



# The role of chiral mixing in dilepton production

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# Why chiral mixing?

- Q. Do we see any signal of chiral symmetry restoration in dilepton measurement?
- Light vector mesons change their properties in hot/dense matter --- χ-sym. restoration?
- □Strategy: vector and axial-vector states
- Axial-vector mesons can show up in vector spectrum in a medium!

<VV>  $\leftarrow$  chiral mixing  $\rightarrow$  <AA>

## My fingers crossed, J-PARC!



#### [Domokos, Harvey ('07)] Direct V-A mixing at finite µB $S_{\text{4dim}} = \int d^4x \left| \frac{1}{2} \left( \partial_\mu \pi \right)^2 - \frac{1}{2} m_\pi^2 \pi^2 - \frac{1}{4} \left( \rho_{\mu\nu} \right)^2 - \frac{1}{4} \left( a_{\mu\nu} \right)^2 \right|$ $\left. + \frac{1}{2}m_{\rho}^{2}\rho_{\nu}^{2} + \frac{1}{2}m_{a}^{2}a_{\mu}^{2} + C\epsilon^{ijk}\left(\rho_{i}\partial_{j}a_{k} + a_{i}\partial_{j}\rho_{k}\right)\right|$ $p_0^2 - |\vec{p}|^2 = \frac{1}{2} \left[ m_\rho^2 + m_{a_1}^2 \pm \sqrt{(m_{a_1}^2 - m_\rho^2)^2 + 16C^2 |\vec{p}|^2} \right]$ 10<sup>0</sup> longitudinal transy: C=0.5 GeV transverse transv: C=1 GeV 2.5 average 10<sup>-1</sup> ImG<sub>v</sub> [Gev<sup>2</sup>] p<sub>0</sub> [GeV] 2 1.5 1 10<sup>-3</sup> Not BW! 0.5 10<sup>-4</sup> 0 1.5 0.5 0 1 0.5 1.5 2 0 2 s<sup>1/2</sup> [GeV] p [GeV]

### [Harada, CS ('09)] Chiral mixing induced from WZW

**Wess-Zumino-Witten term** [Kaiser, Meissner ('90)]

$$\mathcal{L}_{\omega\rho a_{1}} = g_{\omega\rho a_{1}} \epsilon^{\mu\nu\lambda\sigma} \omega_{\mu} \left[ \partial_{\nu} V_{\lambda} \cdot A_{\sigma} + \partial_{\nu} A_{\lambda} \cdot V_{\sigma} \right]$$
$$\langle \omega_{0} \rangle = g_{\omega NN} \cdot n_{B} / m_{\omega}^{2} \qquad C = g_{\omega\rho a_{1}} \cdot g_{\omega NN} \cdot \frac{n_{B}}{m_{\omega}^{2}}$$

 $\Box Mixing strength: C = 0.1 GeV at \rho_0$ 

- AdS/QCD  $\rightarrow$  C = 1 GeV at  $\rho_0 \rightarrow$  vector cond.!?
- Why so large? --- higher-lying states in large Nc cf. VMD in SS  $C_{hQCD} \sim C_{\omega\rho a_1} + \sum C_{\omega^n \rho a_1}$

#### Weak mixing ... No impact?

## A missing piece: xsym. restoration

 $<AA> \rightarrow <VV>$ 

# Chiral restoration vs. mixing

Dispersion relations for small 3-momenta

$$p_0^2 \simeq m_{a_1,\rho}^2 + \left(1 \pm \frac{4C^2}{m_{a_1}^2 - m_{\rho}^2}\right) \bar{p}^2$$

The mixing effect will be enhanced as δm decreases!

