

# Pre-shower detector sensitive to incident photon direction

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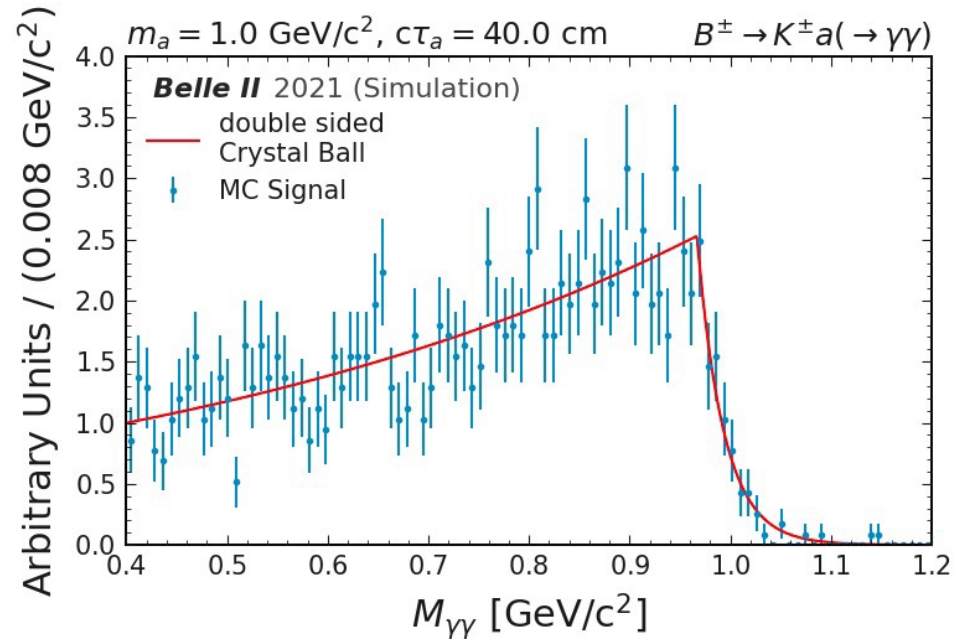
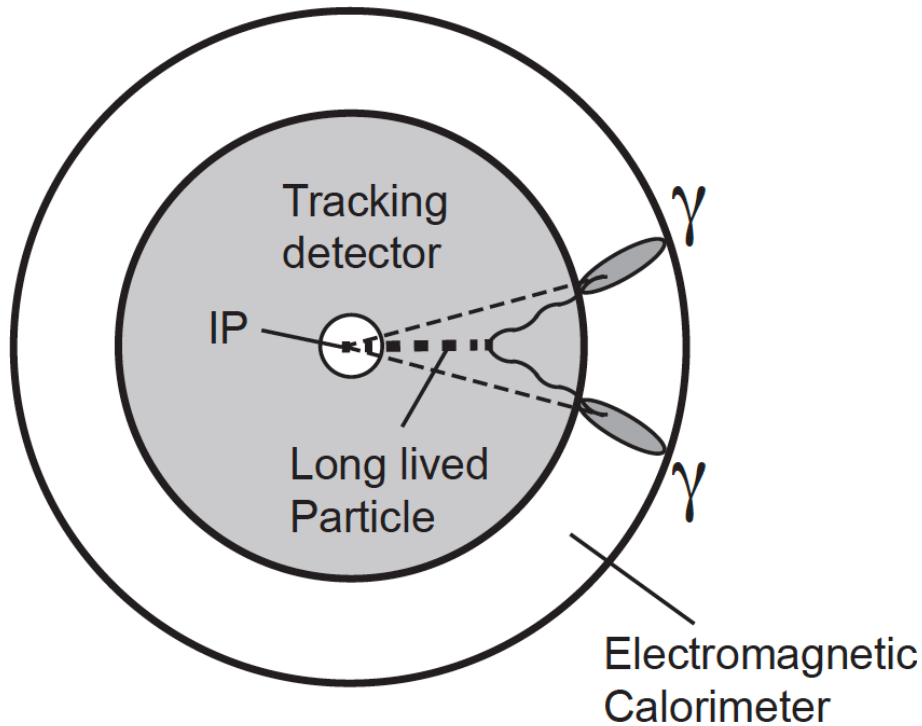
at 3<sup>rd</sup> international workshop for the  
extension of J-PARC

Hadron Experimental Facility

# Motivation

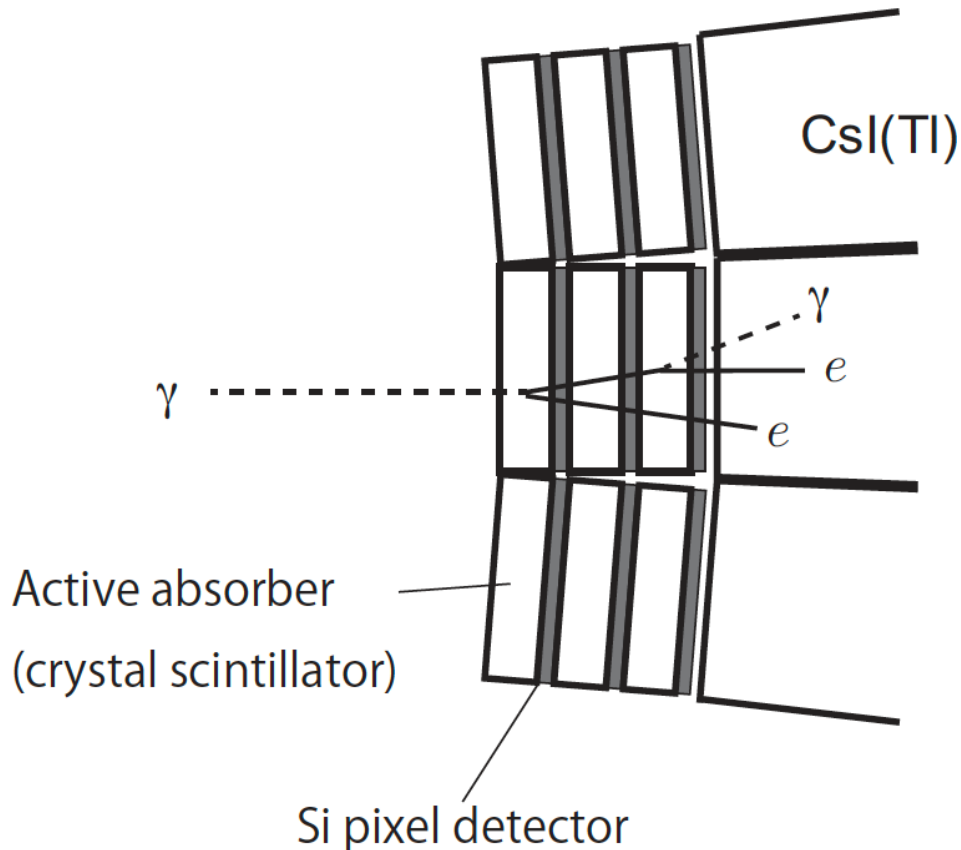
- CsI(Tl) is bright but slow, pile-up noise increase by beam background is seen at SuperKEKB/Belle II.
- Replacing all the crystals too big job to do.
- Does a pre-shower detector help?
  - To mitigate beam background effect (soft photon  $\sim 1\text{MeV}$  dominant) by using heavy and fast material
  - Making it sensitive to the photon incident direction gives a physics advantage to detect a long-lived ALP-like particle.
  - Absorber should also be active not to sacrifice energy resolution and efficiency for low energy photon (down to  $50\text{ MeV} \sim 100\text{ MeV}$ ) in SuperKEKB/Belle II environment.

