

# Neutron removal reactions of Ne, Mg, and Si isotopes near/inside the island of inversion

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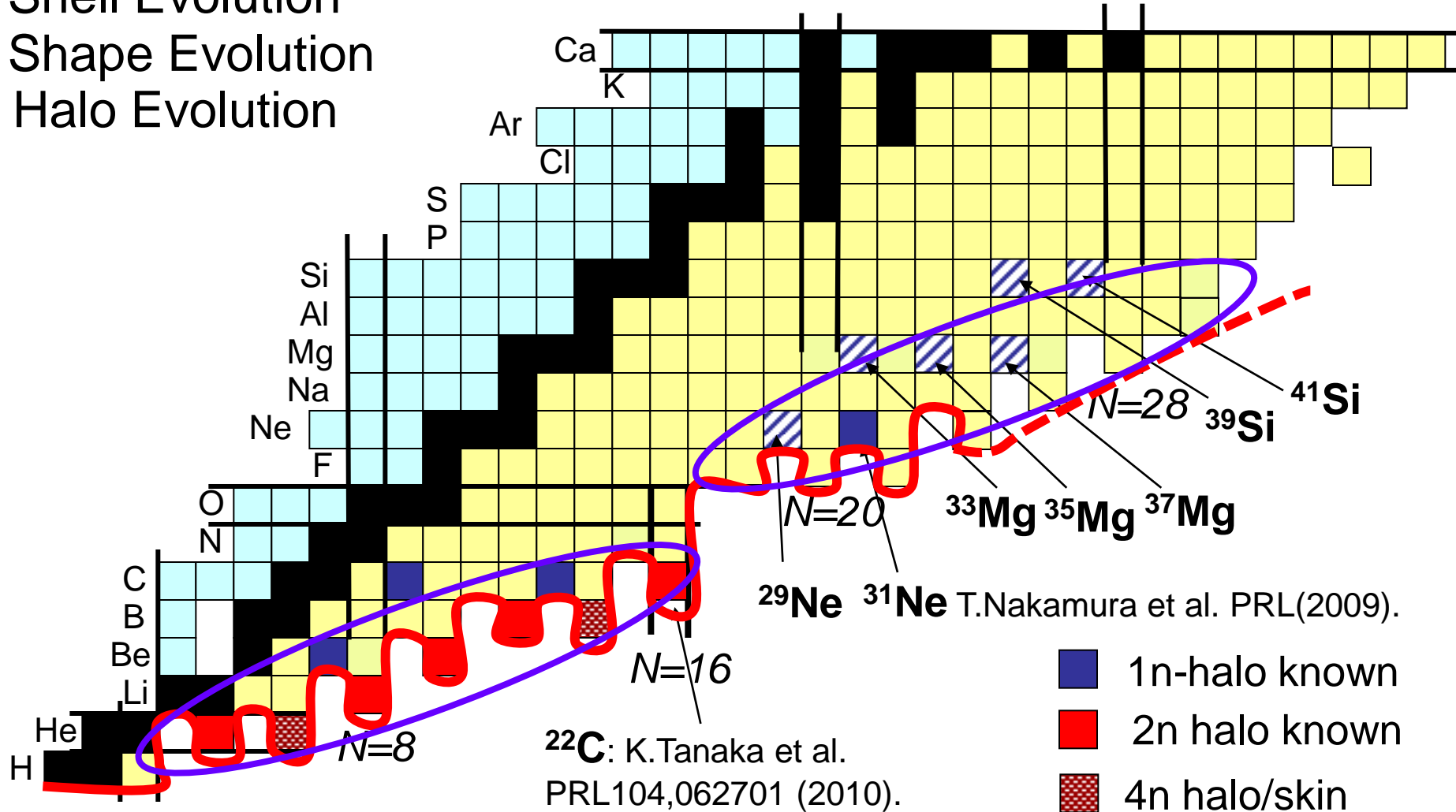
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# Evolution Towards the Stability Limit

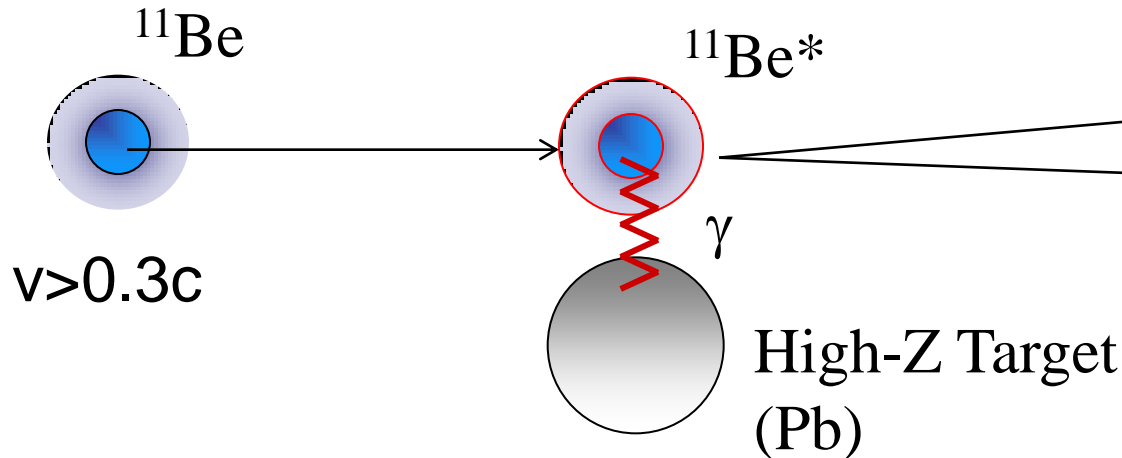
## A 30~40 ( $20 < N < 28$ )

Shell Evolution  
 Shape Evolution  
 Halo Evolution



# Probe-1: Coulomb Breakup

→ Photon absorption of a fast projectile



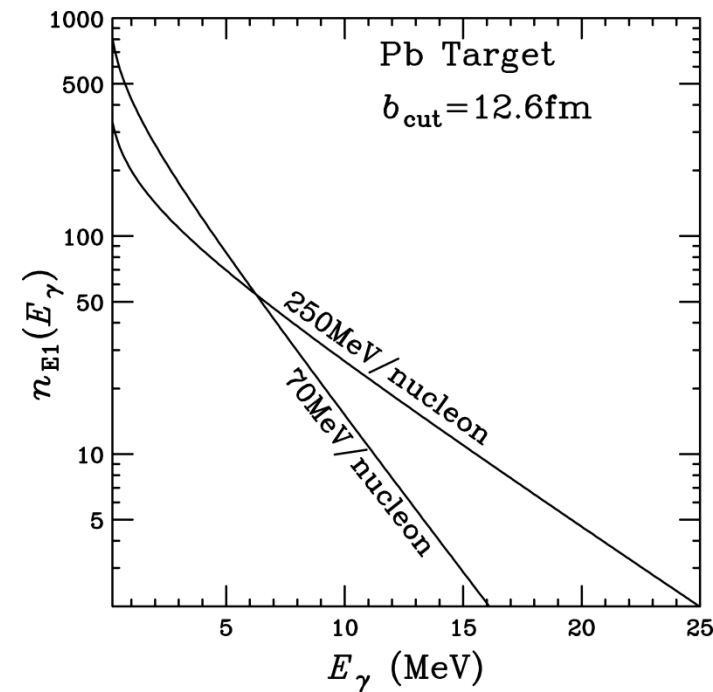
$\vec{P}(n), \vec{P}(^{10}\text{Be})$   
Invariant Mass  
⇒  $E_x, E_{\text{rel}}$

