Vlasov, n = 4 for box-cascade with only nucleons, n = 5 for box-cascade with pion and Δ resonance production, and n = 6 for the prediction of pion ratios for Sn+Sn collisions.

	BUU Type	Code Correspondents	Energy Range [A GeV]	Relativity	Comparisons
	BLOB	P Napolitani M Colonna	0.01-0.5	non-rel	2
	BUU-VM	S. Mallik	0.02-1	rel	3.4.5
10	DIBUU	Y. Kim, S. Jeon, M. Kim, CH. Lee, K. Kim	0.05-2	COV	3
1	GiBUU	J. Weil, T. Gaitanos, K. Gallmeister, U. Mosel	0.05-40	rel/cov	1,2,3,4
E.	IBL	W.J. Xie, F.S. Zhang	0.05-2	rel	2
	IBUU	J. Xu, L.W. Chen, B.A. Li	0.05-2	rel	2,3,4,5
2	LBUU(LHV)	R. Wang, Z. Zhang, L-W. Chen	0.01-1.5	rel	3
F	- pBUU	P. Danielewicz	0.01-12	rel	1,2,3,4,5,6
Ē	PHSD	E. Bratkovskaya, W. Cassing	0.1-200	rel/cov	1,6
	KBUU	1. Galtanos 7. Zhang, C.M. Ko, T. Song	0.05-2	COV	1,2
	SMASH	D. Oliinychenko, H. Elfner, A. Sorensen	0.05-2	COV	2,456
	SME	M Colonna P Nanolitani	0.01-0.5	non-rel	234
	x BUU	Z. Zhang, C.M. Ko	0.01-0.5	non-rel	6
	X	2. 2			
	QMD Type	Code Corespondents	Energy Range [AGeV]	[Relativity	Comparisons
	AMD	A. Ono	0.01-0.3	non-rel	2
	AMD+JAM	N. Ikeno, A. Ono	0.01-0.3	non-rel+re	el 6
	BQMD/IQMD	A. Le Fèvre, J. Aichelin, C. Hartnack, R. Kumar	0.05-2	rel	1,2,6
	CoMD	M. Papa	0.01-0.3	non-rel	2,4
	ImOMD	Y.X. Zhang, N. Wang, Z.X. Li	0.02-0.4	rel	2.3.4
	IOMD-BNU	J. Su, F.S. Zhang	0.05-2	rel	2,3,4,5,6
	IOMD-SINAP	G.O. Zhang	0.05-2	rel	2
	IAM	A. Ono, N. Ikeno, Y. Nara, A. Ohnishi	1-158	rel	4.5
	IOMD 2.0	T. Ogawa, K. Niita, S. Hashimoto, T. Sato	0.01-3	rel	4.5
	LOMD(IOMD-IMP)	Z.O. Feng, H.G. Cheng	0.01-10	rel	2,3,4,5
	TuOMD/dcOMD	D. Cozma	0.1-2	rel	1.2.3.4.5.6
	UrQMD	Y. J. Wang, Q. F. Li, Y. X. Zhang	0.05-200	rel	1,2,3,4,6

Chack for updates

Pion production in heavy-ion collisions near threshold energies

PHYSICAL REVIEW LETTERS 126, 162701 (2021)

Systematics of pion emission in heavy ion collisions in the 1 A GeV regime

ScienceDirect

Nuclear Physics A 781 (2007) 459-508

FOPI Collaboration

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FOPI Collaboration / Nuclear Physics A 781 (2007) 459-508

Probing the Symmetry Energy with the Spectral Pion Ratio

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CSR能区的高重子密度核物质研究进展

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Progress of study on the properties of nuclear matter with high barvon density at CSR energy region

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This review is focused on the low energy part of the OCD phase diagram, the properties of the nuclea baryon density which could be experimentally studied at the large scale facility for nuclear physics in China, such as the split of nucleon effect mass, symmetric energy of high baryon density nuclear matter, the properties of hadrons in the nuclear medium. The properties of nuclear matter with high baryon density and the resonance states of hadron have been studied theoretically and some physical observables related to the properties of nuclear matter with high baryon density are pointed out under the heavy ion collisions with the energy at the region of 0.5-1.2 GeV/u. The existing experimental setup at HIRFL-CSR has been improved and a test experiment has been performed. The results show the ability to explore the properties of the nuclear matter with high baryon density. Our studies are helpful to design and built a professional spectrometer for the low energy part of the QCD phase diagram in the near future.