# Limitation of mass determination for long-lived ALP-like in $\gamma\gamma$



Since very poor sensitivity for incident photon direction by  $5.5 \times 5.5 \times 30 \text{ cm}^3$  CsI(Tl) crystals (in total 8736 pieces), limitation of mass determination for long-lived ALP-like particles.

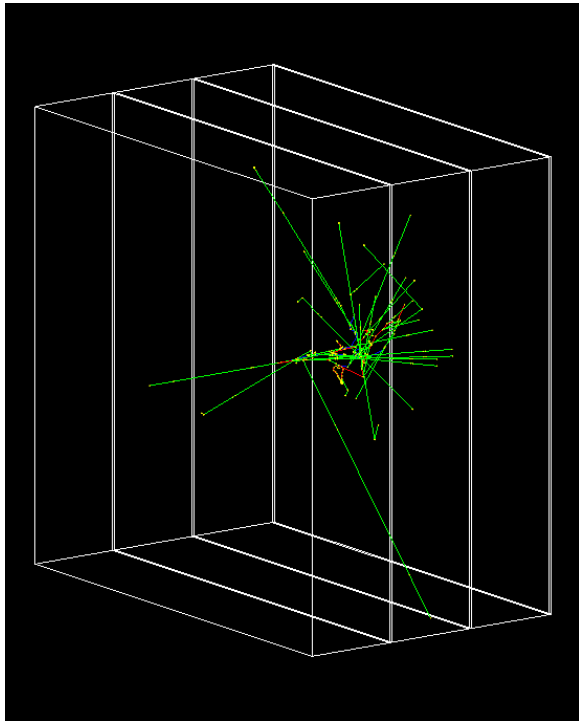
# A model of pre-shower detector



- Active absorber made by a crystal scintillator to be active, thin/compact photon sensor is necessary.
- To detect  $e^+/e^-$  passage in the shower, pixel detector inserted.
- One layer has  $1 X_0$  thickness.
- Think about 3 layers case.

# To start to look for a possibility

- On a PC (CentOS7 Linux, i7=8cores, 8GB memory) in Nara, GEANT4 v10.6 (latest) has been installed.
- Modify a sampling calorimeter example (B4a).



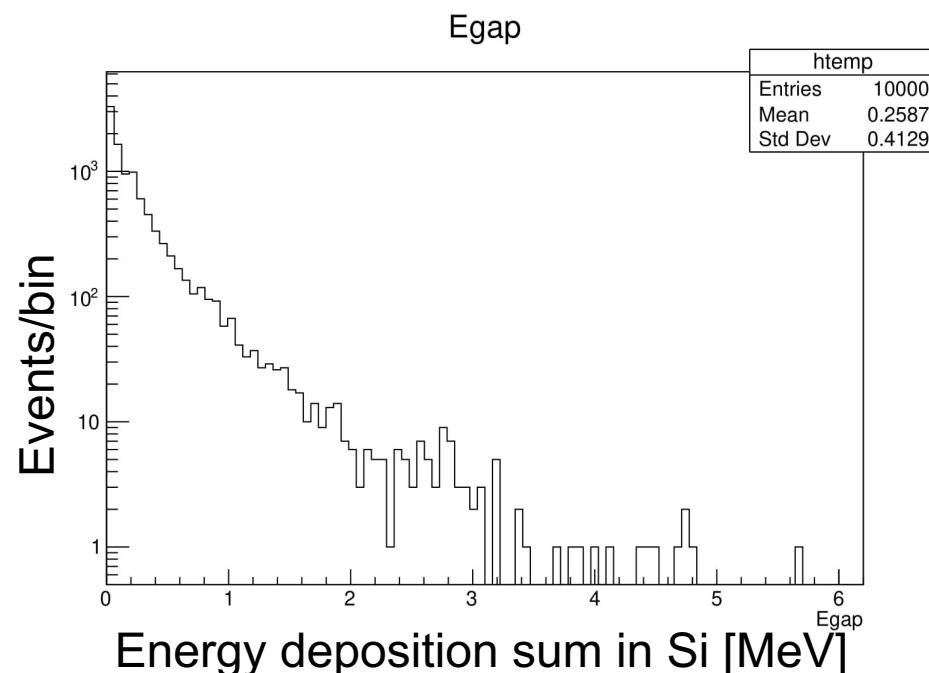
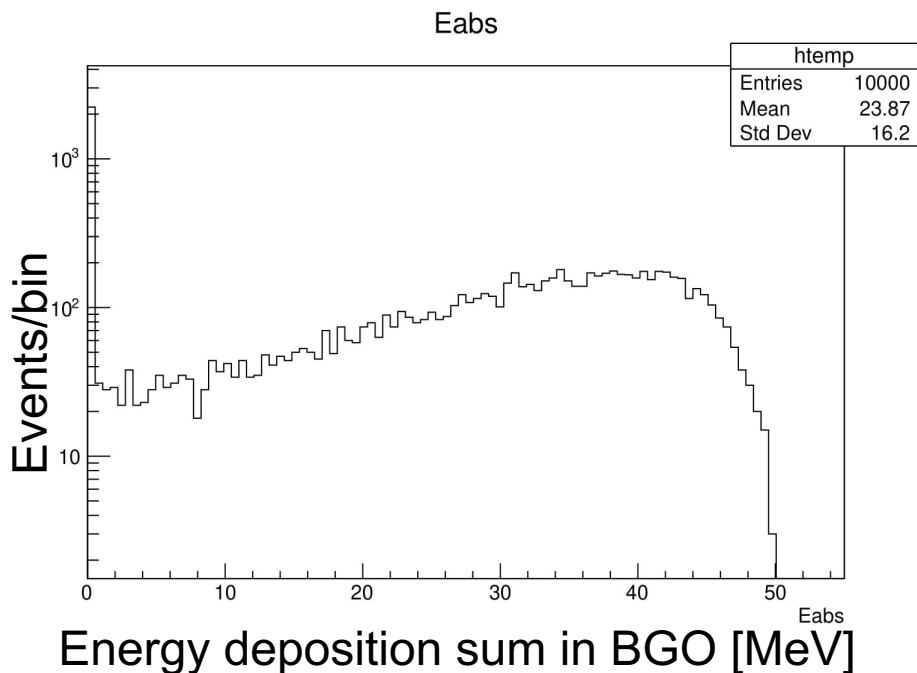
55mm × 55mm cross section.