Equivalent Photon Method

$$\frac{d\sigma_{CB}}{dE_x} = \frac{16\pi^3}{9\hbar c} N_{E1}(E_x) \frac{dB(E1)}{dE_x}$$

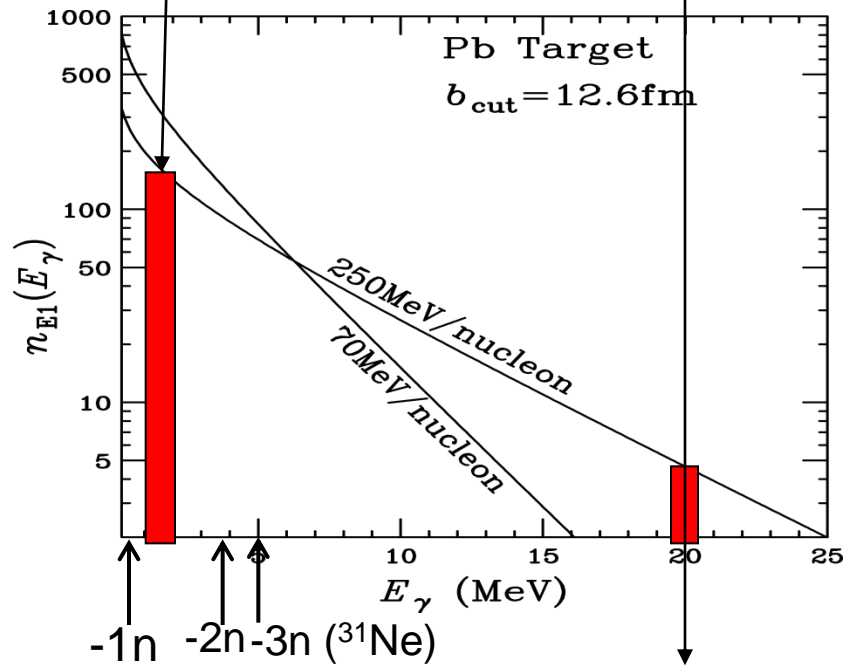
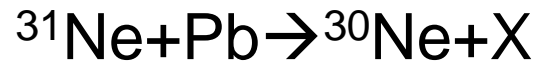
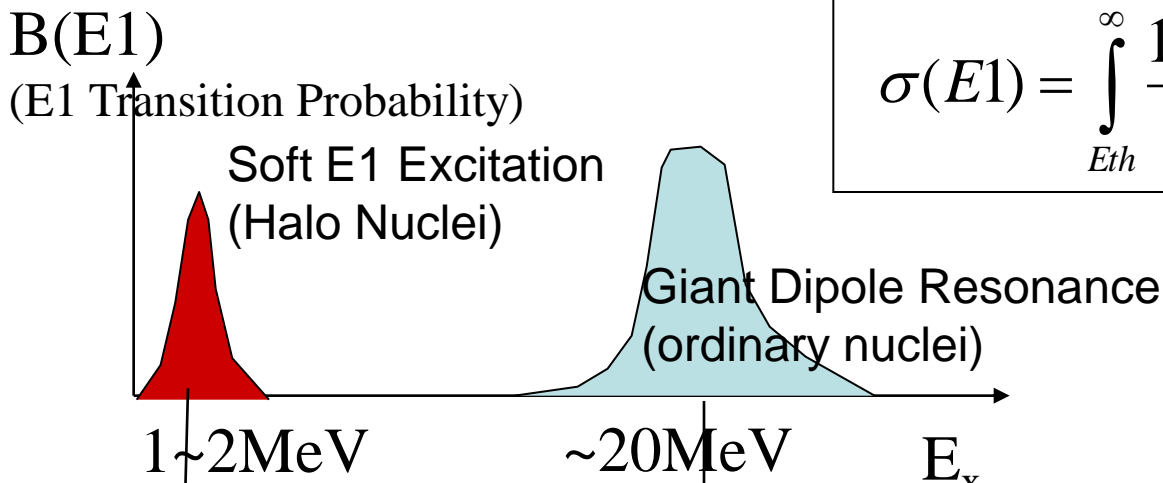
Cross section = (Photon Number) x (Transition Probability)

C.A. Bertulani, G. Baur, Phys. Rep. 163,299(1988).



# Inclusive Coulomb Breakup

$$\sigma(E1) = \int_{E_{th}}^{\infty} \frac{16\pi^3}{9\hbar c} N_{E1}(E_x) \frac{dB(E1)}{dE_x} dE_x$$