QCD phase diagram, resonant state of nucleon, nuclear matter with high baryon densit

PACS: 25.70.Lm, 25.75.Dw, 21.65.+f

doi: 10.1360/SSPMA-2019-0046

25

-10

30

2.5 F

2.0

-0.5

-1.0 4.1.

(c)



II. Transport approach for pion production (LQMD)



Heavy-ion collisions (5 MeV – 5 GeV/nucleon) and hadron induced reaction (p, \bar{p} , π , K, e, etc)

- **Lanzhou quantum molecular dynamics** (Skyrme interaction, Walecka model with σ , ω, ρ, δ)
- **Isospin physics at intermediate energies** (constraining nuclear symmetry energy at sub- and suprasaturation densities in HICs and probing isospin splitting of nucleon effective mass from HICs)
- In-medium properties of hadrons in dense nuclear matter from heavy-ion
 collisions (extracting optical potentials, i.e., Δ(1232), N*(1440), N*(1535)), hyperons (Λ,Σ,Ξ) and mesons (π,K,η,ρ,ω,φ...), hypernucleus dynamics)
- **Nuclear cluster and hypernuclear cluster production** (production cross section, transverse momentum spectra, rapidity distribution, collective flows, e.g., $\Lambda(\Sigma)X$, $\Lambda\Lambda X$, ΞX , $\overline{\Lambda}X(S=1)$)

1. Skyrme energy-density functional (LQMD-Skyrme)

PHYSICAL REVIEW C 84, 024610 (2011)

 $H_B = \sum_i \sqrt{\mathbf{p}_i^2 + \mathbf{m}_i^2} + U_{\text{int}} + U_{\text{mom}}$

Momentum dependence of the symmetry potential and its influence on nuclear reactions

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$$U_{loc} = \int V_{loc}(\rho(\mathbf{r})) \, d\mathbf{r} \quad V_{loc}(\rho) = \frac{\alpha}{2} \frac{\rho^2}{\rho_0} + \frac{\beta}{1+\gamma} \frac{\rho^{1+\gamma}}{\rho_0^{\gamma}} + E_{sym}^{loc}(\rho)\rho\delta^2 + \frac{g_{sur}}{2\rho_0}(\nabla\rho)^2 + \frac{g_{sur}^{lso}}{2\rho_0} [\nabla(\rho_n - \rho_p)]^2,$$

$$U_{mom} = \frac{1}{2\rho_0} \sum_{i,j,j\neq i} \sum_{\tau,\tau'} C_{\tau,\tau'} \delta_{\tau,\tau_i} \delta_{\tau',\tau_j} \int \int \int d\mathbf{p} \, d\mathbf{p}' \, d\mathbf{r} \, f_i(\mathbf{r},\mathbf{p},t) \\ \times \left[\ln(\epsilon(\mathbf{p}-\mathbf{p}')^2+1) \right]^2 f_j(\mathbf{r},\mathbf{p}',t).$$

$$E_{sym}(\rho) = \frac{1}{3} \frac{\hbar^2}{2m} \left(\frac{3}{2} \pi^2 \rho\right)^{2/3} + E_{sym}^{loc}(\rho) + E_{sym}^{mom}(\rho).$$

 $E_{sym}^{loc}(\rho) = \frac{1}{2} C_{sym}(\rho/\rho_0)^{\gamma_s} \qquad E_{sym}^{loc}(\rho) = a_{sym}(\rho/\rho_0) + b_{sym}(\rho/\rho_0)^2.$ Equation of State of Dense Nuclear Matter at RIBF and FRIB, May 23-26, 2023

Table 1: The parameters and properties of isospin symmetric EoS used in the LQMD model at the density of 0.16 fm^{-3} .

