Nuclear Physics

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Fundamental interactions

Elementary particles

Composite particles

Nuclear Physics is not Particle Physics, not Condensed Matter Physics

Interaction? Effective interaction ? Correlations ? Isospin, Density, temperature dependences ? Surface boundary, non-linear, finite system Collective motions

- Q. 1 Life time of neutron?
- Q. 2 Age of universe is 13.7B Years after BigBang. At present, there are neutrons in materials. Why?
- Q. 3 Spin-parity for ground state of deuteron?
- Q. 4 Limits of existence of nuclei?
- Q. 5 Magic numbers of nuclei?
- Q. 6 Size of nuclei?
- Q. 7 Collective motions of nuclei?
- Q. 8 How and where elements around us have been created ?

Exploration of the Limit of Existence



Nuclear Collective Motion



Solar Abundance of Elements



T.Motobayashi

Gold







Exploration of the Limit of Existence



New frameworks for the new region of nuclear chart



To write up new text book: Exotic phenomena, Systematics, etc. Isospin-, density-dependences of effective interactions, nucleon-corrections Microscopic system (nuclei) to Macroscopic system (neutron stars)

Liberation from Stable Region and Exotic Nuclei

Shell Evolution : magicity loss and new magicity







Dynamics of new "material" : Neutron-skin(halo)



New quantum objects with two surfaces neutron skin Skin thickness? Density distribution? Role of skin in reactions? Pairing in skin? di-neutrons? Exotic modes of skin?

RIBF provides data for nuclei far from the stability line

Challenges in establishing new frame work of nuclear physics

Challenge for r-process path and explosion in supernovae

Synthesis up to U (r-process) unknown neutron-rich nuclei theoretical predictions only

Necessary of experimental investigation for nuclear properties of heavy and neutron-rich nuclei Mass, life-time, decay mode



Explosion mechanism of supernova No explosion in theoretical works

Outer clast of neutron star

Necessary of experimental study for Equation-of-State for nuclear matter



1987A

Challenge to investigate EOS of neutron matter

from nuclei to neutron stars



3NF

T=3/2 channels? density dependence?

Elastic d+p for T=1/2 Nuclear structure in very neutron-rich nuclei for T=3/2? Heavy-ion Collisions to achieve ρ~2-3ρ0 ?

³P₂ correlation

pairing gap? Density dependence?

????

Role of di-neutron in skin? : collectivity, transfer reactions

RIKEN RI Beam Factory (RIBF)



Intense (80 kW max.) H.I. beams (up to U) of 345AMeV at SRC Fast RI beams by projectile fragmentation and U-fission at BigRIPS Operation since 2007

Exploration of the Limit of Existence



Large-scaled Facilities in the world



RI beam production via in-flight method





World's First and Strongest K2600MeV Superconducting Ring Cyclotron

400 MeV/u Light-ion beam 345 MeV/u Uranium beam

World's Largest Acceptance 9 Tm Superconducting RI beam Separator

~250-300 MeV/nucleon RIB



Identification of 45 New Neutron-Rich Isotopes Produced by In-Flight Fission of a ²³⁸U Beam at 345 MeV/nucleon



T. Ohnishi, et al., JPSJ 79, 073201 (2010).

Nov., 2008 Averaged beam intensity ~0.2 pnA Maximum intensity 0.4 pnA

Mn (Z=25) to Ba (Z=56) Covered by three Brho settings Be and Pb targets Total dose 1-2x10¹⁴ for each Brho setting

Even Z

Odd Z

150

Yield rates reasonably reproduced by LISE++



Press-Conference on June 8th, 2010

June 8, 2010 RIKEN

Scientists discover 45 new radioisotopes in 4 days



2004
2003

> 2002

2001

2000

1999

BigRIPS has found 45 new radioisotopes

The chart of the known nuclides has been extended significantly by physicists in Japan, who have discovered 45 new neutron-rich isotopes. The nuclei were spotted at the RIKEN laboratory by smashing a powerful beam of heavy ions into beryllium and lead targets.