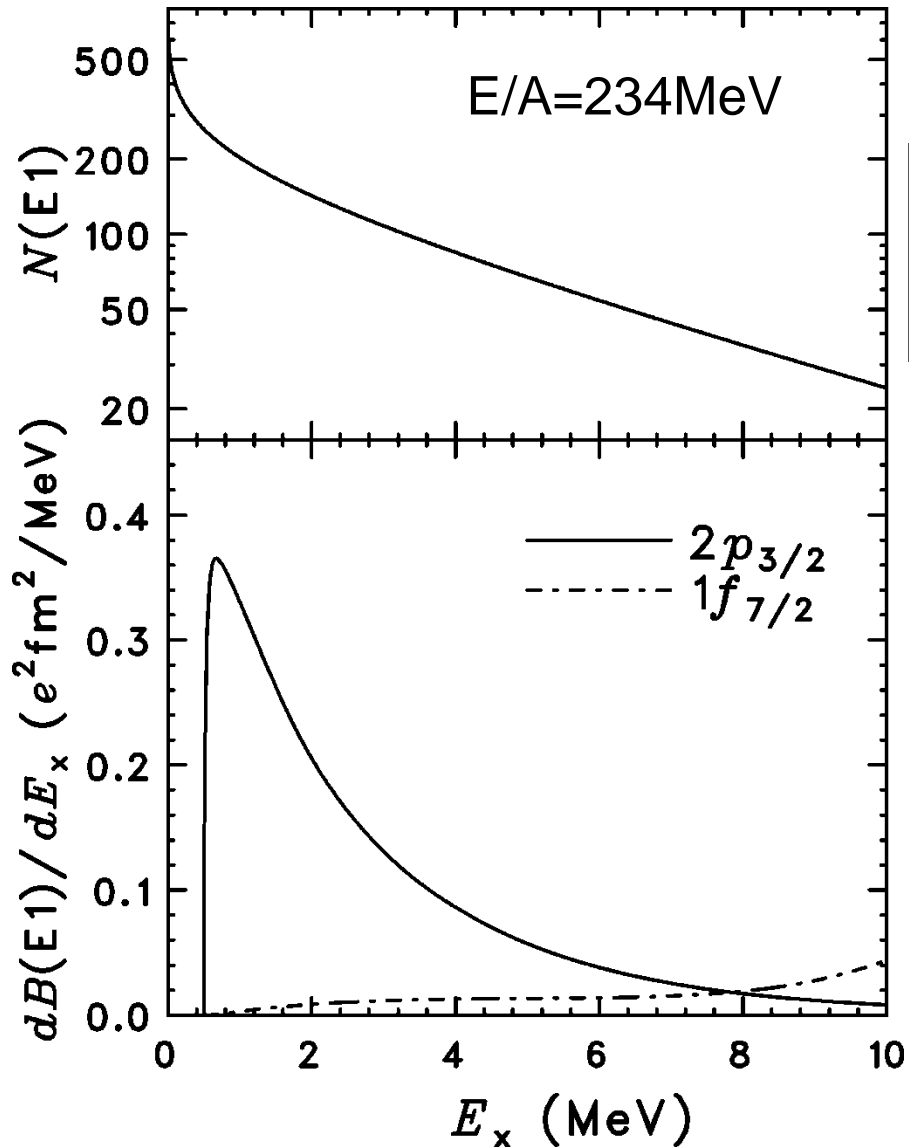


$\sigma(E1) \sim 0.5\text{--}1\text{b}$

$< \sim 0.1\text{b}$

*Halo/Non Halo can be distinguished only from the inclusive cross section !*

$^{31}\text{Ne}$  (N=21)  $S_n=0.29(1.64)\text{MeV}$

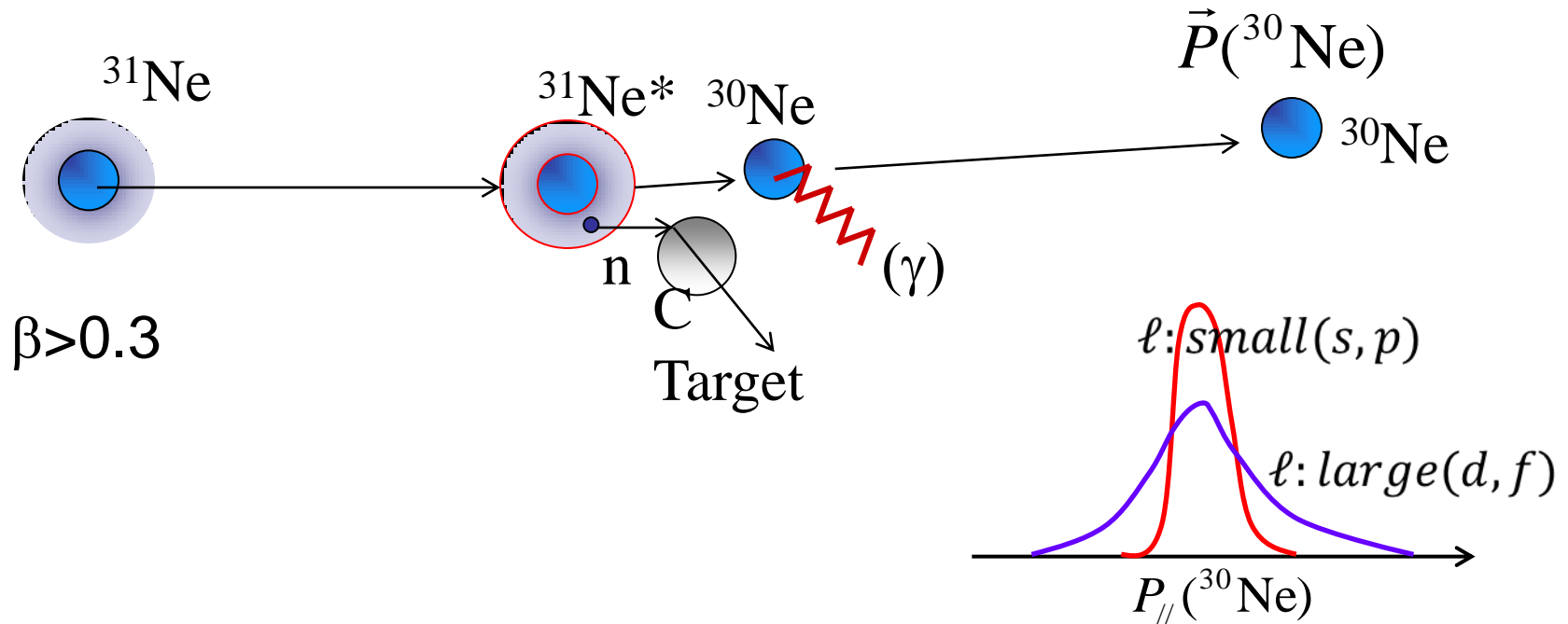


$$\sigma(E1) = \int_{E_{th}}^{\infty} \frac{16\pi^3}{9\hbar c} N_{E1}(E_x) \frac{dB(E1)}{dE_x} dE_x$$

$$\begin{aligned} \frac{dB(E1)}{dE_x} &\propto \left| \langle \mathbf{q} | \frac{Z}{A} \mathbf{r} Y^1_m | \Phi_{gs} \rangle \right|^2 \\ &= \Sigma C^2 S \left| \langle \Phi_f | \frac{Z}{A} \mathbf{r} Y^1_m | \phi_{nlj} \rangle \right|^2 \end{aligned}$$