Parameters	$\alpha \ ({\rm MeV})$	β (MeV)	γ	C_{mom} (MeV)	$\epsilon \; (c^2/MeV^2)$	m_{∞}^{*}/m	K_{∞} (MeV)
PAR1	-215.7	142.4	1.322	1.76	5×10^{-4}	0.75	230
PAR2	-226.5	173.7	1.309	0.	0.	1.	230



2. Covariant energy-density functional (LQMD-RMF)

$$\begin{split} L &= \bar{\psi} [i\gamma_{\mu}\partial^{\mu} - (M_N - g_{\sigma}\varphi - g_{\delta}\vec{\tau}\cdot\vec{\delta}) - g_{\omega}\gamma_{\mu}\omega^{\mu} - g_{\rho}\gamma_{\mu}\vec{\tau}\cdot\vec{b}^{\mu}]\psi \\ &+ \frac{1}{2}(\partial_{\mu}\varphi\partial^{\mu}\varphi - m_{\sigma}^2\varphi^2) - U(\varphi) + \frac{1}{2}(\partial_{\mu}\vec{\delta}\partial^{\mu}\vec{\delta} - m_{\sigma}^2\vec{\delta}^2) \\ &+ \frac{1}{2}m_{\omega}^2\omega_{\mu}\omega^{\mu} - \frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{1}{2}m_{\rho}^2\vec{b}_{\mu}\vec{b}^{\mu} - \frac{1}{4}\vec{G}_{\mu\nu}\vec{G}^{\mu\nu} \end{split}$$

$$F_{\mu\nu} = \partial_{\mu}\omega_{\nu} - \partial_{\nu}\omega_{\mu},$$

$$G_{\mu\nu} = \partial_{\mu}\vec{b}_{\nu} - \partial_{\nu}\vec{b}_{\mu},$$

$$U(\varphi) = \frac{g_2}{3}\varphi^3 + \frac{g_3}{4}\varphi^4$$

Energy density functional

$$\varepsilon = \sum_{i=n,p} 2 \int \frac{d^3k}{(2\pi)^3} \sqrt{k^2 + M_i^{*2}} + \frac{1}{2}m_\sigma^2 \varphi^2 + U(\varphi) + \frac{1}{2}m_\omega^2 \omega_0^2 + \frac{1}{2}m_\rho^2 b_0^2 + \frac{1}{2}m_\delta^2 \delta_3^2$$

Temporal evolution in phase space

$$\begin{split} \dot{\mathbf{x}} &= \frac{\mathbf{P}_{\mathbf{i}}^{*}}{p_{\mathbf{0}}^{*}} + \sum_{i \neq j}^{N} \{ \frac{g_{v}^{2}}{2m_{v}^{2}} z_{j}^{*\mu} u_{i,\mu} B_{i} B_{j} \frac{\partial \rho_{ij}}{\partial \mathbf{p}_{\mathbf{i}}} + \frac{g_{v}^{2}}{2m_{v}^{2}} z_{i}^{*\mu} u_{j,\mu} B_{i} B_{j} \frac{\partial \rho_{ji}}{\partial \mathbf{p}_{\mathbf{i}}} + \frac{g_{v}^{2}}{2m_{v}^{2}} z_{j}^{*\mu} \rho_{ji} B_{i} B_{j} \frac{\partial u_{i,\mu}}{\partial \mathbf{p}_{\mathbf{i}}} \\ &+ z_{j}^{*\mu} \frac{B_{i} B_{j} \bar{g}_{v}^{2}}{2m_{v}^{2}} [\frac{\rho_{ij}}{1 - p_{T,ij}^{2}/\Lambda_{v}^{2}} \frac{\partial u_{i,\mu}}{\partial \mathbf{p}_{\mathbf{i}}} + \frac{u_{i,\mu}}{1 - p_{T,ij}^{2}/\Lambda_{v}^{2}} \frac{\partial \rho_{ij}}{\partial \mathbf{p}_{\mathbf{i}}} + u_{i,\mu} \rho_{ij} \frac{\partial [1/(1 - p_{T,ij}^{2}/\Lambda_{v}^{2})]}{\partial \mathbf{p}_{\mathbf{i}}}] \\ &+ z_{i}^{*\mu} \frac{B_{i} B_{j} \bar{g}_{v}^{2}}{2m_{v}^{2}} [\frac{u_{j,\mu}}{1 - p_{T,ji}^{2}/\Lambda_{v}^{2}} \frac{\partial \rho_{ji}}{\partial \mathbf{p}_{\mathbf{i}}} + u_{j,\mu} \rho_{ji} \frac{\partial [1/(1 - p_{T,ji}^{2}/\Lambda_{v}^{2})]}{\partial \mathbf{p}_{\mathbf{i}}}] \\ &+ z_{i}^{*\mu} \frac{B_{i} B_{j} \bar{g}_{v}^{2}}{2m_{v}^{2}} [\frac{u_{j,\mu}}{1 - p_{T,ji}^{2}/\Lambda_{v}^{2}} \frac{\partial \rho_{ji}}{\partial \mathbf{p}_{\mathbf{i}}} + u_{j,\mu} \rho_{ji} \frac{\partial [1/(1 - p_{T,ji}^{2}/\Lambda_{v}^{2})]}{\partial \mathbf{p}_{\mathbf{i}}}] \\ &+ z_{i}^{*\mu} \frac{B_{i} B_{j} \bar{g}_{v}^{2}}{2m_{v}^{2}} [\frac{u_{j,\mu}}{1 - p_{T,ji}^{2}/\Lambda_{v}^{2}} \frac{\partial \rho_{ji}}{\partial \mathbf{p}_{\mathbf{i}}} + u_{j,\mu} \rho_{ji} \frac{\partial [1/(1 - p_{T,ji}^{2}/\Lambda_{v}^{2})]}{\partial \mathbf{p}_{\mathbf{i}}}] \\ &- \frac{m_{j}^{*}}{p_{j}^{*0}} \frac{\partial S_{j}}{\partial \mathbf{p}_{\mathbf{i}}} - \frac{m_{i}^{*}}{p_{i}^{*0}} \frac{\partial S_{j}}{\partial \mathbf{p}_{\mathbf{i}}} \}, \quad \text{Equation of State of Dense Nuclear Matter at RIBF} \text{RIBF}} \text{Auter at RIBF}}$$