One layer consists of:

- 11.2mm thick BGO (active absorber, corresponding to  $1 X_0$ , thickness must be similar in other scintillator case such as LYSO, etc. )
- 0.3mm thick Si (to detect for  $e^+$  or  $e^-$  passage)

An event with 200 MeV  $\gamma$  incident

# Low energy, at 50 MeV $\gamma$

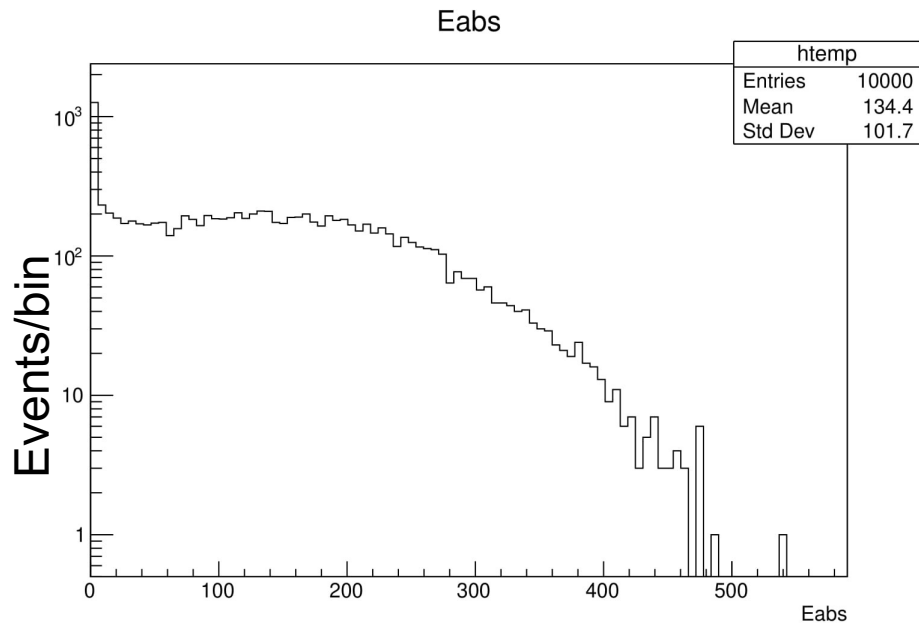


25% of events went through without any interaction.

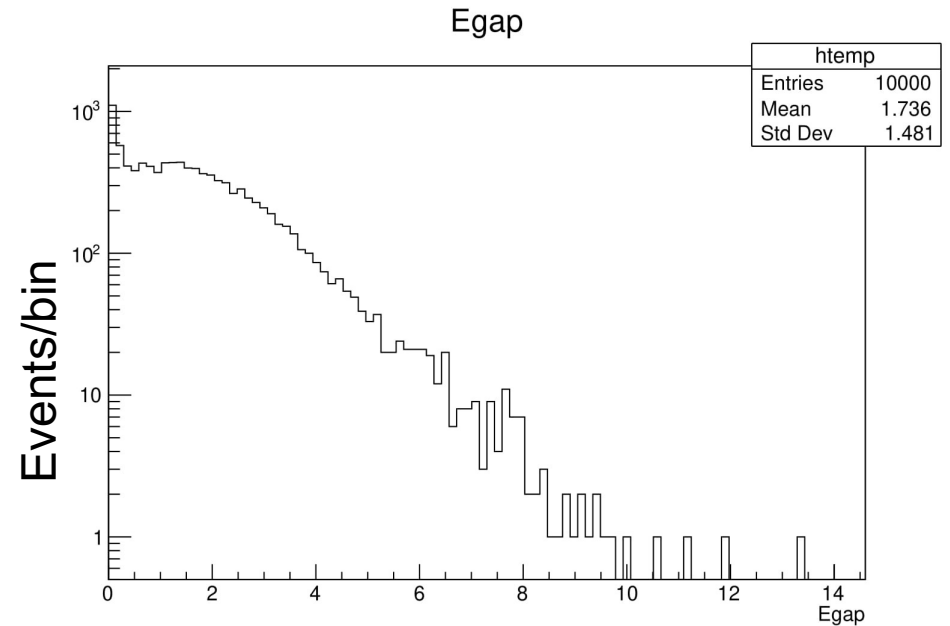
While, significant portion of events lost most of the initial energy, active absorber part should have good energy resolution.

One  $e^+$  or  $e^-$  deposits  $\sim 0.1$  MeV, average number of those is  $\sim 2.6$ .

# 1 GeV $\gamma$ case



Energy deposition sum in BGO [MeV]



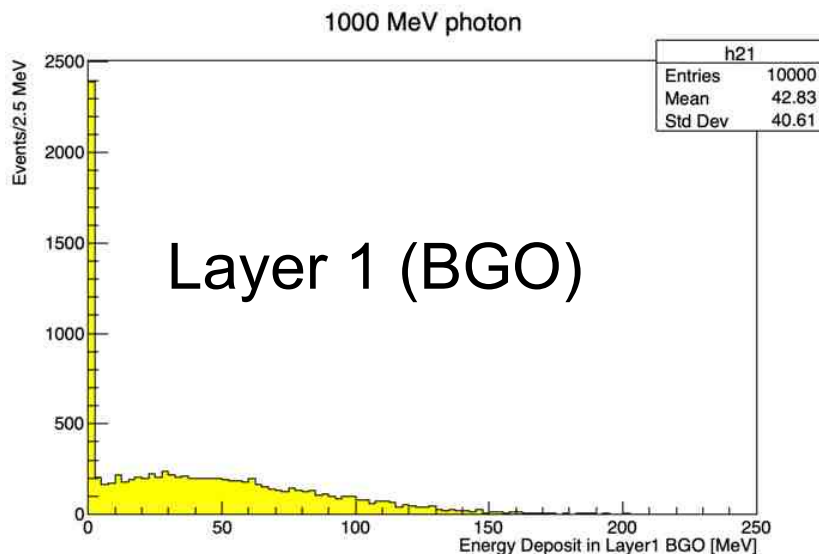
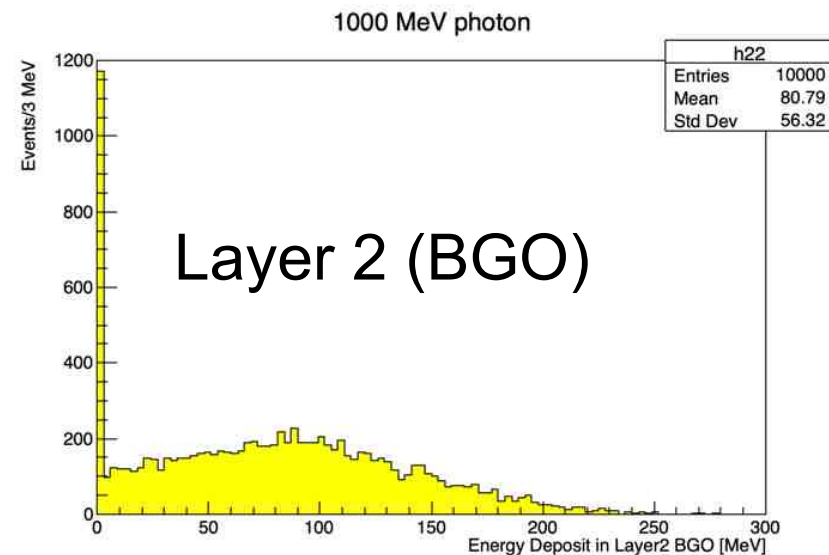
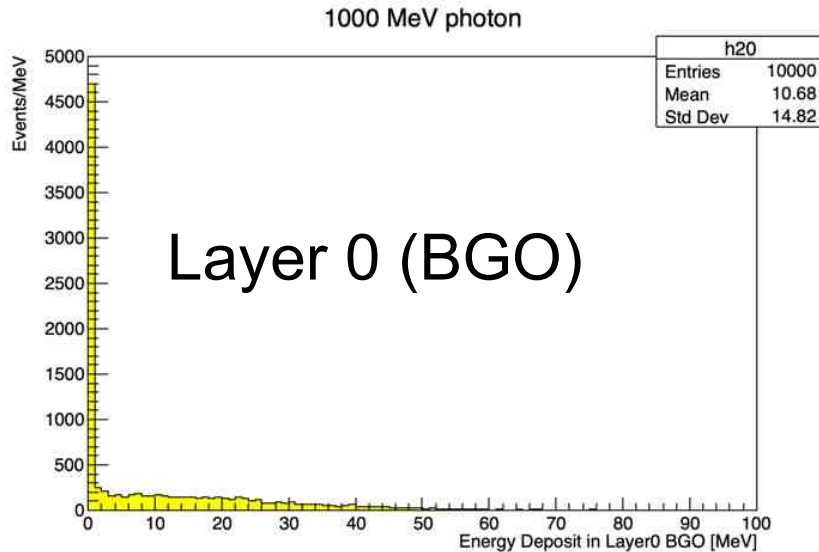
Energy deposition sum in Si [MeV]