Half-Lives of Very Neutron-Rich NucleiS. Nishimura et al.(Kr,Rb,Sr,Y,Zr,Nb,Mo,Tc) around 2nd R-Process PeakS. Nishimura et al.







8 hour data acquisition
T1/2 data of 38 isotopes including first data for 18 isotopes
FRDM may underestimate Q-value for Zr and Nb by 1 MeV at A~110
More rapid flow in the rapid neutron-capture process than expected S. Nishimura et al., PRL 106 (11) 052502

1/3 ~ 1/2 Shorter Half-lives of Zr and Nb (A~110)





Faster r-process synthesis in supernova explosion ? —new half-life data for 18 neutron-rich nuclei

	2
超新星爆発で重元素 再現	
超新星爆発で重い元素が作られる様子	
を地上で再現したところ、一部の元素は	
これまでの理論よりも多く作られること	
がわかった。理化学研究所が発表した。	
太陽系にあるいくつかの元素は理論予想	由子星きえ
より10倍多く、どう作られたのか20年間	性核市たた 子が心とと
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素の工場だった可能性が高まった。	に過たが新変剰鉄有星わとな力爆
理研の西村俊二先任研究員らは、埼玉	るなどだ。帰る
県和光市の大強度重イオン加速器施設	て超だは
「RIBF」で、ウランを光速の70%に加	い新け「重要」
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ウランが分裂してできた元素のうち、18	(」と置た。成小話程。の野すをこ謎
種類は世界で初めて寿命が調べられたと	晋。解れを 史)明か少
いう。(フィジカル・レビュー・レターズ)	050

	日で合成 同で合成 一時 の一部のを示したのは世界初 その一部の言元素合成過 差による言元素合成過 差による言元素合成過 差による言元素合成過 その一部の言元素合成過 たちる記が有力だ。 にとうの超新星爆発でで ときの超新星爆発をご にとるの指新星爆発でで ときの超新星爆発でで ときの超新星爆発でで ときの超新星爆発でで ときの超新星爆発でで	平成23年(3 中子 星 きえは ウ 生 稼 中たた、 う 自電	2011年 ムの実験で得られた。超新 レイアンディアングロアンディングロアンディングロアングロアングログラング しんしょう しょうしん しょうしょう しょうしょうしょう しょうしょう しょうしょう しょうしょうしょう しょうしょう ひょうしょう ひょうしょう ひょうしょう ひょうしょう ひょうしょう しょうしょう ひょうしょう ひょうしょう ひょうしょう ひょうしょう しょうしょう しょうしょう ひょうしょう しょうしょう しょうしょう ひょうしょう ひょうしょう ひょうしょう しょうしょう しょう	2 合取れてご思いが、理ビを、 速を 7 で、一部の元素は従来の理 太陽 第 7 で、一部の元素は従来の理 太陽	月曜日超新星爆発によって鉄よ星爆発の玉	9 重元素は短時間
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DayOne Experiments in Dec., 2008 -The first data in the "island-of-inversion" -



Coordinated by Aoi



A new candidate of halo nuclei ³¹Ne via Coulomb breakup

Nakamura et al., PRL 103, 262501(2009)



> Total interaction cross sections for the neutron-rich Ne isotopes

Takechi, Otsubo et al., Niigata 2010 symposium

Spectroscopy of ³²Ne and the "island-of-inversion" E(2+) = 722 keVDoornenbal, Scheit et al. PRL 103, 032501 (2009) New states in ^{31,32,33}Na PRC 81, 041305R (2010)



The First RI Beam Experiment (1974)



Development of Nuclear Physics and Applications



New Devices of RIBF

To maximize the potentials of intense RI beams available at RIBF



Recoiled proton tagged knockout reaction for He-8 at RIPS

PKU-RIKEN-IMP-TITech-Seaul

Ye et al., July 2009







"Quasi-elastic scattering of He-6 from C-12" Lou and Ye et al., PRC 83, 034612 (2011)

Cao and Ye et al., in preparation

Challenge

Action

Discussion

Enjoy