Si-Na Wei, Zhao-Qing Feng, arXiv:2302.09984

TABLE I: Parameter sets for RMF. The saturation density ρ_0 is set to be 0.16 fm^{-3} . The binding energy of saturation density is $E/A - M_N = -16$ MeV. The isoscalar-vector ω and isovector-vector ρ masses are fixed to their physical values, $m_{\omega} = 783$ MeV and $m_{\rho} = 763$ MeV. The remaining meson mass m_{σ} is set to be 550 MeV.

model	g_{σ}	g_{ω}	$g_2 (fm^{-1})$	g_3	$g_{ ho}$	g_{δ}	K (MeV)) $E_{sym}(\rho_0) (MeV)$	$L \ (\rho_0) (MeV)$
set1	8.145	7.570	31.820	28.100	4.049	-	230	31.6	85.3
set2	8.145	7.570	31.820	28.100	8.673	5.347	230	31.6	109.3
set3	8.145	7.570	31.820	28.100	11.768	7.752	230	31.6	145.0

Symmetry energy in LQMD-RMF

$$E_{sym} = \frac{1}{6} \frac{k_F^2}{E_F^*} + \frac{1}{2} \left[f_\rho - f_\delta \left(\frac{M^*}{E_F^*} \right) \right] \rho$$

$$f_{
ho,\delta} = g_{
ho,\delta}/m_{
ho,\delta}$$



Equation of State of Dense Nuclear Matter at RIBF and FRIB, May 23-26, 2023

Collective flows of protons and flow difference with LQMD-RMF



3. Pion-nucleon potential

$$\omega_{\tau_z}(\rho, \vec{p}) = \omega_{isoscalar}(\rho, \vec{p}) + C^{\pi}_{iso}\tau_z \delta\left(\frac{\rho}{\rho_0}\right)^{\gamma_{\pi}}$$

 $C_{\pi} = \rho_0 \hbar^3 / (4 f_{\pi}^2) = 36$ MeV,

 $τ_z$ =1, 0, -1 for $π^-$, $π^0$ and $π^+$ **Evaluation of ω_{isoscalar} 1) phenomenological ansatz from pionic atom** C. Gale and J. Kapusta, Phys. Rev. C 35, 2107 (1987); C. Fuchs et al., Phys. Rev. C 55, 411 (1997); Z. Q. Feng and G. M. Jin, Phys. Rev. C 82, 044615 (2010))

2) Δ -hole model

L. Xiong, C. M. Ko, V. Koch, Phys. Rev. C 47, 788 (1993); C. Fuchs et al., Phys. Rev. C 55, 411 (1997)

 $\omega_{\text{isoscalar}}(\mathbf{p}_i, \rho_i) = S_{\pi}(\mathbf{p}_i, \rho_i) \omega_{\pi\text{-like}}(\mathbf{p}_i, \rho_i)$

