CSM formation via super-Eddington accretion in double WD system as Progenitors of Type la supernovae

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Abstract



Light curve *****Assumption:

The SNe Ia immediately occurs after the WD merger (Violent merger scenario; e.g., Dan+12).

5. LC synthesis setup: SN Ejecta + CSM

• We synthesize the LCs of SNe Ia with our CSM model using STELLA code [Blinnikov+ 98, 00, 06].

• We adopt the W7 model for the ejecta of SNe Ia [e.g., Nomoto+ 84].



6. Synthesized LCs & observations possible DD candidate: over-luminous Ia SNe

• An initial flash (~0.1-1.0 days) appears in their LCs due to dissipation of the ejecta kinetic energy by the CSM.

Some of our synthesized LCs match well with the observational data of some over-luminous Ia.



7. Summary of our results

• The lifetime of the DD system, after which the MT exceeds super-Eddington accretion rate, is $\sim 0.1 - 1.0$ yr.

• WD binary systems, which are candidate for progenitors of SNe Ia, could form CSM.

• The timescale of initial flash in the synthesized opt. LCs is ~0.1-1.0 days $(\approx V_{\rm CSM} = 1000 \,\rm km \, s^{-1}).$

• The LCs of over-luminous Ia SNe 2020hvf, 2021zny and 2022ilv, with an initial flash in their LCs, can be explained by our CSM model.

=> consistent with the DD & violent merger scenarios.