11% of events went through without any interaction.

Significant portion of events' energy loss is a few  $\times 100$  MeV.

One  $e^+$  or  $e^-$  deposits  $\sim 0.1$  MeV, average number of those is  $\sim 17$ .

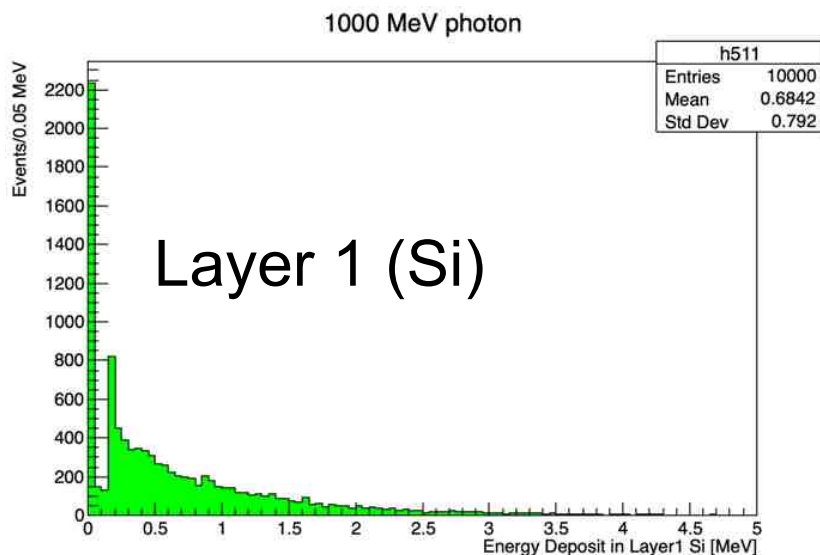
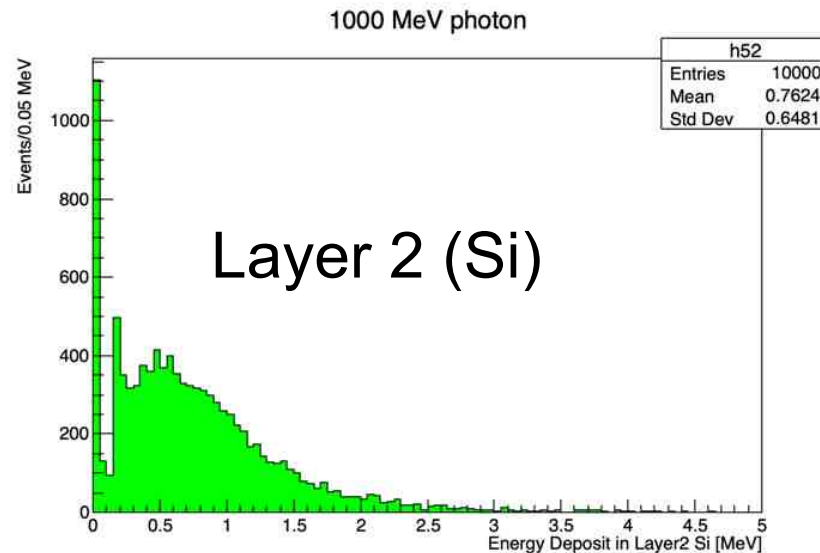
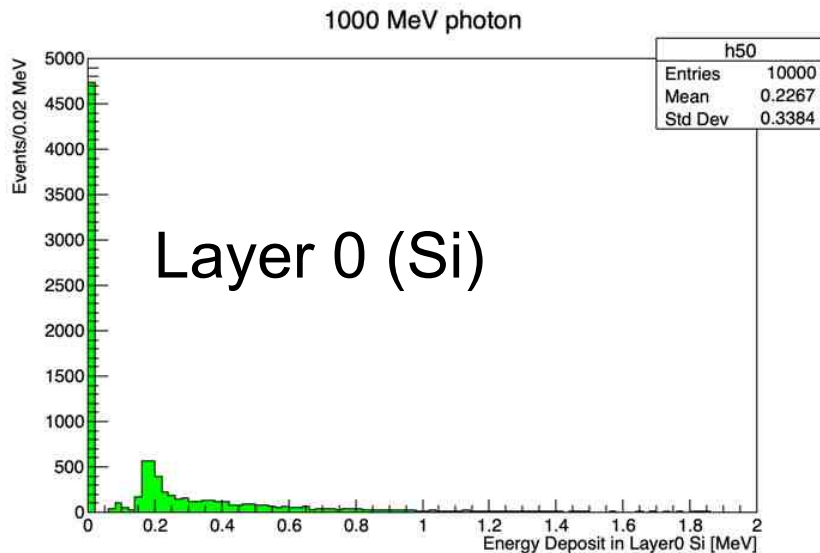
# Shower development, 1 GeV



- ~half of events don't interact in the first one  $X_0$ .
- Making first part  $2 X_0$  thick may make sense, as did in the CMS endcap pre-shower (though absorber is tungsten, inactive in their case).