+ $S_{\Delta}(\mathbf{p}_i,\rho_i)\omega_{\Delta-\text{like}}(\mathbf{p}_i,\rho_i)$

Equation of State of Dense Nuclear Matter at RIBF and FRIB, May 23-26, 2023

Z. Q. Feng et al., Phys. Rev. C 92 (2015) 044604





4. Energy conservation for pion production Shedding light on the pion production in heavy-ion collisions for constraining the high-density symmetry energy

May 23-26, 2023

$$\begin{split} &\sqrt{m_{\Delta}^{2} + \mathbf{p}_{\Delta}^{2}} + U_{\Delta}(\rho_{\Delta}, \mathbf{p}_{\Delta}) \\ &= \sqrt{m_{N}^{2} + (\mathbf{p}_{\Delta} - \mathbf{p}_{\pi})^{2}} \\ &+ U_{N}(\rho_{N}, \mathbf{p}_{N}) + \omega_{\pi}(\rho_{\pi}, \mathbf{p}_{\pi}) + V_{\pi N}^{\text{Coul}}, \\ &U_{\Delta^{-}} = U_{n}, U_{\Delta^{++}} = U_{p}, \\ &U_{\Delta^{+}} = \frac{1}{3}U_{n} + \frac{2}{3}U_{p}, \\ &U_{\Delta^{0}} = \frac{1}{3}U_{p} + \frac{2}{3}U_{n}. \end{split}$$

The Coulomb potential only exist for the charged pairs of

 $\triangle^0 \leftrightarrow \pi^- + p \text{ and } \triangle^{++} \leftrightarrow \pi^+ + p.$

Heng-Jin Liu, Hui-Gan Cheng, and Zhao-Qing Feng* School of Physics and Optoelectronics, South China University of Technology, Guangzhou 510640, China (Dated: March 19, 2023) arXiv:2302.02131, PLB (submitted)



5. hadron-hadron scattering channels

Dynamics of strangeness production in heavy-ion collisions near threshold energies

Zhao-Qing Feng^{*} and Gen-Ming Jin

π and resonances (Δ (1232), N*(1440), N*(1535), ...) production:

 $NN \leftrightarrow N\Delta, NN \leftrightarrow NN^*, NN \leftrightarrow \Delta\Delta, \Delta \leftrightarrow N\pi,$ $N^* \leftrightarrow N\pi, NN \leftrightarrow NN\pi(s - state), N^*(1535) \leftrightarrow N\eta$

Collisions between resonances, NN*↔N∆, NN*↔NN*

Strangeness channels:

$$\begin{array}{l} BB \rightarrow BYK, BB \rightarrow BBK\overline{K}, B\pi(\eta) \rightarrow YK, YK \rightarrow B\pi, \\ B\pi \rightarrow NK\overline{K}, Y\pi \rightarrow B\overline{K}, \quad B\overline{K} \rightarrow Y\pi, \quad YN \rightarrow \overline{K}NN, \\ BB \rightarrow B\Xi KK, \overline{K}B \leftrightarrow K\Xi, YY \leftrightarrow N\Xi, \overline{K}Y \leftrightarrow \pi\Xi. \end{array}$$

Reaction channels with antiproton:

$$\overline{p}N \to \overline{N}N, \ \overline{N}N \to \overline{N}N, \ \overline{N}N \to \overline{B}B, \ \overline{N}N \to \overline{Y}Y$$

$$\overline{N}N \to \text{annihilation}(\pi, \eta, \rho, \omega, K, \overline{K}, K^*, \overline{K}^*, \phi)$$

Statistical model with SU(3) symmetry for annihilation (E.S. Golubeva et al., Nucl. Phys. A 537,9393 (1992)) of Dense Nuclear Matter at RIBF and FRIB, May 23-26, 2023



III. π^*/π^* ratio for constraining the high-density symmetry energy

1.6

1.5

¹³²Sn+¹²⁴Sn, b=1 fm

E/A=200 MeV



2.1

Kinetic energy spectra of π/π^+ ratio

Bao-An Li, Phys. Rev. Lett. 88, 192701 (2002) M. B. Tsang et al., Phys. Rev. C 95, 044614 (2017)