# Probe-2: Nuclear Breakup

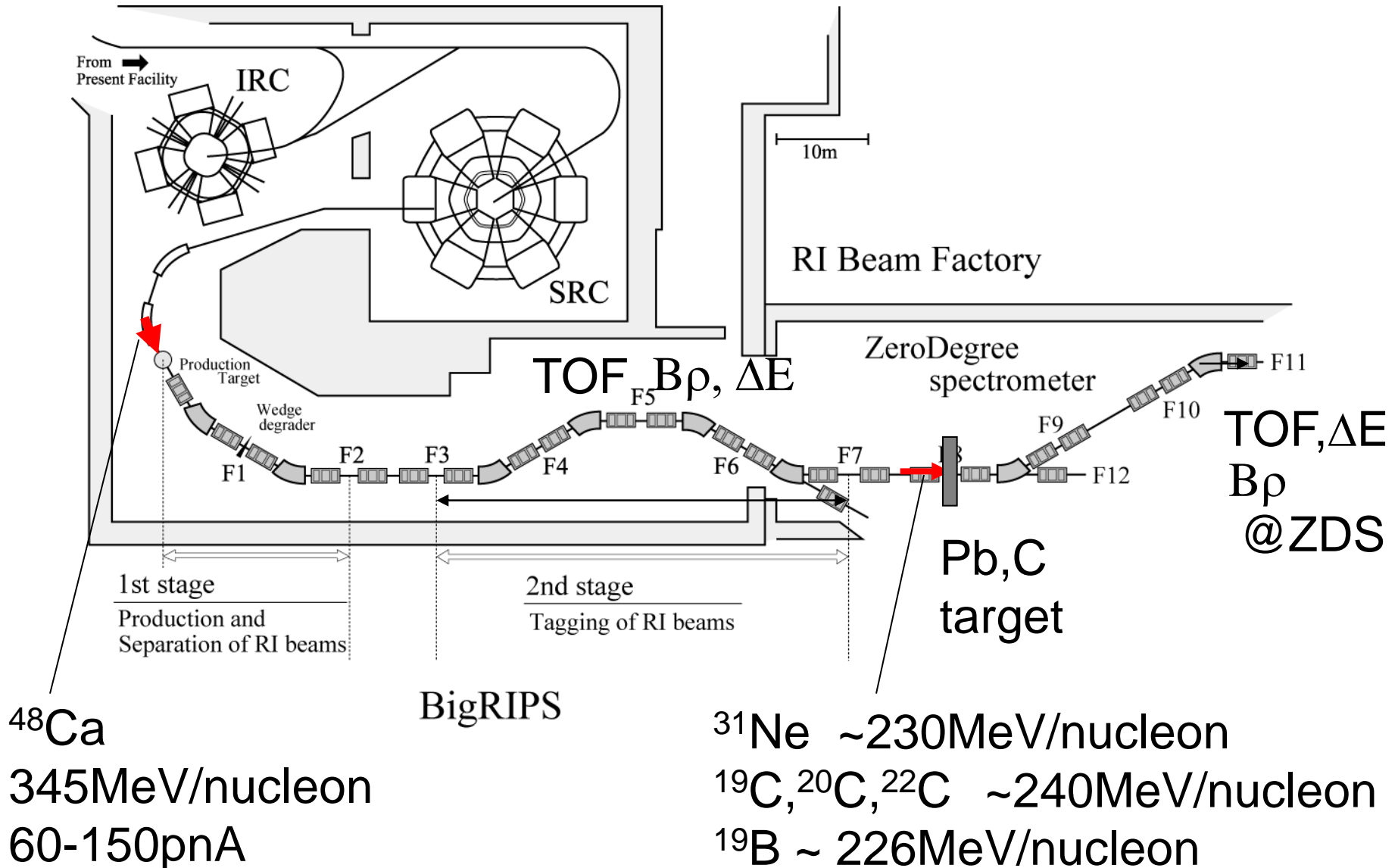
→ e.g. 1n knockout reaction of  $^{31}\text{Ne}$



- $\gamma$  ray in coincidence →  $^{30}\text{Ne}(2^+) / ^{30}\text{Ne}(0^+)$  Contribution
- $\sigma_{-1n}$  and  $P_{//}$  distribution →  $\ell$  of valence n, configuration

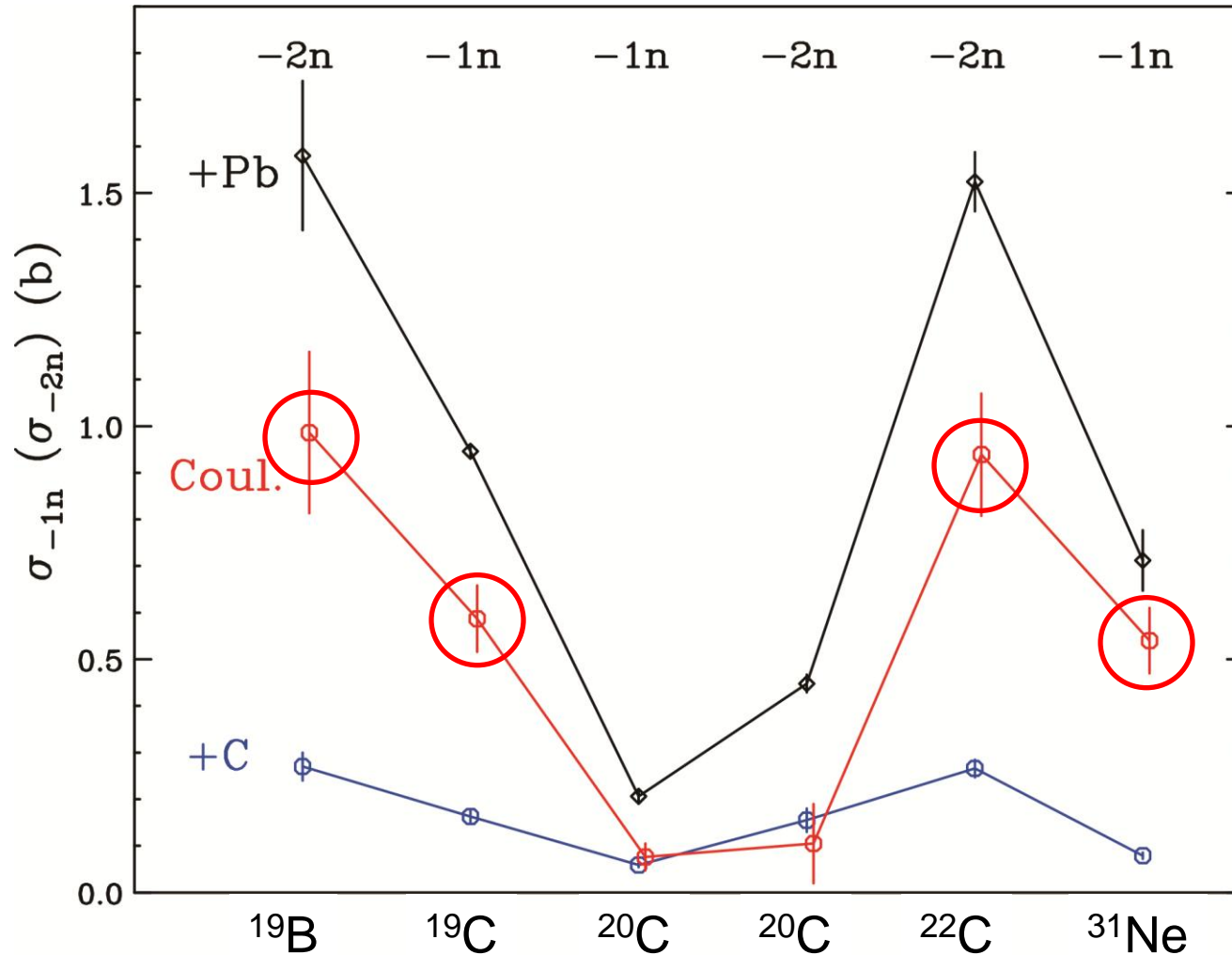
*Theory: Eikonal Approximation*

# Experiment at BigRIPS & ZDS at RIBF



# RESULTS: 1n (or 2n) removal cross section

→ Coulomb breakup cross section



Evidence for  
 2n Halo in  $^{22}\text{C}$   
 1n Halo in  $^{31}\text{Ne}$   
 c.f.  $^{19}\text{C}$  (known halo)