- ≻In-medium δm
- ➤In-medium mixing C
- ← Quark-nucleon hybrid model[NS: Marczenko et al. (19,20)]

#### Spectral function at T = 50 MeV ImG[GeV<sup>-2</sup>] ImG[GeV<sup>-2</sup>] ImG[GeV<sup>-2</sup>] 0.1 0.1 0.1 0.01 0.01 0.01 .001 0.001 0.001 0.0001 0001 0.0001 √s [GeV] √s [GeV] 0 0.250.50.75 250.50 0 0 250.50 ➤ Near µc Low µ ImG[GeV<sup>-2</sup>] ImG[GeV<sup>-2</sup>] ImG[GeV<sup>-2</sup>] 0.1 0.1 0.1 0.01 0.01 0.01 0.001 0.001 .001 0001 0.0001 0.0001 √s [GeV] √s [GeV] 0 0.250.50.75 1 1.251.5 0 0.250.50.75 1 1.251 0 0.250.50.75 251.5

(top) chiral restoration (bottom) no restoration--- longitudinal --- transverse --- average



R. Ejima, P. Gubler, C. Sasaki and K. Shigaki, in preparation

# TOWARD A MEASUREMENT OF CHIRAL RESTORATION AT J-PARC

For more details, talk by Ren Ejima

# φ meson in nuclear matter

 $\Box$ No  $\phi$ N resonances, but the kaon cloud. □ Kaon in nuclear matter: Kaplan, Nelson (86)  $m_K^* = \left[m_K^2 - a_K \rho_S + (b_K \rho)^2\right]^{1/2} + b_K \rho,$  $m_{\bar{K}}^* = \left[m_{K}^2 - a_{\bar{K}}\rho_S + (b_{K}\rho)^2\right]^{1/2} - b_{K}\rho,$  $b_K = 3/(8f_\pi^2)$   $a_K = a_{\bar{K}} = \Sigma_{KN}/f_\pi^2$ Li, Lee, Brown (97): kaon production in Ni+Ni at 1 & 1.8 A GeV  $a_K \approx 0.22 \text{ GeV}^2 \text{ fm}^3 \text{ and } a_{\bar{K}} \approx 0.45 \text{ GeV}^2 \text{ fm}^3$ 



 Strong evidence of partial (~30%) restoration in pionic atoms [Nishi et al., Nature Physics, 2023]
 Density-induced chiral mixing in broken phase
 More structure & their shift due to f<sup>\*</sup><sub>π</sub> in SF

# E16 experiment at J-PARC



Measurements of spectral change of vector mesons in nuclei

□ Proton beam at 30-50 GeV







Run 1 (Dec 2024): 15k φ mesons Run 2 (?): 69k φ mesons



# **Dilepton production**

InvMassDist = 
$$\int \left[ \int \frac{dN}{d\vec{p}d\rho dt} \frac{d\vec{p}}{dp} d\rho dt + \int \frac{Bkg(s,p)}{dp} dp \right] g(m-s) ds$$

Spectral Fx Kinematic dist Background Detector responce





# Statistical significance



### SUMMARY

# Final remarks

- Parity doubling of hadrons as signatures of chiral symmetry restoration in a medium
- Density-induced chiral mixing in cold dense matter
  - Estimated signatures at J-PARC E16 experiment (p+Pb) via dilepton production
  - Run-2 adequate

#### BACKUP

# Vector-current correlator $G_V^L = \left(\frac{g_{\rho}}{m_{\rho}}\right)^2 \frac{-s}{D_V}, \quad G_V^T = \left(\frac{g_{\rho}}{m_{\rho}}\right)^2 \frac{-sD_A + 4C^2\bar{p}^2}{D_V D_A - 4C^2\bar{p}^2},$ $D_{V,A} = s - m_{\rho,a_1}^2 + im_{\rho,a_1}\Gamma_{\rho,a_1}(s),$

Im and Γ: *in-medium* masses and widths
Strategy of an illustrative computation:

- Modify only mass and width of axial-vector states.
- Set G\_A equal to G\_V at CSR, according to  $\Gamma_a 1 = \Gamma(a 1 \rightarrow \rho \pi) + \delta \Gamma(f_p i) \rightarrow \Gamma_p$

#### Mass difference vs. mixing : T=50 MeV





## Spectral function of pmeson

■ At chiral crossover with p = 0.1, 0.5, 1.0 GeV■ S-wave vs. p-wave states ■ CSR (m\_- → m\_+) vs. no CSR (m\_- ≠ m\_+)





#### [Li, Lee, Brown (97)] [Chung, Ko, Li (98)] Kaon and anti-kaon



#### Int.over p > 0.5 GeV Dilepton rates at T=50 MeV



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