Pion production near threshold energies in heavy-ion collisions



RBUU: PRL97 (2006) 202301

1.2

 $\boldsymbol{E}_{beam}\left(AGeV\right)$

1.4

1.8

21

1.6

🔻 🗸 DDF

● NL

🗕 🖬 NLp

♦ • ♦ ΝLρδ

A-A NLDDp



Threshold effects of \triangle production by the relativistic Vlasov-Uehling-Uhlenbeck (RVUU)

Taesoo Song, Che Ming Ko, Phys. Rev. C 91, 014901 (2015)





Jun Hong and P. Danielewicz, Phys. Rev. C 90, 024605 (2014)

pBUU with γ being the stiffness parameter of SE



Temporal evolution of production rate and density profiles of particles in collisions of ¹⁹⁷Au+¹⁹⁷Au

Zhao-Qing Feng, Nuclear Science and Techniques 29 (2018) 40. (Invited review) (arXiv:1802.10294)



Equation of State of Dense Nuclear Matter at RIBF and FRIB,

May 23-26, 2023

Pion production in isotopic nuclear reactions by LQMD-RMF (Si-Na Wei, Zhao-Qing Feng, arXiv:2302.09984, PRC, no threshold correction!) The data are taken from J. Estee et al. (SπRIT Collaboration), Phys. Rev. Lett. 126, 162701 (2021).





Σ -/ Σ + and Ξ -/ Ξ - for constraining the high-density SE

Zhao-Qing Feng, arXiv: 2303.04415, Phys. Lett. B (submitted)



The STAR Collaboration, arXiv:2211.16981 (PRL, accepted)





Multi-strangeness hypernuclide production

H.G. Cheng, Z. Q. Feng, Phys. Lett. B 824 (2022) 136849





TABLE I. Comparison between cross sections of double lamda hypernuclei calculated with $r_0 = 3.5$ fm for Λ in ¹⁹⁷Au + ¹⁹⁷Au and ⁴⁰Ca + ⁴⁰Ca collisions at 3A GeV

Hypernuclei	Cross sections (mb)					
	$^{197}Au + ^{197}Au$	$^{40}Ca + ^{40}Ca$				
$^{4}_{\Lambda\Lambda}\mathrm{H}$	$2.6 imes10^{-2}$	$1.0 imes 10^{-4}$				
$^{4}_{\Lambda\Lambda}$ He	$1.0 imes10^{-2}$	$\sim 10^{-5}$				
$^{5}_{\Lambda\Lambda}H$	$5.9 imes10^{-3}$	$\sim 10^{-5}$				
$^{5}_{\Lambda\Lambda}\mathrm{He}$	$5.1 imes10^{-3}$	$\sim 10^{-5}$				
$^{5}_{\Lambda\Lambda}$ Li	$1.4 imes 10^{-3}$	$\sim 10^{-6}$				
$^{6}_{\Lambda\Lambda}\mathrm{He}$	$2.2 imes 10^{-3}$	$\sim 10^{-6}$				
Iclear MATER at RIBE and FE	$6.8 imes10^{-4}$	$\lesssim 10^{-6}$				
JUIG AL MALLOL AL NIDE ANG FR	чь,	17				

May 23-26, 2023

IV. Summary and perspective

- A soft symmetry energy with the slope parameter of $L(\rho_0) = 42\pm25$ MeV by using the standard error analysis within the range of 1σ is obtained by analyzing the experimental data from the S π RIT collaboration in reactions of ¹⁰⁸Sn+¹¹²Sn and ¹³²Sn+¹²⁴Sn at 270 MeV/nucleon.
- > The Σ^{-}/Σ^{+} ratio depends on the stiffness of symmetry energy, in particular at the beam energy below the threshold value (E_{th}=1.58 GeV), i.e., the kinetic energy spectra of the single ratios, excitation functions and energy spectra of the double ratios in the isotopic reactions of ¹⁰⁸Sn + ¹¹²Sn, ¹¹²Sn + ¹¹²Sn, ¹²⁴Sn + ¹²⁴Sn and ¹³²Sn + ¹²⁴Sn. The double strangeness ratio Ξ^{-}/Ξ^{0} weakly depends on the symmetry energy because of the hyperon-hyperon collision mainly contributing the Ξ production below the threshold energy (E_{th} = 3.72 GeV).
- Nuclear clusters and hypernuclear clusters for probing the high-density symmetry energy in the near future, t/³He, t/³He etc. May 23-26, 2023



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Thank you for your attention

Equation of State of Dense Nuclear Matter at RIBF and FRIB, May 23-26, 2023