2n Halo in  $^{19}\text{B}$  :  
 First Measurement

$$\sigma(E1) = \sigma(\text{Pb}) - \Gamma \sigma(\text{C})$$

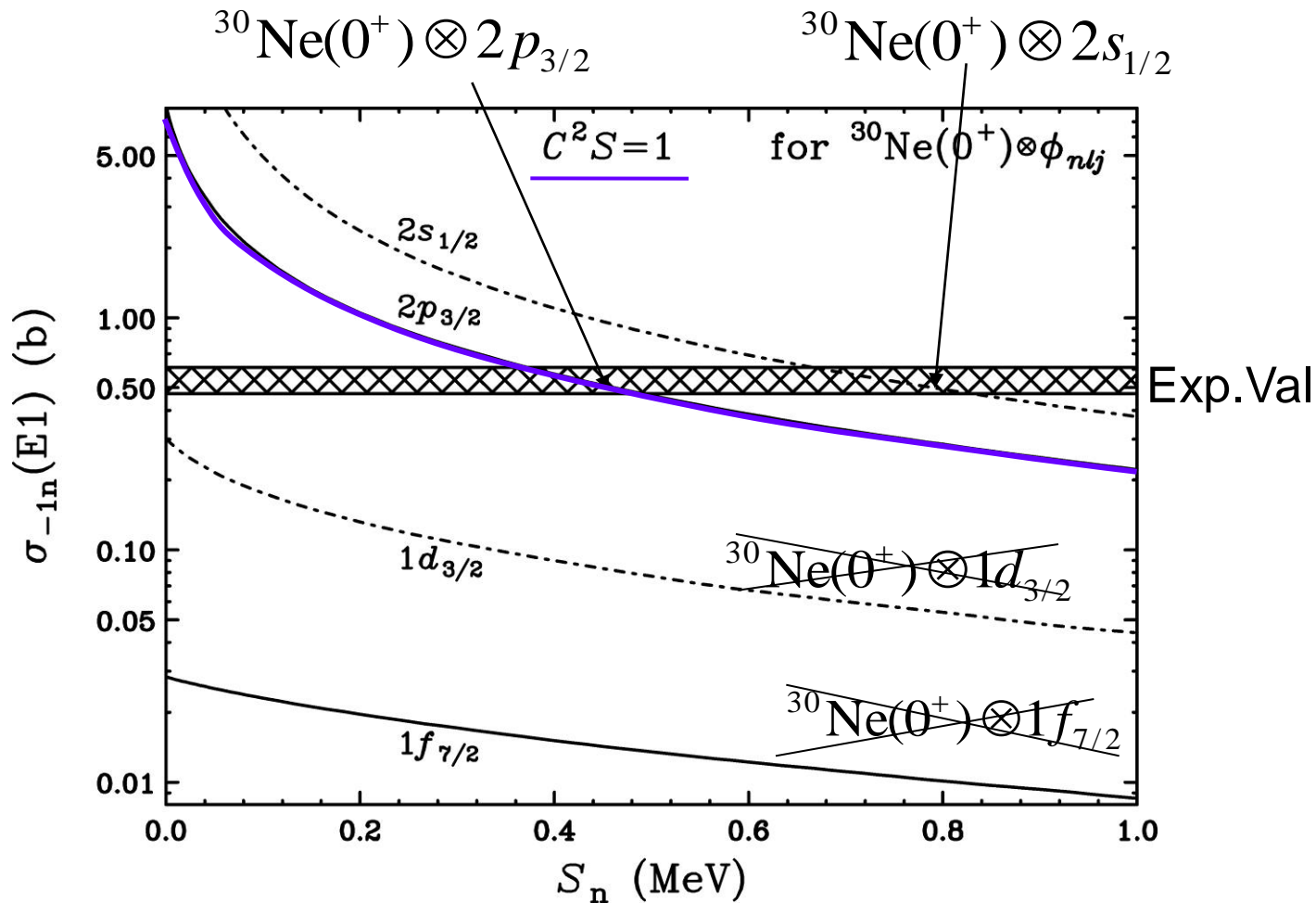
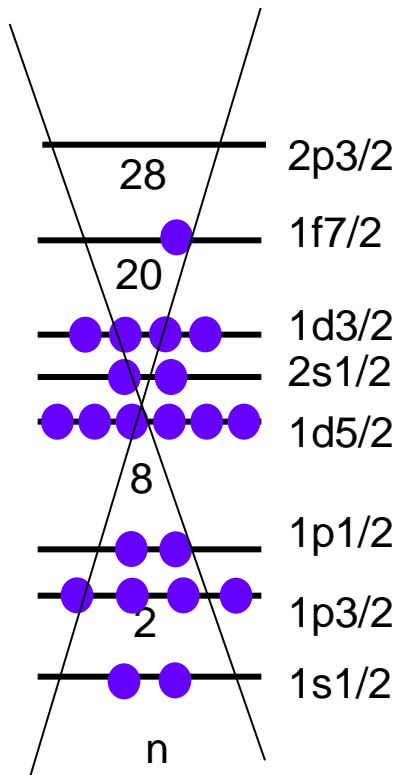
$$\Gamma \approx 1.7 - 2.6$$

c.f.  $^{22}\text{C}$ : Reaction cross section, K.Tanaka et al.PRL104,062701 (2010).

$^{31}\text{Ne}$ : Reaction cross section, M. Takechi et al.PLB707, 357 (2012).



# $^{31}\text{Ne}$ (N=21) Shell Configuration



$2p_{3/2}$  or  $2s_{1/2}$  Low-L orbits  $\rightarrow$  Large E1  $\rightarrow$  1n-halo structure of  $^{31}\text{Ne}$   
 $^{30}\text{Ne}(0^+) \times 1f_{7/2}$  Excluded  $\rightarrow$  Shell gaps(20,28) vanish at  $^{31}\text{Ne}$   
 $\rightarrow$  Island of inversion

*Still Unknown:  $S_n$ /Configuration Mixing  $C^2S < 1$*

# How the $^{31}\text{Ne}$ g.s. is made of ?

→  $\gamma$  and Nuclear Breakup data

Monte Carlo Shell Model

by Utsuno, Otsuka

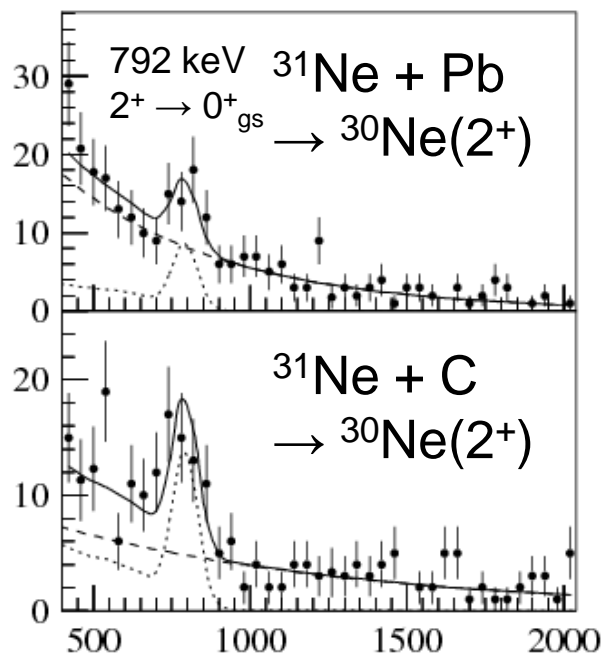
→  $3/2^-$  g.s.