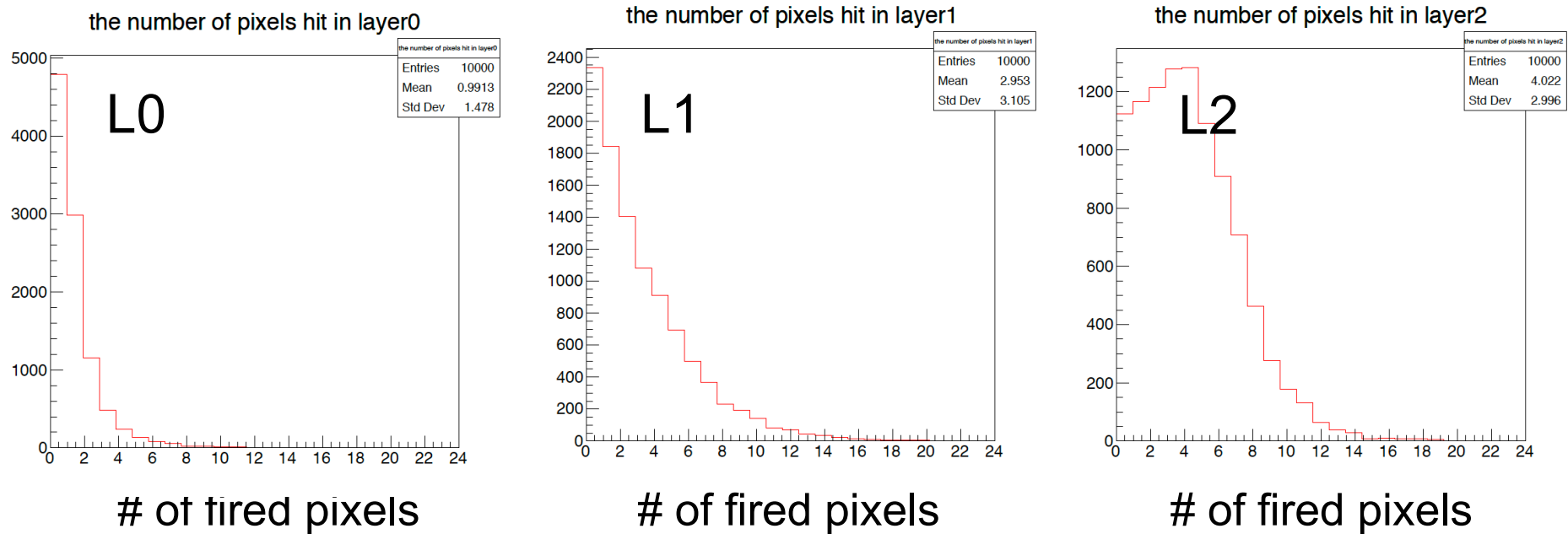


# Shower development, 1 GeV (cont.)



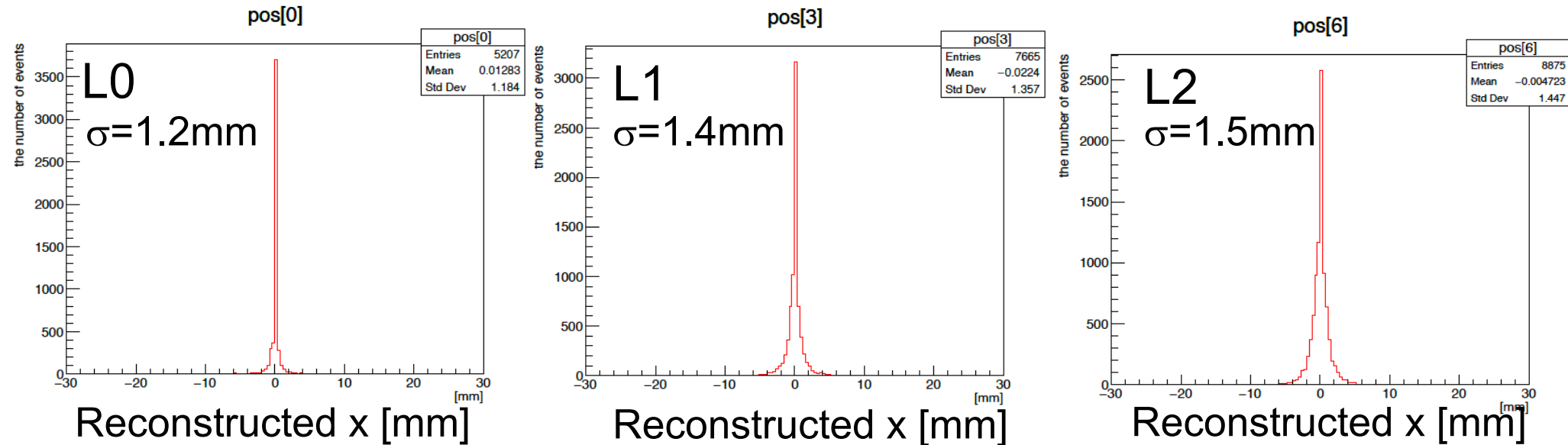
- ~80% of events have hits in 2 layers (or 3 layers) of Si.
- It looks worth to try see if we can reconstruct the direction of the incident  $\gamma$  by dividing Si into proper size pixels

# 1 GeV $\gamma$ , $1 \times 1\text{mm}^2$ each



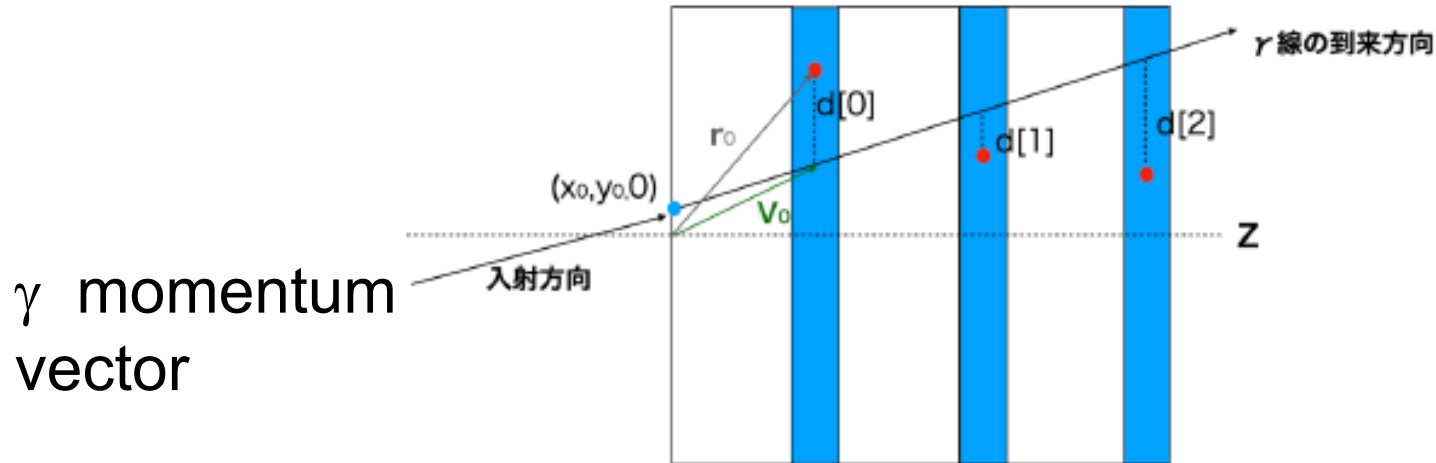
Just normal incident at the center of the detector.  
Looks fine enough to try to reconstruct  $\gamma$  direction.