$$^{30}\text{Ne}(0^+) \otimes 2p_{3/2} \quad C^2S = 0.12$$

$$^{30}\text{Ne}(2^+) \otimes 2p_{3/2} \quad C^2S = 0.27$$

$$^{30}\text{Ne}(2^+) \otimes 1f_{7/2} \quad C^2S = 0.25$$

*Possible Large Proportion for the  $^{30}\text{Ne}(2^+)x(nlj)$  Configuration*



$$^{31}\text{Ne} + \text{Pb} \rightarrow ^{30}\text{Ne}(0^+) : 0.515(103) \text{ barn}$$

$$^{31}\text{Ne} + \text{Pb} \rightarrow ^{30}\text{Ne}(2^+) : 0.197(79) \text{ barn}$$

$$\rightarrow \sigma(E1; 0^+) = 0.45(11) \text{ barn}$$

(~90% of Total  $\sigma(E1) = 0.54(7)$  barn)

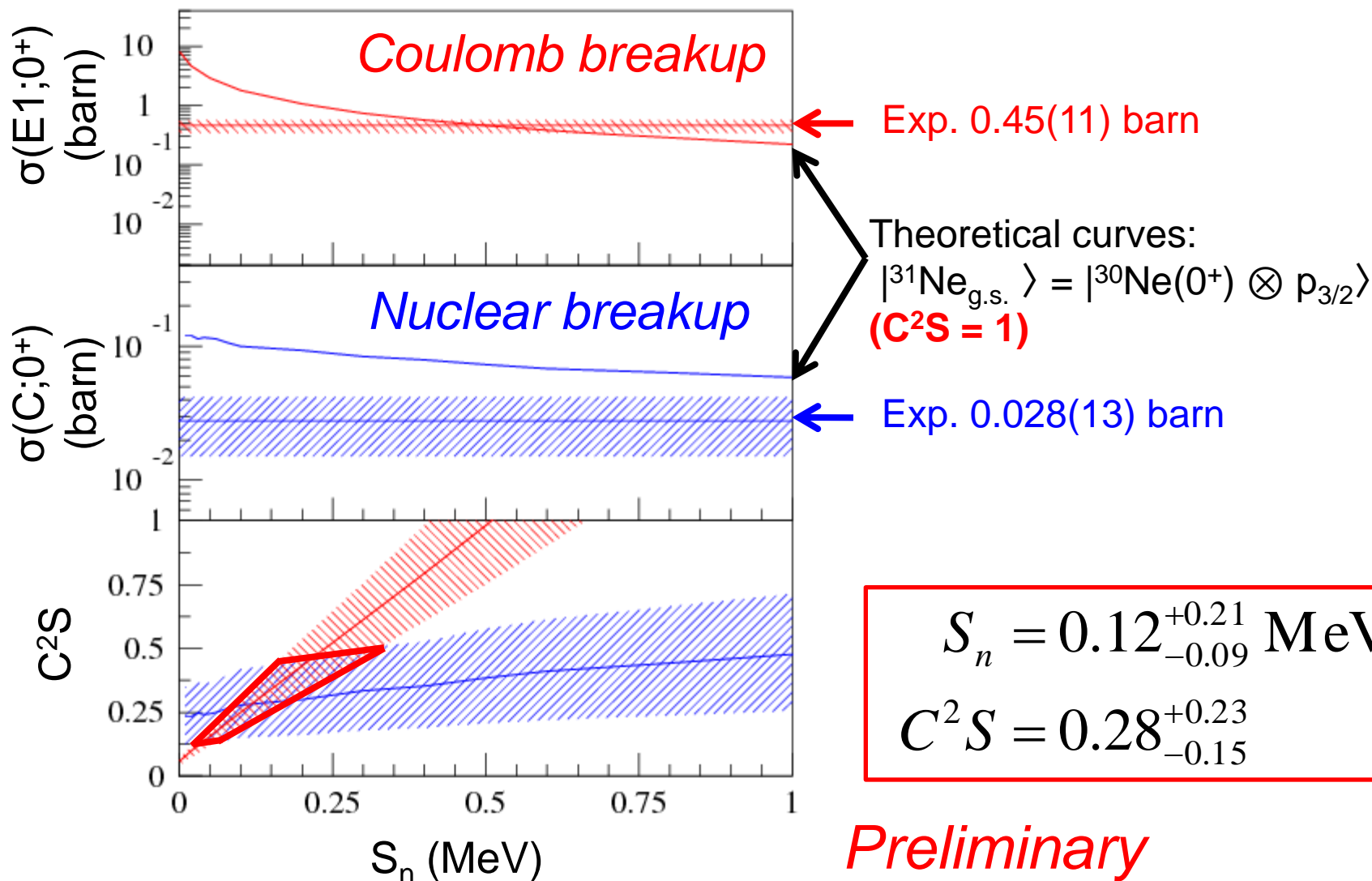
$$^{31}\text{Ne} + \text{C} \rightarrow ^{30}\text{Ne}(0^+) : 0.028(13) \text{ barn} \quad \sim 35\%$$

$$^{31}\text{Ne} + \text{C} \rightarrow ^{30}\text{Ne}(2^+) : 0.051(12) \text{ barn} \quad \sim 65\%$$

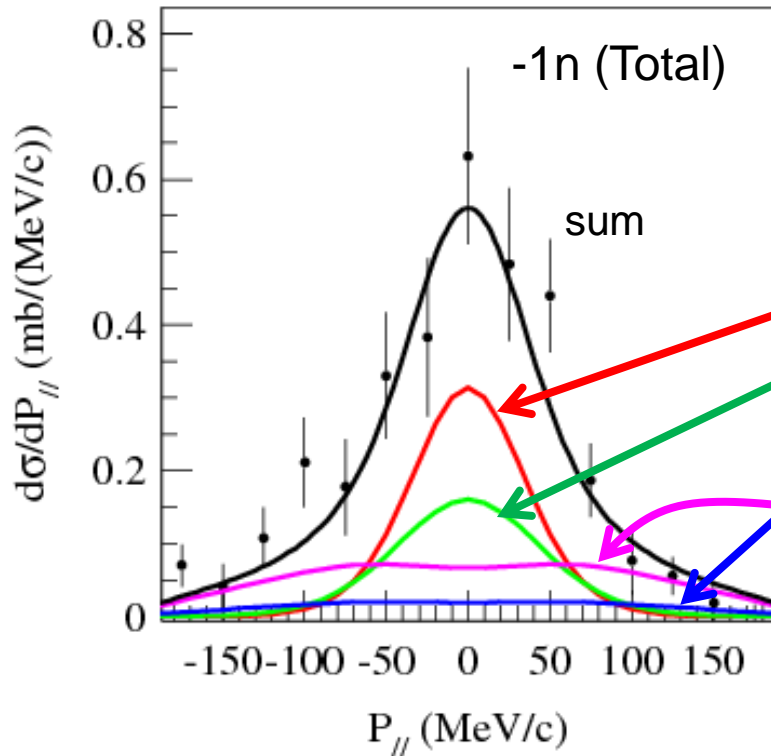


# Estimation of $C^2S$ & $S_n$

Channel:  $^{31}\text{Ne} \rightarrow ^{30}\text{Ne}(0^+)$



# Momentum distribution of $^{30}\text{Ne}$ fragment



Resolution( $\sigma$ )= 24 MeV/c  
 $\Gamma$  (FWHM) =73(11) MeV/c

$S_n = 0.12$  MeV

state	C <sup>2</sup> S	$\sigma$ (barn)	
$ ^{30}\text{Ne}(0^+) \otimes 2p_{3/2}\rangle$	0.28	0.028	← Exp.
$ ^{30}\text{Ne}(2^+) \otimes 2p_{3/2}\rangle$	0.27	0.018	
$ ^{30}\text{Ne}(2^+) \otimes 1f_{7/2}\rangle$	0.25	0.0058	
$ ^{30}\text{Ne}(4^+) \otimes 1f_{7/2}\rangle$	1.06	0.021	
total		0.073 b (0.079(7)b)	

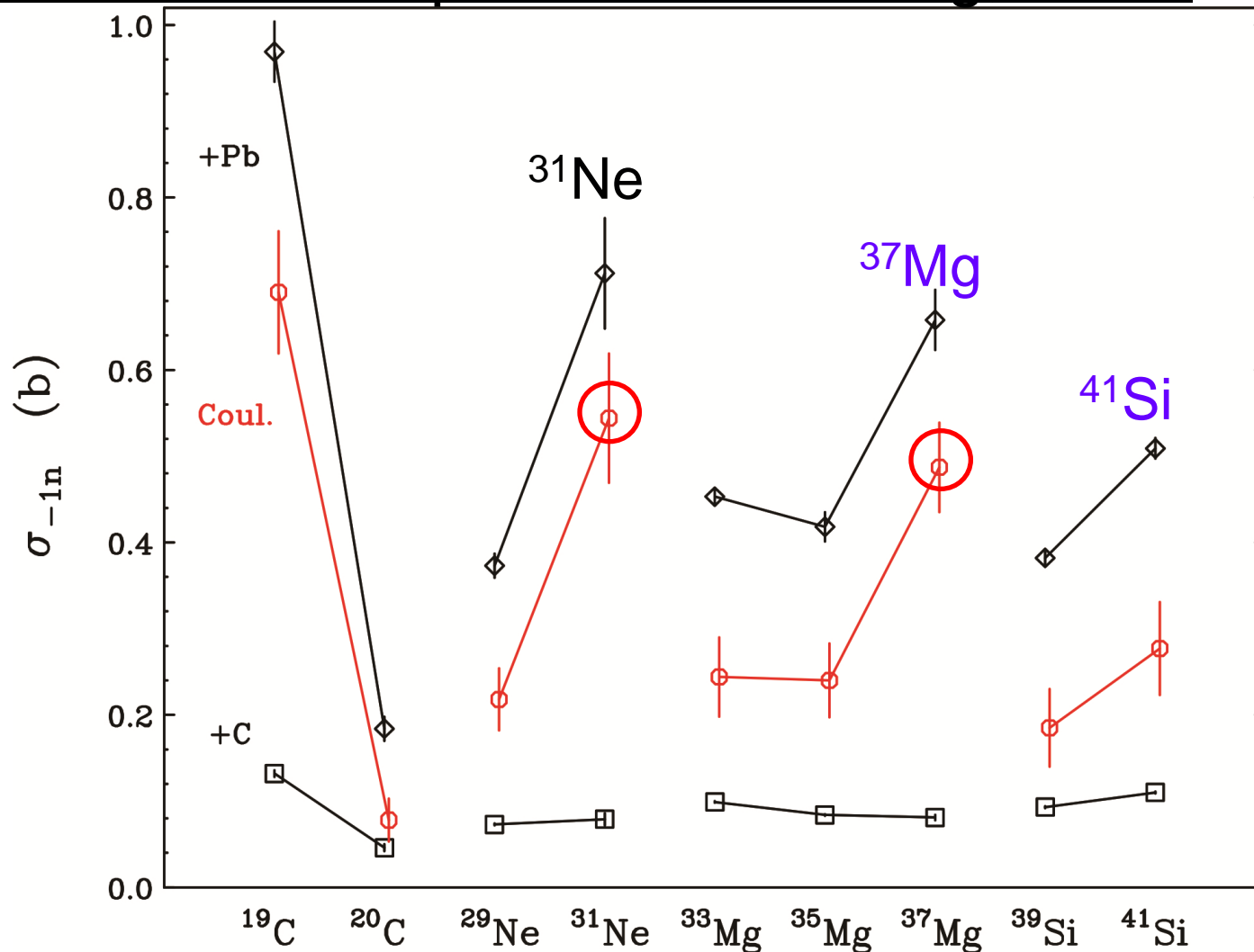
Shell model calculation  
 by using SDPF-M interaction

Preliminary

“p-wave neutron halo composed of two components”

Results of  $^{29}\text{Ne}$ ,  $^{33,35,37}\text{Mg}$ ,  $^{39,41}\text{Si}$

# Coulomb Breakup of $^{29}\text{Ne}$ , $^{33,35,37}\text{Mg}$ , $^{39,41}\text{Si}$



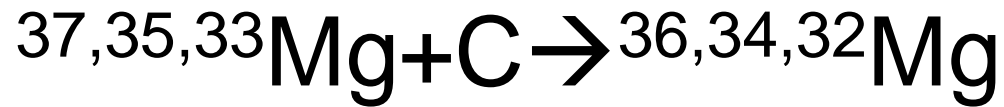
$^{37}\text{Mg}$ :  $\sigma(\text{E1})=490(50)$  mb (*preliminary*)

$\rightarrow$   $^{37}\text{Mg}$ : Possible 1n-Halo State

c.f.  $^{31}\text{Ne}$   $\sigma(\text{E1})=540(70)$  mb

*preliminary*

# Momentum Distribution

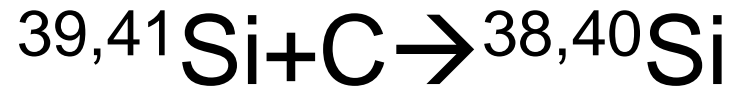


Curves: Fitted with Eikonal Calculation assuming  
gs(Odd Mg 7/2- or 3/2-) to gs (Even Mg 0+) transition

N.B. In reality these two  
components should not be directly added,  
but the p and f ratio could be estimated.

Mixture of  $p_{3/2}$  and  $f_{7/2}$   
 $\sigma_{-1n}(f_{7/2}) \sim \sigma_{-1n}(p_{3/2})$