# 1 GeV $\gamma$ , $1 \times 1\text{mm}^2$ pixel



In this simulation, fired pixels position are weighted by the energy deposition. In reality, binary readout of pixel information is very likely  $\rightarrow$  further study item.

# Incident angle reconstruction



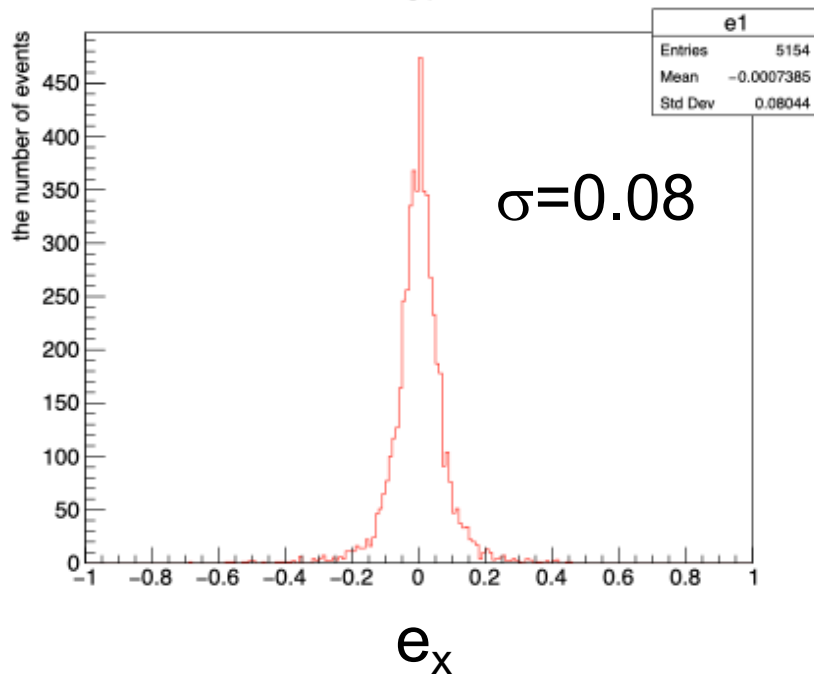
Straight-line fit to reconstruct  $\gamma$  momentum vector direction.

The parameters obtained by the fit are:

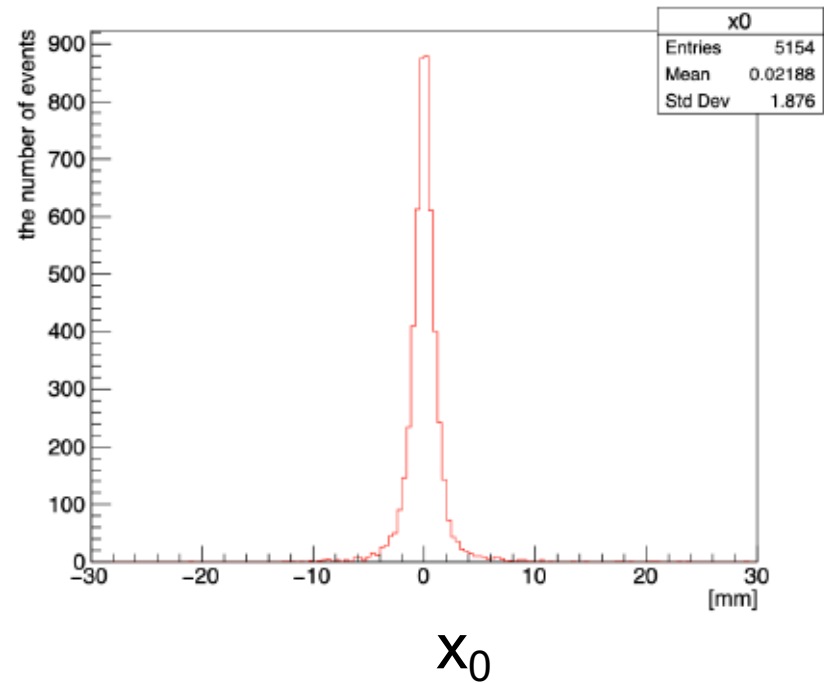
- unit vector of  $\gamma$  direction,  $(e_x, e_y, e_z)$
- incident point  $(x_0, y_0)$

# Angle and position resolutions

Normal incident 1 GeV  $\gamma$  at the center of the detector.

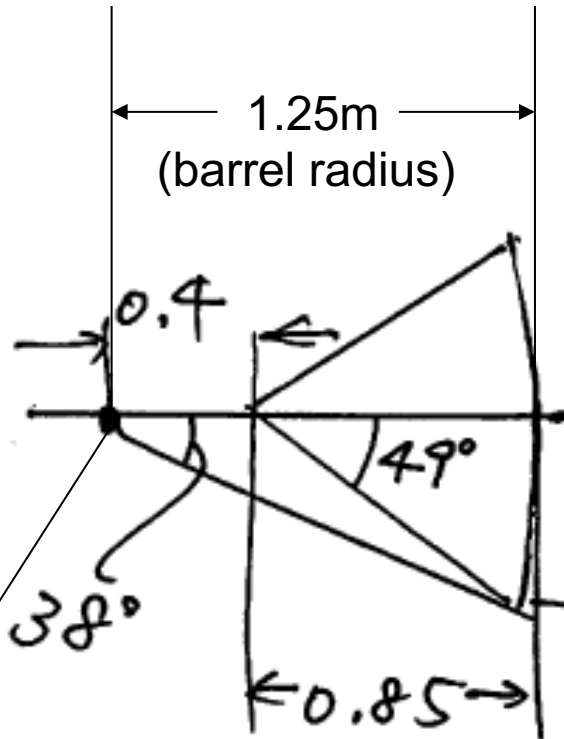


Since  $e_z \sim 1$ , incident angle resolution is estimated to be  $\sim 80$  mrad.



incident position resolution is estimated to be  $\sim 1.9$  mm.

# A simple estimation



B at rest,  $B \rightarrow Ka$ ,  $a \rightarrow \gamma\gamma$ ,  
 $m_a = 2\text{GeV}$ , 40 cm flight

When  $a \rightarrow \gamma\gamma$  symmetric energy, opening angle between them  $\sim 98^\circ$ .

The angle between  $\gamma$  direction and the straight line toward IP is  $49^\circ - 38^\circ = 190 \text{ mrad}$ ,  $\sim 2\sigma$  separation.

In  $e^+e^- \rightarrow \gamma a$ ,  $a \rightarrow \gamma\gamma$ ,  $m_a = 4\text{GeV}$ , decaying after 40 cm flight case, similar separation from the  $\gamma_s$  from the IP.