# Momentum Distribution *preliminary*

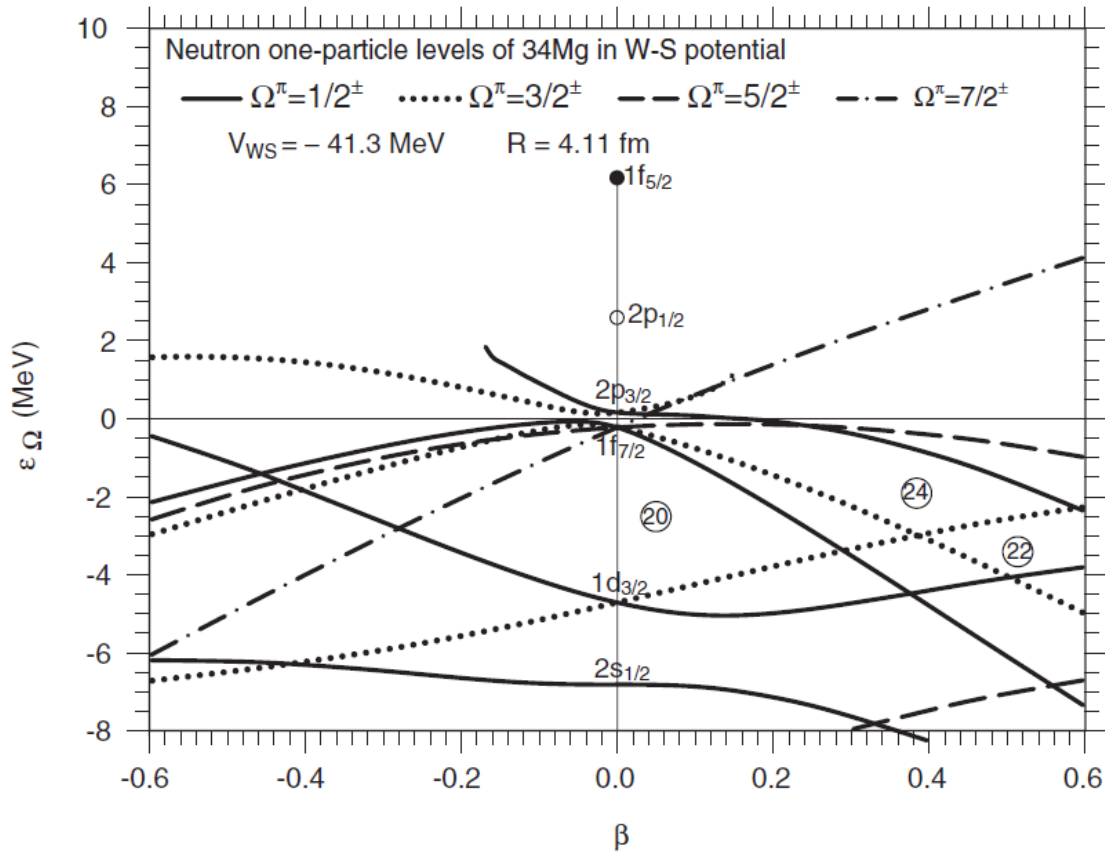


Mixture of  $s_{1/2}$  and  $d_{3/2}$   
 $\sigma_{-1n}(d_{3/2}) > \sigma_{-1n}(s_{1/2})$

Mixture of  $p_{3/2}$  and  $f_{7/2}$   
 $\sigma_{-1n}(f_{7/2}) > \sigma_{-1n}(p_{3/2})$



# Preliminary Interpretation by Nilsson Model



$^{29}\text{Ne}, ^{33}\text{Mg}, ^{35}\text{Mg}, ^{37}\text{Mg}, ^{39}\text{Si}, ^{41}\text{Si}$   
( $N=19, 21, 23, 25, 25, 27$ )

eg.  $\beta=0.2$

$^{29}\text{Ne}(N=19) 3/2^+ (\Omega=3/2^+)$   
 $^{31}\text{Ne}(N=21) 3/2^- (\Omega=1/2^-)$   
 $^{33}\text{Mg}(N=21) 3/2^- (\Omega=1/2^-)$   
 $^{35}\text{Mg}(N=23) 3/2^- (\Omega=3/2^-)$   
 $^{37}\text{Mg}(N=25) 5/2^- (\Omega=5/2^-)$   
 $^{39}\text{Si} (N=25) 5/2^- (\Omega=5/2^-)$   
 $^{41}\text{Si} (N=27) 3/2^- (\Omega=1/2^-)$

I.Hamamoto PRC85, 064329 (2012).

$^{37}\text{Mg}(N=25)$

most probably  $3/2^- \rightarrow \beta > 0.25$

$^{33,35,37}\text{Mg}$  seems to have similar behavior except that  $^{37}\text{Mg}$  is like halo.  $2^+$  energy of the core is similar, too.

# Summary

## Coulomb/Nuclear Breakup of $^{31}\text{Ne}$ (inclusive)

Combinatorial Analysis  $\rightarrow$  Sn  $\sim 0.12\text{MeV}$ , C<sup>2</sup>S( $p_{3/2} \times ^{30}\text{Ne}(0^+)$ )  $\sim 0.27$

Mixed with  $p_{3/2} \times ^{30}\text{Ne}(2^+)$   
 $f_{7/2} \times ^{30}\text{Ne}(2^+)$

## Coulomb/Nuclear Breakup of $^{29}\text{Ne}$ , $^{33,35,37}\text{Mg}$ , $^{39,41}\text{Si}$ (inclusive)

$^{37}\text{Mg}$  ---- Large E1 Cross Section 0.49(5) barn  $\rightarrow$  Halo Like Structure

Momentum Distribution of fragment with C target  $\rightarrow$

$^{33,35,37}\text{Mg}$  Mixture of  $p_{3/2}$  and  $f_{7/2}$  configuration  $\sigma(f_{7/2}) \sim \sigma(p_{3/2})$

$^{39,41}\text{Si}$   $\sigma(f_{7/2}) > \sigma(p_{3/2})$

$^{29}\text{Ne}$   $\sigma(d_{3/2}) > \sigma(s_{1/2})$

---"Island of inversion"  $\rightarrow$  Halo with Deformed Core

(How do we call this?)

Deformed Halo ?

Deformed-core coupled halo?)

# Collaborators

## Inclusive Coulomb Breakup of $^{31}\text{Ne}$ and $^{22}\text{C}$ ( $^{31}\text{Ne}$ Coulomb BU: PRL103,262501(2009))

**T.Nakamura**, **N.Kobayashi**, Y.Kondo, Y.Satou, N.Aoi, H.Baba, S.Deguchi, N.Fukuda, J.Gibelin, N.Inabe, M.Ishihara, D.Kameda, Y.Kawada, T.Kubo, K.Kusaka, A.Mengoni, T.Motobayashi, T.Ohnishi, M.Ohtake, N.A.Orr, H.Otsu, T.Otsuka, A.Saito, H.Sakurai, E. Simpson, S.Shimoura, T.Sumikama, H.Takeda, E.Takehita, M.Takechi, S.Takeuchi, K.Tanaka, K.N.Tanaka, N.Tanaka, Y.Togano, J.A. Tostevin, Y.Utsuno, K. Yoneda, A.Yoshida, K.Yoshida,

## Inclusive Coulomb/Nuclear Breakup of $^{29}\text{Ne}$ , $^{33,35,37}\text{Mg}$ , $^{39,41}\text{Si}$

**N.Kobayashi**, **T.Nakamura**, Y.Kondo, Y.Satou, N.Aoi, H.Baba, R. Barthelemy, S.Deguchi, M. Famiano, N.Fukuda, J.Gibelin, Lee Giseung, N.Inabe, M.Ishihara, D.Kameda, R.Kanungo, Y.Kawada, T.Kubo, M. Matsushita, T. Motobayashi, T.Ohnishi, K. Nikolski, N.A.Orr, H.Otsu, T. Otsuka, T. Sako, H.Sakurai, Lee H. Sang, T.Sumikama, K. Sunji, H.Takeda, K. Takahashi, S.Takeuchi, N.Tanaka, R. Tanaka, Y.Togano, Y. Utsuno, K. Yoneda

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