# More thoughts

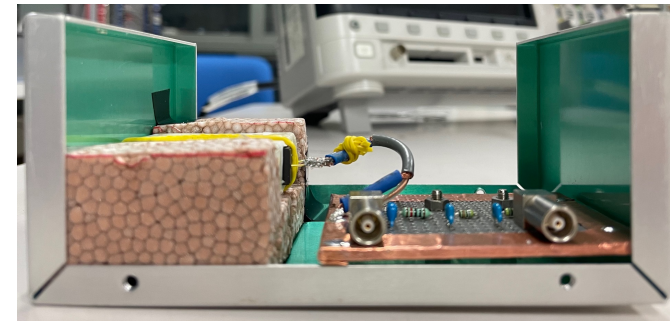
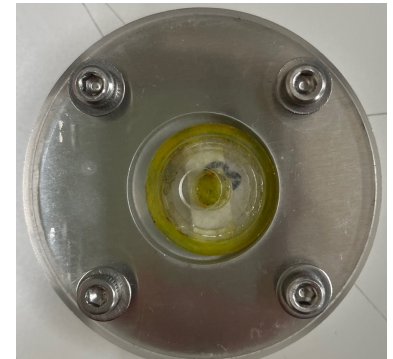
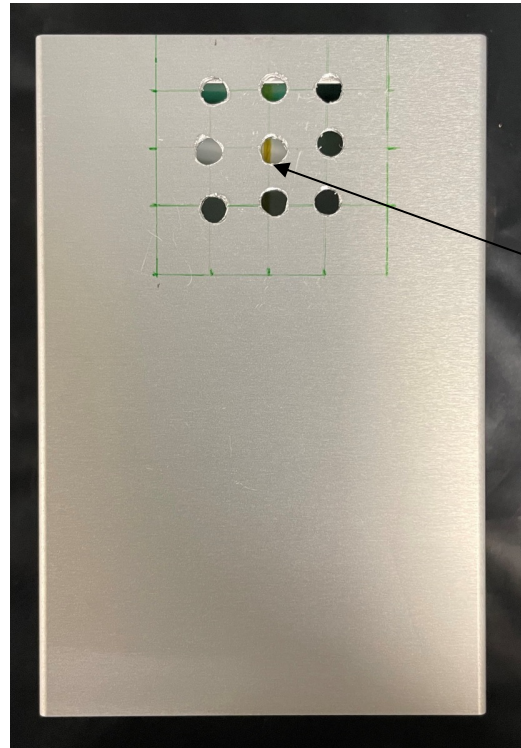
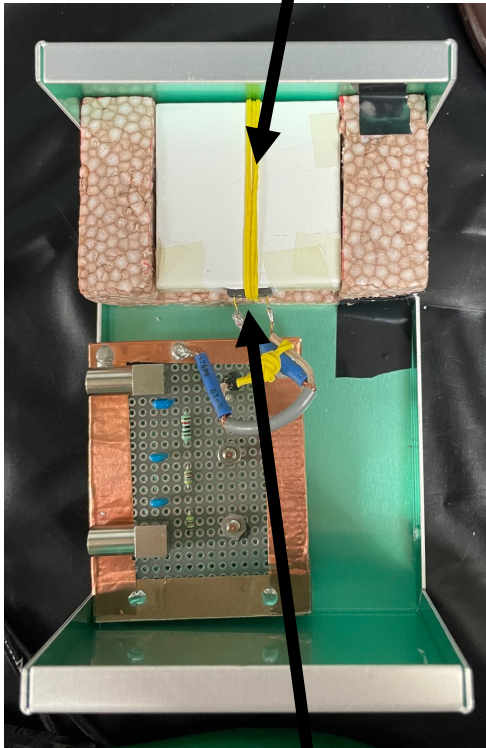
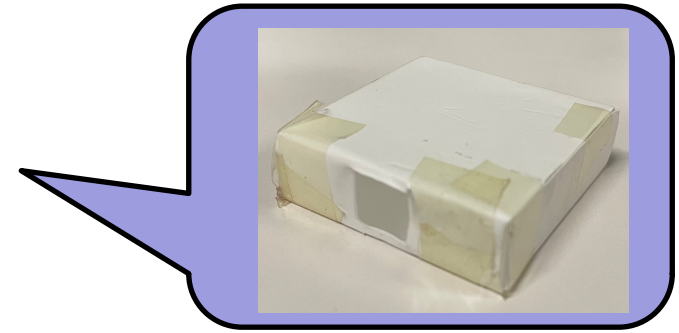
- 0.08 rad angular resolution looks limited by the pixel size (1mm) and absorber thickness(12mm).
- While it may be limited by the multiple scattering during passing through the absorber ( $1X_0$ ).
- Need simulation with finer pixel pitch → on going.
- As for the pixel sensor solution(s), we need experts' help: Sol, MAPS, etc. could be candidate(s).
- All these pixel detector solutions are binary readout of pixel hits, need to be studied by simulation.

# Light output measurement

Crystal scintillator(BGO or FastLGSO)

$4 \times 4 \times 1.2 \text{ cm}^3$

Wrapped by GoreTex(reflector)

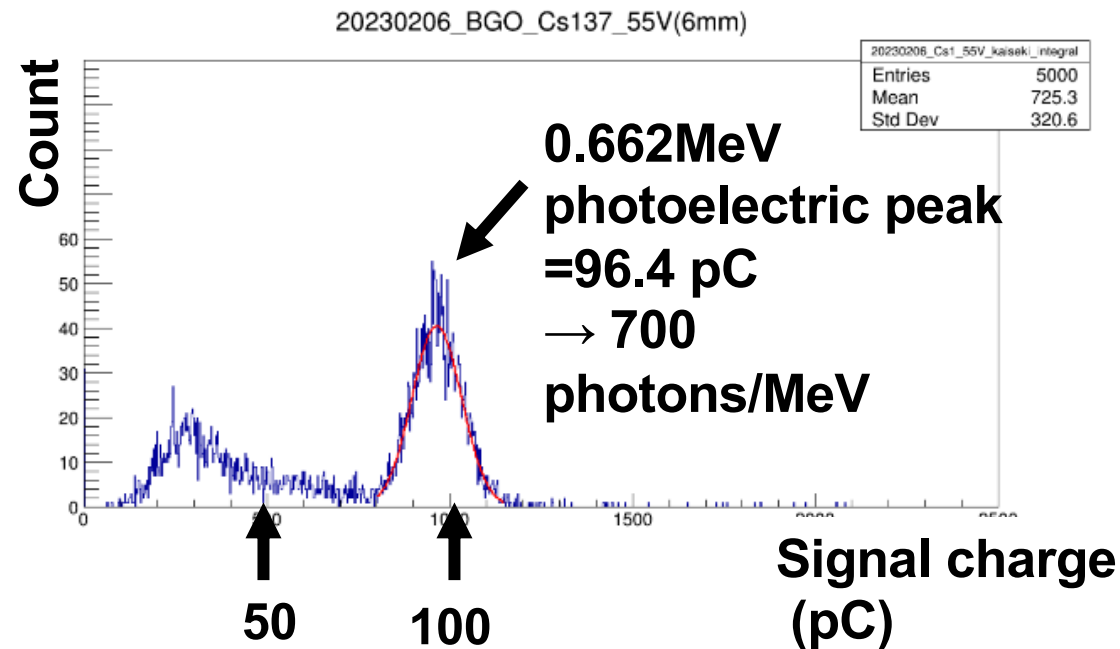
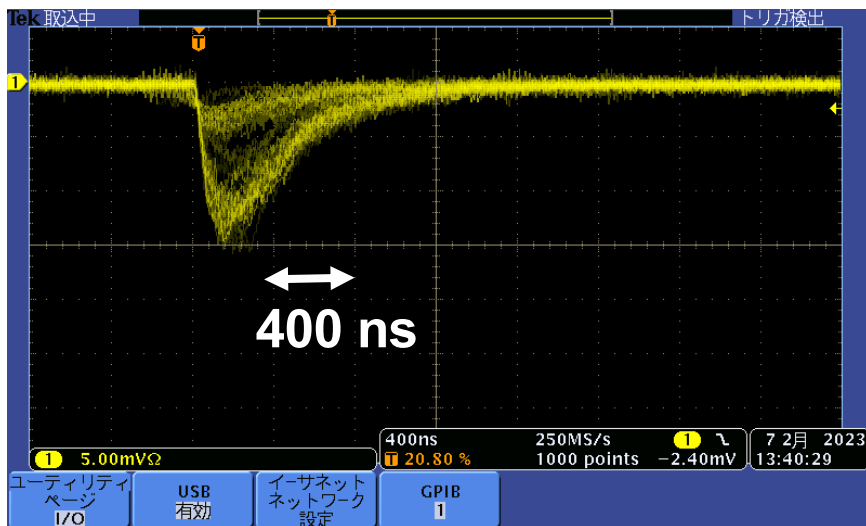


$6 \times 6 \text{ mm}^2$  MPPC(S13360-6050CS)

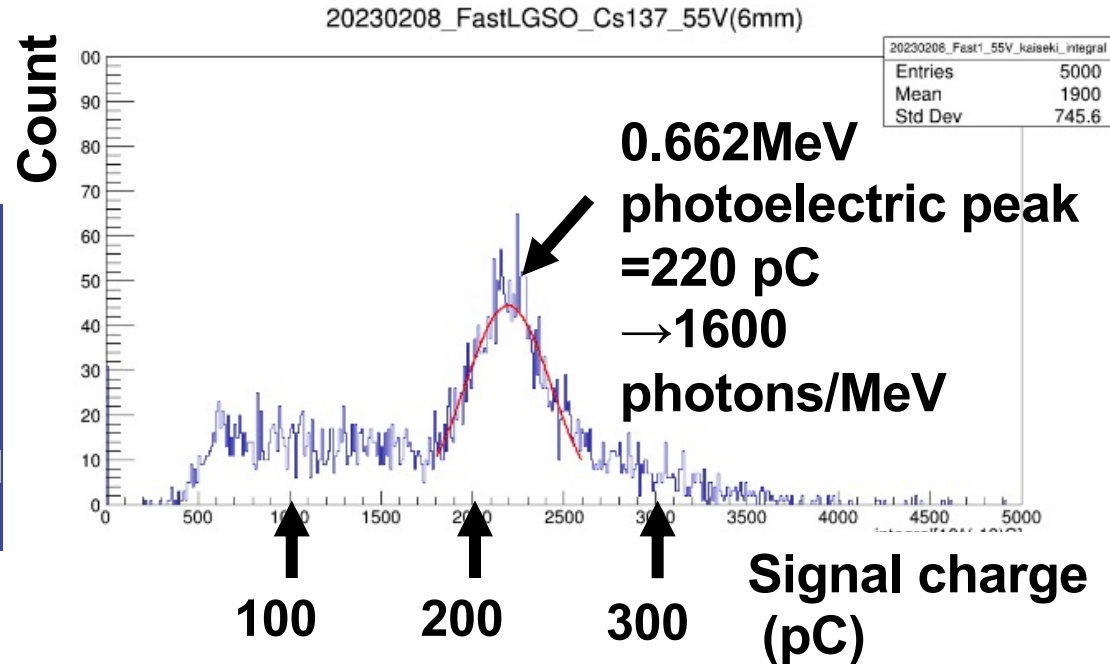
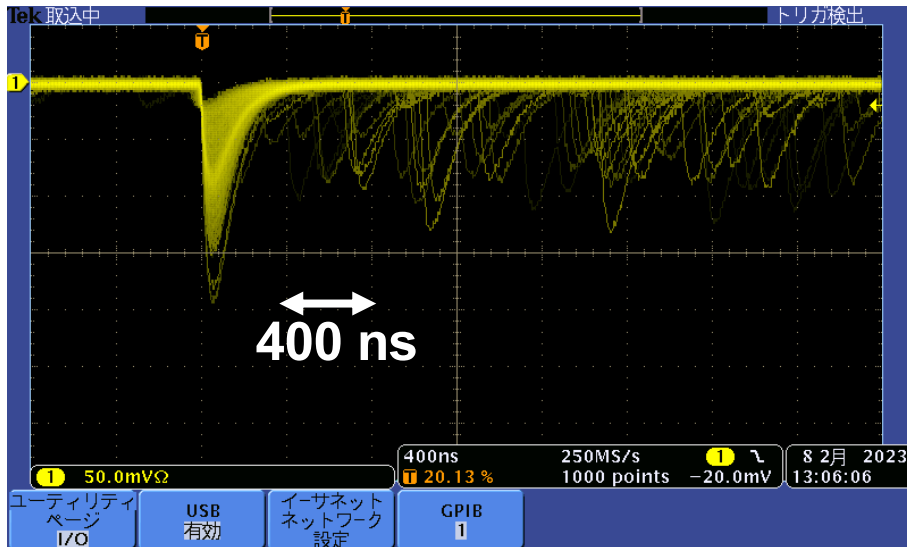


# Light output (BGO)

Signal read out by DPO3034 digital oscilloscope.  
Waveform is taken by 250Msa/s, 1000 points, numerical  
summing up them gives the signal charge.



# Light output (Fast LGSO)



- Higher light output, faster scintillation decay time than BGO, as expected.
- S13360-6050CS has 14400 pixels, measurement is still safe from the saturation due to the number of pixels.

# Remarks

- In actual Belle II detector, the TOP counter is just in front of the CsI(Tl) calorimeter.
  - Criticism “There is no space to put it”.
- Still worth to pursuit
  - Figuring out how much physics sensitivity gain.
  - Similar detector configuration is also being considered in other experiments, EIC, etc.
- Entire project is pretty large
  - Both crystal scintillation detector (active absorber) and silicon pixel detector ( $e^+/e^-$  passage detector) are necessary to be handled.
  - Needs interested friends involvement.

# Acknowledgement and reference

- Tsukasa Konno for graduation thesis FY2019
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- Hina Tagashira for work as master student
- Iori Okada for graduation thesis FY2022
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